**BRITISH GEOLOGICAL SURVEY** 

# Eskdalemuir Observatory

# Monthly Magnetic Bulletin

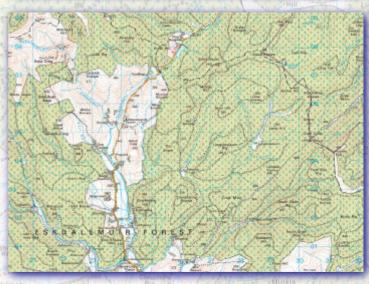
October 2016

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#### ESKDALEMUIR OBSERVATORY MAGNETIC DATA

#### 1. Introduction

Eskdalemuir observatory is one of three geomagnetic observatories in the UK operated and maintained by the British Geological Survey (BGS).

This bulletin is published to provide rapid access to the provisional geomagnetic observatory results. The information is freely available for personal, academic, educational and non-commercial research or use. Magnetic observatory data are presented as a series of plots of one-minute, hourly and daily values, followed by tabulations of monthly values, reports of rapid variations and geomagnetic activity indices. The operation of the observatory and presentation of data are described in the rest of this section.

Enquiries about the data should be addressed to:

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Internet: www.geomag.bgs.ac.uk

#### 2. Position

The observatory is situated on a rising shoulder of open moorland in the upper part of the valley of the White Esk River in the Southern Uplands of Scotland. The observatory co-ordinates are:

Geographic: 55° 18' 50.4"N 356° 47' 38.4"E Geomagnetic: 57° 33' 07"N 083° 44' 06"E Height above mean sea level: 245 m

The geographical coordinates are measured by a handheld GPS device, which uses WGS84 as the reference coordinate system. The height above MSL is determined from the best available contour The geomagnetic co-ordinates maps. approximations, calculated using the 12th generation International Geomagnetic Reference Field (IGRF) at epoch 2016.5. On-line access to models (including IGRF), charts and navigational data are available at

www.geomag.bgs.ac.uk/data\_service/models\_compass/home

#### 3. The Observatory Operation

#### **3.1 GDAS**

The observatory operates under the control of the Geomagnetic Data Acquisition System (GDAS), which was developed by BGS staff, installed in 2002, and became fully operational in January 2003. The data acquisition software, running on QNX operated computers, controls the data logging and the communications.

There are two sets of sensors used for making magnetic measurements. A tri-axial linear-core fluxgate magnetometer, manufactured by the Danish Meteorological Institute, is used to measure the variations in the horizontal (H) and vertical (Z) components of the field. The third sensor is oriented perpendicular to these, and measures variations, which are proportional to the changes in declination (D). Measurements are made at a rate of 1 Hz.

In addition to the fluxgate sensors there is a proton precession magnetometer (PPM) making measurements of the absolute total field intensity (F) at a rate of 0.1Hz.

The raw unfiltered data are retrieved automatically via Internet connections to the BGS office in Edinburgh in near real-time. The fluxgate data are filtered to produce one-minute values using a 61-point cosine filter and the total field intensity samples are filtered using a 7-point cosine filter. The one-minute values provide input for various data products, available on-line at

www.geomag.bgs.ac.uk/data\_service/home

#### 3.2 Back-up Systems

There are two other fully independent identical systems, GDAS 2 and GDAS 3, operating at the observatory. The data from these are also processed in near real-time and used for quality control purposes. They are also used to fill any gaps or replace any corrupt values in the primary system, GDAS 1.

#### 3.3 Absolute Observations

The GDAS fluxgate magnetometers accurately measure variations in the components of the geomagnetic field, but not the absolute magnitudes. Two sets of absolute measurements of the field are made manually once per week. A fluxgate sensor mounted on a theodolite is used to determine *D* and inclination (*I*); the GDAS PPM measurements,

with a site difference correction applied, are used for *F*. The absolute observations are used in conjunction with the GDAS variometer measurements to produce a continuous record of the absolute values of the geomagnetic field elements as if they had been measured at the observatory reference pillar.

#### 4. Observatory Results

The data presented in the bulletin are in the form of plots and tabulations described in the following sections.

#### 4.1 Absolute Observations

The absolute observation measurements made during the month are tabulated. Also included are the corresponding baseline values, which are the differences between the absolute measurements and the variometer measurements of D, H and Z (in the sense absolute—variometer). These are also plotted (markers) along with the derived preliminary daily baseline values (line) throughout the year. Daily mean differences between the measured absolute F and the F computed from the baseline corrected H and Z values are plotted in the fourth panel (in the sense measured—derived). The bottom panel shows the daily mean temperature in the fluxgate chamber.

#### 4.2 Summary magnetograms

Small-scale magnetograms are plotted which allow the month's data to be viewed at a glance. They are plotted 16 days to a page and show the one-minute variations in D, H and Z. The scales are shown on the right-hand side of the page. On disturbed days the scales are multiplied by a factor, which is indicated above the panel for that day. The variations are centred on the monthly mean value, shown on the left side of the page.

#### 4.3 Magnetograms

The daily magnetograms are plotted using oneminute values of D, H and Z from the fluxgate sensors, with any gaps filled using back-up data. The magnetograms are plotted to a variable scale; scale bars are shown to the right of each plot. The absolute level (the monthly mean value) is indicated on the left side of the plots.

#### **4.4 Hourly Mean Value Plots**

Hourly mean values of *D*, *H* and *Z* for the past 12 months are plotted in 27-day segments corresponding to the Bartels solar rotation number. Magnetic disturbances associated with active regions and/or coronal holes on the Sun may recur after 27 days: the same is true for geomagnetically quiet intervals. Plotting the data in this way

highlights this recurrence. Diurnal variations are also clear in these plots and the amplitude changes throughout the year highlight the seasonal changes. Longer term secular variation is also illustrated.

Full lists of the UK observatory hourly mean values from 1983 to the present day are available at <a href="https://www.geomag.bgs.ac.uk/data\_service/data/obs\_data/hourly\_means">www.geomag.bgs.ac.uk/data\_service/data/obs\_data/hourly\_means</a>

#### 4.5 Daily and Monthly Mean Values

Daily mean values of D, H, Z and F are plotted throughout the year. In addition, a table of monthly mean values of all the geomagnetic elements is provided. These values depend on accurate specification of the fluxgate sensor baselines. It is anticipated that these provisional values will not be altered by more than a few nT or tenths of arcminutes before being made definitive at the end of the year.

#### 4.6 Rapid Variations

Charged particles stream from the Sun in the solar The solar wind interacts with the geomagnetic field to create a cavity, the magnetosphere, in which the field is confined. When a region of enhanced velocity and/or density in the solar wind arrives at the dayside boundary of the magnetosphere (at about 10 earth radii) the boundary is pushed towards the Earth. Currents set up on the boundary of the magnetosphere can cause an abrupt change in the geomagnetic field measured on the ground and this is recorded on observatory magnetograms as a sudden impulse (si). If, following an si, there is a change in the rhythm of activity, the si is termed a storm sudden commencement (ssc). A classical magnetic storm exhibiting initial, main and recovery phases (shown by, for instance, the Dst ring current index) can often occur after a ssc, in which case the start of the storm is taken as the time of the ssc.

Solar flares, seen at optical wavelengths as a sudden brightening of a small region of the Sun's surface, are also responsible for increased X-ray emissions. These X-rays cause increased ionisation in the ionosphere, which leads to absorption of short-wave radio signals. A solar flare effect (*sfe*), or "crochet", may be observed on a magnetogram during geomagnetically quiet times. It is a relatively short-term change (tens of minutes) to the normal diurnal variation and can vary in size (tens of nT) depending on local time (LT), geomagnetic latitude and solar zenith angle.

#### 4.7 Local geomagnetic activity indices

The Observatory K index. This summarises geomagnetic activity at an observatory by assigning a code, an integer in the range 0 to 9, to each 3-hour Universal Time (UT) interval. The index for each 3-hour UT interval is determined from the maximum range in H or D (scaled in nT), with allowance made for the regular (undisturbed) diurnal variation. The conversion from range to an index value is made using a quasi-logarithmic scale, with the scale values dependent on the geomagnetic latitude of the observatory. The lower bounds (in nT) for the classification of each period at Eskdalemuir are:

0	1	2	3	4	5	6	7	8	9
0	8	15	30	60	105	180	300	500	750

The *K* index retains the LT and seasonal dependence of activity associated with the position of the observatory. The 3-hourly *K* indices for the month are tabulated and also plotted as a histogram. All UK observatory *K* indices are available at

www.geomag.bgs.ac.uk/data\_service/data/magnetic\_indices/k\_indices

#### 4.8 Global geomagnetic activity indices

The aa index. A number of 3-hour geomagnetic indices are computed by combining K indices from networks of observatories to characterise global activity levels and to eliminate LT and seasonal effects. The simplest of these is the aa index, computed using the K indices from two approximately antipodal observatories: Hartland in the UK and Canberra in Australia. The aa index is calculated from linearisations of the Hartland and Canberra K indices, and has units of nT. The 3hourly aa indices are tabulated along with the daily mean value of aa (denoted Aa), the mean values of aa for the intervals 00-12UT (Aaam) and 12-24UT  $(Aa_{nm})$  and the monthly mean value. The 3-hourly aa indices for the month are also plotted as a histogram.

Although the *aa* index is based on data from only two observatories, provided averages over 12 hours or longer are used, the index is strongly correlated with the *ap* and *am* indices, which are derived using data from more extensive observatory networks.

The *aa* indices listed in this bulletin are available at <a href="https://www.geomag.bgs.ac.uk/data\_service/data/magneticondices/aaindex">www.geomag.bgs.ac.uk/data\_service/data/magneticondices/aaindex</a> as well as the full data set from 1868.

Definitive *aa* are published by the International Service for Geomagnetic Indices, LATMOS, 4 Avenue de Neptune, F-94107 Saint Maur Cedex, France.

#### 5. Conditions of Use

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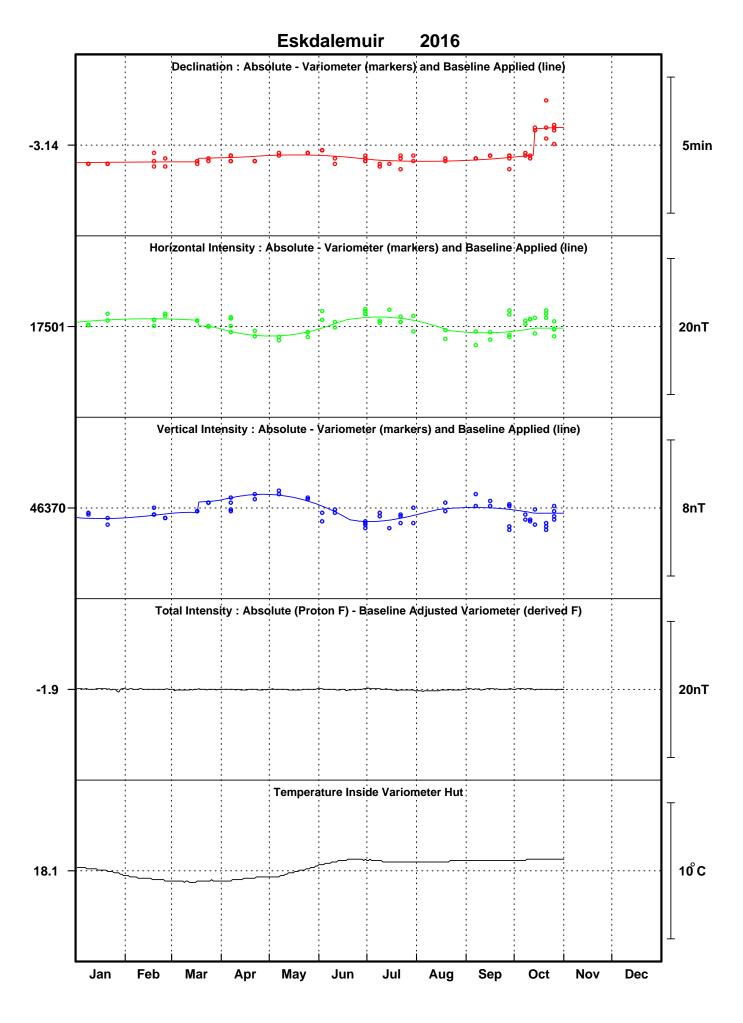
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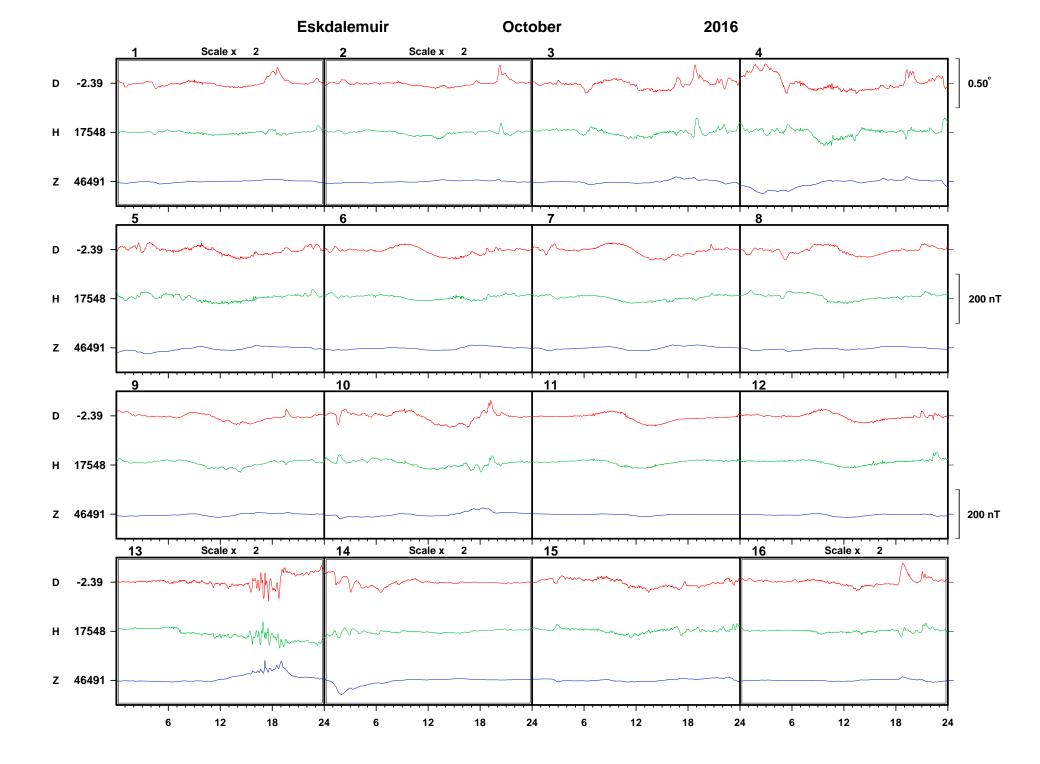
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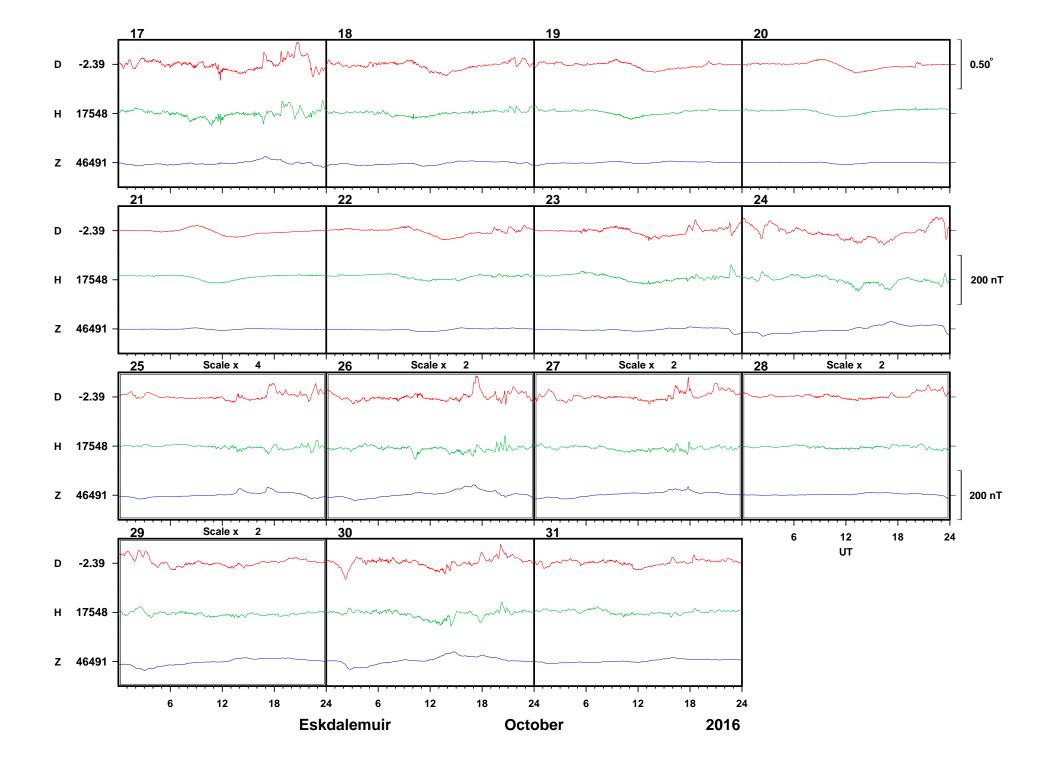
# ESKDALEMUIR OBSERVATORY

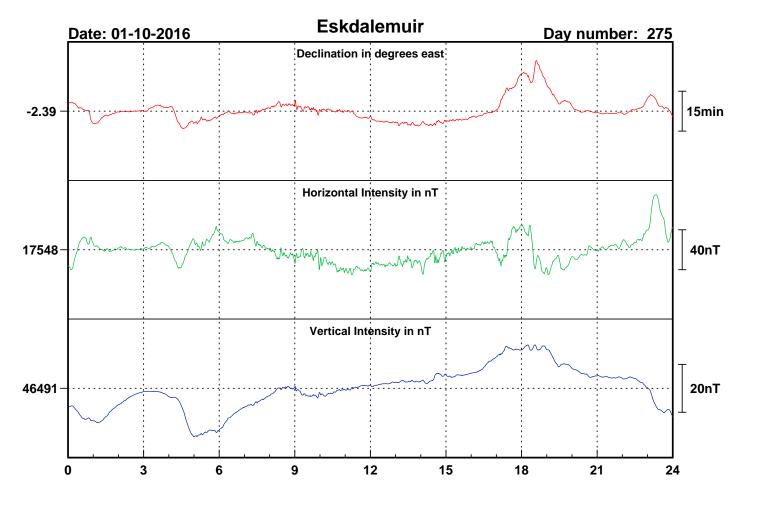
# ABSOLUTE OBSERVATIONS

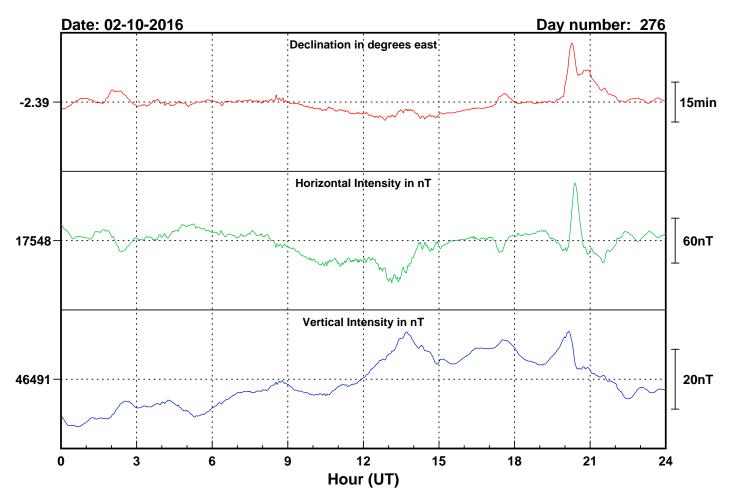
			Declination		Inclination		Total Field		Horizontal Intensity		Vertical Intensity		
Date	Day Number	Time (UT)	Absolute (°)	Baseline (°)	Time (UT)	Absolute (°)	Site difference (nT)	Absolute corrected (nT)	Absolute (nT)	Baseline (nT)	Absolute (nT)	Baseline (nT)	Observer
07-Oct-16	281	10:43	-2.3532	-3.1467	10:55	69.3317	1.9	49678.9	17534.5	17501.8	46481.5	46369.2	СР
07-Oct-16	281	11:03	-2.3590	-3.1450	11:13	69.3327	1.9	49676.6	17532.9	17502.3	46479.7	46368.9	CP
10-Oct-16	284	12:36	-2.4705	-3.1467	12:46	69.3335	1.9	49680.2	17533.5	17502.5	46483.3	46368.8	AS
10-Oct-16	284	12:55	-2.4820	-3.1483	13:04	69.3350	1.9	49679.9	17532.2	17502.5	46483.5	46368.9	AS
13-Oct-16	287	13:18	-2.4328	-2.0633	13:29	69.3642	1.9	49709.5	17519.0	17502.7	46520.1	46368.6	СР
13-Oct-16	287	13:37	-2.4578	-2.0650	13:47	69.3663	1.9	49714.5	17519.0	17500.4	46525.4	46369.5	CP
20-Oct-16	294	10:20	-2.3672	-2.0633	10:38	69.3272	1.9	49683.9	17539.9	17503.3	46484.9	46368.5	TT
20-Oct-16	294	11:10	-2.4064	-2.0700	11:32	69.3280	1.9	49680.7	17538.2	17503.8	46482.1	46368.3	KL
20-Oct-16	294	11:50	-2.4124	-2.0467	12:07	69.3278	1.9	49680.7	17538.3	17502.7	46482.1	46368.7	VF
25-Oct-16	299	10:14	-2.4339	-2.0650	10:27	69.3167	1.9	49686.8	17549.5	17502.2	46484.3	46368.9	TT
25-Oct-16	299	10:43	-2.4002	-2.0617	10:53	69.3290	1.9	49687.7	17539.8	17501.0	46488.9	46369.4	TT
25-Oct-16	299	13:38	-2.4611	-2.0733	13:47	69.4487	1.9	49745.9	17463.1	17500.0	46580.0	46369.7	TT
25-Oct-16	299	13:54	-2.3822	-2.0633	14:03	69.4215	1.9	49771.1	17494.1	17501.1	46595.3	46369.1	ТТ

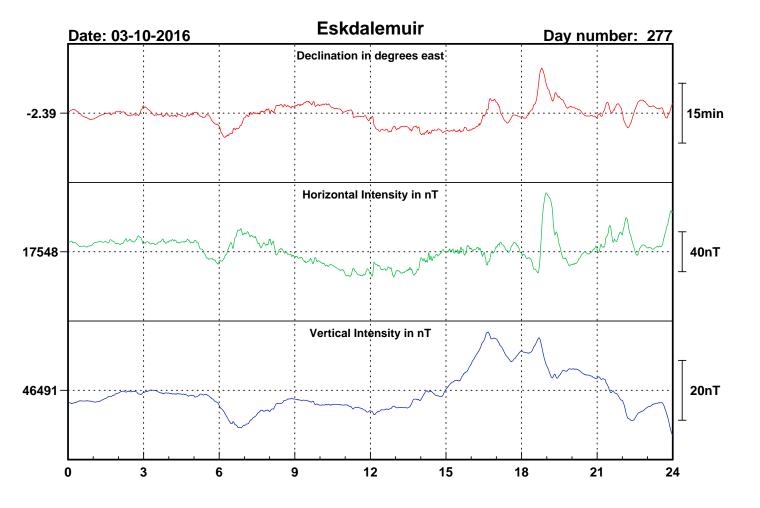


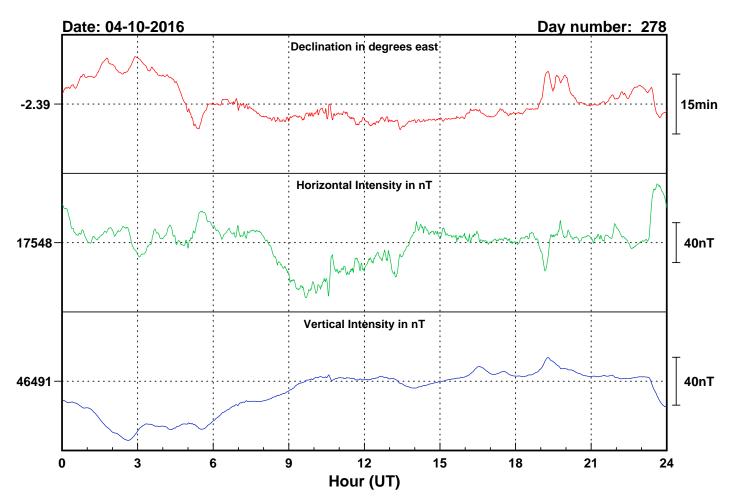


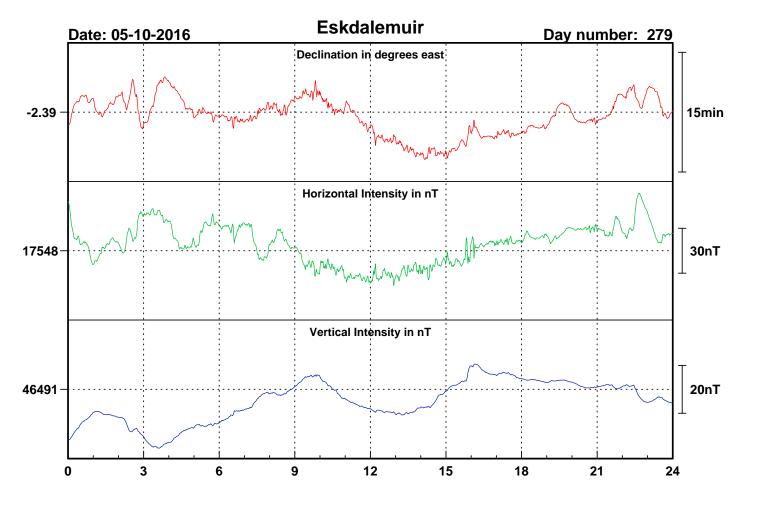


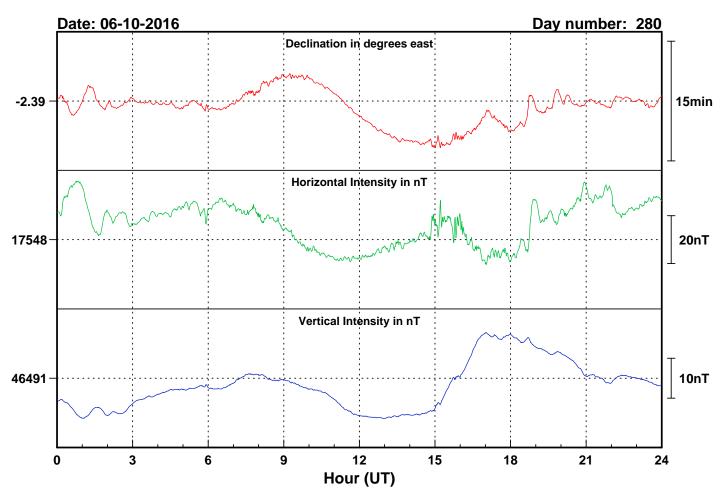


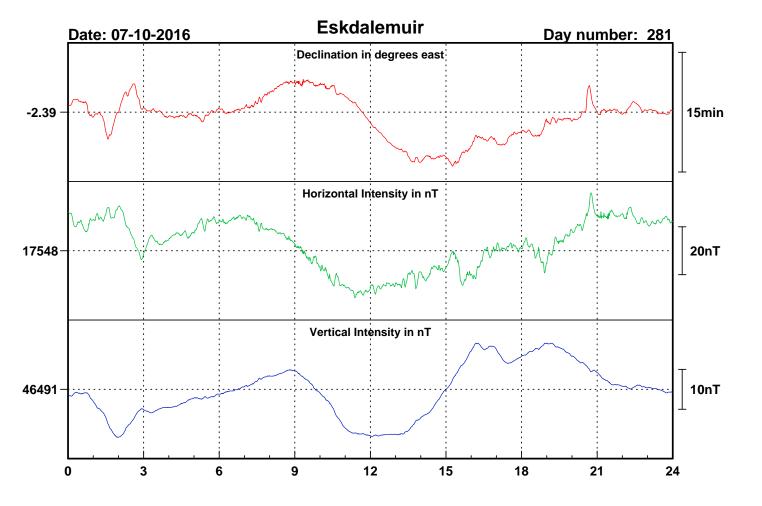


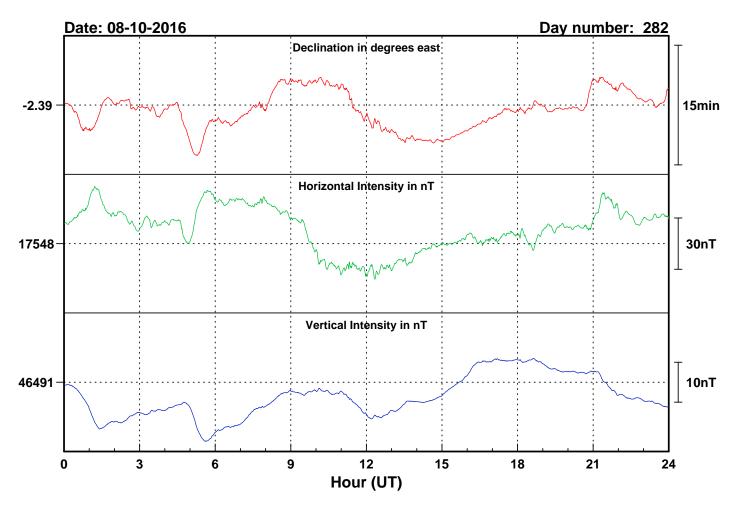


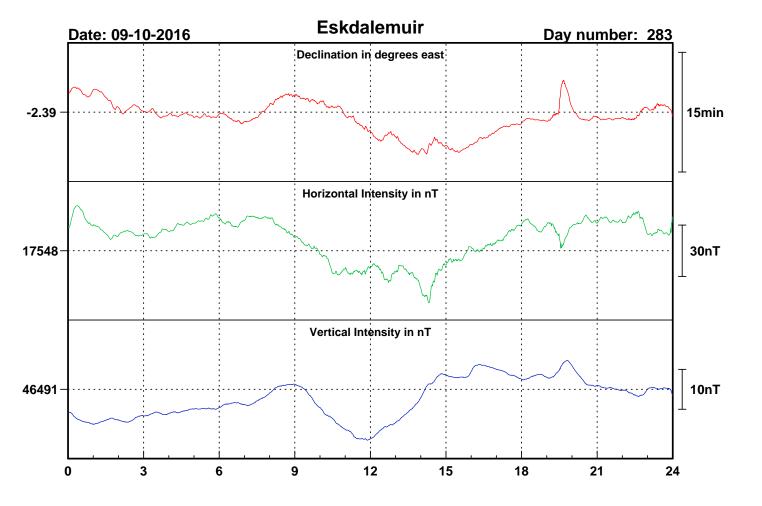


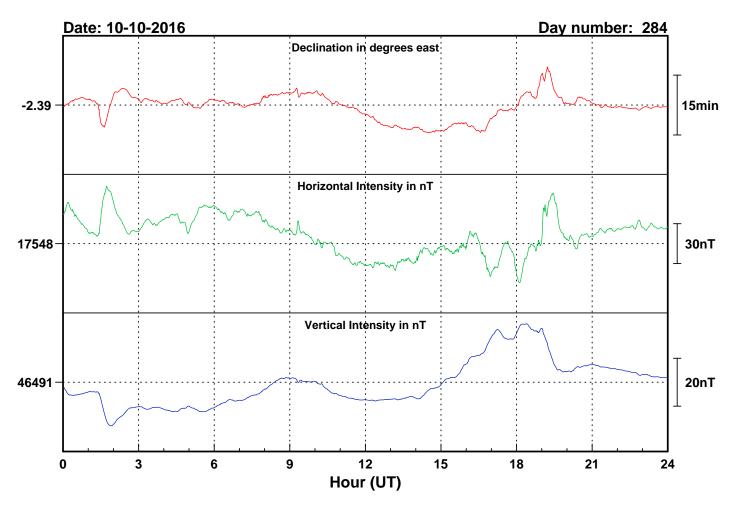


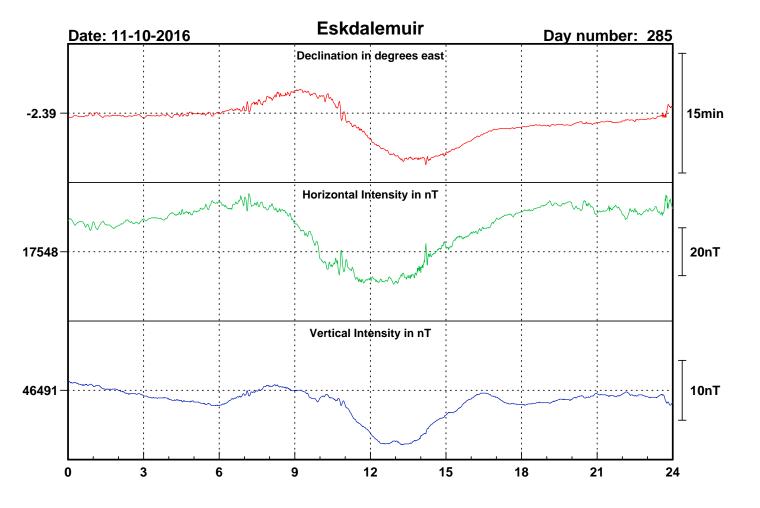


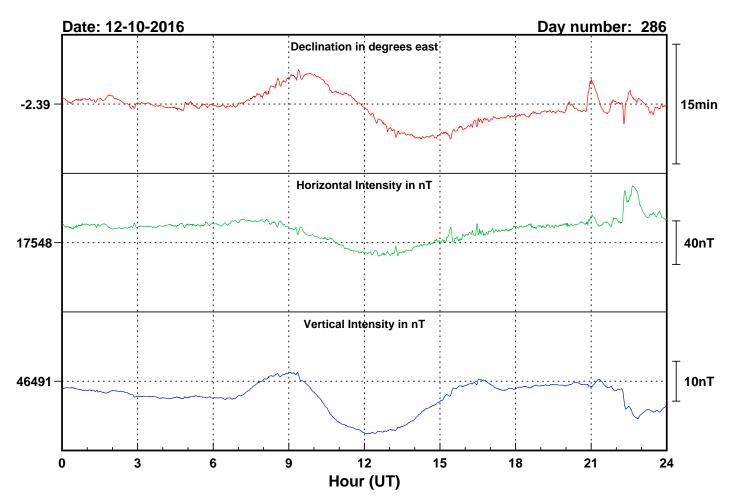


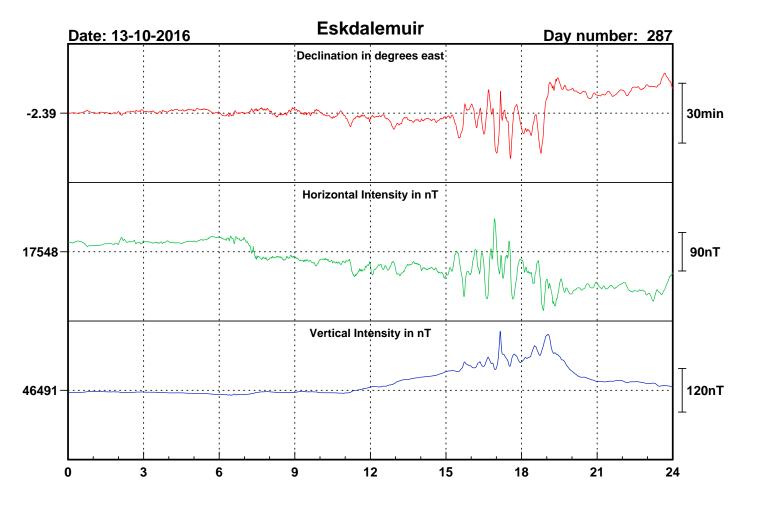


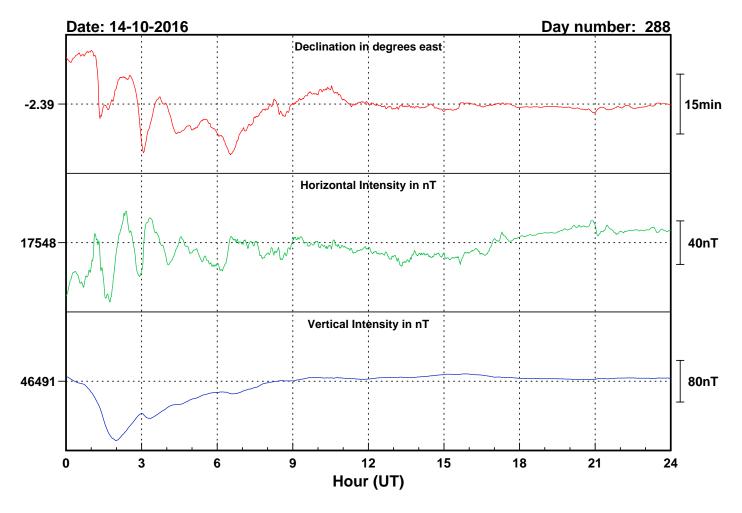


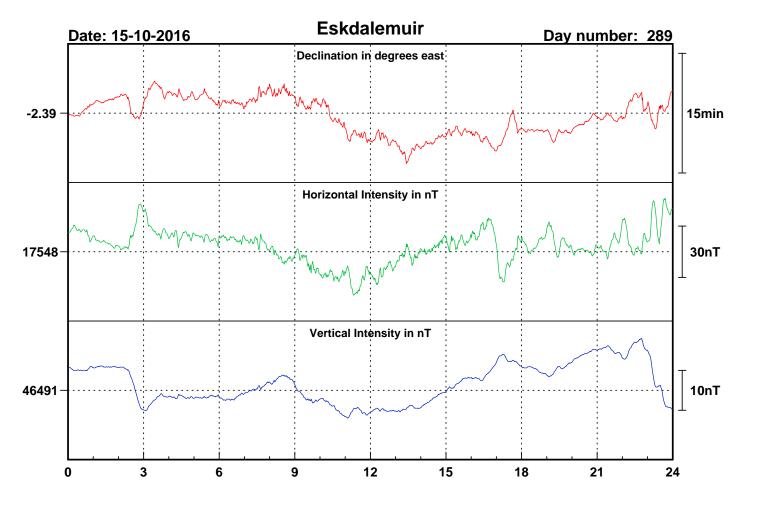


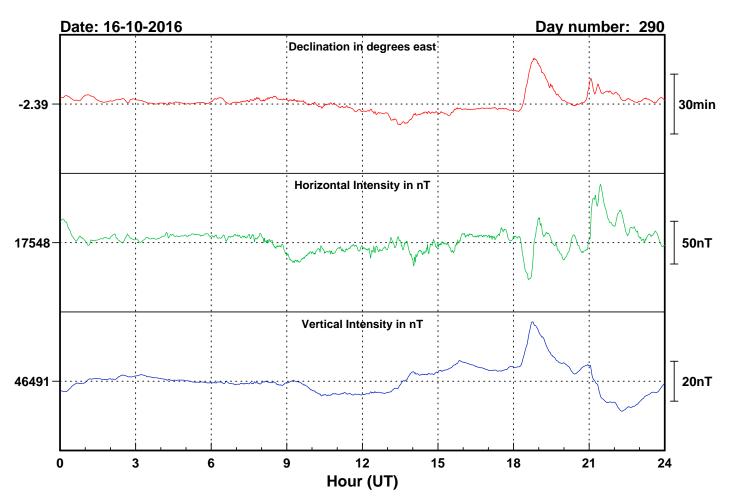


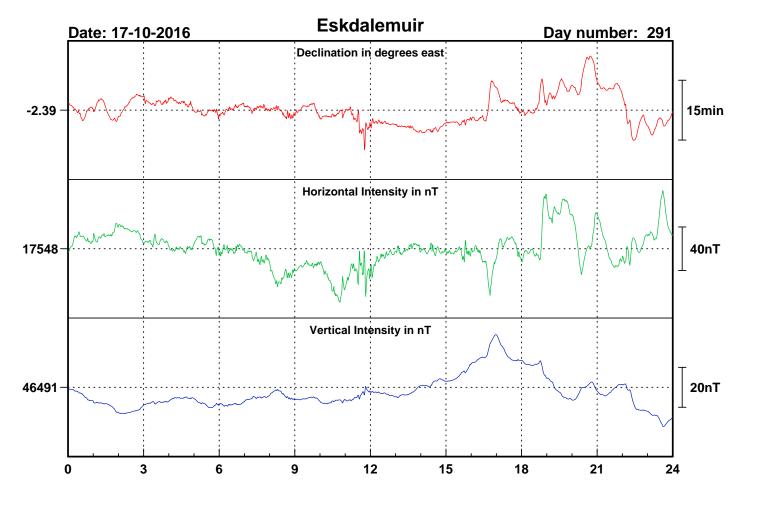


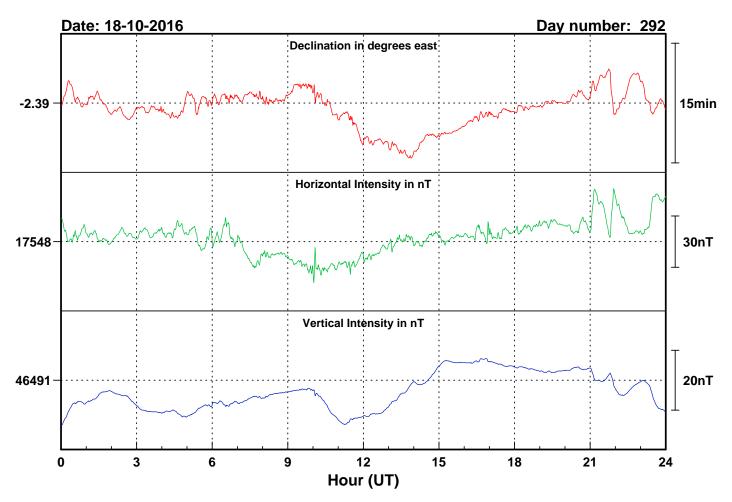


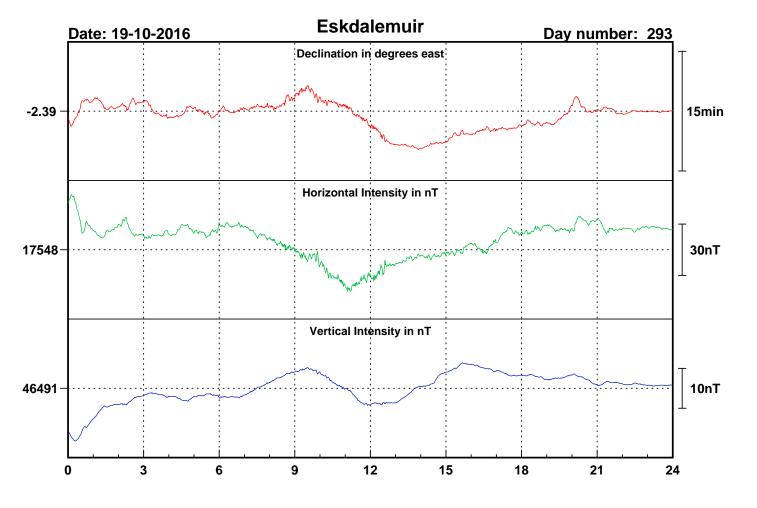


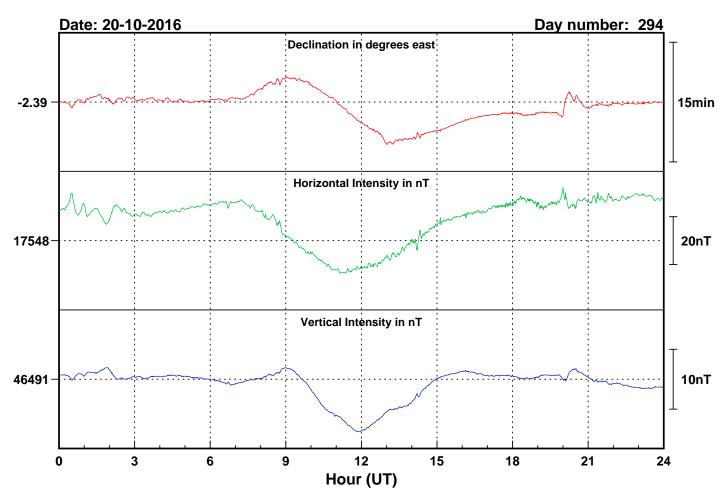


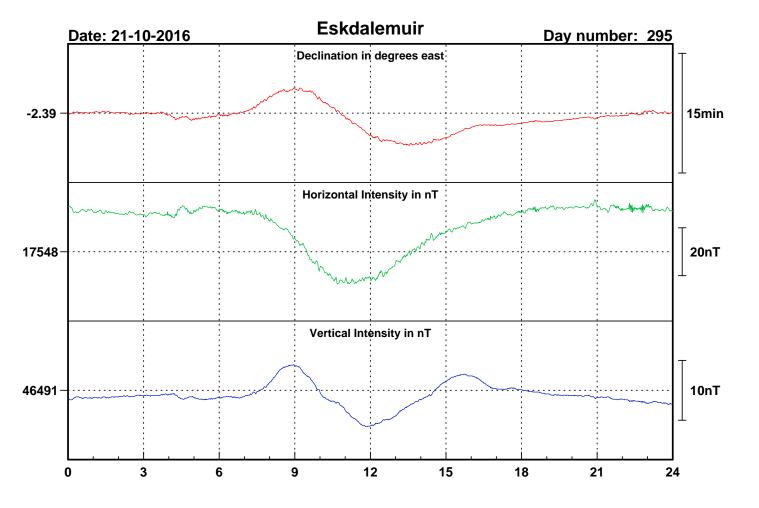


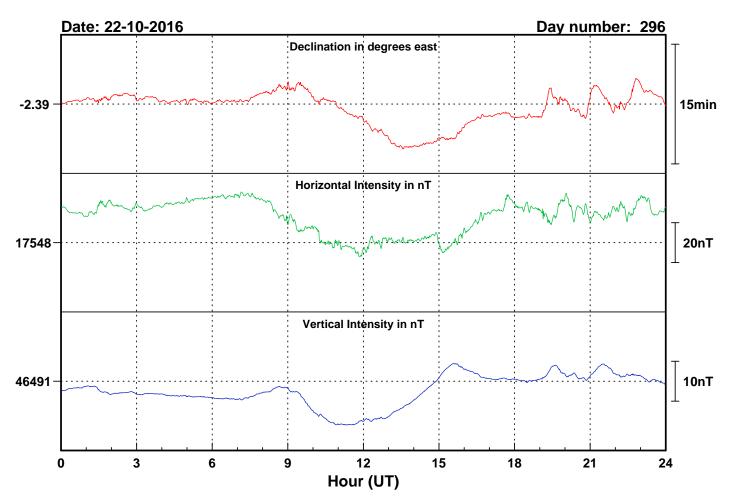


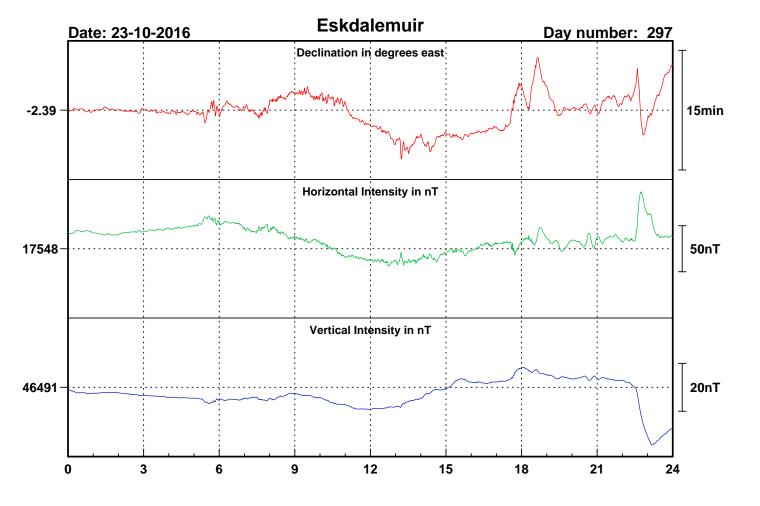


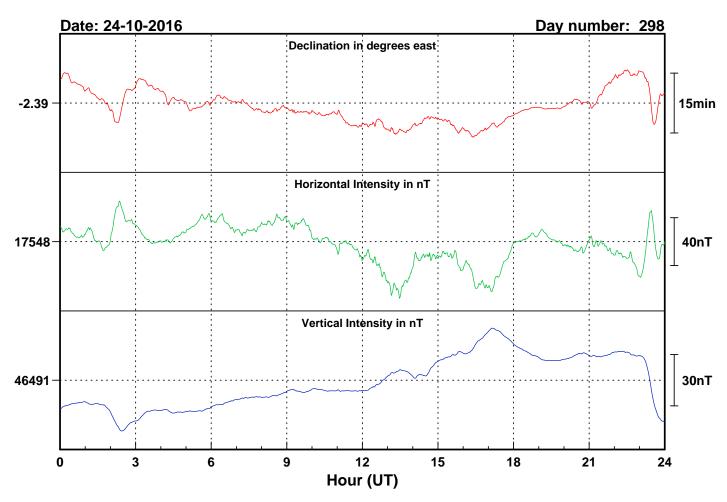


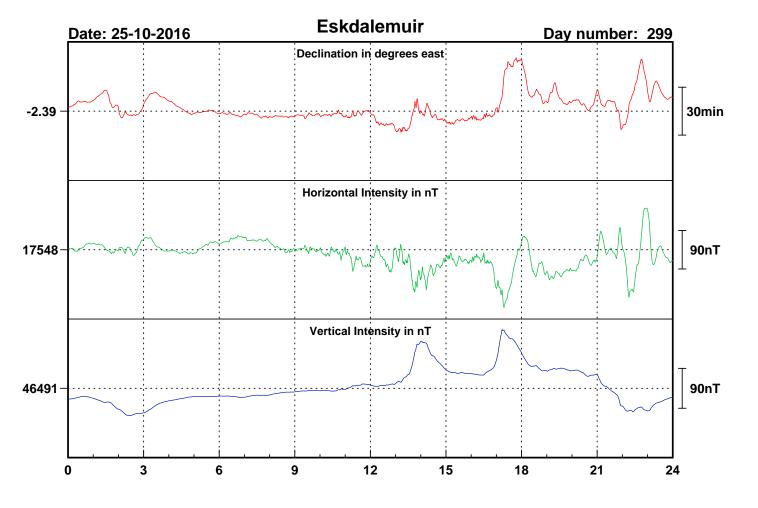


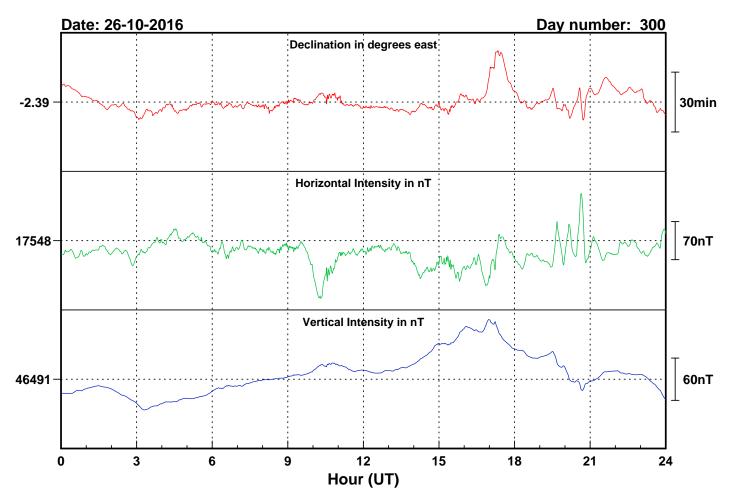


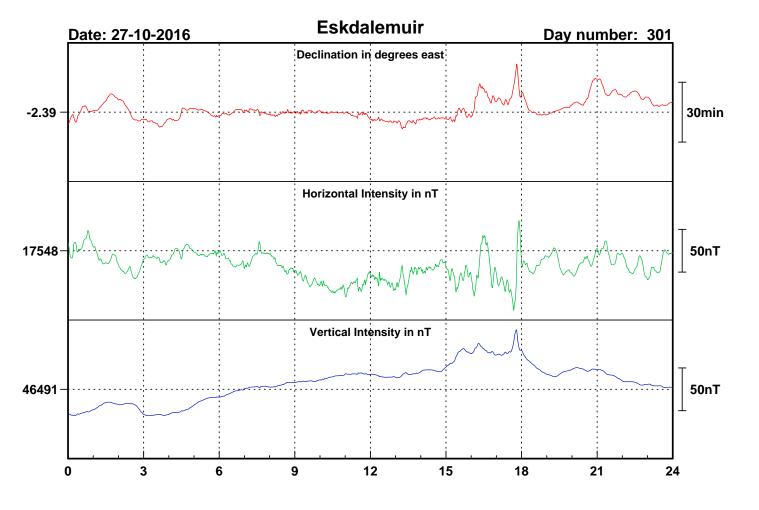


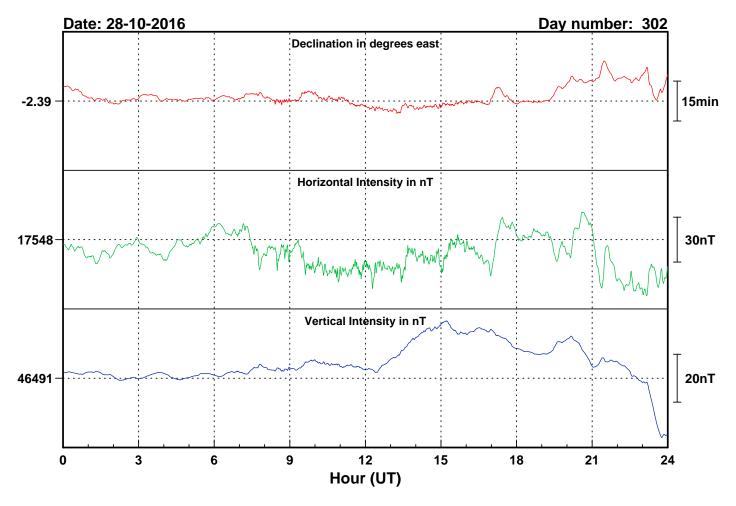


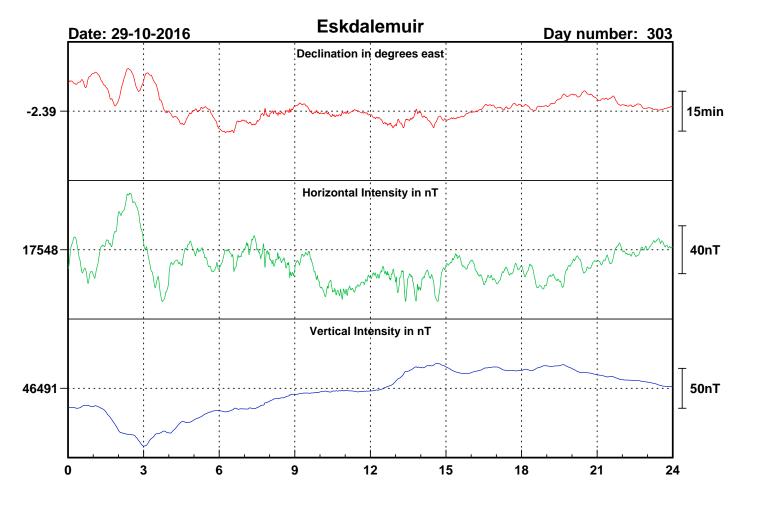


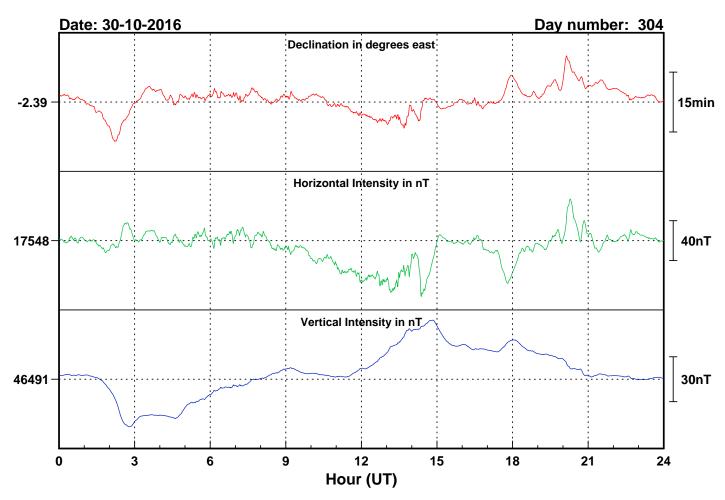


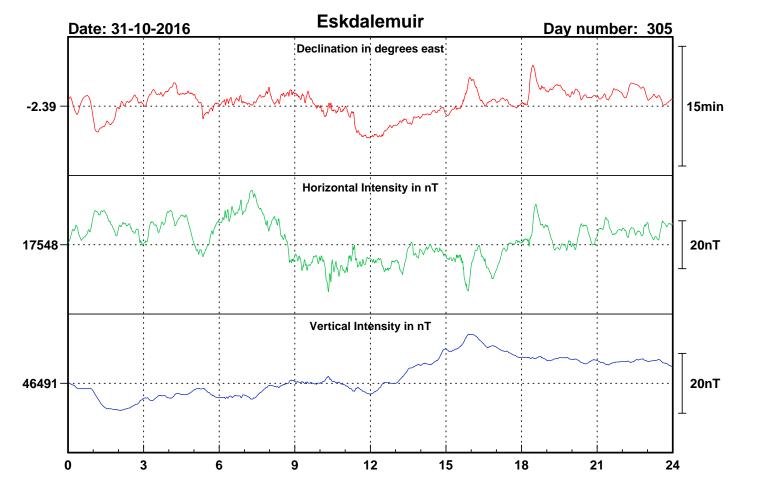




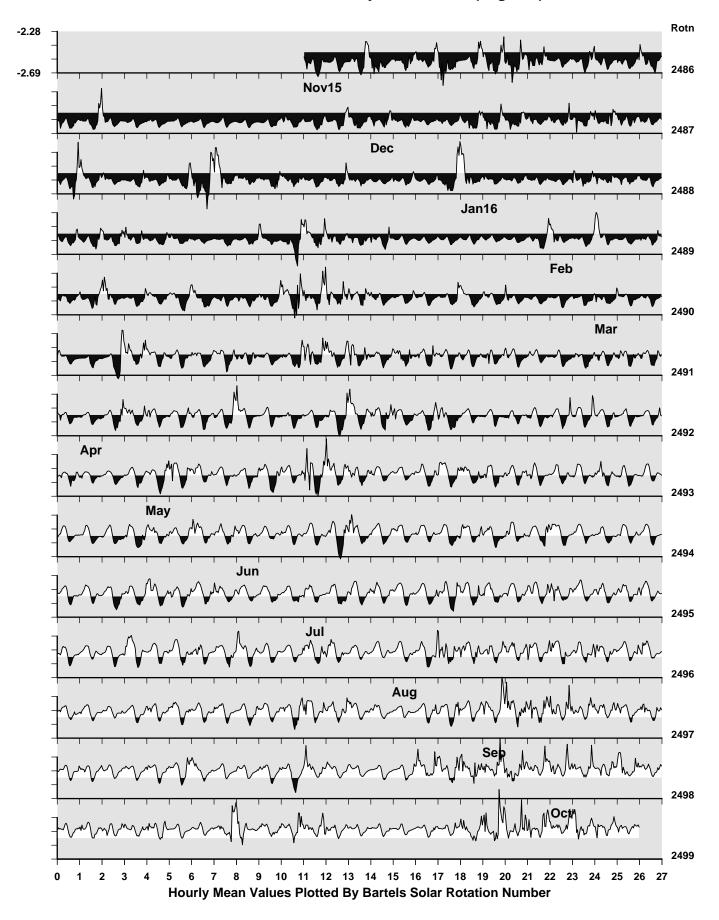




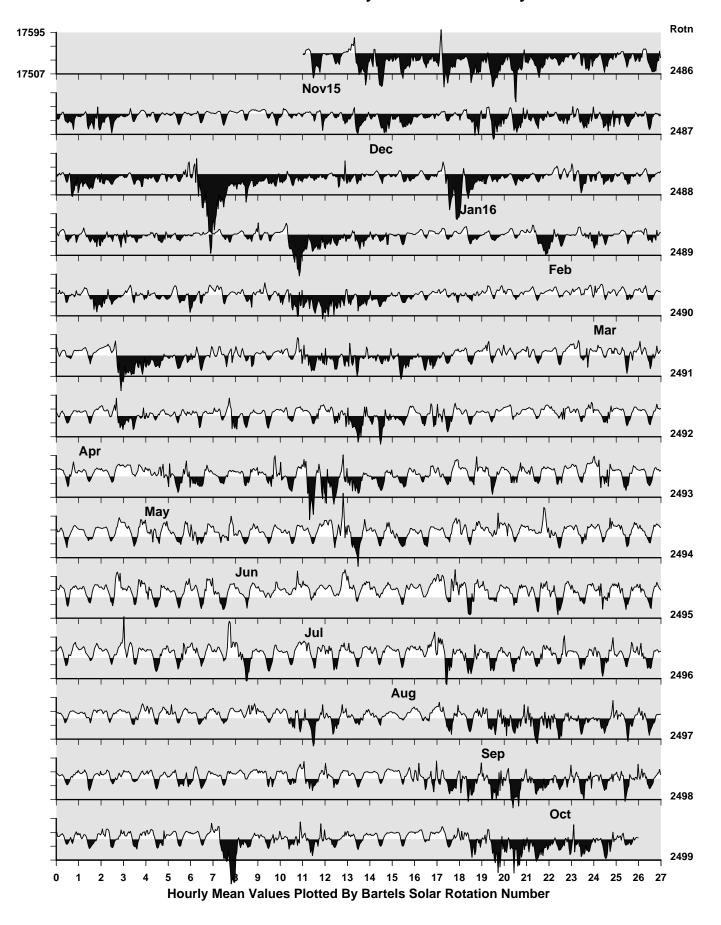




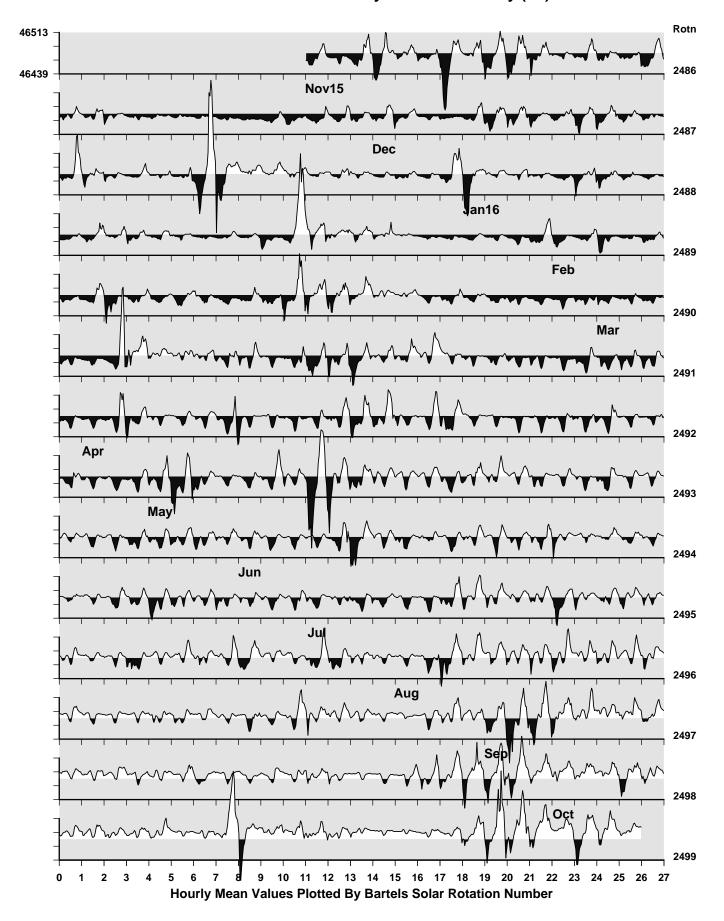
# **Eskdalemuir Observatory: Declination (degrees)**

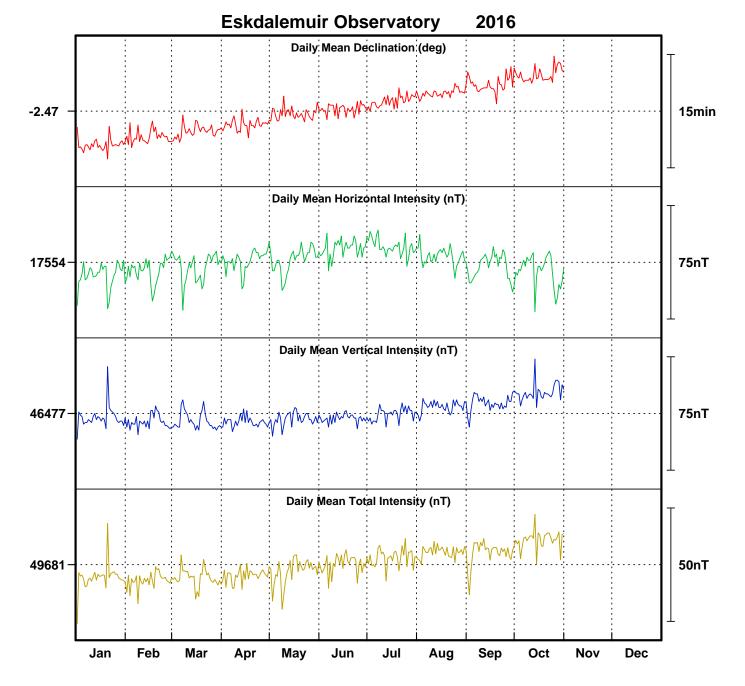


# Eskdalemuir Observatory: Horizontal Intensity fhTŁ



# Eskdalemuir Observatory: Vertical Intensity (nT)





# Monthly Mean Values for Eskdalemuir Observatory 2016

Month	D	H	I	X	Y	Z	F
January	-2° 32.7′	17545 nT	69° 19.1′	17528 nT	-779 nT	46474 nT	49676 nT
February	-2° 31.7′	17548 nT	69° 18.8′	17531 nT	-774 nT	46472 nT	49675 nT
March	-2° 30.9′	17550 nT	69° 18.7′	17533 nT	-770 nT	46472 nT	49676 nT
April	-2° 30.1′	17555 nT	69° 18.3′	17538 nT	-766 nT	46471 nT	49677 nT
May	-2° 28.6′	17555 nT	69° 18.4′	17539 nT	-759 nT	46473 nT	49679 nT
June	-2° 28.0′	17563 nT	69° 17.9′	17547 nT	-756 nT	46474 nT	49682 nT
July	-2° 26.9′	17564 nT	69° 17.9′	17548 nT	-750 nT	46477 nT	49685 nT
August	-2° 25.9′	17557 nT	69° 18.4′	17542 nT	-745 nT	46482 nT	49687 nT
September	-2° 24.7′	17550 nT	69° 18.9′	17535 nT	-738 nT	46484 nT	49687 nT
October	-2° 23.5′	17548 nT	69° 19.3′	17533 nT	-732 nT	46491 nT	49692 nT

# Note

i. The values shown here are provisional.

#### ESKDALEMUIR RAPID VARIATIONS

#### SIs and SSCs

Date	Time (UT)	Type	Quality	H (nT)	D (min)	Z (nT)
12-10-16	22 13	SSC	В	28.1	-2.66	-4.9
25-10-16	09 21	SSC	C	7.2	-1.62	-

#### **Notes:**

An asterisk (\*) indicates that the principal impulse was preceded by a smaller reversed impulse.

The quality of the event is classified as follows:

A = very distinct

B = fair, ordinary, but unmistakable

C = doubtful

The amplitudes given are for the first chief movement of the event.

#### **SFEs**

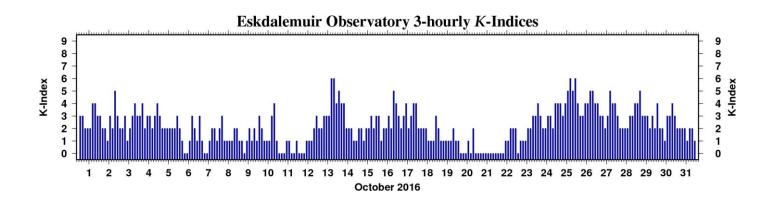
Date		<b>Universal Time</b>		H (nT)	D (min)	Z (nT)
	Start	Maximum	End			
None						

#### **Note:**

The amplitudes given are for the first chief movement of the event.

# INDICES OF GEOMAGNETIC ACTIVITY

	K - INDICES FOR THREE-HOUR INTERVAL									
Day	00-03	03-06	06-09	09-12	12-15	15-18	18-21	21-24		
1	3	3	2	2	2	4	4	3		
2	3	2	2	1	3	2	5	3		
3	2	2	3	1	2	3	4	3		
4	3	4	2	3	3	2	3	4		
5	3	2	2	2	2	2	2	3		
6	2	1	0	0	1	3	2	1		
7	3	1	0	0	1	2	2	1		
8	2	3	1	1	1	1	2	2		
9	1	1	0	1	2	1	2	1		
10	3	2	1	1	1	3	4	1		
11	0	0	0	1	1	0	0	1		
12	0	0	0	1	1	1	2	3		
13	2	2	3	3	3	6	6	4		
14	5	4	4	2	2	2	1	1		
15	2	2	1	2	2	3	2	3		
16	3	1	2	2	3	2	5	4		
17	3	2	3	4	2	3	4	4		
18	2	2	2	2	1	1	1	3		
19	2	1	1	1	1	1	2	1		
20	1	0	0	0	1	0	2	0		
21	0	0	0	0	0	0	0	0		
22	0	0	0	1	1	2	2	2		
23	0	1	1	1	2	2	3	3		
24	4	3	2	2	3	3	2	4		
25	4	4	3	4	5	6	5	6		
26	4	3	3	4	4	5	5	4		
27	4	3	3	2	3	5	4	4		
28	3	2	2	2	2	3	3	4		
29	4	5	3	3	3	2	3	2		
30	4	2	2	1	3	3	4	3		
31	2	2	2	2	1	2	2	1		



The aa Index

Date	Day	3-hourly <i>aa-</i> indices								Aaam	$Aa_{pm}$	Aa
01-10-16	275	20	32	16	24	24	67	67	37	23.1	48.6	35.9
02-10-16	276	20	16	16	20	59	24	80	45	18.0	52.2	35.1
03-10-16	277	16	24	24	8	24	24	37	24	18.0	27.4	22.7
04-10-16	278	37	45	24	46	59	16	45	37	38.1	39.4	38.8
05-10-16	279	24	24	24	24	24	16	12	20	24.1	18.0	21.1
06-10-16	280	12	8	12	2	5	24	24	12	8.5	16.3	12.4
07-10-16	281	20	8	8	5	12	16	16	5	10.2	12.3	11.2
08-10-16	282	20	24	12	24	12	8	12	12	20.0	10.9	15.4
09-10-16	283	9	5	8	8	24	8	12	8	7.5	12.9	10.2
10-10-16	284	20	16	12	8	12	32	45	8	13.9	24.3	19.1
11-10-16	285	2	5	5	8	5	2	5	8	5.2	5.1	5.2
12-10-16	286	8	5	5	8	2	8	12	32	6.5	13.6	10.0
13-10-16	287	16	24	59	59	67	80	115	37	39.5	74.8	57.2
14-10-16	288	58	37	37	32	12	16	8	5	41.3	10.3	25.8
15-10-16	289	24	16	16	24	16	32	12	32	20.1	23.1	21.6
16-10-16	290	24	8	24	24	32	24	67	37	20.1	40.0	30.1
17-10-16	291	24	32	46	59	16	59	45	37	40.2	39.4	39.8
18-10-16	292	12	16	16	16	16	12	8	20	15.0	13.9	14.5
19-10-16	293	12	12	12	12	8	12	12	5	12.0	9.2	10.6
20-10-16	294	8	2	2	2	5	2	12	2	3.8	5.5	4.6
21-10-16	295	2	5	2	2	2	2	2	2	3.1	2.5	2.8
22-10-16	296	8	5	8	12	12	24	12	12	8.2	15.0	11.6
23-10-16	297	2	24	16	12	24	24	24	33	13.7	26.3	20.0
24-10-16	298	37	24	24	24	46	32	24	37	27.4	34.8	31.1
25-10-16	299	37	37	32	81	137	115	101	115	46.8	116.8	81.8
26-10-16	300	59	46	46	102	59	136	67	67	63.0	82.1	72.5
27-10-16	301	45	45	32	32	32	80	67	45	38.8	56.1	47.4
28-10-16	302	20	16	32	24	24	32	45	45	23.1	36.8	29.9
29-10-16	303	45	80	59	32	46	16	32	16	54.1	27.5	40.8
30-10-16	304	45	24	24	38	46	32	37	16	32.9	32.8	32.8
31-10-16	305	12	16	24	24	12	16	12	8	19.1	11.9	15.5
									Mo	onthly Mea	an Value	26.7

Notes

i. The units of the *aa* index are nT.

- ii. The 3-hour aa values are rounded to the nearest integer. Where aa = \*.5, aa is rounded down.
- iii. Daily values ( $Aa_{am}$ ,  $Aa_{pm}$  and Aa) are computed from aa values of original resolution.
- iv. The monthly mean value is computed from the daily mean values, Aa.
- v. Definitive *aa* indices are derived and published by the International Service for Geomagnetic Indices.

