

1 **Downward-propagating eruption following vent unloading implies no**
2 **direct magmatic trigger for the 2018 lateral collapse of Anak Krakatau**

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45 **Abstract**

46 The lateral collapse of Anak Krakatau volcano, Indonesia, in December 2018,
47 highlighted the potentially devastating impacts of volcanic edifice instability. The
48 trigger for the Anak Krakatau collapse remains obscure; the volcano had been
49 erupting for the previous six months, and although failure was followed by intense
50 explosive activity, it is the period immediately prior to collapse that is potentially key
51 in providing identifiable, pre-collapse warning signals. Here, we integrate physical,
52 microtextural and geochemical characterisation of tephra deposits spanning the
53 collapse period. We demonstrate that the first post-collapse eruptive phase (erupting
54 juvenile clasts with a low microlite areal number density and relatively large
55 microlites, reflecting crystal growth) is best explained by instantaneous unloading of
56 a relatively stagnant upper conduit. This was followed by the second post-collapse
57 phase, on a timescale of hours, which tapped successively deeper portions of the
58 plumbing system, reflected in highly anorthitic microlite populations, alongside higher
59 calculated ascent velocities and decompression rates, within the post-collapse
60 tephra-stratigraphy. This implies downward propagating destabilisation of the
61 magma storage system, as a response to collapse, rather than pre-collapse magma
62 ascent triggering failure. Importantly, this suggests that the collapse was a
63 consequence of longer-term processes linked to edifice growth and instability, and
64 that no indicative changes in the magmatic system could have signalled the potential
65 for incipient failure. Therefore, monitoring efforts may need to focus on integrating
66 short- and long-term edifice growth and deformation patterns to identify increased
67 susceptibility to lateral collapse. The post-collapse eruptive pattern also suggests a
68 magma pressurisation regime that is highly sensitive to surface-driven perturbations,
69 which led to elevated magma fluxes after collapse and rapid edifice regrowth. Not

70 only does rapid regrowth potentially obscure evidence of past collapses, but it also
71 emphasises the finely balanced relationship between edifice loading and crustal
72 magma storage.

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74 **Keywords:** volcanic lateral collapse; sector collapse; magma ascent; microlites;
75 Anak Krakatau; rapid decompression; post-collapse volcanism

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77 **Graphical abstract:**

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93 **1. Introduction**

94 Landslides generated by volcanic flank failure are significant hazards and can
95 cause destructive tsunamis in island settings. Due to their relatively low frequency,
96 very few lateral collapses have been well observed, and there remains limited
97 understanding of the precursors to collapse. Determining the trigger(s) for collapse is
98 often difficult, as edifice instability can develop from a range of factors including
99 gravitational and structural weaknesses, shallow magmatic intrusions and
100 hydrothermal alteration (e.g., Siebert et al., 1987; López and Williams, 1993;
101 McGuire, 2003). However, it is important to constrain the factors driving failure, and
102 particularly whether shifts in magmatic activity are implicated, in order to develop a
103 systematic monitoring framework that could potentially identify signals of incipient
104 collapse.

105 The lateral collapse of Anak Krakatau, Indonesia, on 22 December 2018, which
106 induced a tsunami causing over 400 fatalities on surrounding coastlines (Grilli et al.,
107 2019), both highlights the impacts of such events and provides an opportunity to
108 better understand the processes that lead to volcanic edifice failure. Anak Krakatau
109 is a small stratovolcano within the Sunda Strait, positioned on the NE edge of the
110 1883 Krakatau caldera (Deplus et al., 1995). The volcano emerged above sea-level
111 in 1927, developing rapidly through frequent eruptions to a pre-collapse height of
112 330 m (Grilli et al., 2019). Collapse of its SW flank, with a volume of 0.18-0.31 km³
113 (Hunt et al., in press), occurred six months into an eruption phase characterised by
114 Strombolian, Vulcanian and effusive activity. Although this activity was typical of
115 previous decades (e.g., Abdurrachman et al., 2018), it involved a relatively elevated
116 magma flux (Walter et al., 2019). Infrared data from the Moderate Resolution
117 Imaging Spectroradiometer (MODIS) revealed the volcano's highest thermal levels

118 since MODIS measurements began in 2000. However, these were particularly
119 elevated during late September 2018 (Walter et al., 2019), rather than showing a
120 clear temporal progression that culminated in collapse. It thus remains unclear if the
121 collapse was directly initiated by a discrete and identifiable shift in magmatic activity,
122 or if collapse resulted from longer-term growth and destabilisation. What is clear is
123 that the collapse was accompanied (as far as the temporal resolution limits of
124 geophysical observations can constrain) and/or immediately followed by intense
125 explosive activity. The collapse cut the active conduit beneath sea-level, resulting in
126 Surtseyan eruptions that produced extensive ash deposits, rapidly buried the
127 collapse scar, fed convective atmospheric plumes reaching 16-18 km (Prata et al.,
128 2020), and involved a higher magma flux than anything recorded in recent decades
129 (Gouhier and Paris, 2019). Here, we seek to determine the role of magmatic activity
130 in the 2018 Anak Krakatau collapse, specifically addressing whether the intense
131 accompanying volcanism was a driver or a consequence of edifice failure, by
132 reconstructing magma ascent conditions spanning the syn- and post-collapse period.
133 This is important not only to understand this particular event, but more generally to
134 identify causes of edifice failure at both active and inactive volcanoes, to develop
135 approaches for monitoring edifice stability, and to understand the relationship
136 between surface mass redistribution and magma ascent behaviour (cf. Petrone et
137 al., 2009; Watt, 2019).

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139 **2. 2018 eruption and collapse observations**

140 Observations in the period spanning the collapse provide important constraints for
141 understanding eruptive behaviour, ash dispersal, and characteristics of the post-

142 collapse tephra-stratigraphy. All times stated in this section are Western Indonesian
143 Time (WIB; UTC + 7 hours).

144 The 2018 eruption began in late June (PVMBG, 2018). Discharge rates,
145 calculated from MODIS data, indicate that the magma flux peaked in September,
146 gradually waning after October. Intensified eruptions on 22 December produced
147 another peak in activity, but discharge rates had reached comparable or higher
148 levels on ten previous occasions between June and December (Walter et al., 2019).
149 Fishermen, familiar with the area, were located near the island during the collapse,
150 and reported that the activity on 22 December had increased but was not unusual
151 (Perttu et al., 2020). Infrasound signals suggest intense activity started at ~13:00 on
152 22 December, an interpretation confirmed by eyewitness reports (from the Javan
153 coast) of audible and visible Strombolian eruptions. The infrasound signals were
154 similar in intensity to those from elevated eruption phases in 1999 and October 2018
155 (Perttu et al., 2020).

156 Although the June-December 2018 activity was not atypical in style, it is estimated
157 to have added 54 million tons of rock onto Anak Krakatau's central cone and
158 southern flank (Walter et al., 2019). SW deformation of the flank is identifiable from
159 Interferometric Synthetic Aperture Radar (InSAR) prior to June 2018, and at an
160 accelerated rate thereafter (Walter et al., 2019). Fissure development and fumarolic
161 activity within the pyroclastic cone supports these observations of gradual
162 deformation (Hunt et al., in press).

163 On 22 December 2018, plumes extending into the cloud base were observed from
164 14:30, with pulsatory Strombolian explosions peaking at 18:30 and a white plume
165 descending to the shoreline, implying lava effusion (Perttu et al., 2020). The same
166 authors report that fishermen decided the island was too dangerous to return to after

167 19:00 and observed lightning in the Strombolian plume at 20:00; and that two high-
168 frequency seismic signals between 19:50 and 20:00 are consistent with small-scale
169 slope failures. In the hour before the collapse, Darwin Volcano Ash Advisory Centre
170 (VAAC) registered a final pre-collapse plume at 20:10 (Perttu et al., 2020). A
171 cessation in audible explosions in the 30 minutes before the collapse broadly
172 coincides with an infrasound signal pause of a few minutes (Walter et al., 2019),
173 suggesting a pre-collapse break in surface activity (Perttu et al., 2020).

174 The main collapse occurred in a single stage of movement as indicated by
175 tsunami observations and modelling (Grilli et al., 2019; timed at 20:55-57),
176 eyewitness accounts (Perttu et al., 2020), as well as seismic (20:55:49) and
177 infrasound signals (20:55:51) (Walter et al., 2019; Perttu et al., 2020; Ye et al., 2020;
178 all consistent with a SW-moving landslide source). There is no evidence for an
179 unusually large explosion preceding collapse (Perttu et al., 2020), although the main
180 failure is bracketed by relatively high-frequency infrasound signals, two minutes
181 before and 1.5 minutes afterwards. The former, at 20:54, is also observed in seismic
182 records; Walter et al. (2019) interpret it as a possible explosion or earthquake signal
183 that may have triggered the main collapse, whereas Perttu et al. (2020) conclude
184 that it is consistent with a smaller-scale slope failure.

185 A brightness temperature reduction from Himawari-8 satellite data (11.2 μm
186 channel) at ~20:55 indicates a volcanic cloud reaching ~16 km, coincident with the
187 collapse (Gouhier and Paris, 2019; Perttu et al., 2020). This signal has a sharp onset
188 and peaks at ~21:00, forming an ice-rich but ash-poor cloud advected to the SW
189 (Prata et al., 2020) with ash emission lasting ~40 minutes (described as a blast-like
190 explosion by Gouhier and Paris, 2019). Gouhier and Paris (2019) derive a higher
191 mass-eruption rate for this specific plume (9×10^5 kg/s) than for the subsequent

192 sustained phase (5×10^5 kg/s). Perttu et al. (2020) identify two further discrete
193 explosive pulses at 22:25 and 22:55, progressing into sustained activity until 12:05
194 on 28 December. This supports gradual development of an ice-rich convective plume
195 from 22:30, identified by Prata et al. (2020), reaching sustained levels by 01:00 on
196 23 December, at heights of 16-18 km. The SW ice-rich plume formed above an ash-
197 laden weak column, generated by intense Surtseyan eruptions that advected
198 eastwards at low altitudes (documented in aerial photographs on 23 December; Grilli
199 et al., 2019; Prata et al., 2020; Hunt et al., in press). This low-altitude, ash-rich plume
200 deposited tephra on Panjang island, 2 km east of Anak Krakatau (Fig.1), causing
201 severe vegetation damage, with fine ash reaching the Javan coast (authors'
202 observations). Intense Surtseyan activity lasted for two weeks, divided into three
203 phases by Gouhier and Paris (2019) (22-27 December, 28-29 December, 3-6
204 January 2019) and two phases by Perttu et al. (2020) (22-28 December; 30
205 December to early January 2019).

206 Current observations, outlined above, demonstrate that the pre-collapse eruptive
207 activity was intense but not unusual. There is no evidence of activity strongly
208 accelerating in the hours before the collapse; behaviour on 22 December
209 represented a renewal of vigorous eruptions, but output had peaked three months
210 beforehand. A powerful explosive eruption accompanied the collapse, and was
211 distinct from the sustained activity that followed within a few hours, with infrasound
212 signals, satellite observations and aerial photographs suggesting an immediate
213 switch from Strombolian to Surtseyan behaviour as water infiltrated the vent.

214

215 **3. Tephra deposits**

216 Ash samples (Supplementary Table 1) were collected from four different localities
217 on Panjang and Sertung, islands respectively east and west of Anak Krakatau
218 (Fig.1b). Difficulties with accessing higher ground meant that only one sample was
219 collected from Sertung, at site U23-2 (6.38 km SW of the vent) (Fig.1c). This was
220 collected in healthy forest, on level ground, at a ~100 m altitude and beyond the
221 tsunami inundation limit. The dark volcanic ash sample was found at the surface as
222 a 1-cm thick structureless layer mixed with leaf litter and above an organic soil,
223 consistent with a fall deposit. U23-2 is in line with the south-westward dispersal of
224 the high-level, ice-rich plume described above, but not with the NE/E-advected low-
225 level plume. Visual and satellite observations show no evidence of post-collapse
226 vegetation damage on Sertung. In contrast, significant ash deposition stripped
227 leaves and branches from the mature forest on Panjang (S.1), consistent with
228 observations of ash-laden plumes drifting over the island for several days after the
229 collapse. Samples were collected at multiple sites on Panjang, displaying a well-
230 bedded ash stratigraphy consistent with fall deposition, exceeding 20 cm in places
231 (excluding remobilised surface deposits). At site U10, a flat, open area ~50 m from
232 the shore on north Panjang, estimated to lie ~5 m above high-tide level, and 4.09 km
233 from the vent, these deposits directly overlie a pumice-rich sand layer mixed with
234 sparse marine shells, deposited on top of an organic soil (Fig.1a). A comparable ash
235 stratigraphy was observed and sampled at other sites across Panjang (S.2), but with
236 the pumice layer absent.

237 Samples from U23-2 and U10 were selected for further textural and geochemical
238 analysis. All analytical methods are described in Appendix 1. For comparison, an
239 older ash sample (KRA-233), from a May 1997 eruption, was also analysed, and was
240 collected from a fall deposit on Sertung (Fig.1b), 3.80 km from the vent.

241 4. Results

242 4.1. Physical overview of eruptive products

243 The physical features (i.e. grain size, componentry, grain exterior surfaces) of the
244 1997 ash (KRA-233) provide an insight into the products of pre-collapse
245 (Strombolian) magma ascent and fragmentation conditions, which can be contrasted
246 with the December 2018 samples (Fig.1a and c) to evaluate changes in eruptive
247 behaviour. KRA-233 is a poorly sorted ash (1.2 ϕ) with a unimodal grain size
248 distribution peak of 2-3 ϕ and a fine ash content of 13% (>4 ϕ or <63 μm ; Wohletz,
249 1983). The ash grains are predominantly glassy, black and angular (see
250 Supplementary Table 2 for component descriptions).

251 U23-2 is a 1 cm thick, structureless and poorly sorted (1.2 ϕ) fine ash-fall deposit
252 with a unimodal grain size of 2-3 ϕ . The deposit comprises highly glossy, fresh
253 glassy grains, which are angular and have fluidal, frothy or pitted surface textures.
254 These juvenile clasts can be divided into black and brown types; the latter are more
255 open-textured, and this clast type is rare in KRA-233, in contrast with all 2018
256 samples. The unimodal characteristics and homogeneous physical appearance of
257 U23-2 suggests it is the product of a single depositional event, rather than an
258 amalgamated deposit from pre- or post-collapse activity (i.e. an upwind equivalent of
259 the Panjang Surtseyan deposits). The high angularity (Fig.2d) and narrow grain-size
260 range of U23-2 contrasts with the U10 samples (Fig.1a and c), as do several
261 characteristics discussed in later sections.

262 U10 is a well-stratified sequence that lies beneath a >6-cm thick structureless
263 surface layer of remobilised ash (U10-11) and can be subdivided into 8 distinct ash

264 units (some comprising multiple layers; U10-3 to U10-10, with a total thickness of 21
265 cm overlying a pumiceous unit (U10-2). The ash units are characterised by moderate
266 to poor sorting ($0.9-1.6 \phi$) and unimodal grain-size peaks at $2-3 \phi$ with a fine ash
267 content ranging from ~8 to 33%. At the base of U10 (U10-1), a black structureless
268 organic soil, rich in ash and rootlets, is inferred to derive from pre-2018 Anak
269 Krakatau activity. The overlying unit, U10-2, is poorly sorted (1.4ϕ), structureless
270 and marked by erosional contacts with a maximum thickness of 11 cm. Sub-
271 angular/sub-rounded cream pumice fragments dominate the layer (52%), and are
272 assumed to originate from the 1883 Krakatau eruption based on appearance (cf.
273 Carey et al., 2001). U10-2 also contains minor (<4%) marine biogenic material (e.g.,
274 gastropod shells, sponge spicules) (Fig.2c). Based on these characteristics and an
275 absence of this layer at more elevated sites further inland (i.e., NP1 and NP2; S.2),
276 U10-2 is interpreted as a tsunami deposit resulting from the 2018 landslide. U10-3
277 (U10-3A and -3B) is a thin (0.5 cm) and slightly indurated purple ash with an oxidised
278 yellow-brown crust, distinct in colour and with by far the highest proportion of altered
279 and lithic grains (Fig.2e; Supplementary Table 2) of any studied samples. U10-3 is
280 also the only U10 unit that can be well-correlated, based on its distinctive colour, with
281 other 2018 ash exposures on N Panjang (i.e., NP2-1; S.2). Above this, U10-4 is an
282 inverse graded, planar bedded brown ash (35 mm) with a very fine top. Like U23-2,
283 this ash is dominated by black and brown juvenile clasts, but these have a duller and
284 less angular appearance. Grain characteristics are similar throughout the overlying
285 sequence. Above U10-4, U10-5 (25 mm) and U10-6 (28 mm) are very fine (modal
286 peak, $3-4 \phi$) brown ash beds, distinct from the rest of the sequence in displaying
287 weak cross-stratification that becomes more developed near the top of the layers; we
288 infer these may reflect deposition from basal surges rather than fallout. U10-7 to

289 U10-10 return to parallel planar bedded structures (Fig.1a). U10-7 (black medium
290 ash; 25 mm) and U10-8 (a similar deposit; 4 mm) show slight normal grading and
291 U10-9 (27-55 mm in total; up to 10 individual beds on a few-mm scale) and U10-10
292 (a 2 mm very fine basal brown ash overlain by three normally graded beds; 32 mm)
293 display alternate fine and very fine ash size grades. Based on its multi-bedded
294 characteristics and aerial observations of N/NE-directed ash-rich plumes (PVMBG,
295 2018; Prata et al., 2020), we infer that the U10-3 – U10-10 sequence represents
296 deposition from pulsatory post-collapse Surtseyan activity between 22 December
297 and early January 2019, though we cannot correlate exact dates with individual
298 layers.

299

300 4.2. Exterior surfaces of grains

301 Scanning electron microscope (SEM) secondary-electron images of mounted ash
302 grains were used to examine micro-scale features and potential differences in
303 fragmentation modes between samples. Many grains from all samples display brittle
304 features including stepped (Fig.2g) and conchoidal fractures, as well as river-line
305 patterns that indicate fracturing under mixed-mode stresses (Hull, 1999).

306 Additionally, many grains in U23-2 and U10 have secondary minerals (e.g., cubic
307 NaCl) and/or finer particles adhered to their surfaces; in some cases, these finer
308 particles have annealed together, creating irregular moss-like grains (Fig.2f; Büttner
309 et al., 2002). There are also infrequent occurrences (~1%) of ductile features in U23-
310 2 and U10 including Pele's hairs, platy grains, fluidal grains, and grains with molten
311 surfaces. A notable feature of U23-2 is the presence of smooth glass spherules (cf.
312 Genareau et al., 2015), which are not observed in the U10 or KRA-233 samples.

313

314 4.3. Ash morphology

315 Additional morphological and textural analyses of ash grains were undertaken on
316 SEM backscattered-electron images of samples. Liu et al. (2015a) determined that
317 proportions of dense (i.e. vesicle-poor grains, free crystals and lithics) and bubbly
318 grains (vesicular grains/glass shards) can be distinguished using a concavity index
319 (CI) threshold when plotting axial ratio against CI (Fig.3). Concavity index combines
320 solidity (solidity = grain area/grain convex hull) and convexity (convexity = convex
321 hull perimeter/grain perimeter) parameters, while axial ratio compares the major and
322 minor axes of the grains' best fit ellipse (Liu et al., 2015b).

323 U23-2 displays a high proportion of dense grains (79%) to bubbly grains,
324 averaged across three size fractions (250 -180 μm ; 180 - 125 μm ; 125 - 63 μm),
325 characterised by a low concavity index (CI <0.4) (Fig.3a). The dominance of planar,
326 fractured grains supports the high angularity observed from macrocomponentry of
327 U23-2. By comparison, U10-3B and U10-4 (the lowermost 2018 Panjang deposits)
328 have a very different morphology, with higher proportions of bubbly (75 and 70%,
329 respectively) to dense grains across the size fractions (Fig.3b and c). There are
330 slight variations between different size fractions: in U10-3B, the proportion of bubbly
331 grains decreases with size, whereas the CI peaks in the 125-180 μm size fraction of
332 U10-4, which may represent higher glass shard content (Liu et al., 2015a).
333 Morphometric analysis of multiple grain size fractions was not undertaken for other
334 units. However, data from U10-10 (63-125 μm size fraction) suggest comparable
335 features to U10-3B and U10-4, and visual observations throughout the U10
336 sequence imply little variation in physical characteristics.

337

338 4.4. Whole-rock and groundmass glass compositions

339 Both bulk rock (X-ray fluorescence) and groundmass glass (electron probe
340 microanalysis (EPMA)) compositions are summarised in Supplementary Table 3.
341 U23-2 and all analysed U10 samples have a basaltic andesite bulk composition
342 ($\text{SiO}_2 = 54.1\text{--}55.4$ wt.%, $\text{NaO} + \text{K}_2\text{O} = 4.3\text{--}4.4$ wt.%), maintaining the chemical
343 homogeneity of older (1993 to 2017; Gardner et al., 2013) and pre-collapse deposits
344 (July 2018; Walter et al., 2019) ($\text{SiO}_2 = 54.0\text{--}55.8$ wt.%, $\text{NaO} + \text{K}_2\text{O} = 4.2\text{--}4.6$ wt.%).
345 The only unit that displays a subtle difference is U10-3, which has a slightly higher
346 loss on ignition value and elevated SO_3 and Cu contents. Along with its
347 discolouration and lithic abundance (Fig. 2e), this may reflect a higher proportion of
348 hydrothermally altered material in this Surtseyan deposit than later units.

349 Matrix glass compositions (S.3) between KRA-233 (the 1997 ash) and U23-2 are
350 indistinguishable (mostly within the range $\text{SiO}_2 = 59.0\text{--}61.0$ wt.%), but those in the
351 U10 sequence are slightly less evolved with a narrow andesite compositional range
352 ($\text{SiO}_2 = 58.0\text{--}60.5$ wt.%).

353

354 4.5. Phase assemblage

355 EMPA and SEM analyses show that all studied samples contain a microlite (<50
356 μm in size), microphenocryst (~50 μm -0.5 mm in size) and phenocryst (~0.5-1 mm in
357 size) assemblage of plagioclase, orthopyroxene, olivine, clinopyroxene and
358 titanomagnetite within a glassy silicate matrix.

359 Plagioclase is the most abundant phase (~10 to 26 % area) in all samples,
360 forming subhedral to euhedral phenocrysts and microphenocrysts. Both phenocryst
361 types display normal or reverse zoning, resorbed and sieve textures, and amoeboid
362 melt inclusions. Plagioclase microlite morphologies in KRA-233 and U23-2 are
363 dominantly tabular, with rare swallow-tail and skeletal forms observed in U23-2. The

364 U10 samples contain a wider variety of morphologies including tabular, acicular,
365 skeletal, hopper and swallow-tail forms (Fig.4). Moreover, numerous feldspar
366 microlites in U10 are surrounded by a thin Fe-Ti rich compositional boundary layer
367 (Fig.4e), formed only when plagioclase growth rates match or exceed diffusion rates
368 in adjacent melt (Honour et al., 2019). KRA-233 and U23-2 microlites share a similar
369 compositional range of An₄₈ to An₆₈ (An, Anorthite = molar Ca/ (Ca + Na + K); mol
370 %). In contrast, the microlites in the U10 samples extend to An₇₉ (Fig.5a). Microlites
371 in all samples display a common main peak at An₅₆, but U10 samples show two
372 subsidiary peaks at An₆₈ and An₇₅. Microphenocryst and phenocryst core
373 compositions for all samples are more primitive (i.e. more anorthitic; ranges of An₆₂-
374 ₉₁ and An₄₅₋₈₉, respectively) than corresponding rims (An₄₈₋₆₈ and An₅₁₋₇₉,
375 respectively). The rims also show more primitive compositions progressively higher
376 up the U10 stratigraphy (U23-2 is also among the least primitive, but does not extend
377 to anorthite contents as low as KRA-233); core compositions exhibit the same trend,
378 albeit less strongly (Fig.5a). The range of anorthite contents is consistent with data
379 from 1970-2002 lava flows (Camus et al., 1987; Dahren et al., 2012).

380 KRA-233 and U23-2 have minor proportions of mafic minerals (i.e.,
381 orthopyroxene, olivine, clinopyroxene and oxides) (areal mean ~<6%), whereas U10
382 samples contain nearly double the proportion of these minerals (areal mean ~11%).
383 Pyroxene, olivine, and oxide phenocrysts and microphenocrysts in all samples are
384 predominantly euhedral. Microlite morphologies are largely tabular, equant or
385 skeletal. U23-2 is the only sample where polymineralic glomerocrysts (i.e., crystal
386 aggregates of various phases) are observed. Orthopyroxene is more dominant than
387 clinopyroxene, but both are heterogenous in composition, with orthopyroxene
388 spanning the enstatite-pigeonite fields and clinopyroxene the pigeonite-augite fields

389 (Fig. 5c and d). The Mg# (Mg# = molar Mg/(Mg + Fe)) of orthopyroxene (Mg#₆₆₋₇₄)
390 and clinopyroxene (Mg#₆₉₋₇₆) are homogeneous for core-rim pairs, with groundmass
391 values generally falling below Mg#₇₀, consistent with 1990-2002 lava flows (Mg#₆₁₋₇₄,
392 opx; Mg#₇₃₋₇₇, cpx; Dahren et al., 2012). Forsterite content (Fo = molar Mg/(Mg +
393 Fe); mol %) of olivine in phenocryst core-rim pairs (Fo₆₉₋₈₀) and microlites (Fo₆₆₋₇₃) for
394 the U23-2 and U10 samples are slightly more primitive than in rocks erupted in the
395 1970s (Camus et al., 1987) and KRA-233 (Fo₅₆₋₆₈, microlites; Fo₆₂₋₇₃, cores and rims)
396 (Fig.5b).

397

398 4.6. Microtextural observations

399 The U23-2 and U10 samples display distinct stratigraphic differences in the 2D
400 analysis of feldspar populations (Fig.6; S.4). All 2D textural data are summarised in
401 Table 1. Plagioclase areal number density (N_A), which defines the number of
402 feldspars per unit area (mm^{-2}), is lowest in the 1997 KRA-233 sample ($11,194 \text{ mm}^{-2}$).
403 N_A in U23-2 is also relatively low ($12,062 \text{ mm}^{-2}$), whereas U10-4, early in the post-
404 collapse Surtseyan sequence, has the highest density ($44,545 \text{ mm}^{-2}$). U10-3 (both
405 U10-3A and B) is slightly lower than U10-4, and later U10 samples range between
406 $13,342$ and $37,969 \text{ mm}^{-2}$.

407 Plagioclase microlite crystallinity (ϕ) describes the fraction that feldspar microlites
408 occupy within the groundmass, excluding all vesicles, phenocrysts, and pyroxene
409 and Fe-Ti oxide microlites (Hammer et al., 2000). A plot of N_A versus ϕ reveals that
410 U23-2 and KRA-233 have similar characteristics (S.5). Later U10 samples are also
411 similar but with slightly higher N_A and lower ϕ , while U10-4 and, to a lesser extent,
412 U10-3A/B show much higher values of both parameters (S.5). The mean crystal
413 sizes (S_n) of all samples range from ~ 2.56 to $4.56 \mu\text{m}$, with the smallest crystals in

414 U10-3B and the largest crystals in U23-2. A negative correlation is evident between
415 N_A and S_n , indicating that samples either have high numbers of small microlites or
416 the reverse (S.5). Feldspar aspect ratios for U10-4, U10-6, U10-8 and U10-10 are
417 high (5.14 – 7.28), representing more elongated crystal habits, whereas KRA-233,
418 U23-2 and U10-3A/B have lower ratios (4.08 – 4.19), defining more tabular crystals.

419 Vesicle morphologies within grains vary depending on crystallinity. Microlite-poor
420 grains typically exhibit smaller, spherical, and isolated vesicles, whereas microlite-
421 rich grains contain irregular and coalesced vesicles.

422

423 4.6.1. Ascent rates

424 To assess approximate ascent rates, we used the microlite number density (MND)
425 exsolution rate meter of Toramaru et al. (2008). Although this explicitly relates to
426 microlite nucleation depths, it likely reflects a time-averaged ascent rate throughout
427 the conduit (Murch and Cole, 2019). Since we used vesicular, microlite-rich grains,
428 the textures reveal ascent conditions near the conduit walls, providing longer
429 residence times for microlite crystallisation (Taddeucci et al., 2004). Therefore,
430 ascent rates may be considered as minimum values. The main equations from
431 Toramaru et al. (2008) are summarised here. Decompression rates (dP_w/dz) are
432 estimated by:

$$433 \left| \frac{dP_w}{dz} \right| = \frac{c}{b} \left(\frac{Nv}{a} \right)^{\frac{2}{3}} \quad (1)$$

434 where b represents a constant (40 for plagioclase), c is a function of water content
435 (C_w), N_v is microlite number volume, and a is a calculation involving C_w and glass
436 silica content (C_{Si}). Microlite water content was estimated using the plagioclase-melt
437 hygrometer of Putirka (2008), with values ranging between 2.5 and 3.6 ± 1 wt.% for

438 all samples (akin to Dahren et al., 2012). Magmatic ascent velocities (V_n) can then
439 be calculated using the decompression rate:

$$440 \quad V_n = \frac{1}{\rho g} \left| \frac{dP_w}{dz} \right| \quad (2)$$

441 where ρ is the density of vesiculated magma (estimated using porosity values from
442 ImageJ analysis and a 2700 kg m^{-3} magma density), g is gravity, and dP_w/dz is the
443 decompression rate at the nucleation depth of feldspar microlites.

444 KRA-233, the 1997 Strombolian ash, yields a relatively low mean ascent rate of
445 1.04 m s^{-1} . This is similar but slightly higher than that of U23-2, with a mean rate of
446 0.76 m s^{-1} . Early U10 samples show sharply elevated mean values of $\sim 2 \text{ m s}^{-1}$ for
447 U10-3, increasing to 4.45 m s^{-1} for U10-4 and then dropping to lower ascent rates for
448 later U10-samples ($0.99\text{-}1.84 \text{ m s}^{-1}$) (Fig.6).

449

450 **5. Discussion**

451 5.1. Timing of initial deposits

452 Bulk magma compositions at Anak Krakatau from 1993 to 2018 show negligible
453 changes (Gardner et al., 2013; Walter et al., 2019), implying no compositionally
454 different magma input associated with the 2018 lateral collapse. This corresponds
455 with observations that pre-collapse activity was typical of previous eruptions, despite
456 increased intensity (Perttu et al., 2020). Nonetheless, our textural and mineralogical
457 observations reveal that U23-2 is distinct from the U10 sequence, while early U10
458 deposits (U10-3 and U10-4) also differ slightly from overlying units. U23-2 contrasts
459 with U10 (Panjang post-collapse deposits) morphologically, in its predominance of
460 highly angular, dense clasts; it is also less primitive in terms of both glass and
461 mineral compositions, and its microlite crystallinity and estimated mean ascent rate
462 are consistent with relatively slow processes, broadly corresponding with pre-

463 collapse Strombolian conditions (i.e., KRA-233, the 1997 ash). The overall
464 homogeneity of U23-2 in its deposit- and grain-scale characteristics suggests that it
465 cannot be an upwind equivalent of U10, implying derivation from a discrete event.
466 We thus interpret U23-2 as a deposit from the initial explosive pulse that
467 accompanied the lateral collapse at ~20:55, involving the most intense mass
468 eruption rate in the collapse period (Gouhier and Paris, 2019), its plume reaching 16
469 km in altitude and transported SW (e.g. Prata et al., 2020). This 'syn-collapse'
470 explosion lasted ~40 minutes (Gouhier and Paris, 2019). The resulting 1-cm thick
471 deposit on Sertung indicates that although this plume was relatively ash-poor, it was
472 distinct from the ash-depleted, ice-dominated convective plumes sustained over the
473 following days (Prata et al., 2020). The presence of glass spherules, observed only
474 in U23-2 (cf. Genareau et al., 2015), is consistent with material derived from a
475 lightning-rich column (Fig.2f) (a notable feature of the high-level SW plume; Prata et
476 al., 2020).

477 The U10-2 tsunami deposit also provides an important time constraint for the
478 Panjang deposits. Tsunami simulations suggest that maximum wave inundation at
479 U10 occurred ~6.3 minutes after the collapse (S.2c). Our interpretation would
480 suggest that U23-2 and U10-2 were deposited broadly synchronously, and if U23-2
481 was derived from the high level syn-collapse plume it may have an equivalent
482 deposit on Panjang. On Panjang, such a deposit would be thin (given its upwind
483 position) and potentially mixed with or eroded by the U10-2 tsunami deposit. We
484 found no evidence of U23-2-type clasts within U10-2, and not did we find evidence of
485 a discrete U23-2-type layer beneath the basal purple ash at other sites on Panjang.
486 Any such layer would likely be very thin and difficult to distinguish from the

487 underlying ash-rich soil, and we are thus unable to provide any further constraints on
488 the relative timing and depositional pattern of these various units.

489

490 5.2. Unloading effects on microlite textures

491 Edifice destruction causes an instantaneous pressure drop in the underlying
492 magmatic system (Pinel and Jaupart, 2005), which may be manifested in microlite
493 textures that reveal decompression conditions during ascent (e.g., Preece et al.,
494 2013). Variable microlite textures can be attributed to changes in effective
495 undercooling (ΔT , defined as $T_{\text{liquidus}} - T_{\text{magma}}$; Crabtree and Lange, 2011) during
496 crystallisation. Low ΔT and slow decompression produces fewer and larger crystals,
497 developing under a crystal growth-dominated regime. In contrast, a nucleation-
498 dominated regime arises at higher ΔT and faster decompression, forming numerous
499 smaller microlites (e.g., Mollo and Hammer, 2017). Our samples can be divided into
500 three phases that span the collapse. Phase A is recorded by U23-2, representing a
501 crystal growth-dominated process, displaying the lowest areal microlite number
502 density (N_A), more tabular microlites, and the largest mean microlite size (Lofgren,
503 1980). The textural similarities of U23-2 and KRA-233 suggest both samples record
504 steady-state conditions characteristic of the Strombolian feeder system up until the
505 point of collapse (Fig. 7a), with relatively low ascent and decompression rates. U23-2
506 is thus essentially recording the imprint of pre-collapse ascent and crystallisation
507 conditions. The absence of textural disequilibrium in U23-2 suggests that, texturally,
508 it does not record a pressure perturbation associated with collapse, and that it
509 therefore represents magma within the surficial portion of the conduit. Its
510 fragmentation would then be a direct consequence of decompression, as a blast-like
511 response to edifice failure (e.g. Alidibirov and Dingwell, 1996). This notion of 'conduit

512 clearing' is consistent with the timing and short-lived nature of the initial explosive
513 pulse (Gouhier and Paris, 2019; Perttu et al., 2020; Prata et al., 2020), which is
514 indistinguishable from the timing of the collapse itself.

515 The U10-3 and U10-4 deposits represent Phase B, with a sharp increase in N_A
516 and crystallinity, smaller mean microlite sizes, higher aspect ratios, and various
517 disequilibrium morphologies, implying nucleation-dominated crystallisation under
518 high ΔT (Lofgren, 1980). The abundance of acicular, hopper and swallow-tail
519 morphologies in the U10 samples affirm high decompression rates and undercooling
520 (Hammer and Rutherford, 2002; Couch et al., 2003). Rapid degassing associated
521 with open-system decompression of the conduit following collapse may explain the
522 sharp change in crystal development, as T_{liquidus} (and ΔT) would have increased
523 through both decompression and water loss (e.g., Cashman and Blundy, 2000).
524 Destabilisation of a pressurised conduit during the collapse, further unloaded by the
525 initial explosive pulse of U23-2, may have facilitated the rapid acceleration of deeper
526 magma, evident in the increasing ascent rates (and increasing evidence of
527 disequilibrium microlite textures) estimated for U10-3 and U10-4. Rapid,
528 decompression-induced crystallisation of early U10 samples is further supported by
529 high pyroxene and titanomagnetite microlite proportions (Fig.4c and d), as
530 decompression experiments indicate pyroxene and oxide content increases under
531 higher decompression (Szramek et al., 2006).

532 Finally, Phase C is represented by later U10 samples (U10-6, U10-8 and U10-10),
533 which indicate a general decrease in N_A and crystallinity, and an increase in mean
534 microlite size, consistent with a gradual re-stabilisation of the feeder system with
535 reduced magma ascent velocities.

536

537 5.3. Effects of collapse on mineral chemistry

538 Mineral compositions are sensitive to changes in temperature, pressure, and
539 volatile content. Differences in plagioclase anorthite-content between U23-2 and U10
540 support the above textural interpretation, with U10 groundmass compositions (Phase
541 B and C) extending to notably higher values (An_{67-79}) than U23-2 (Phase A), with a
542 discrete calcic population indicative of relatively higher water pressures and
543 temperatures (Fig.5a) (Couch et al., 2003). This could correspond with derivation as
544 entrained crystals from deeper in the plumbing system, since the main microlite
545 population in all samples spans a comparable compositional range (An_{50-63})
546 consistent with crystallisation extending to shallow levels. Plagioclase phenocryst
547 rims exhibit increasing An-content higher up the U10 stratigraphy (Fig.5a),
548 suggesting that most plagioclase crystals equilibrated under differing physio-
549 chemical conditions in each sample and/or were in disequilibrium with the
550 compositionally homogenous melt (S.3) (Petrone et al., 2009). Last-equilibrated
551 depths can be estimated from EPMA data using plagioclase-melt and
552 orthopyroxene-melt barometers (Putirka, 2008) (Supplementary Table 4).
553 Considering the K_D crystal-melt equilibrium criterion (Putirka, 2008; for both
554 plagioclase and orthopyroxene, using average matrix glass compositions), many
555 phenocryst rim compositions fail to meet the threshold. However, from the limited
556 usable dataset ($n=17$), calculated depths (using densities $2320-2800 \text{ kg m}^{-3}$; Kopp et
557 al., 2001) are consistent with previous upper crustal estimates (e.g. Dahren et al.,
558 2012), indicating two main levels at $\sim 4-7 \text{ km}$ (plagioclase) and $8-12 \text{ km}$ (plagioclase
559 and orthopyroxene). Plagioclase phenocryst core compositions are relatively similar

560 across U23-2 and U10, but more anorthitic than those in the 1997 magma (KRA-
561 233). This implies that all the magma erupted in 2018 can be considered a related
562 batch, even if microlite compositions record an overprint of variable ascent
563 conditions. This is also supported by olivine and pyroxene phenocryst core
564 compositions, based on a more limited dataset (Fig. 5b, c, d).

565

566 5.4. Fragmentation and magma-water interaction

567 Ash morphological and surface characteristics in the U23-2 and U10 samples
568 indicate variation in fragmentation modes caused primarily by two brittle
569 mechanisms. The vitric nature, fracture patterns and high angularity of U23-2 relative
570 to U10 (Fig. 2d and e) correspond with the sudden effect of rapid decompression;
571 angularity resulting from a top-down propagating decompression wave, producing
572 brittle fracture and relatively denser textures (Alidibirov and Dingwell, 1996). In U10,
573 blocky and sub-angular morphologies, alongside stepped or river-line fractures on
574 grain surfaces, also signify a dominant brittle fragmentation process. However, the
575 overall increase in vesicularity relative to U23-2 (Fig.3b and c) suggests this was
576 driven by vesicle overpressure, facilitated by rapid ascent and extensive degassing
577 (Cioni et al., 2014; Gouhier and Paris, 2019). Minor ductile features (i.e. Pele's hairs)
578 also emphasise the melt's relatively low viscosity, exiting the vent at high velocity
579 (Büttner et al., 2002).

580 Visual observations and infrasound signals indicate the onset of Surtseyan activity
581 after the collapse (Perttu et al., 2020). Interaction with seawater, leading to further
582 brittle and ductile fragmentation, may have overprinted the primary brittle
583 mechanisms (cf. Liu et al., 2017), although we would expect this to be more limited

584 in U23-2. The high lithic and altered grain content in U10-3 (and NP2-2; S.2)
585 suggests a vent widening stage that enabled greater magma-water interaction,
586 following the unloading-induced explosion of U23-2. Finer tails in the U10 grain-size
587 distributions compared to U23-2 (Fig.1a) may reflect a greater fragmentation
588 efficiency following this vent reconfiguration. Furthermore, the collapse not only
589 exposed the vent to seawater, but also uncovered the subsurface hydrothermal
590 system beneath the SW flank, indicated by the orange seawater plumes evident
591 days after the collapse (S.1b). Excavation of the hydrothermally altered edifice by
592 Surtseyan activity, during the vent widening stage, is apparent from the distinct
593 colouration and chemistry of U10-3.

594

595 5.5. Reconstruction of syn- and post-collapse eruptive activity

596 Immediately after Anak Krakatau's collapse at ~20:55, decompression initiated
597 brittle fragmentation of magma in the shallow conduit, which had been feeding
598 Strombolian eruptions and had ascended under conditions characteristic of the
599 preceding months (Fig.7a). This elicited a highly explosive and short-lived eruption
600 (Phase A), with rapid plume ascent reaching ~16 km (Gouhier and Paris, 2019;
601 Prata et al., 2020) (Fig.7b). Phase A marked the onset of elevated open-system
602 degassing, with SO₂ output from 22 to 28 December estimated at ~98 kt (Gouhier
603 and Paris, 2019).

604 During Phase B, extensive degassing coupled with rapid decompression and
605 tapping of the deeper feeder system, shifted the crystallisation regime towards rapid
606 nucleation of smaller microlites. U10-3 defines the onset of extensive seawater
607 interaction, with a vent widening stage leading to typical Surtseyan activity,
608 producing cock's tail jets and ash-laden low-altitude plumes (Fig.7c) (Prata et al.,

609 2020). Increasing magma ascent rates and vesicularity suggest primary brittle
610 fragmentation driven by vesicle overpressure.

611 In Phase C, magma ascent conditions gradually stabilised, in response to
612 pressure conditions that had been modified both by the collapse and then by the
613 rapid island regrowth that accompanied the post-collapse eruptions (Fig.7d). Magma
614 ascent rates exhibit progressively slower velocities, and degassing reduced by
615 nearly 50% in late December and by 75% in early January 2019 (Gouhier and Paris,
616 2019). Given the very rapid regrowth (cf. Hunt et al., in press), we suggest that the
617 volcano returned to equilibrium conditions (i.e., potentially comparable with the pre-
618 collapse storage state) within a 1-2 week period as activity waned.

619

620 5.6. Implications for determining future collapse events and collapse impacts

621 As there is no evidence for elevated/unusual magma ascent patterns immediately
622 preceding the collapse, it is unlikely that any distinctive magmatic (i.e., volcano
623 seismicity, inflation or degassing) signatures would have been apparent that could
624 have indicated incipient collapse. However, progressive susceptibility of the SW flank
625 to failure was evident from longer-term deformation and growth patterns. Slow
626 lateral deformation of the failed SW flank was identified over 10 years before the
627 collapse (Agustan et al., 2012; Chaussard and Amelung, 2012), with evidence of an
628 increase in the 2018 eruptive period (Walter et al., 2019). Since longer-term
629 deformation datasets are limited, it is difficult to assess whether deformation rates in
630 2018 were substantially different to those during previous eruptions, or whether
631 displacements showed an overall accelerating pattern. Nonetheless, longer-term
632 patterns of Anak Krakatau's edifice growth reflect significant structural instabilities,
633 including the volcano's position on the scarp of the 1883 Krakatau caldera (Deplus et

634 al., 1995); pre-1960 asymmetrical growth of the tuff cone towards the NE, and post-
635 1960 vent migration facilitating SW emplacement of lava deltas (cf. Hunt et al., in
636 press). While eruptions do not always trigger edifice instability, the June-December
637 2018 activity itself is also likely to have played a role in the timing of the collapse by
638 increasing flank loading and potentially increasing pressurisation in the subsurface
639 hydrothermal system (e.g., Reid, 2004). All these instabilities, combined with the
640 2018 eruptive activity, ultimately pre-conditioned the SW flank for its eventual
641 collapse. For future monitoring of edifice stability at Anak Krakatau or elsewhere, an
642 approach integrating short- and long-term edifice growth patterns with flank
643 deformation monitoring (cf. Gonzalez-Santana and Wauthier, 2020), and an
644 improved understanding of edifice material properties, may hold the best prospects
645 for refining forecasts of collapse timing.

646 The Anak Krakatau collapse also reveals the impact of surface-unloading driven
647 disruption on a shallow magmatic system. Although the collapse volume was small
648 relative to historical collapses (e.g., Ritter Island; Karstens et al., 2019), sudden
649 decompression led to a considerable magmatic response and readjustment, causing
650 highly elevated eruption rates. This elevated magma flux rapidly infilled the landslide
651 scar and extended the island coastline (e.g., Novellino et al., 2020; Hunt et al., in
652 press). This highlights that temporarily high post-collapse eruption rates may hinder
653 opportunities to identify and investigate collapse processes by concealing failure
654 scars and stratigraphic records of collapse-associated activity (e.g., Watt, 2019).

655

656 **6. Conclusions**

657 Our physical, microtextural and geochemical analysis of syn- and post-collapse
658 deposits shows no evidence that a change in magmatic conditions or eruptive

659 behaviour preceded the lateral collapse of Anak Krakatau. Instead, the intense,
660 accompanying volcanism is interpreted as a response to collapse and can be divided
661 into three main phases. Phase A involved a syn-collapse eruption triggered by
662 collapse-driven decompression of the shallow conduit, generating a powerful
663 explosive pulse and depositing ash to the SW. Textures in this ash record pre-
664 collapse ascent conditions, excluding a direct magmatic trigger for the collapse, and
665 suggesting the collapse resulted from structural and gravitational instabilities arising
666 from patterns in edifice development.

667 Post-collapse eruptive activity in Phase B reflects deeper tapping of successive
668 magma batches from the depressurised conduit, with extensive degassing and
669 accelerating ascent rates. Gradual re-stabilisation of conduit conditions occurs in
670 Phase C, as rapid edifice regrowth led to waning activity.

671 The 2018 collapse highlights that lateral collapses are not necessarily directly
672 triggered by immediate shifts in magmatic behaviour, which standard monitoring
673 techniques could have potentially detected. Therefore, effective volcanic monitoring
674 and forecasting of such collapse events may need to focus on identifying areas with
675 increased susceptibility to failure, as signalled by changing edifice growth patterns
676 and flank deformation; this will be particularly relevant for the future growth of Anak
677 Krakatau.

678

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690

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843

844 **Figure captions**

845 **Figure 1**

846 U10 and U23-2 stratigraphic sections and grain size data. **a** The physical
847 characteristics of the U10 sequence on northern Panjang (image, sedimentary log
848 and grain size distributions (sieved and laser diffraction data)). **b** Inset map showing
849 location of Anak Krakatau and sample localities (main sites of U23-2 and U10 =
850 yellow circles; additional sites of NP1 and NP2 = dark blue pentagons). **c** The

851 physical characteristics of U23-2 from southern Sertung (sedimentary log and grain
852 size distribution (sieved and laser diffraction data)). White triangles mark samples
853 analysed texturally and black diamonds mark samples analysed geochemically (XRF
854 and/or EPMA).

855

856 **Figure 2**

857 Component analysis and optical/secondary-electron SEM images. **a** Quantitative
858 componentry of a 1997 sample (KRA-233) and the December 2018 samples (U23-2
859 and U10) using the 500 μm -1 mm size fraction, with the samples listed in
860 chronological order. **b** KRA-233 deposit, showing mostly black juvenile clasts.
861 **c** U10-2 tsunami deposit, dominated by 1883 Krakatau pumice clasts. **d** U23-2 ash
862 deposit. Note the very high grain angularity and glassy nature of the grains. **e** U10-
863 3B deposit, displaying more sub-angular grain morphologies and altered grains. **f**
864 Secondary-electron (SE) SEM images of U23-2 ash grains with brittle or ductile
865 fragmentation features. Left image displays a Pele's hair and a moss-like grain. Right
866 image highlights a glass spherule. **g** SE SEM images of U10-6 ash grains with brittle
867 or ductile fragmentation features. Left image shows a grain with stepped fractures.
868 Right image displays a grain with a molten surface.

869

870 **Figure 3**

871 Graphs showing shape analysis of ash particles from U23-2 (**a**), U10-3B (**b**) and
872 U10-4 (**c**) at three different grain size fractions (63-125 μm , 125-180 μm & 180-250
873 μm). Concavity index (CI) is plotted against axial ratio, with dashed line (CI=0.4)
874 marking the threshold between dense and bubbly grains, after Liu et al. (2015a).

875 Binary images of examples of dense and bubbly grains (vesicular and glass shards)
876 from each deposit are labelled on each diagram.

877

878 **Figure 4**

879 Backscattered-electron (BSE) SEM images of vesicular scoria with plagioclase
880 feldspar microlite textures. **a** Tabular microlites in KRA-233. **b** Tabular and swallow-
881 tail microlites in U23-2. **c** Tabular and acicular microlites in U10-3B. **d** Tabular,
882 hopper and acicular microlites in U10-4. **e** Acicular microlites with prominent Fe-Ti
883 rich compositional boundary layers in U10-6. **f** Acicular, tabular, hopper and swallow-
884 tail microlites in U10-10.

885

886 **Figure 5**

887 EPMA chemical compositions of plagioclase, olivine, and pyroxene shown for KRA-
888 233, U23-2 and U10. **a** Anorthite contents in microlites, microphenocrysts and
889 phenocrysts of plagioclase. **b** Forsterite contents in microlites, microphenocrysts and
890 phenocrysts of olivine. **c** and **d** Mg# for orthopyroxene (**c**) and clinopyroxene (**d**) in
891 microlites, microphenocrysts and phenocrysts. Kernel density estimates are plotted
892 on top of each diagram illustrating main distributions across all samples (and
893 distinguishing KRA-233 from the 2018 samples in plagioclase).

894

895 **Figure 6**

896 Temporal variations of N_A (**a**), crystallinity (**b**), mean microlite size (**c**), aspect ratio
897 (**d**) and mean ascent rate (**e**) (grey bars represent ascent rate ranges using range of
898 water content values estimated from plagioclase-melt hygrometry of Putirka (2008))

899 for plagioclase feldspar microlites in vesicular microlite-rich scoriae from KRA-233
900 (1997) and December 2018 samples (U23-2 and U10).

901

902 **Figure 7**

903 Conceptual model for syn- and post-collapse volcanic processes at Anak Krakatau
904 from 22 Dec. 2018 to early Jan. 2019. **a** Pre-collapse: open, steady-state conditions,
905 typical of the Strombolian feeder system reflected in low ascent velocities and
906 microlite textures of KRA-233 and U23-2. P_{c1} represents an average pre-collapse
907 conduit pressure. **b** Phase A: Lateral collapse and unloading causing downward
908 propagating decompression and an intense explosion from unloading of the surficial
909 conduit; limited seawater interaction, and deposition of U23-2. **c** Phase B:
910 Destabilisation, decompression and deeper tapping of the conduit facilitating fast
911 ascent of U10-3 (Surtseyan vent widening) and U10-4 (sustained Surtseyan activity).
912 $P_{c2} < P_{c1}$ represents Phase B with conditions of a highly depressurised conduit (P_{c2})
913 relative to pre-collapse conduit pressure (P_{c1}). **d** Phase C: System gradually re-
914 stabilised with ascent characterised by lower velocities and decompression rates
915 (U10-6 to U10-10). $P_{c3} > P_{c2}$ represents Phase C conduit pressure re-stabilising
916 following Phase B eruptions and partial edifice regrowth.

917

918 **Table captions**

919 **Table 1**

920 Textural characteristics of plagioclase microlites in vesicular microlite-rich scoriae
921 from Anak Krakatau, with estimated ascent rates (KRA-233: 1997 Strombolian
922 sample; U23-2 and U10- : 2018 samples; n denotes the number of crystals
923 analysed).

924 **Supplementary Material**

925 Supplementary Table 1. Sample list with location and analyses.

926 Supplementary Table 2. Summary of clast component types (evaluated on 500 µm
927 to 1 mm sieved size fraction).

928 Supplementary Table 3. Compositional analyses (XRF & EPMA) for Anak Krakatau
929 2018 and 1997 tephra samples.

930 Supplementary Table 4. Temperatures and pressures calculated using
931 thermobarometers of plagioclase and orthopyroxene (Putirka, 2008).

932

933 Appendix 1. Analytical methods.

934

935 Supplementary Figures (S.1 to S.5)

936 S.1 Sentinel 2 L1C satellite images of the Krakatau archipelago in infrared based on
937 bands 8, 4 & 3. Dense vegetation is highlighted in red and tephra deposition is
938 shown in dark grey. **a** Pre-collapse image showing extensive fresh ash
939 deposition/lava emplacement on Anak Krakatau and potential minor deposition on
940 Panjang taken on 16 Nov. 2018. **b** Post-collapse image of significant ash deposition
941 and vegetation loss on Anak Krakatau and Panjang (image taken on 10 Jan. 2019).
942 White arrow highlights plume of reddish-orange (infrared = turquoise) water
943 emanating from uncovered hydrothermal system off the SW island coastline.

944

945 S.2 **a** Dec. 2018 stratigraphies on Panjang (P) and Sertung (S). Black background
946 shading indicates the only well-correlated layer (U10-3 and NP2-1) throughout the
947 sections based on lithological characteristics (i.e., 0.5 cm-thick layer of purple
948 indurated ash; see right-hand side image in **b**). Colours and grain-size are based on
949 field descriptions. White triangles mark samples analysed texturally and black
950 diamonds mark samples analysed geochemically (XRF and/or EPMA). **b** Left image:
951 close-up of U10 sequence. Right image: close-up of layer U10-3 showing the yellow
952 oxidised crust at the surface of a purple indurated ash immediately overlying the
953 pumiceous tsunami deposit (U10-2). **c** Tsunami simulation at ~380s displaying

954 maximum envelope for tsunami wave inundation at site U10 (~7 m wave height).
955 Simulation used the 2D FUNWAVE model at 50 m resolution (Grilli et al. 2019).

956

957 S.3 Major element compositions of matrix glasses (**a**, **b**, **c**, **d**, and **e**). **f** Whole-rock
958 total alkali vs. silica (TAS) diagram of 2018 products (U23-2, U10, NP1), plotted
959 alongside previously published data of other recent eruptive products (Walter et al.
960 2019; *includes data from Gardner et al. 2013).

961

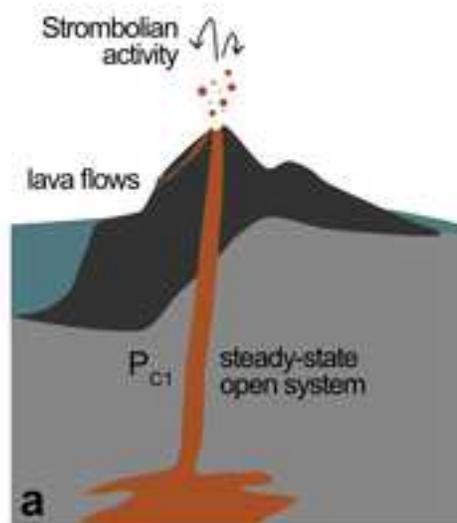
962 S.4 Examples of manually outlined plagioclase feldspar microlites (black) used for
963 the 2D textural analysis.

964

965 S.5 Batch microlite textural parameters for KRA-233, U23-2 and U10 samples. **a**
966 Areal feldspar microlite number density ($NA \text{ mm}^{-2}$) vs. groundmass feldspar microlite
967 crystallinity (φ). **b** Areal feldspar microlite number density vs. mean microlite size
968 ($S_n, \mu\text{m}$). **c** Areal feldspar microlite number density vs. aspect ratio.

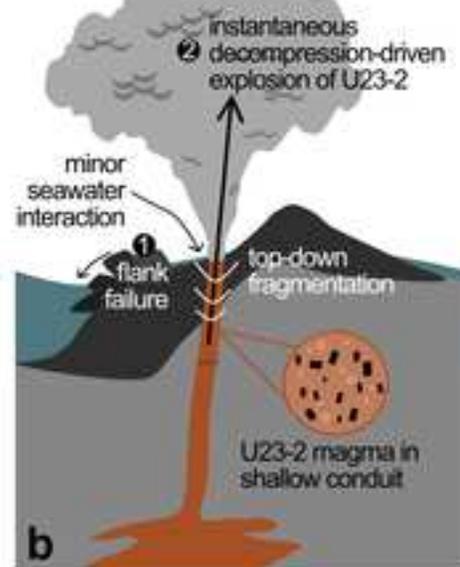
PRE-COLLAPSE
22 Dec. 18

SW



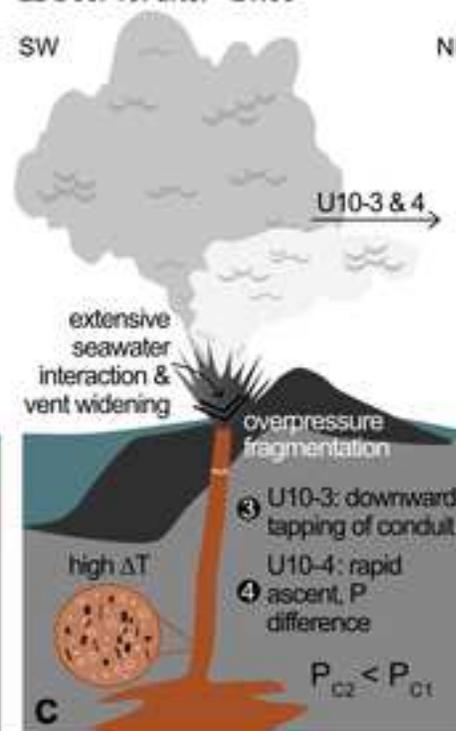
SYN-COLLAPSE: PHASE A
22 Dec. 18: ~20:55

NE SW



POST-COLLAPSE: PHASE B
22 Dec. 18: after ~21:35

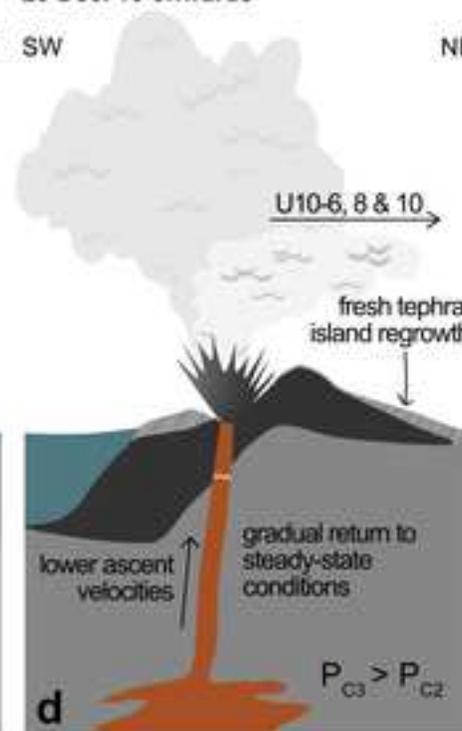
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POST-COLLAPSE: PHASE C
23 Dec. 18 onwards

NE SW

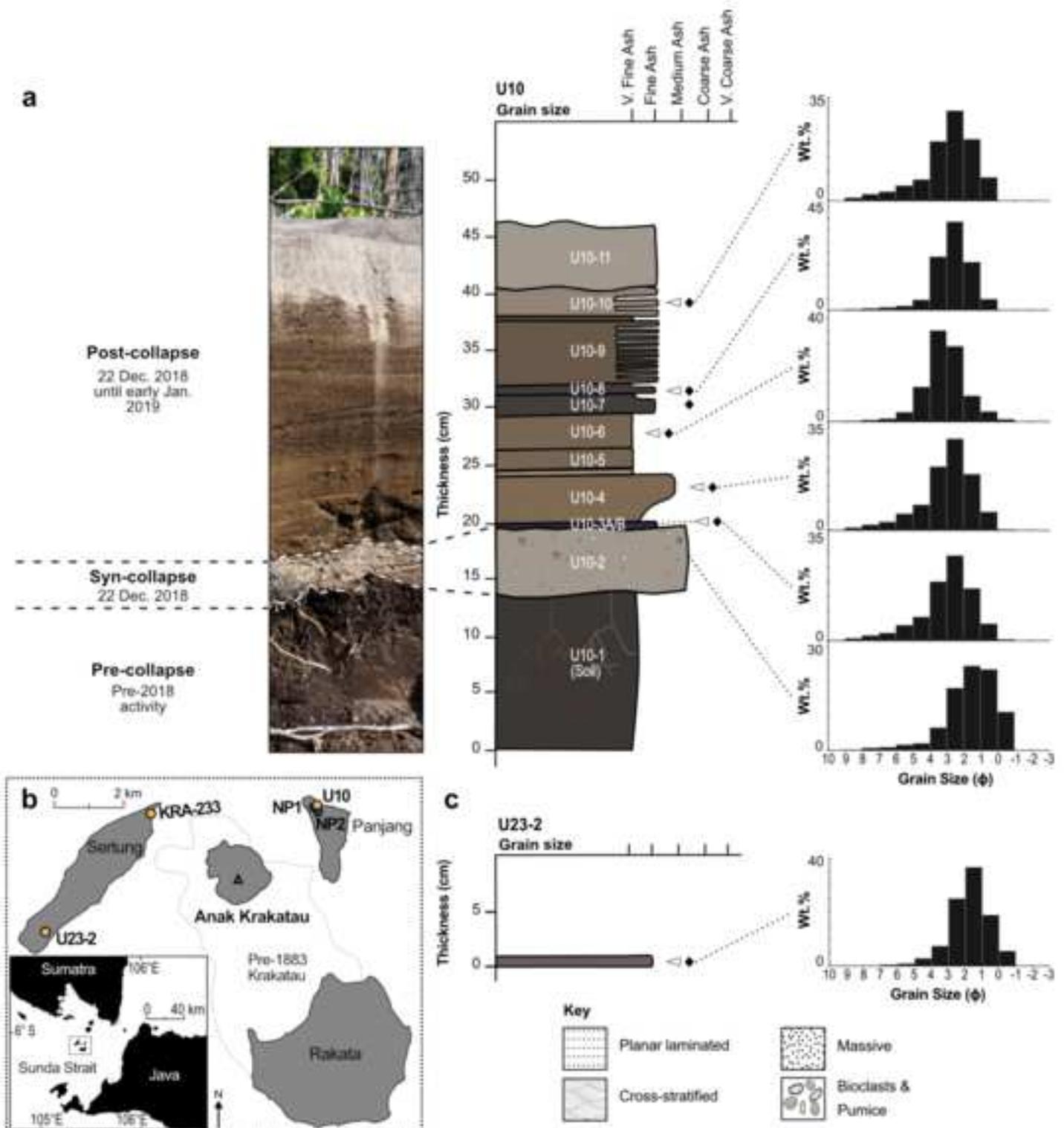
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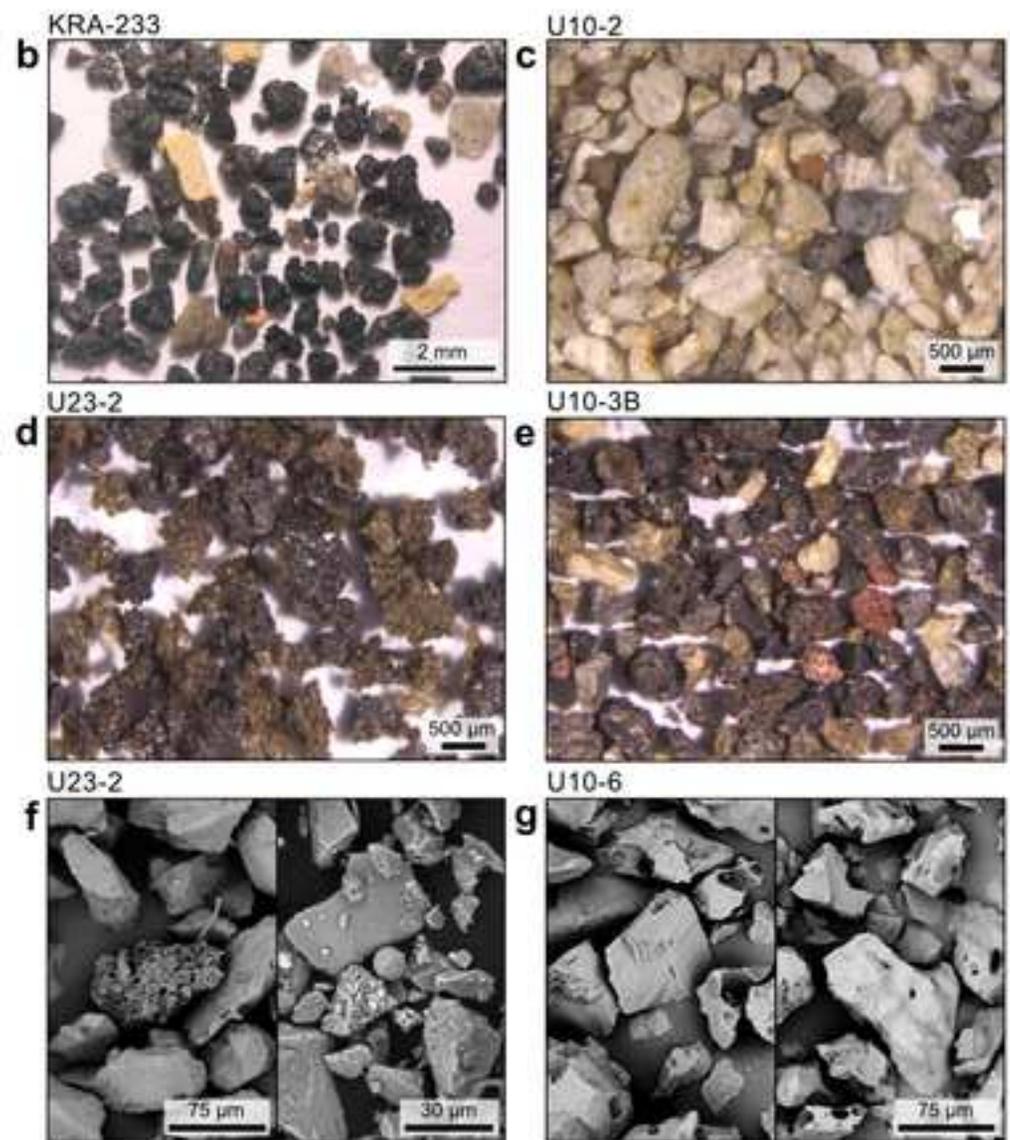
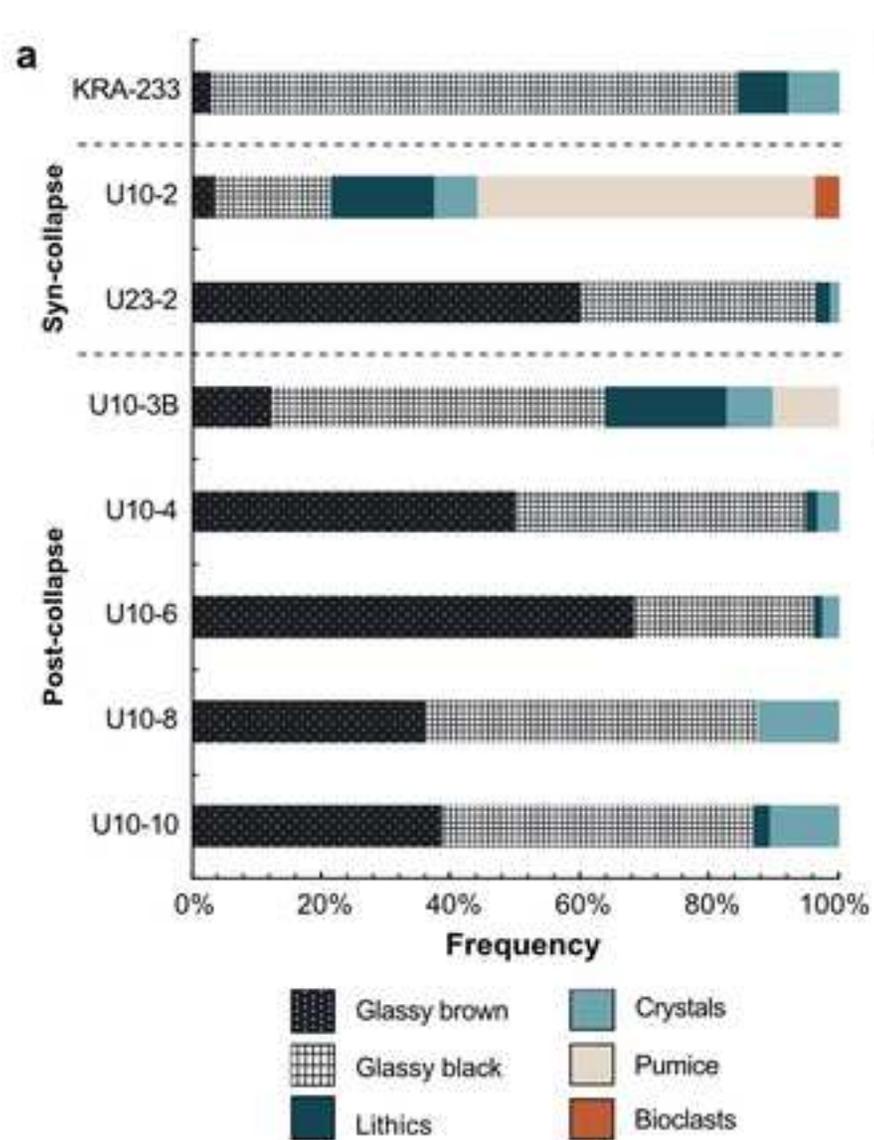


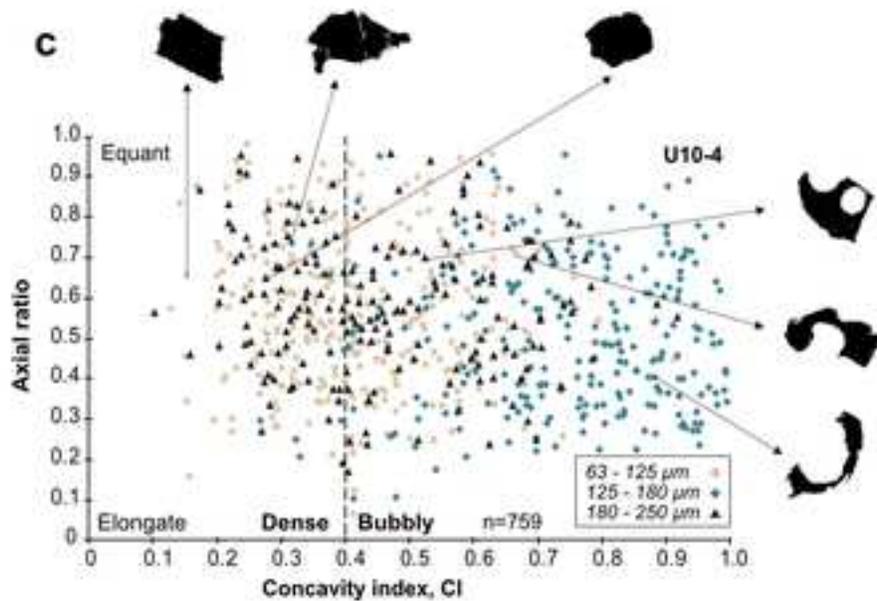
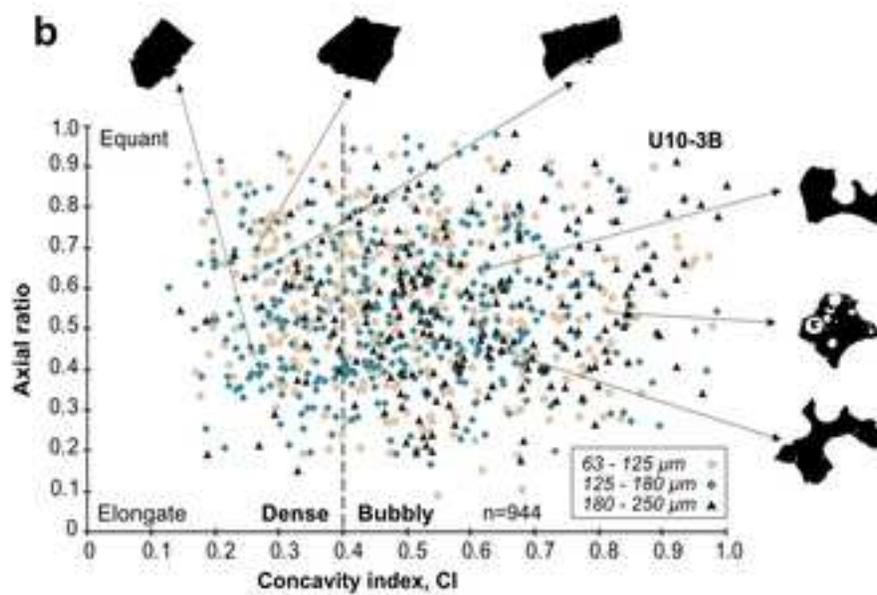
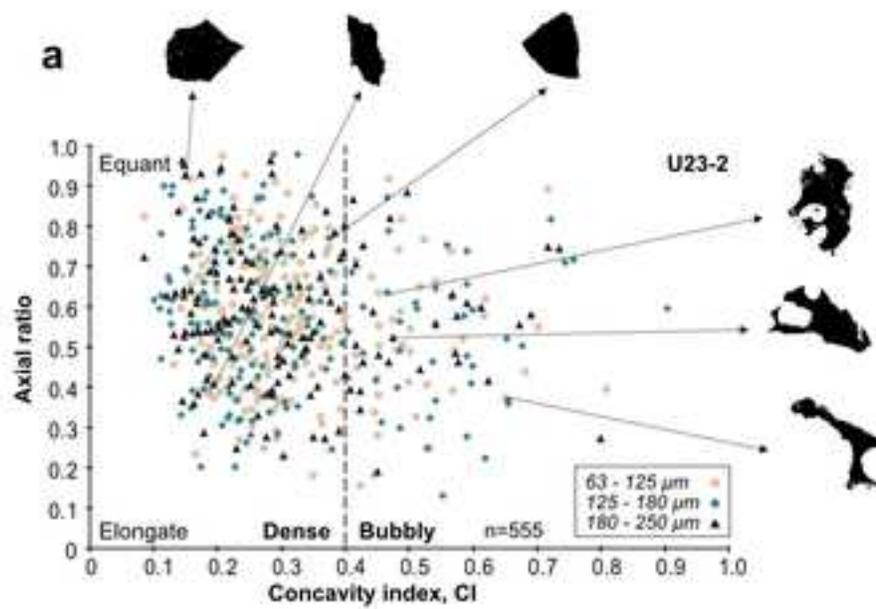
P_c = relative conduit pressure

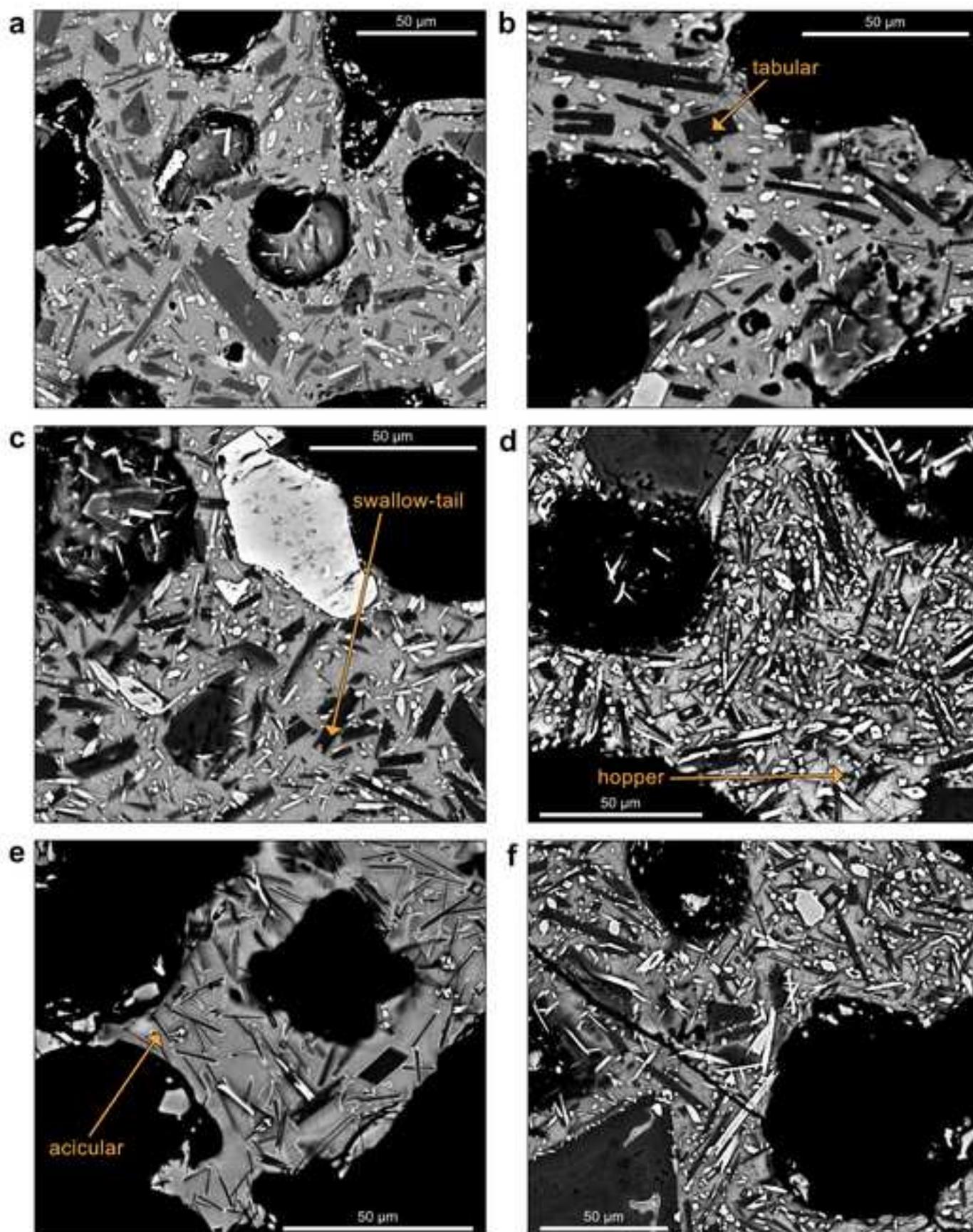
ΔT = undercooling

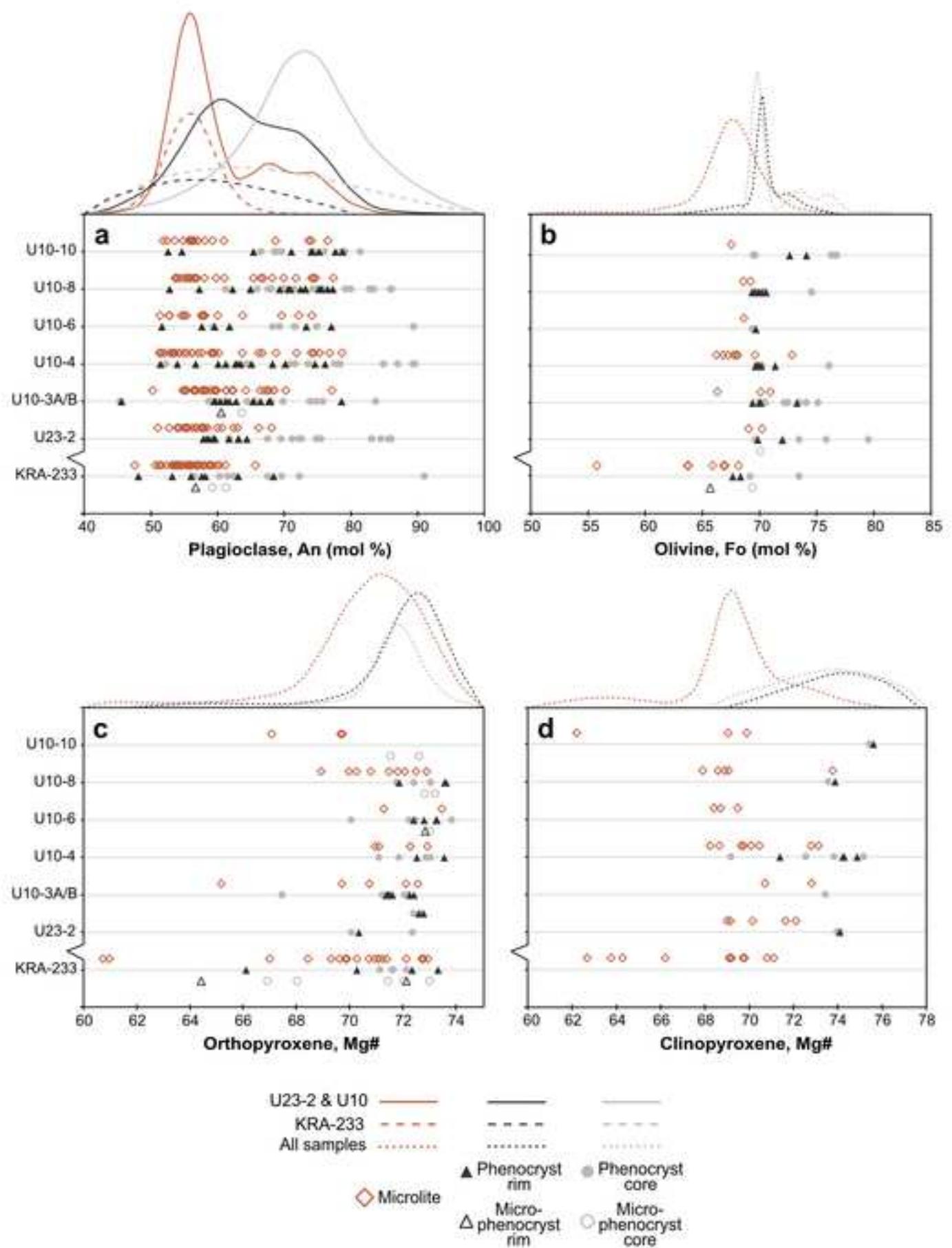
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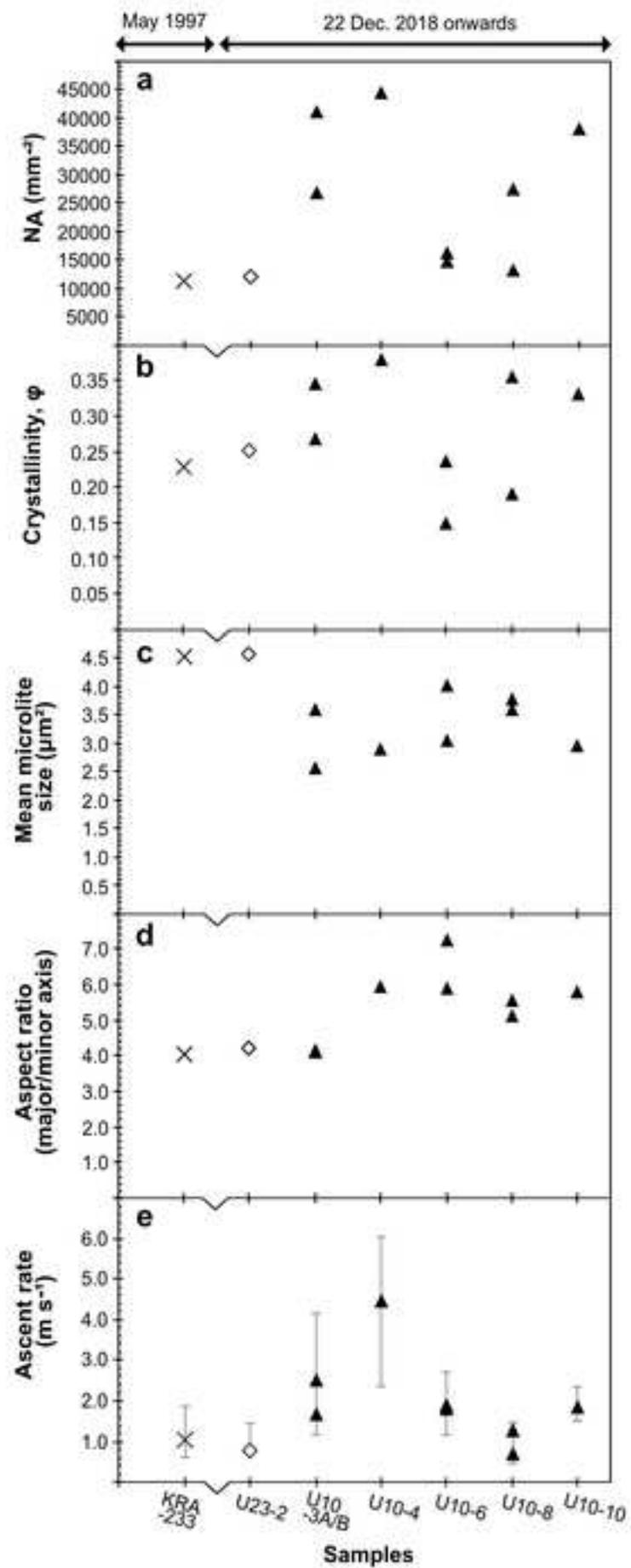






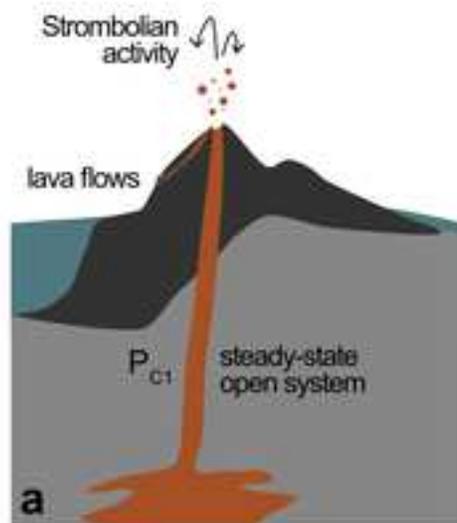






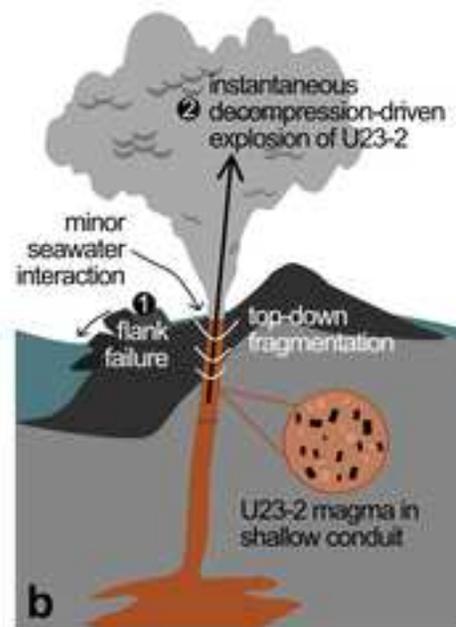
PRE-COLLAPSE
22 Dec. 18

SW



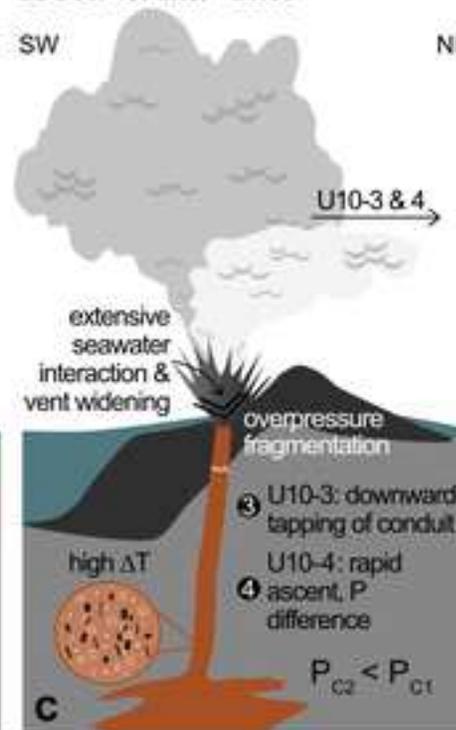
SYN-COLLAPSE: PHASE A
22 Dec. 18: ~20:55

NE SW



POST-COLLAPSE: PHASE B
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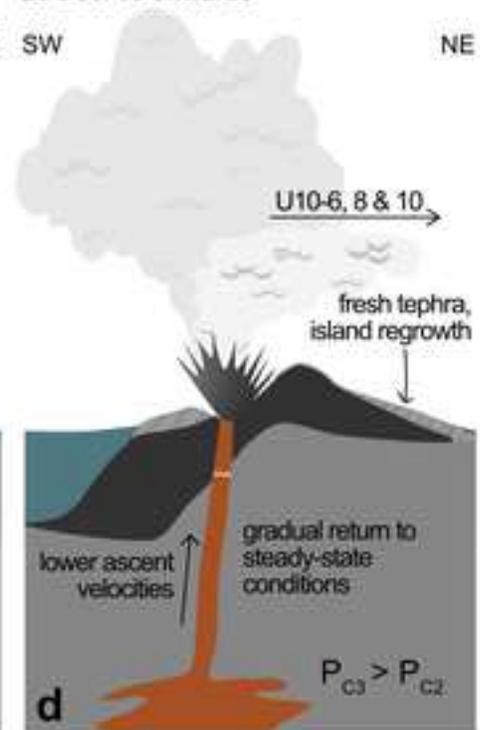
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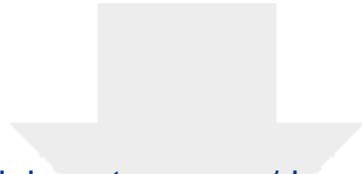


POST-COLLAPSE: PHASE C
23 Dec. 18 onwards

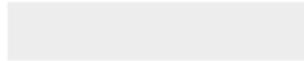
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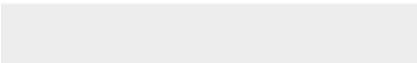
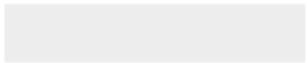


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FIG1_FINALhighres.tiff



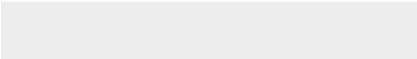
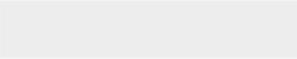


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FIG2_FINAL_highres.tiff



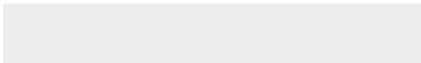


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FIG4_FINAL_highres.tiff





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FIG7_FINAL_highres.tiff



Sample	2D areal measurements							3D volumetric measurements	
	n	Image area (μm^2)	N_A (mm^{-2})	Plag. area %	Plag. area fraction, ϕ	Mean crystal size, S_n (μm^2)	Aspect ratio (major/minor axis)	N_V (mm^{-3})	Mean ascent rate (m s^{-1})
KRA-233	403	35561	11194	16.28	0.23	4.54	4.08	2466412	1.04
U23-2	784	65328	12062	14.59	0.25	4.56	4.19	2644251	0.76
U10-3A	914	33612	26882	24.10	0.35	3.58	4.14	7501290	1.69
U10-3B	1110	27164	41111	22.93	0.27	2.56	4.09	16074394	2.49
U10-4	980	22475	44545	23.40	0.38	2.92	5.93	15278638	4.45
U10-6_1	716	43873	16273	10.43	0.15	3.04	7.28	5349256	1.88
U10-6_2	774	53431	14604	13.56	0.24	4.03	5.91	3619626	1.81
U10-8_1	577	21186	27476	25.63	0.36	3.60	5.14	7631114	1.26
U10-8_2	507	37593	13342	16.17	0.19	3.78	5.55	3527747	0.71
U10-10	1215	32271	37969	21.56	0.33	2.96	5.79	12843341	1.83



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**Supplementary material for review only (e.g., accepted
"in press" reference files)**
supplementarymaterials_FINALV4.pdf

Declaration of interests

The authors declare that they have no known competing financial interests or personal relationships that could have appeared to influence the work reported in this paper.

The authors declare the following financial interests/personal relationships which may be considered as potential competing interests:

CRediT authorship contribution statement

KSC led the analysis, and with SFLW developed the approach and methodology, interpretation and writing, with contributions from MC, ALMN and SLE. All authors contributed to the final interpretation and text. Initial project development was led by SFLW, MC, SLE, with input throughout the data collection and analysis stage from ALMN, MA, MEMN, DRT, SNC, AN, JH, SD, CH, IAK and NK.