

BRITISH GEOLOGICAL SURVEY

Lerwick Observatory

Monthly Magnetic Bulletin

December 2015

15/12/LE



**British
Geological Survey**

NATIONAL ENVIRONMENT RESEARCH COUNCIL

LERWICK OBSERVATORY MAGNETIC DATA

1. Introduction

Lerwick observatory is one of three geomagnetic observatories in the UK operated and maintained by the British Geological Survey (BGS).

This bulletin is published to provide rapid access to the provisional geomagnetic observatory results. The information is freely available for personal, academic, educational and non-commercial research or use. Magnetic observatory data are presented as a series of plots of one-minute, hourly and daily values, followed by tabulations of monthly values, reports of rapid variations and geomagnetic activity indices. The operation of the observatory and presentation of data are described in the rest of this section.

Enquiries about the data should be addressed to:

Geomagnetism Team
Earth Hazards and Systems
British Geological Survey
Murchison House, West Mains Road
Edinburgh EH9 3LA
Scotland, UK

Tel: +44 (0) 131 667 1000
Fax: +44 (0) 131 668 2683
E-mail: enquiries@bgs.ac.uk
Internet: www.geomag.bgs.ac.uk

2. Position

The observatory is situated on a ridge of high ground about 2.5 km to the SW of the port of Lerwick in Shetland. The observatory co-ordinates are:

Geographic: $60^{\circ}08'18.0''N$ $358^{\circ}49'02.4''E$

Geomagnetic: $61^{\circ}49'05''N$ $088^{\circ}45'25''E$

Height above mean sea level: 85 m

The geographical coordinates are measured by a handheld GPS device, which uses WGS84 as the reference coordinate system. The height above MSL is determined from the best available contour maps. The geomagnetic co-ordinates are approximations, calculated using the 12th generation International Geomagnetic Reference Field (IGRF) at epoch 2015.5. On-line access to models (including IGRF), charts and navigational data are available at

www.geomag.bgs.ac.uk/data_service/models_compass/home

3. The Observatory Operation

3.1 GDAS

The observatory operates under the control of the Geomagnetic Data Acquisition System (GDAS), which was developed by BGS staff, installed in 2002, and became fully operational in January 2003. The data acquisition software, running on QNX operated computers, controls the data logging and the communications.

There are two sets of sensors used for making magnetic measurements. A tri-axial linear-core fluxgate magnetometer, manufactured by the Danish Meteorological Institute, is used to measure the variations in the horizontal (H) and vertical (Z) components of the field. The third sensor is oriented perpendicular to these, and measures variations, which are proportional to the changes in declination (D). Measurements are made at a rate of 1 Hz.

In addition to the fluxgate sensors there is a proton precession magnetometer (PPM) making measurements of the absolute total field intensity (F) at a rate of 0.1Hz.

The raw unfiltered data are retrieved automatically via Internet connections to the BGS office in Edinburgh in near real-time. The fluxgate data are filtered to produce one-minute values using a 61-point cosine filter and the total field intensity samples are filtered using a 7-point cosine filter. The one-minute values provide input for various data products, available on-line at

www.geomag.bgs.ac.uk/data_service/home

3.2 Back-up Systems

There are two other fully independent identical systems, GDAS 2 and GDAS 3, operating at the observatory. The data from these are also processed in near real-time and used for quality control purposes. They are also used to fill any gaps or replace any corrupt values in the primary system, GDAS 1.

3.3 Absolute Observations

The GDAS fluxgate magnetometers accurately measure variations in the components of the geomagnetic field, but not the absolute magnitudes. Two sets of absolute measurements of the field are made manually once per week. A fluxgate sensor mounted on a theodolite is used to determine D and

inclination (I); the GDAS PPM measurements, with a site difference correction applied, are used for F . The absolute observations are used in conjunction with the GDAS variometer measurements to produce a continuous record of the absolute values of the geomagnetic field elements as if they had been measured at the observatory reference pillar.

4. Observatory Results

The data presented in the bulletin are in the form of plots and tabulations described in the following sections.

4.1 Absolute Observations

The absolute observation measurements made during the month are tabulated. Also included are the corresponding baseline values, which are the differences between the absolute measurements and the variometer measurements of D , H and Z (in the sense absolute–variometer). These are also plotted (markers) along with the derived preliminary daily baseline values (line) throughout the year. Daily mean differences between the measured absolute F and the F computed from the baseline corrected H and Z values are plotted in the fourth panel (in the sense measured–derived). The bottom panel shows the daily mean temperature in the fluxgate chamber.

4.2 Summary magnetograms

Small-scale magnetograms are plotted which allow the month's data to be viewed at a glance. They are plotted 16 days to a page and show the one-minute variations in D , H and Z . The scales are shown on the right-hand side of the page. On disturbed days the scales are multiplied by a factor, which is indicated above the panel for that day. The variations are centred on the monthly mean value, shown on the left side of the page.

4.3 Magnetograms

The daily magnetograms are plotted using one-minute values of D , H and Z from the fluxgate sensors, with any gaps filled using back-up data. The magnetograms are plotted to a variable scale; scale bars are shown to the right of each plot. The absolute level (the monthly mean value) is indicated on the left side of the plots.

4.4 Hourly Mean Value Plots

Hourly mean values of D , H and Z for the past 12 months are plotted in 27-day segments corresponding to the Bartels solar rotation number. Magnetic disturbances associated with active

regions and/or coronal holes on the Sun may recur after 27 days: the same is true for geomagnetically quiet intervals. Plotting the data in this way highlights this recurrence. Diurnal variations are also clear in these plots and the amplitude changes throughout the year highlight the seasonal changes. Longer term secular variation is also illustrated.

Full lists of the UK observatory hourly mean values from 1983 to the present day are available at www.geomag.bgs.ac.uk/data_service/data/obs_data/hourly_means

4.5 Daily and Monthly Mean Values

Daily mean values of D , H , Z and F are plotted throughout the year. In addition, a table of monthly mean values of all the geomagnetic elements is provided. These values depend on accurate specification of the fluxgate sensor baselines. It is anticipated that these provisional values will not be altered by more than a few nT or tenths of arcminutes before being made definitive at the end of the year.

4.6 Rapid Variations

Charged particles stream from the Sun in the solar wind. The solar wind interacts with the geomagnetic field to create a cavity, the magnetosphere, in which the field is confined. When a region of enhanced velocity and/or density in the solar wind arrives at the dayside boundary of the magnetosphere (at about 10 earth radii) the boundary is pushed towards the Earth. Currents set up on the boundary of the magnetosphere can cause an abrupt change in the geomagnetic field measured on the ground and this is recorded on observatory magnetograms as a sudden impulse (*si*). If, following an *si*, there is a change in the rhythm of activity, the *si* is termed a storm sudden commencement (*ssc*). A classical magnetic storm exhibiting initial, main and recovery phases (shown by, for instance, the Dst ring current index) can often occur after a *ssc*, in which case the start of the storm is taken as the time of the *ssc*.

Solar flares, seen at optical wavelengths as a sudden brightening of a small region of the Sun's surface, are also responsible for increased X-ray emissions. These X-rays cause increased ionisation in the ionosphere, which leads to absorption of short-wave radio signals. A solar flare effect (*sfe*), or "crochet", may be observed on a magnetogram during geomagnetically quiet times. It is a relatively short-term change (tens of minutes) to the normal diurnal variation and can vary in size

(tens of nT) depending on local time (LT), geomagnetic latitude and solar zenith angle.

4.7 Local geomagnetic activity indices

The Observatory K index. This summarises geomagnetic activity at an observatory by assigning a code, an integer in the range 0 to 9, to each 3-hour Universal Time (UT) interval. The index for each 3-hour UT interval is determined from the maximum range in H or D (scaled in nT), with allowance made for the regular (undisturbed) diurnal variation. The conversion from range to an index value is made using a quasi-logarithmic scale, with the scale values dependent on the geomagnetic latitude of the observatory. The lower bounds (in nT) for the classification of each period at Lerwick are:

0	1	2	3	4	5	6	7	8	9
0	10	20	40	80	140	240	400	660	1000

The K index retains the LT and seasonal dependence of activity associated with the position of the observatory. The 3-hourly K indices for the month are tabulated and also plotted as a histogram. All UK observatory K indices are available at

www.geomag.bgs.ac.uk/data_service/data/magnetic_indices/k_indices

4.8 Global geomagnetic activity indices

The aa index. A number of 3-hour geomagnetic indices are computed by combining K indices from networks of observatories to characterise global activity levels and to eliminate LT and seasonal effects. The simplest of these is the aa index, computed using the K indices from two approximately antipodal observatories: Hartland in the UK and Canberra in Australia. The aa index is calculated from linearisations of the Hartland and Canberra K indices, and has units of nT. The 3-hourly aa indices are tabulated along with the daily mean value of aa (denoted Aa), the mean values of aa for the intervals 00-12UT (Aa_{am}) and 12-24UT (Aa_{pm}) and the monthly mean value. The 3-hourly aa indices for the month are also plotted as a histogram.

Although the aa index is based on data from only two observatories, provided averages over 12 hours or longer are used, the index is strongly correlated with the ap and am indices, which are derived using data from more extensive observatory networks.

The aa indices listed in this bulletin are available at www.geomag.bgs.ac.uk/data_service/data/magnetic_indices/aaindex as well as the full data set from 1868.

Definitive aa are published by the International Service for Geomagnetic Indices, LATMOS, 4 Avenue de Neptune, F-94107 Saint Maur Cedex, France.

5. Conditions of Use

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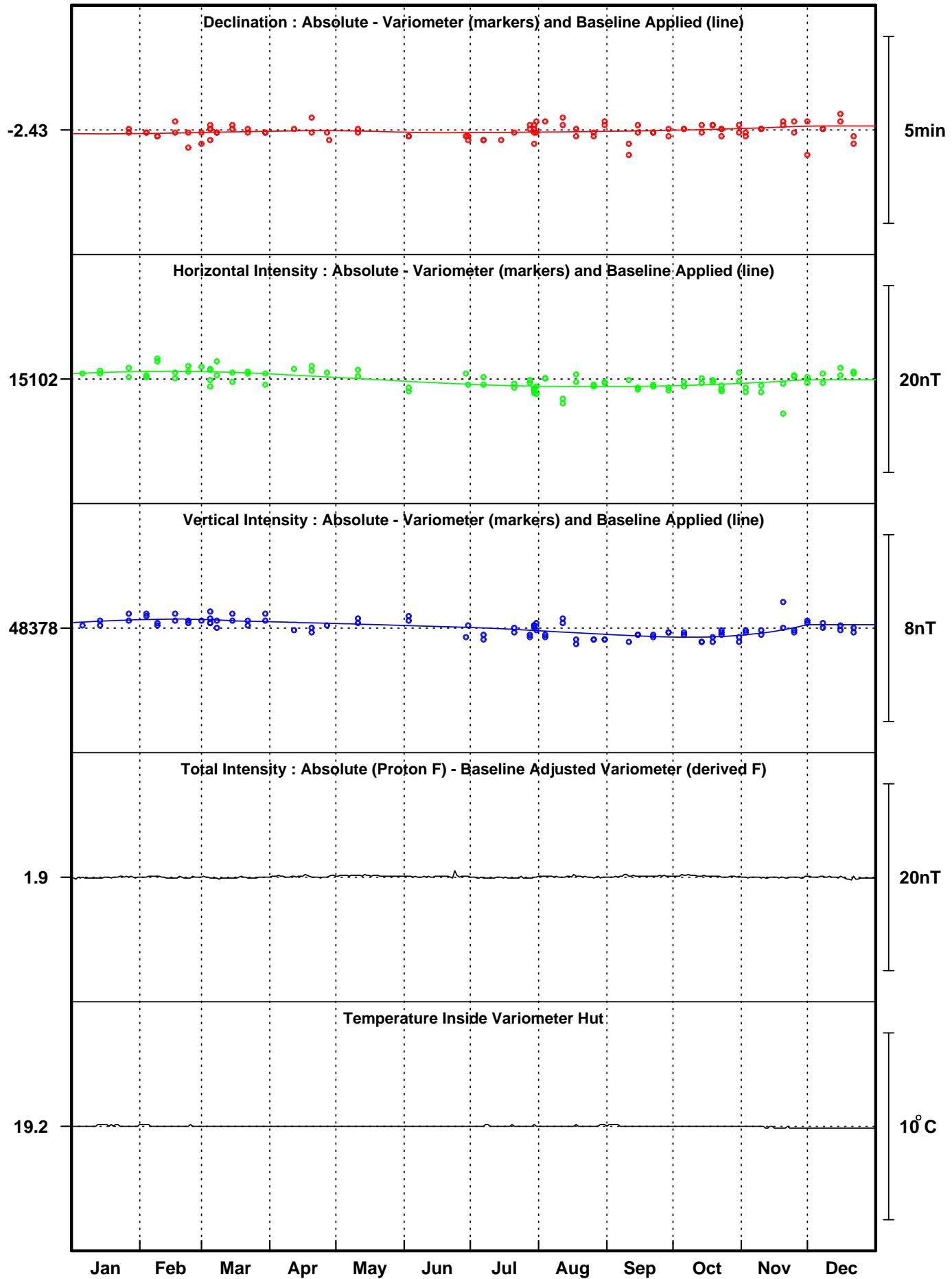
Commercial users can contact the geomagnetism team for information on the range of applications and services offered. Full contact details are available at www.geomag.bgs.ac.uk/contactus/staff

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LERWICK OBSERVATORY

ABSOLUTE OBSERVATIONS

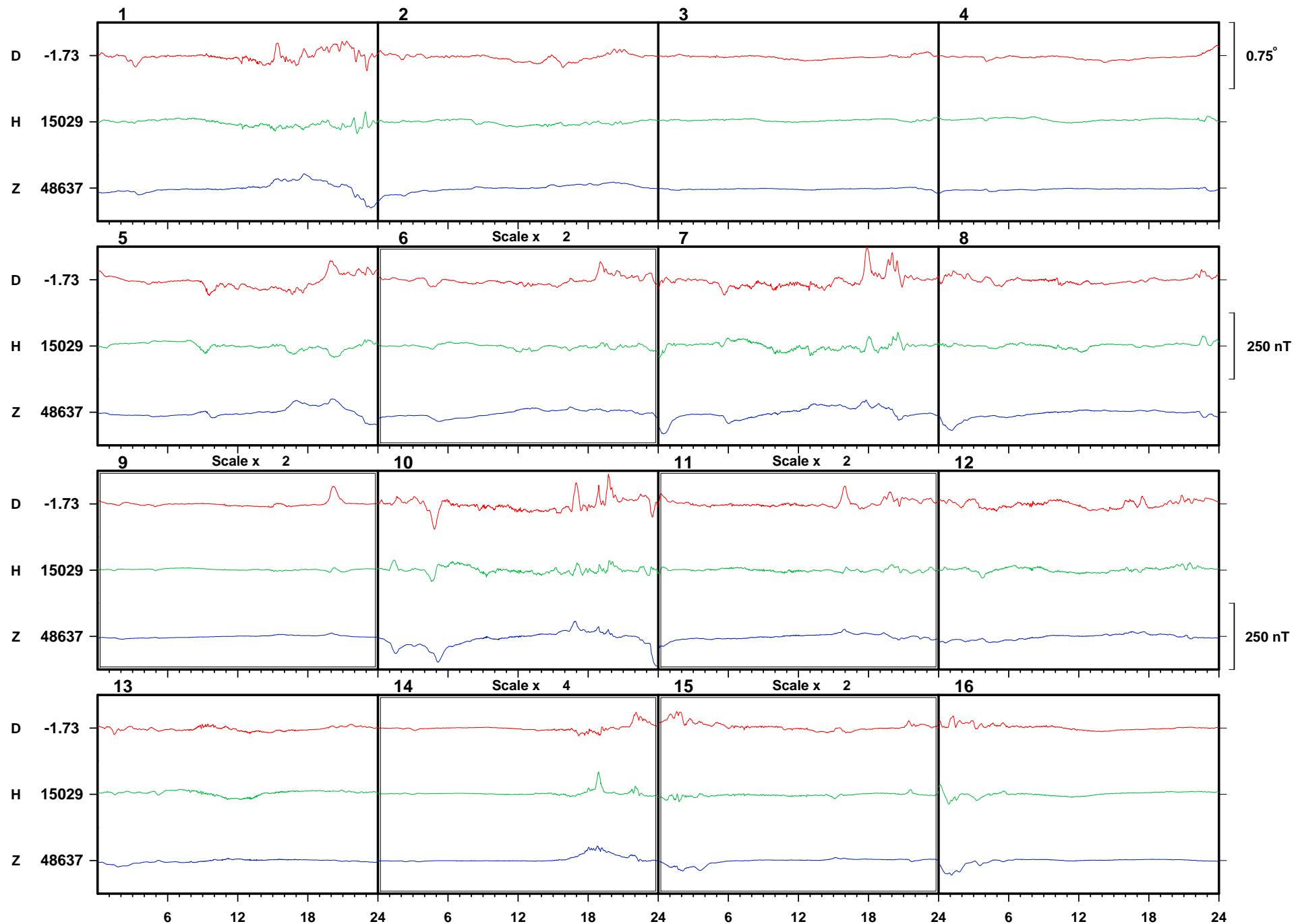
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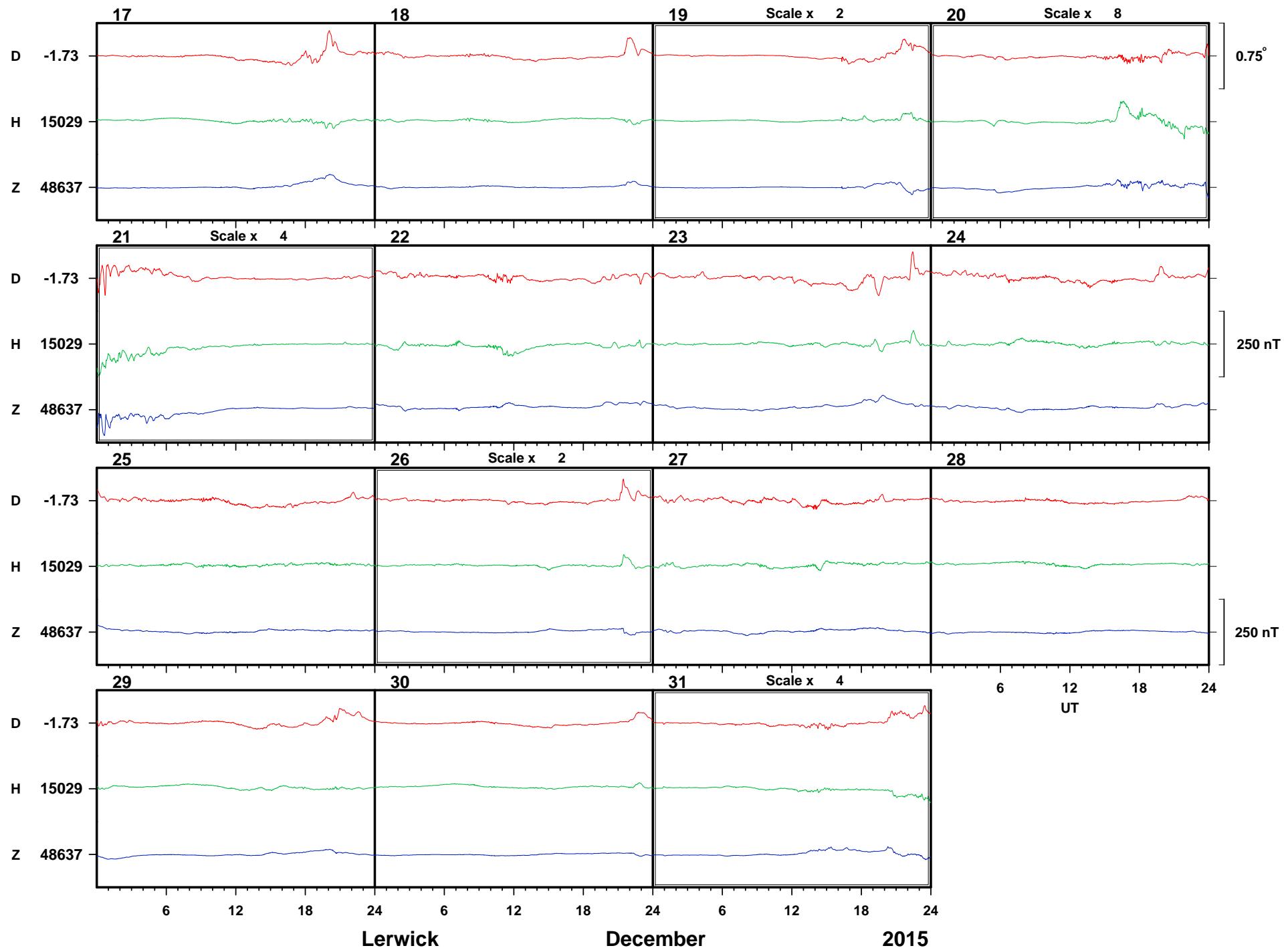


Lerwick

December

2015





Lerwick

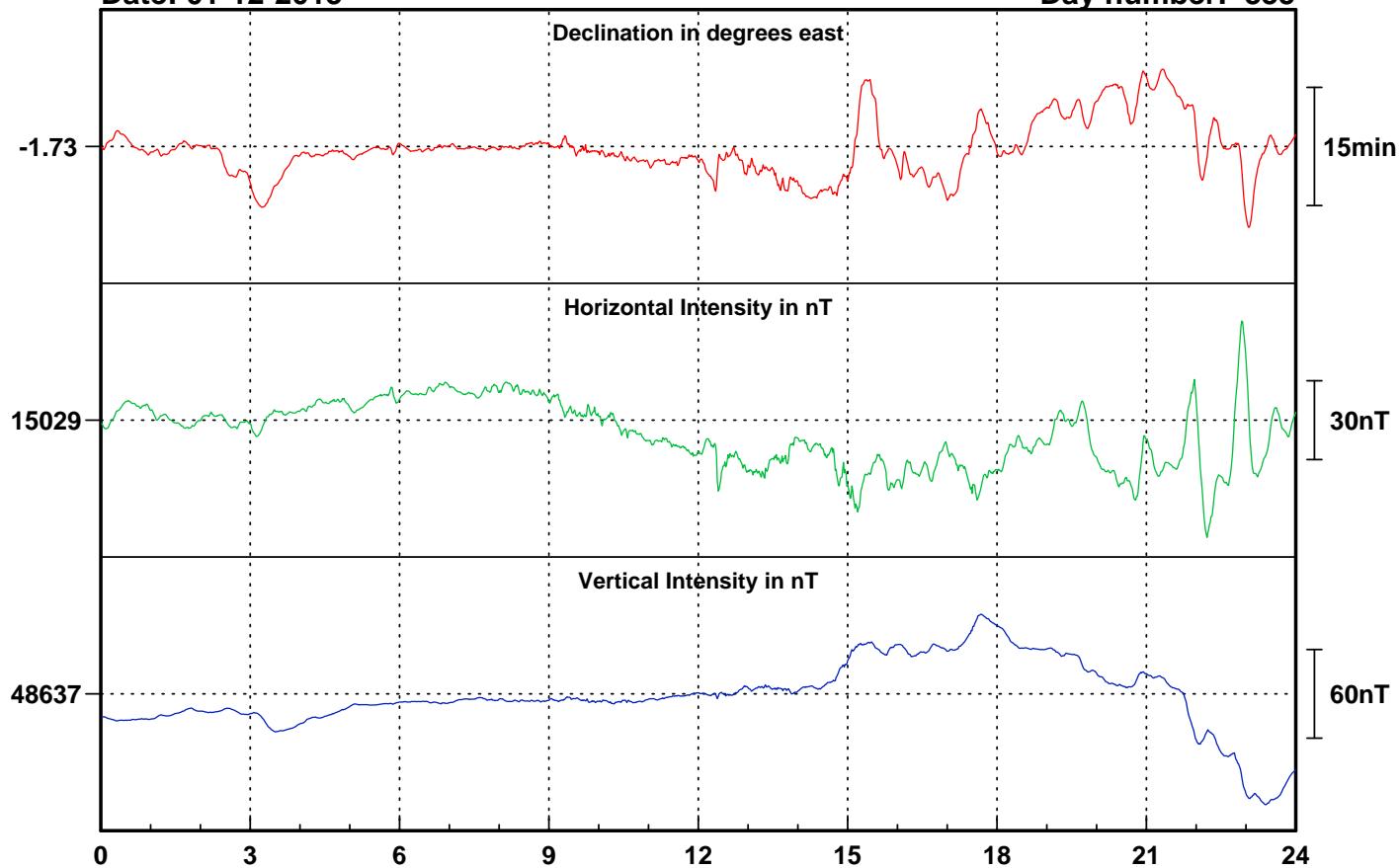
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2015

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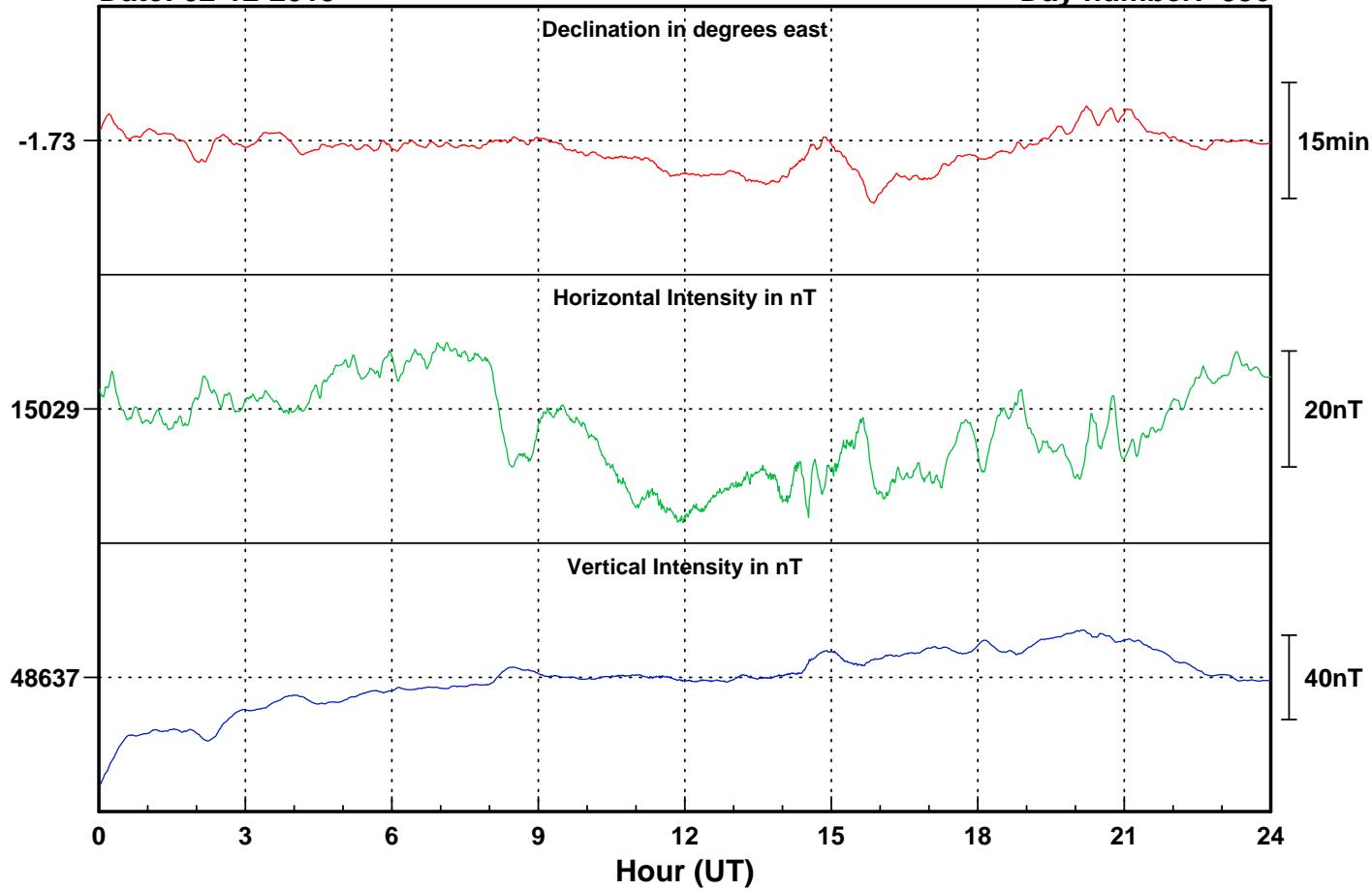
Lerwick

Day number: 335



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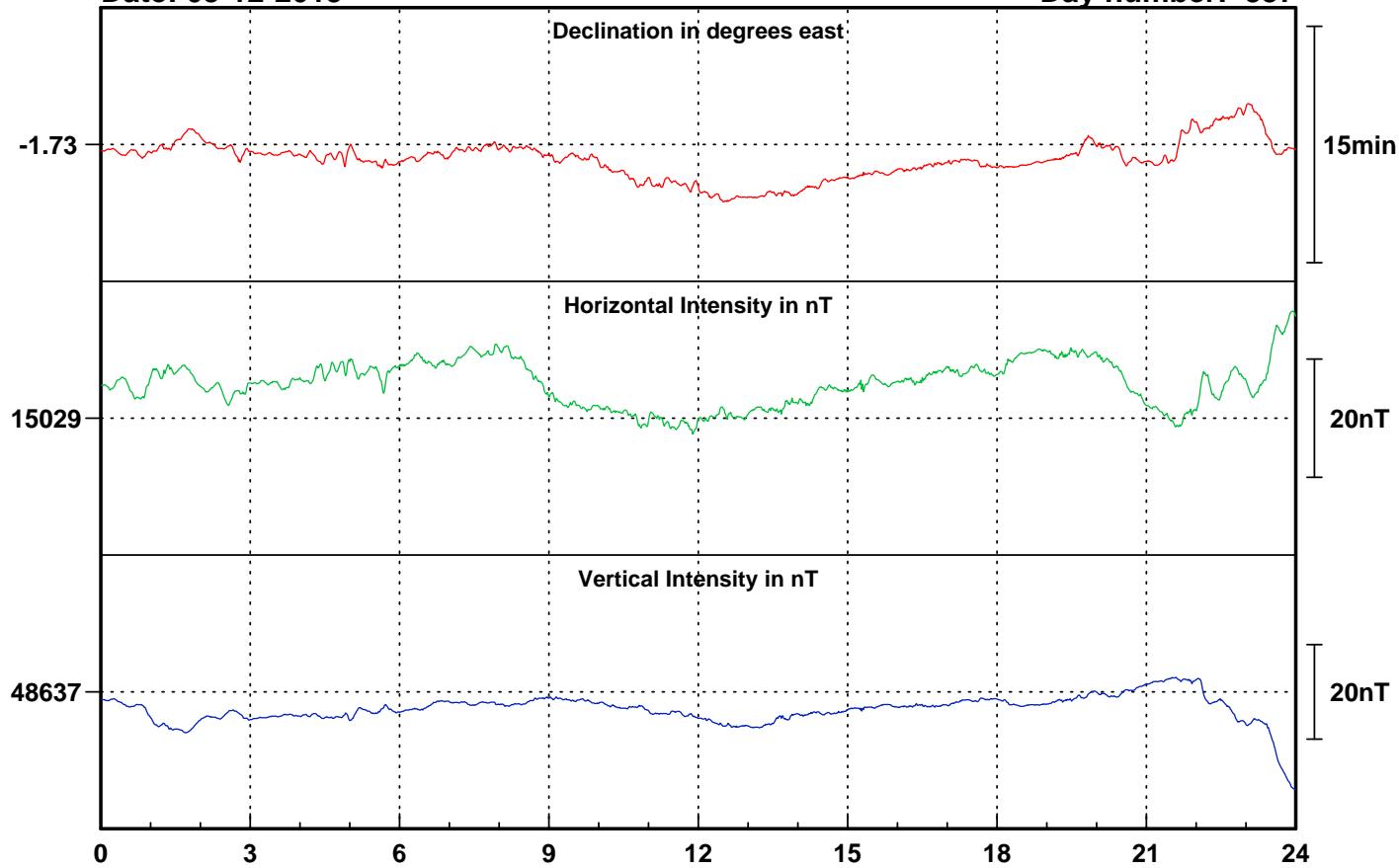
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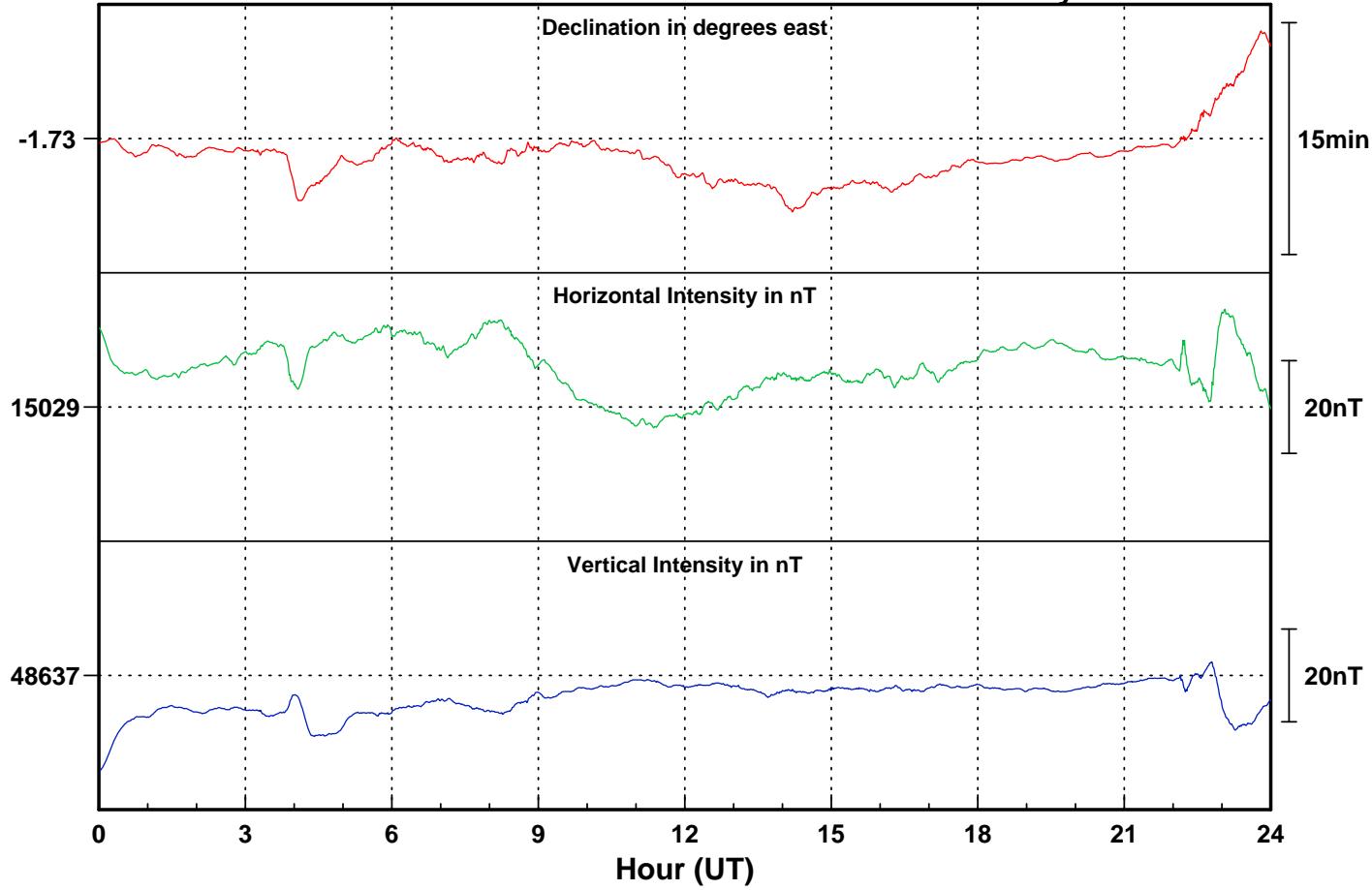
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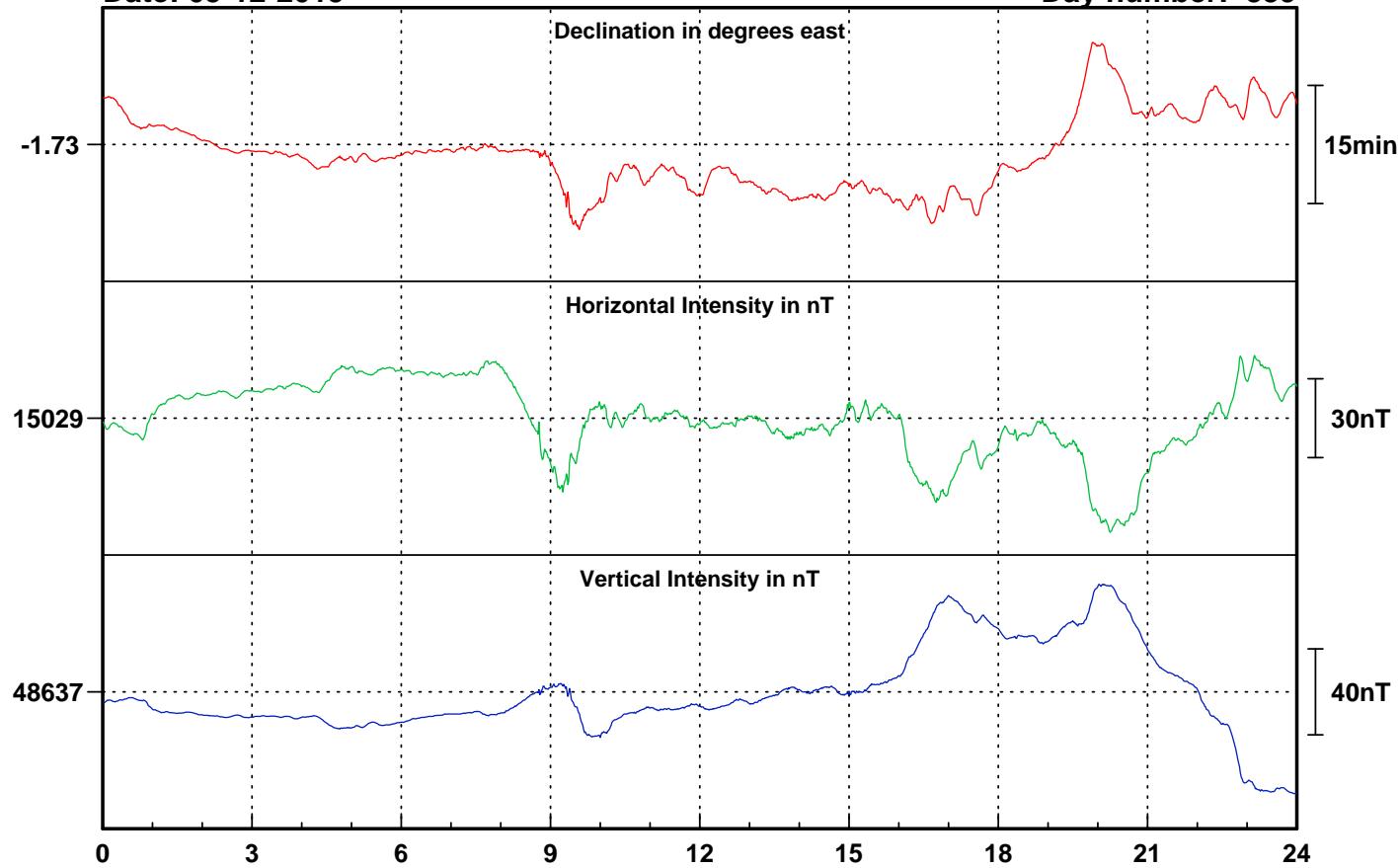
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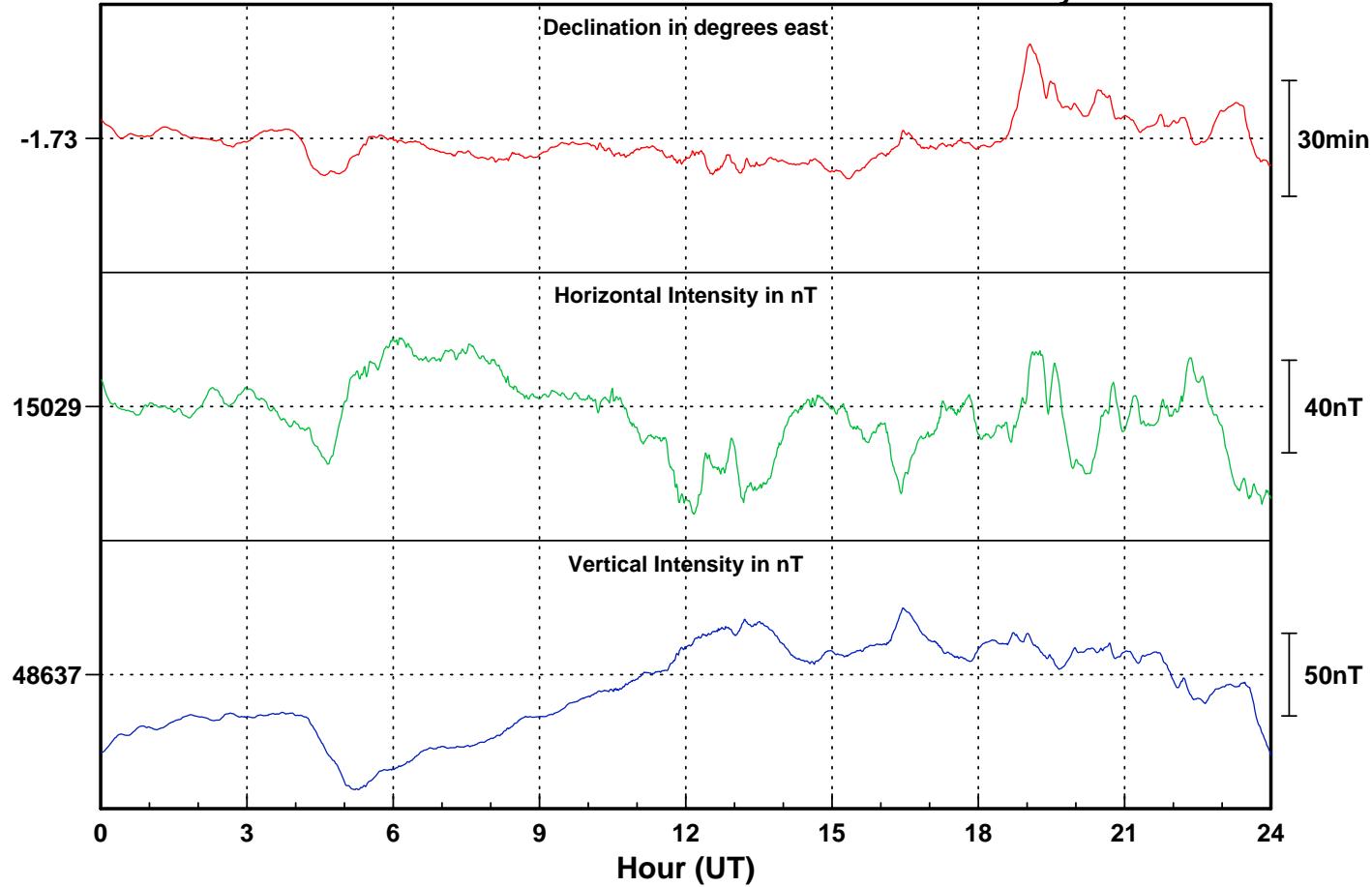
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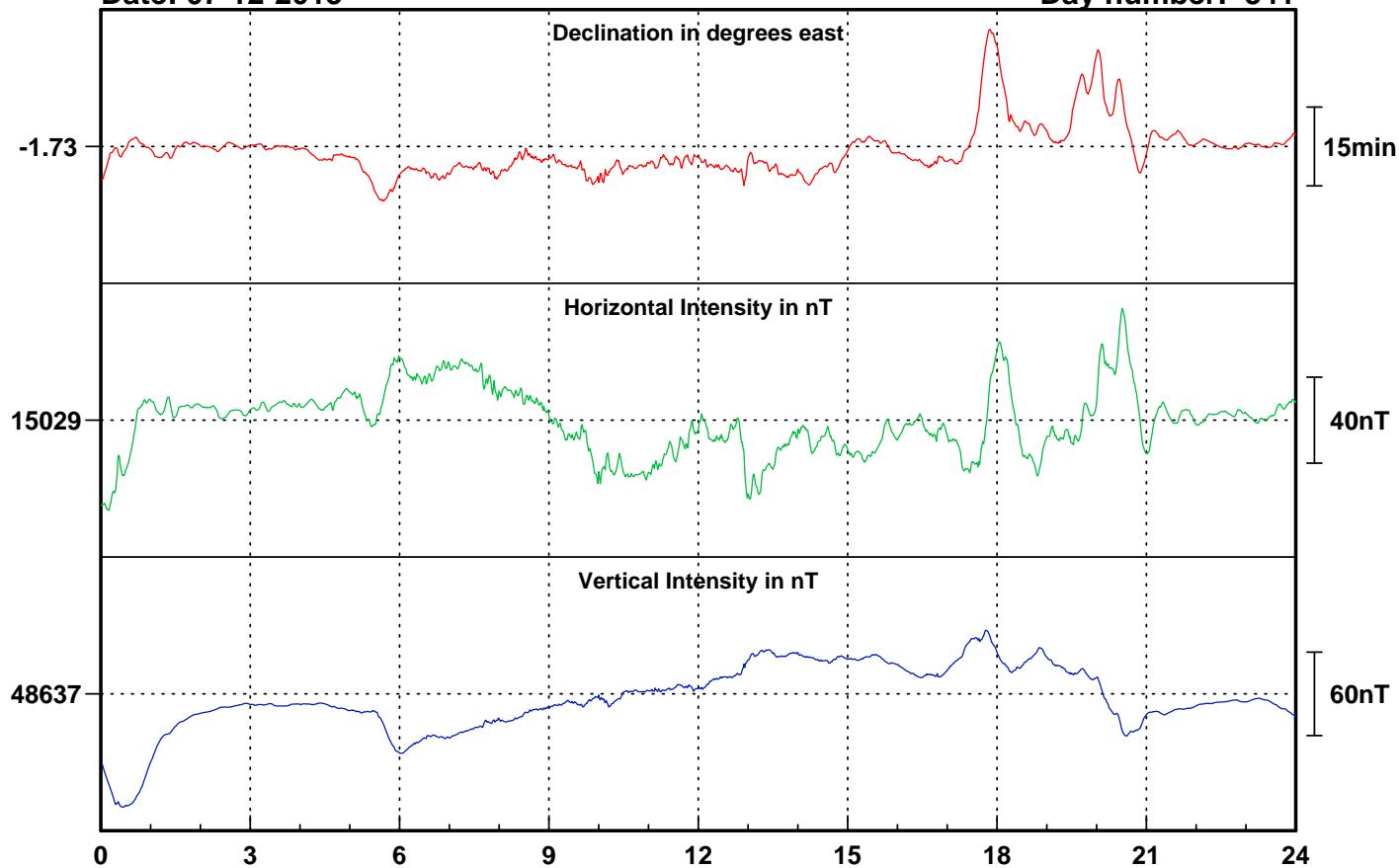
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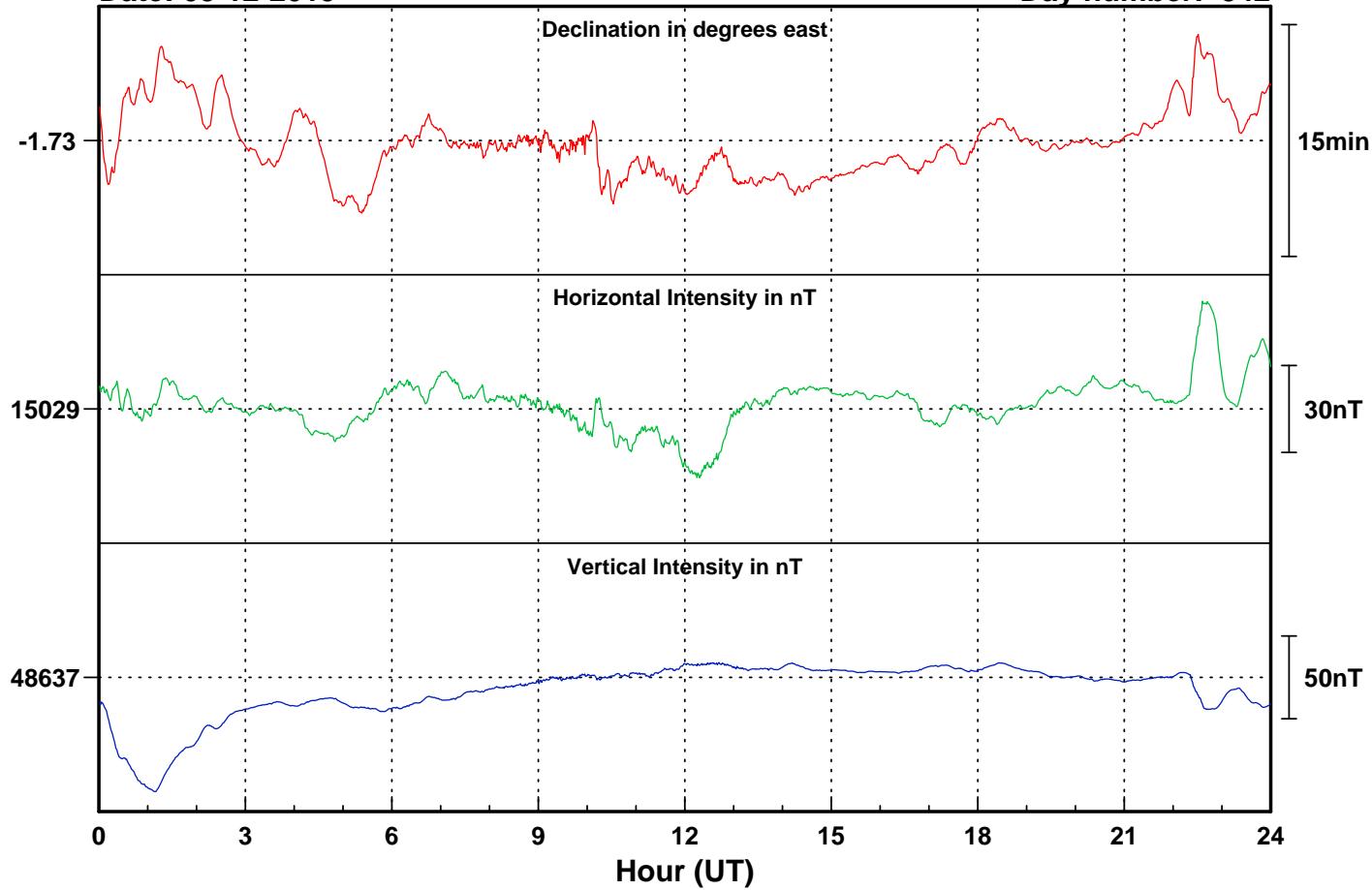
Lerwick

Day number: 341



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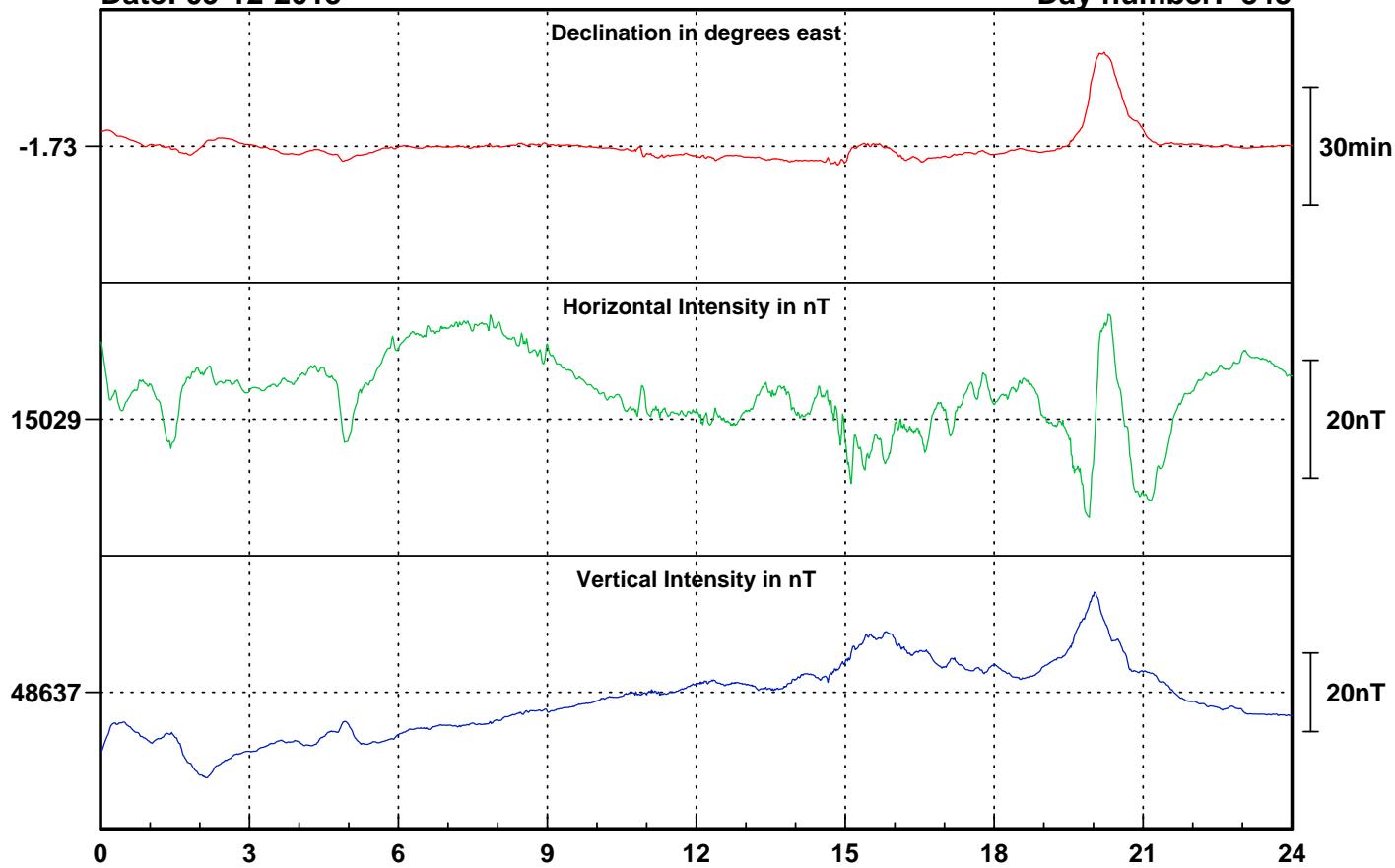
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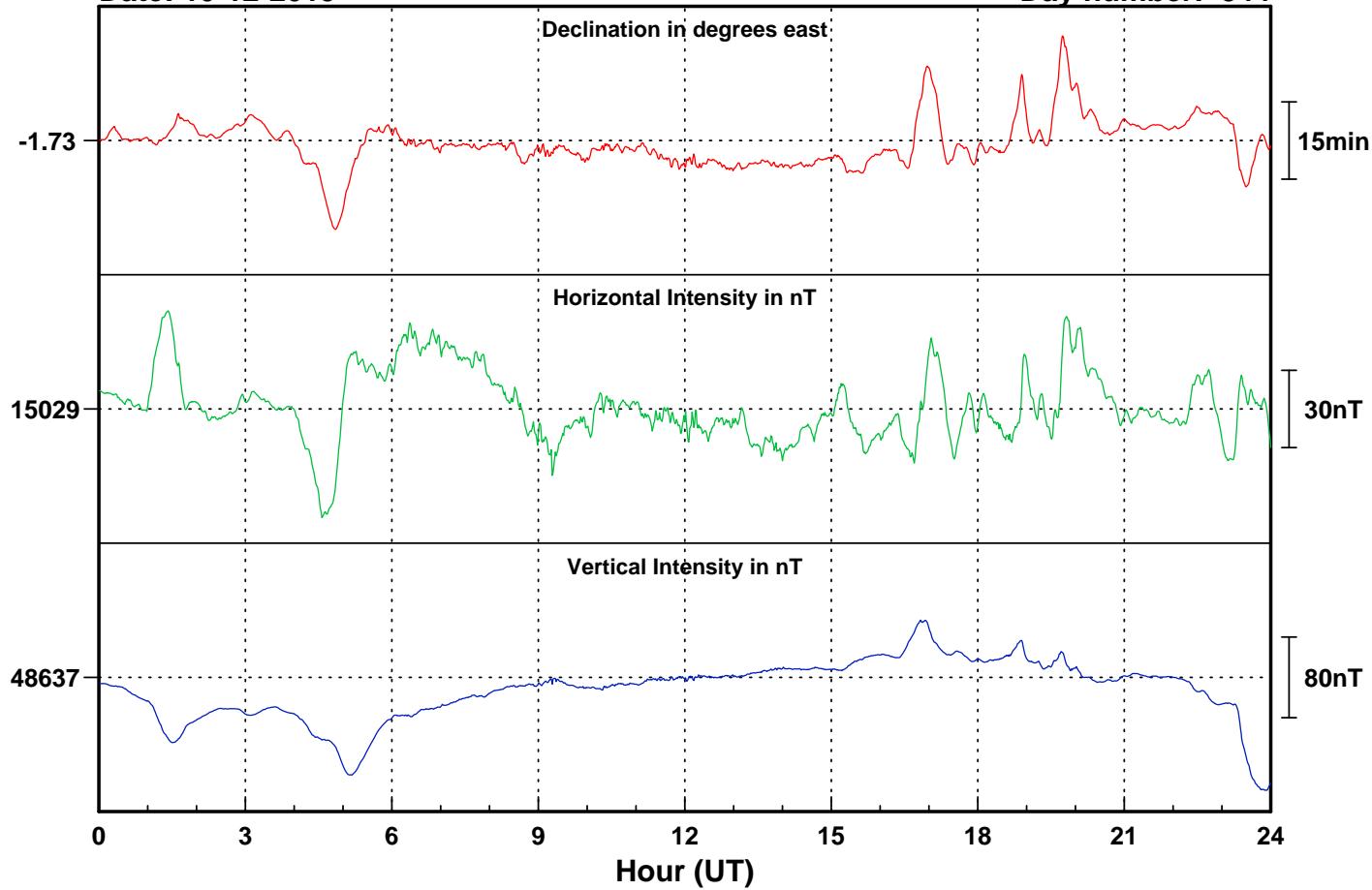
Lerwick

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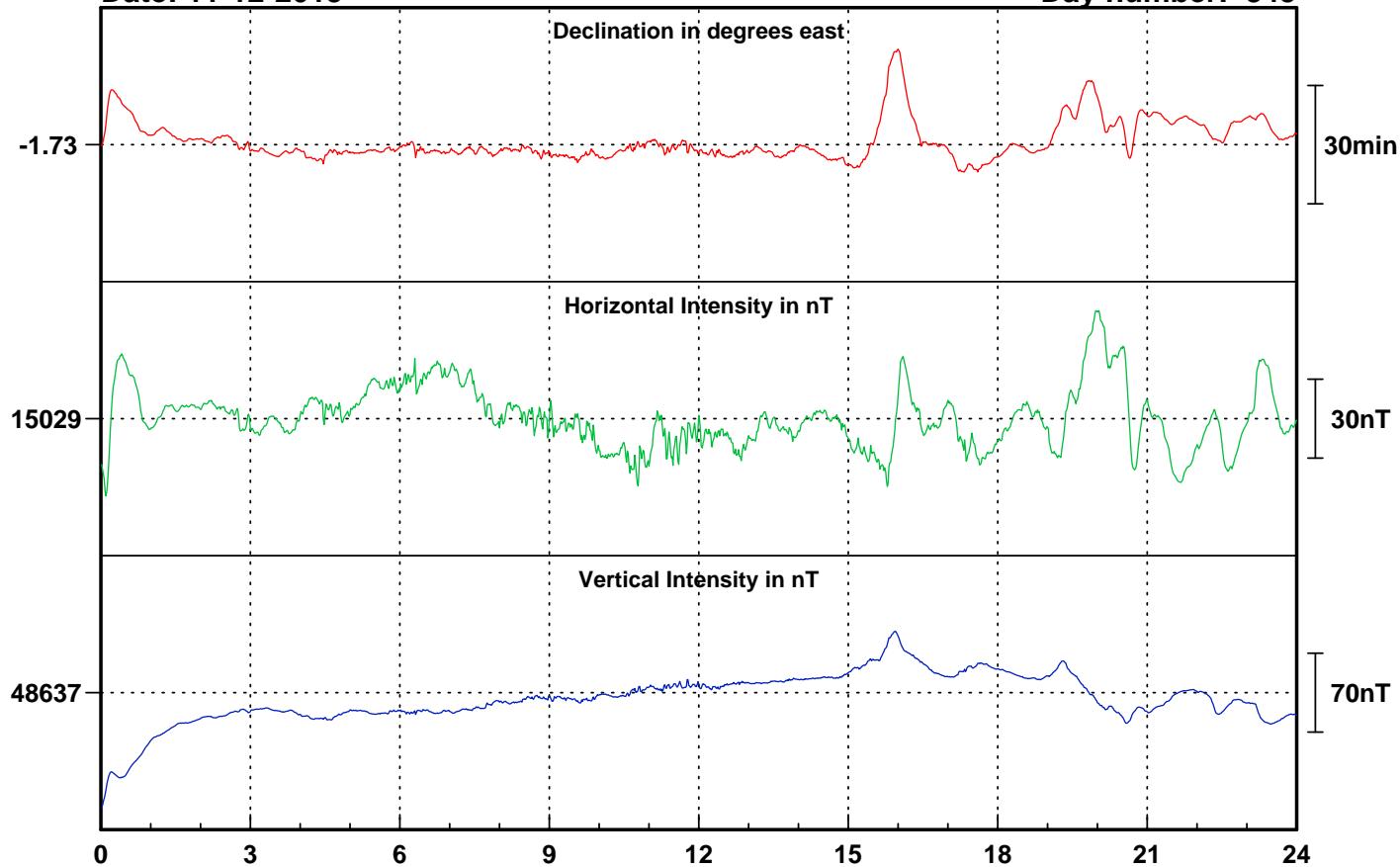
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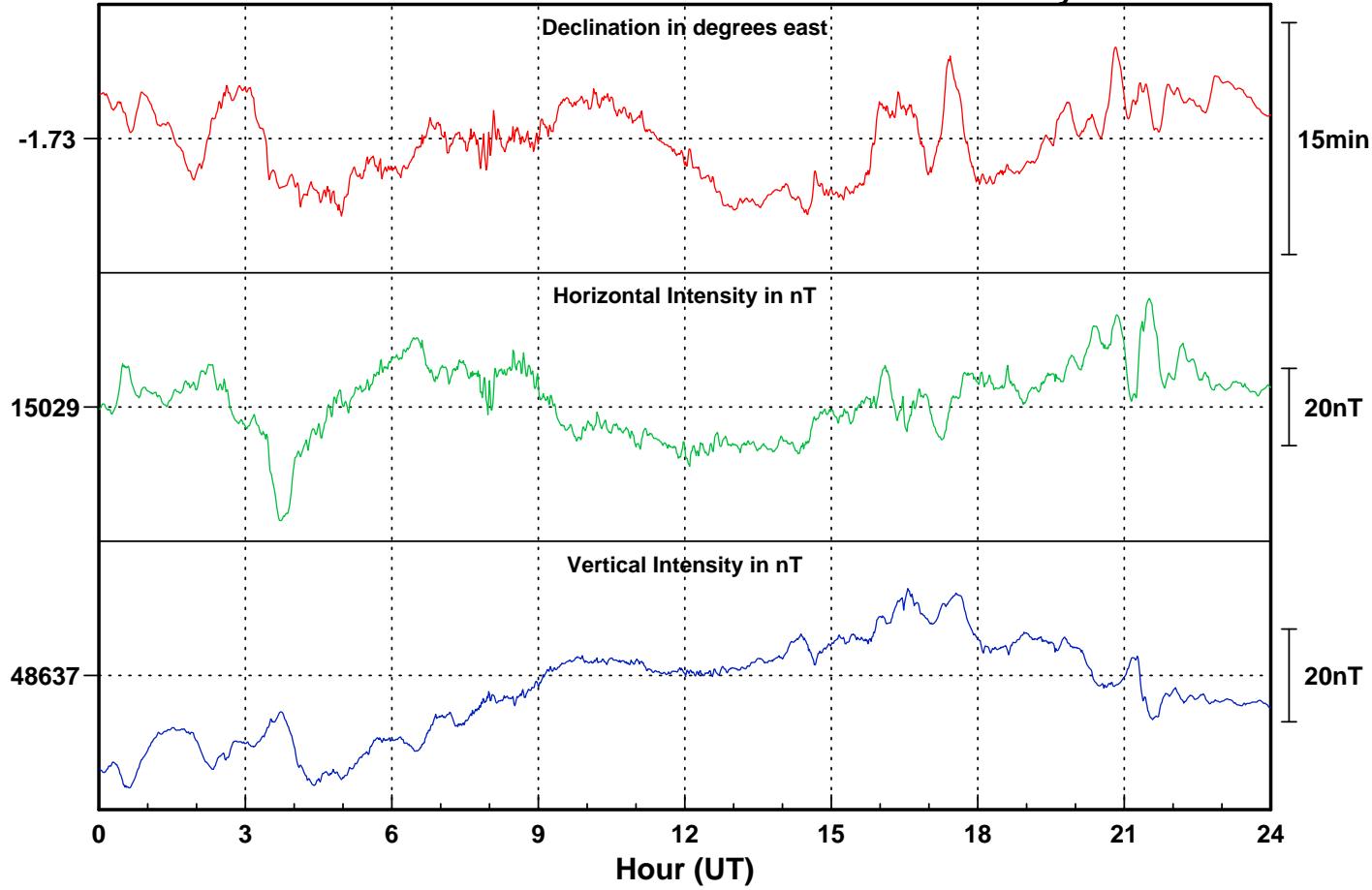
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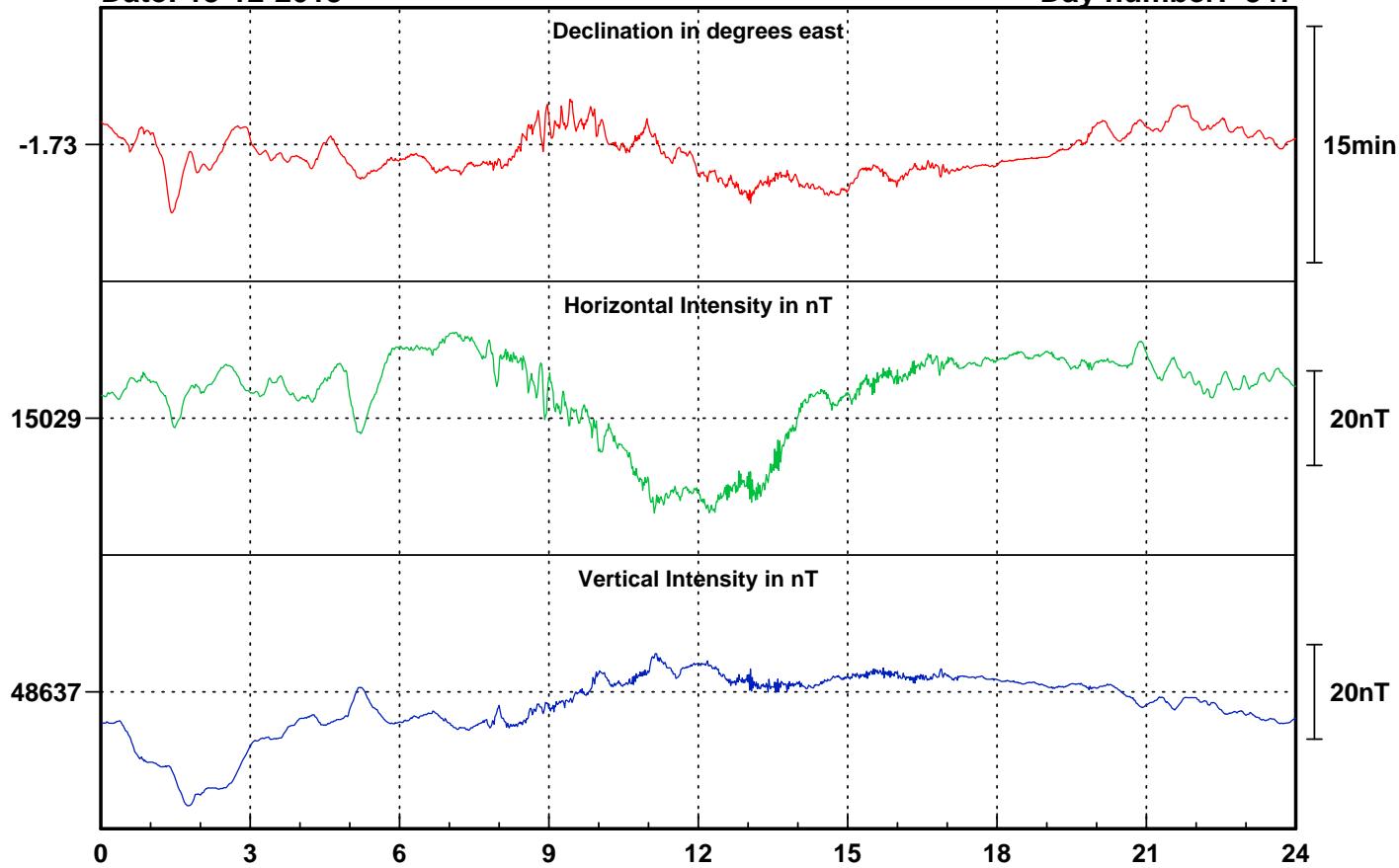
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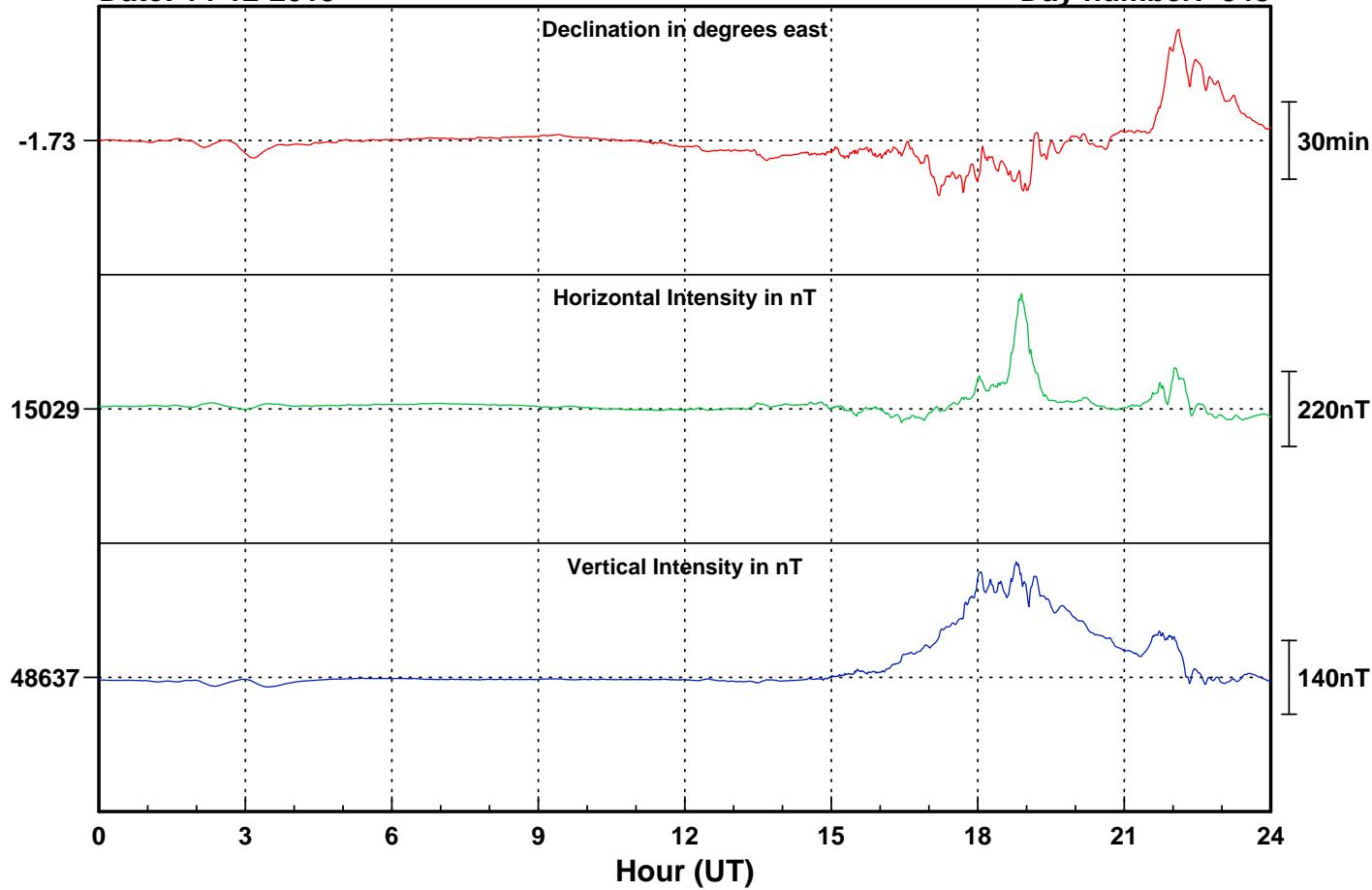
Lerwick

Day number: 347



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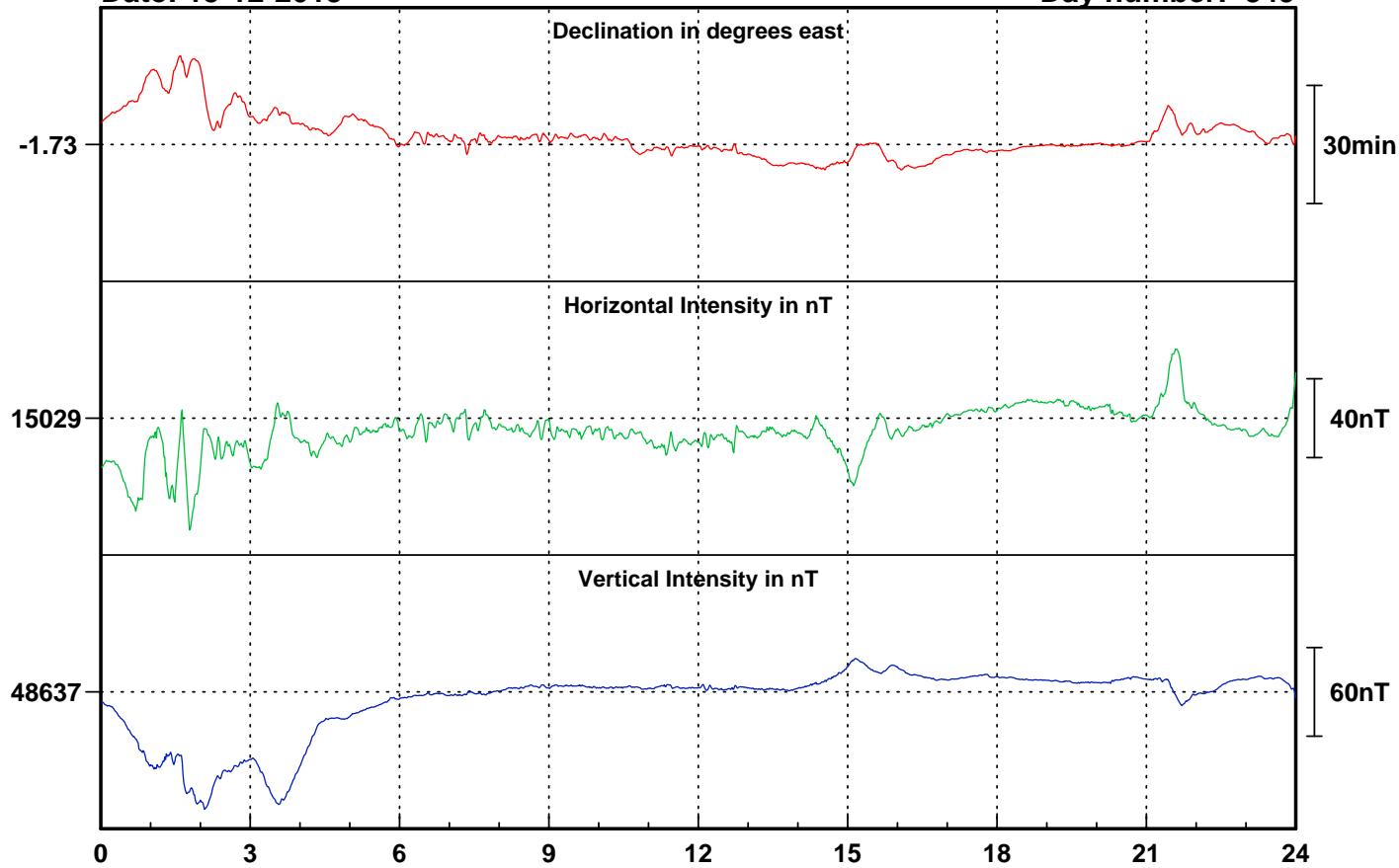
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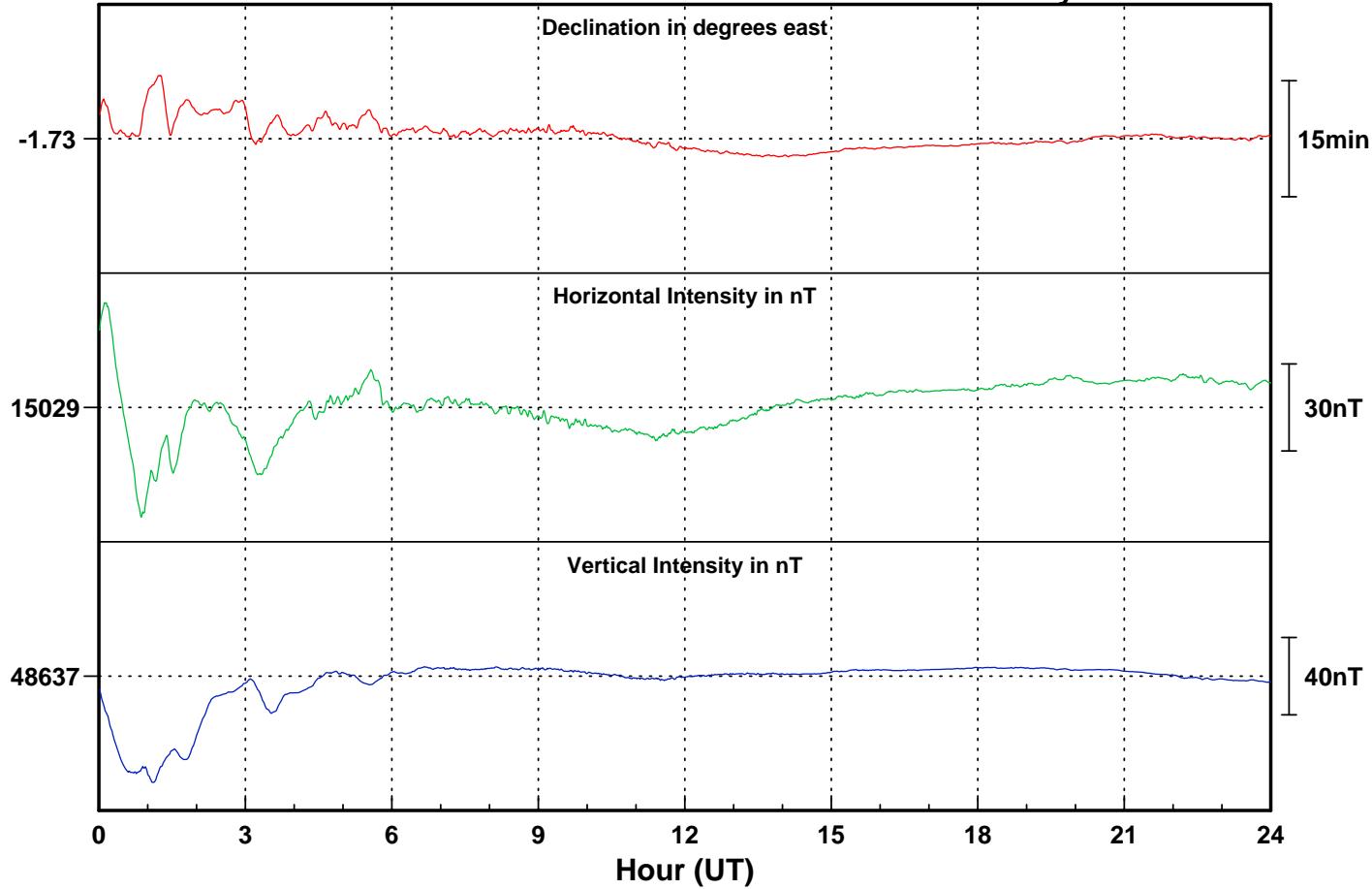
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Day number: 349



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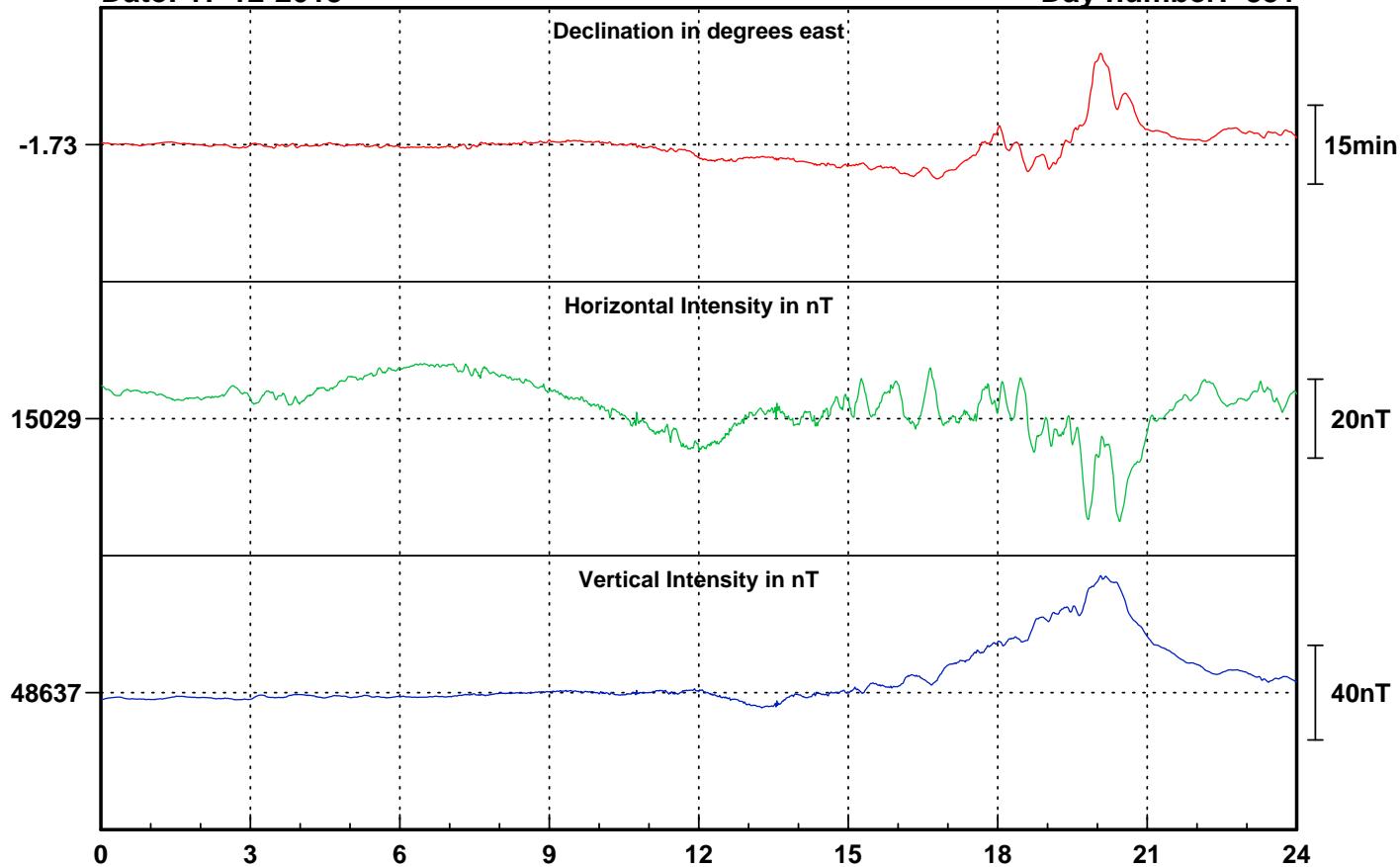
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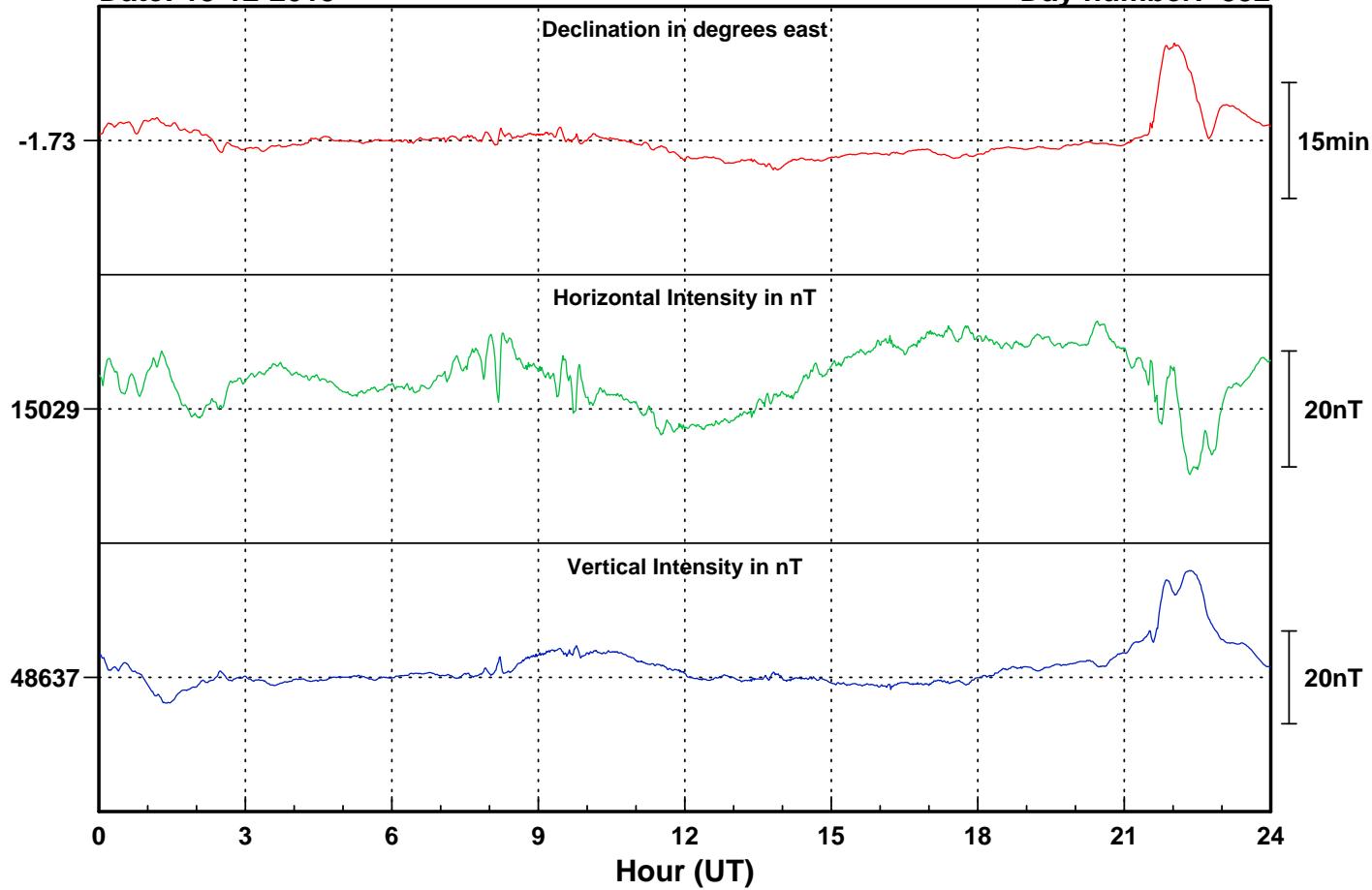
Lerwick

Day number: 351



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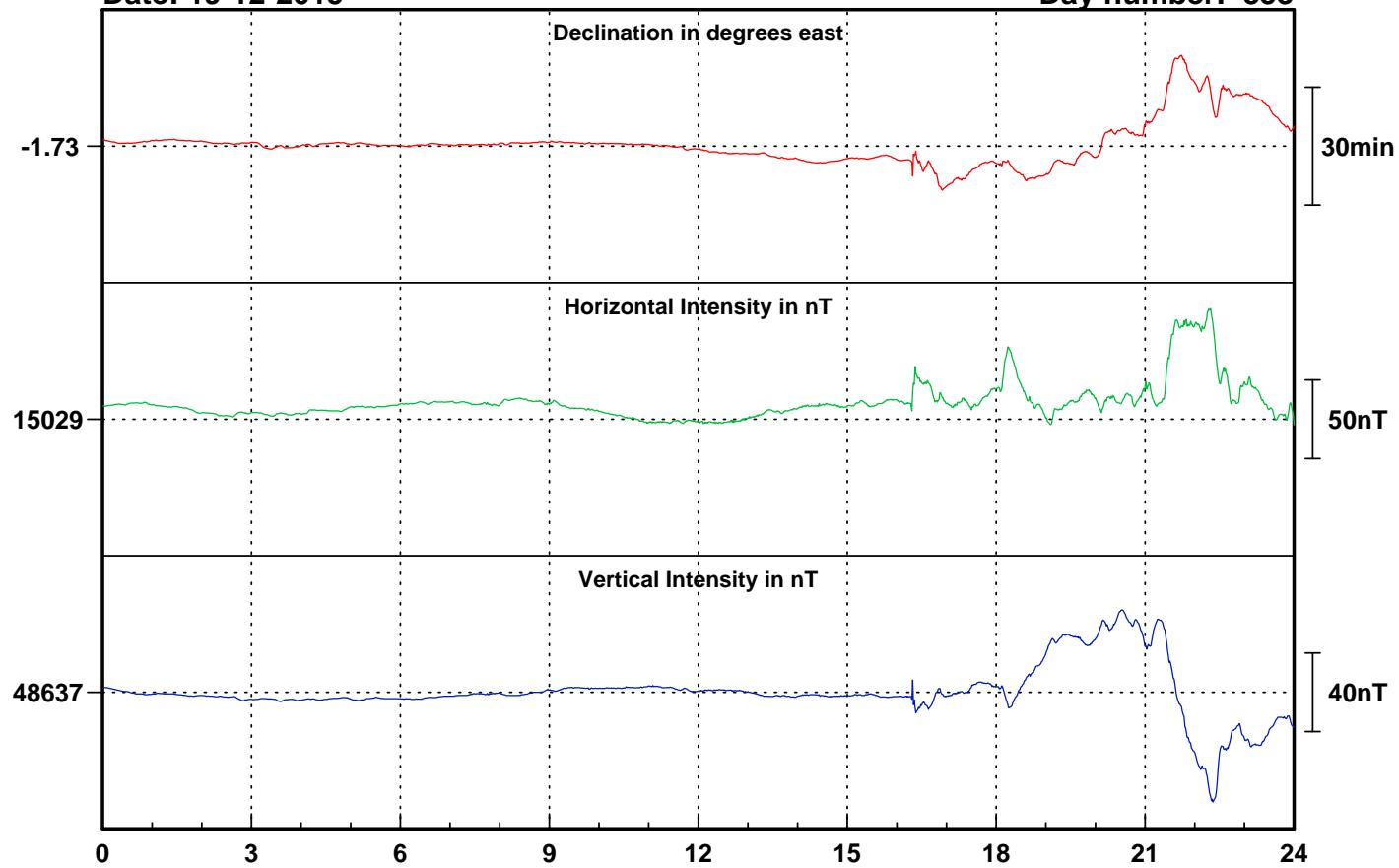
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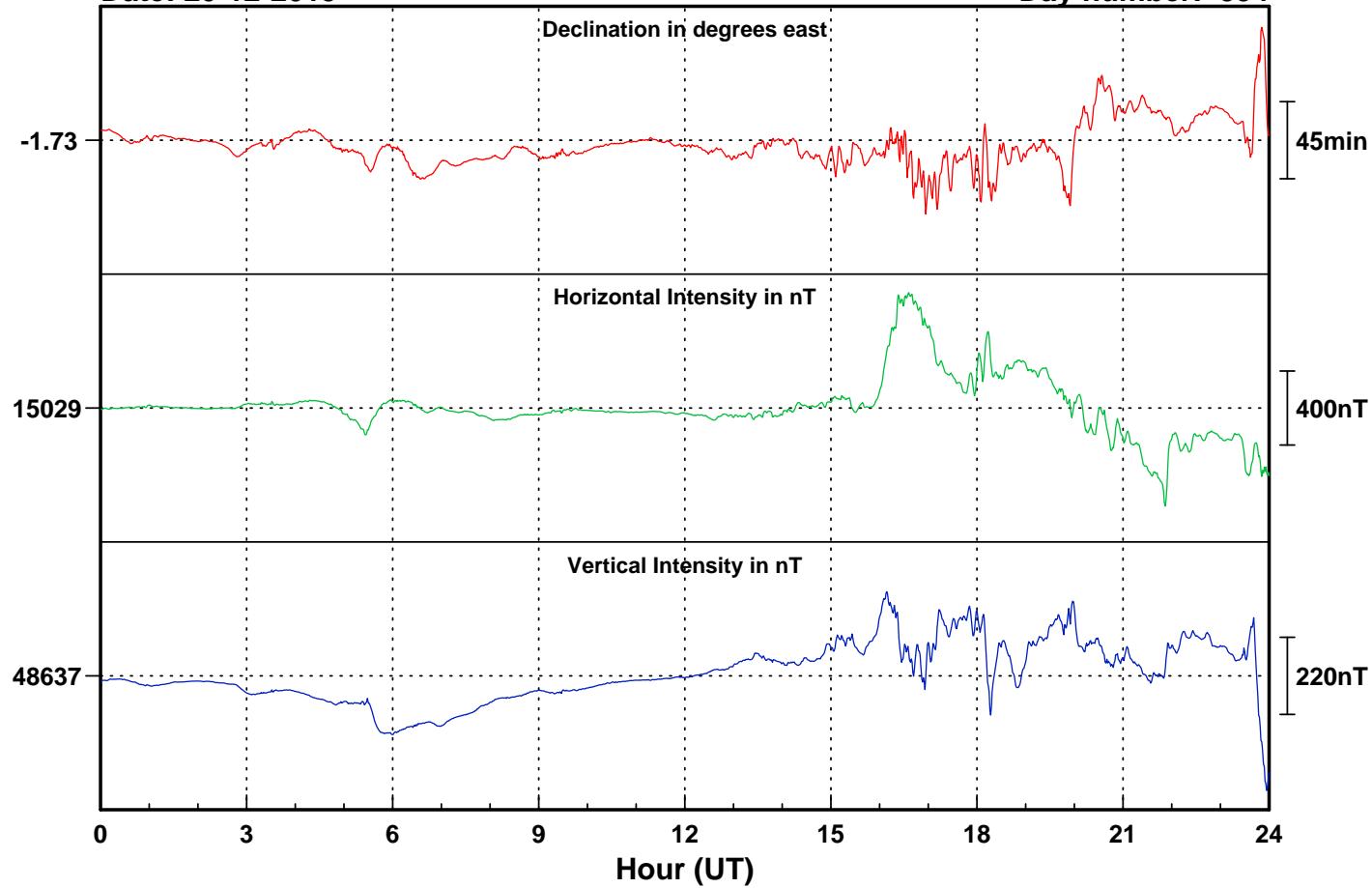
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Day number: 353



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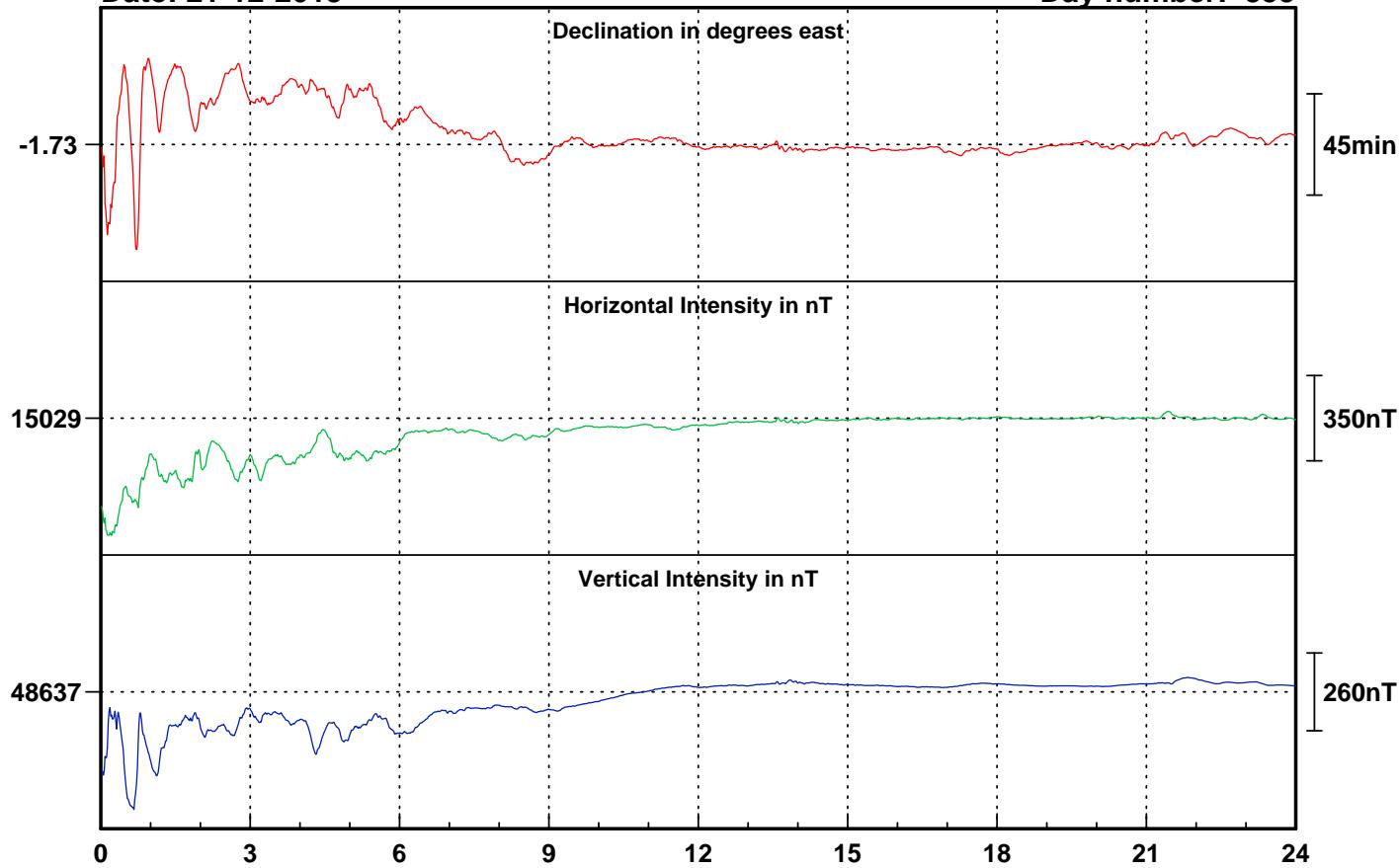
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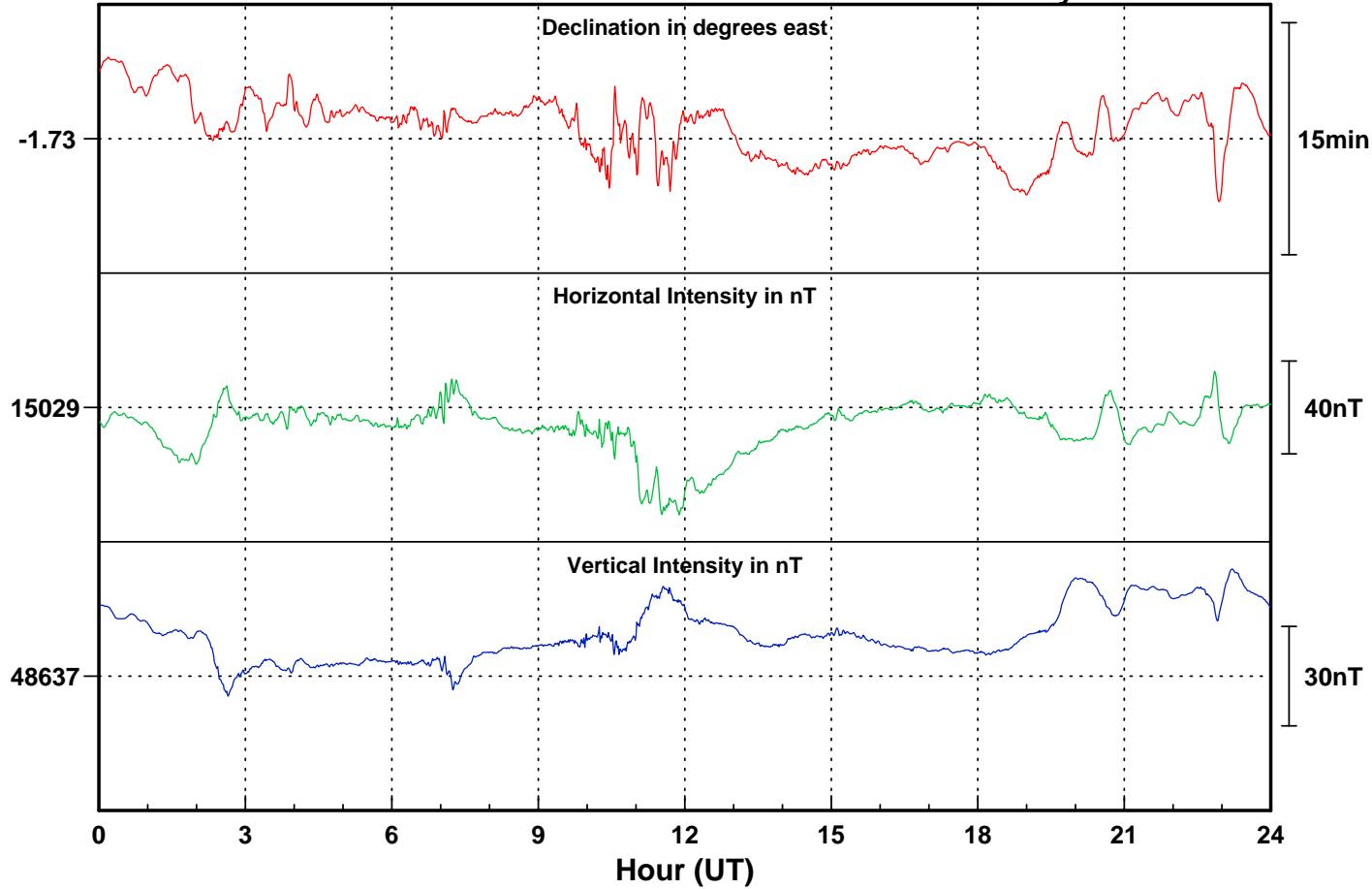
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Day number: 355



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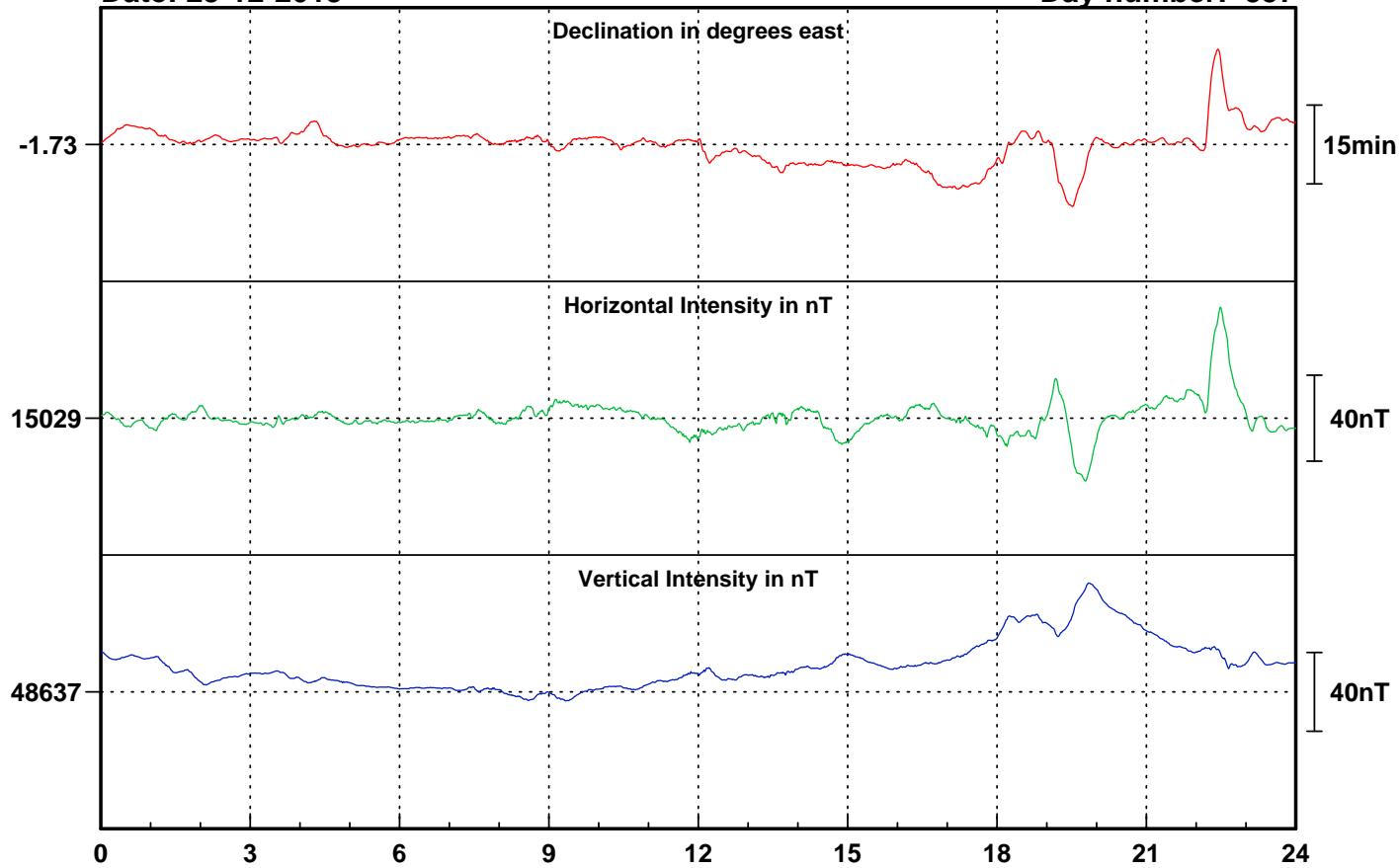
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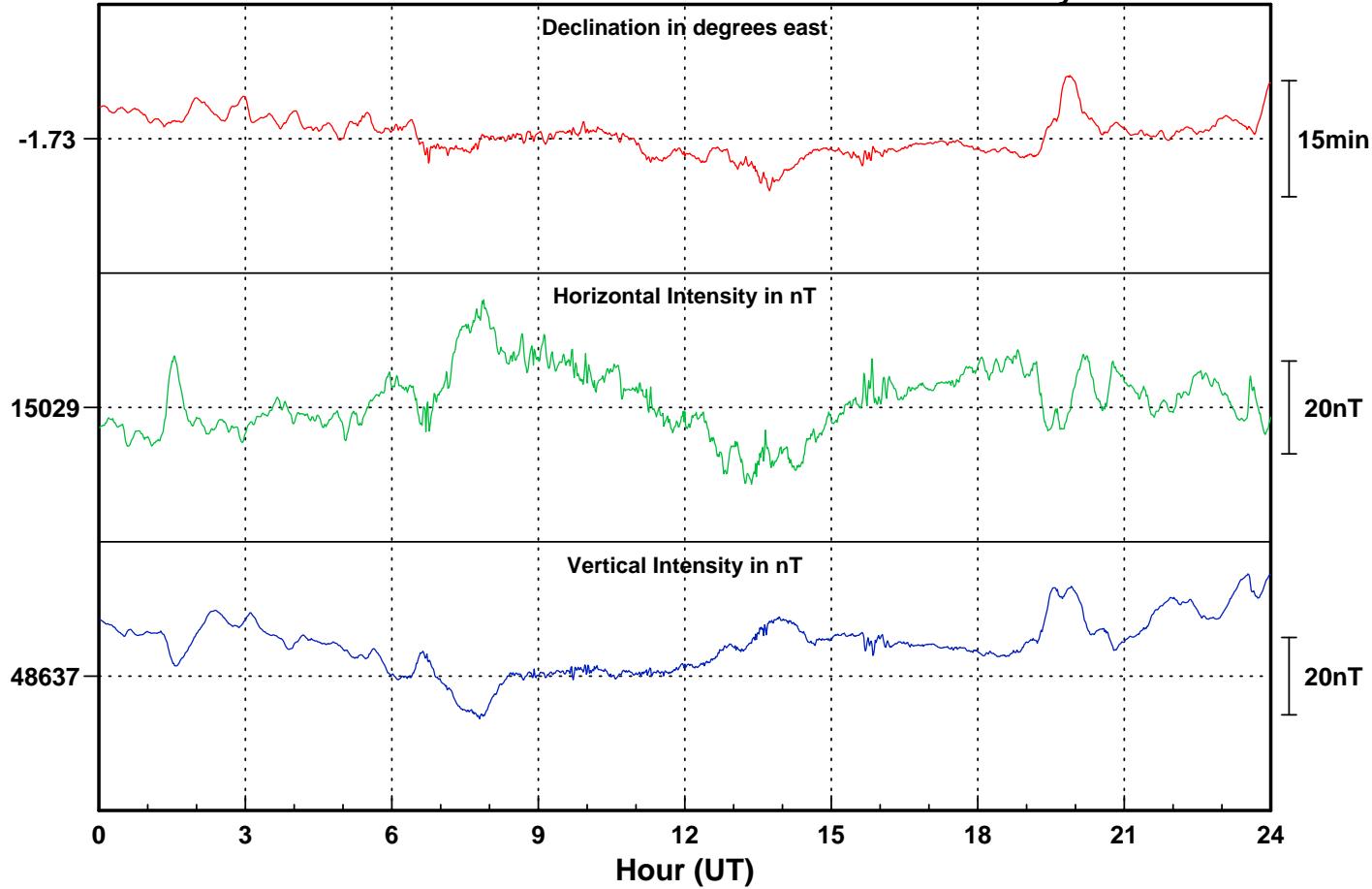
Lerwick

Day number: 357



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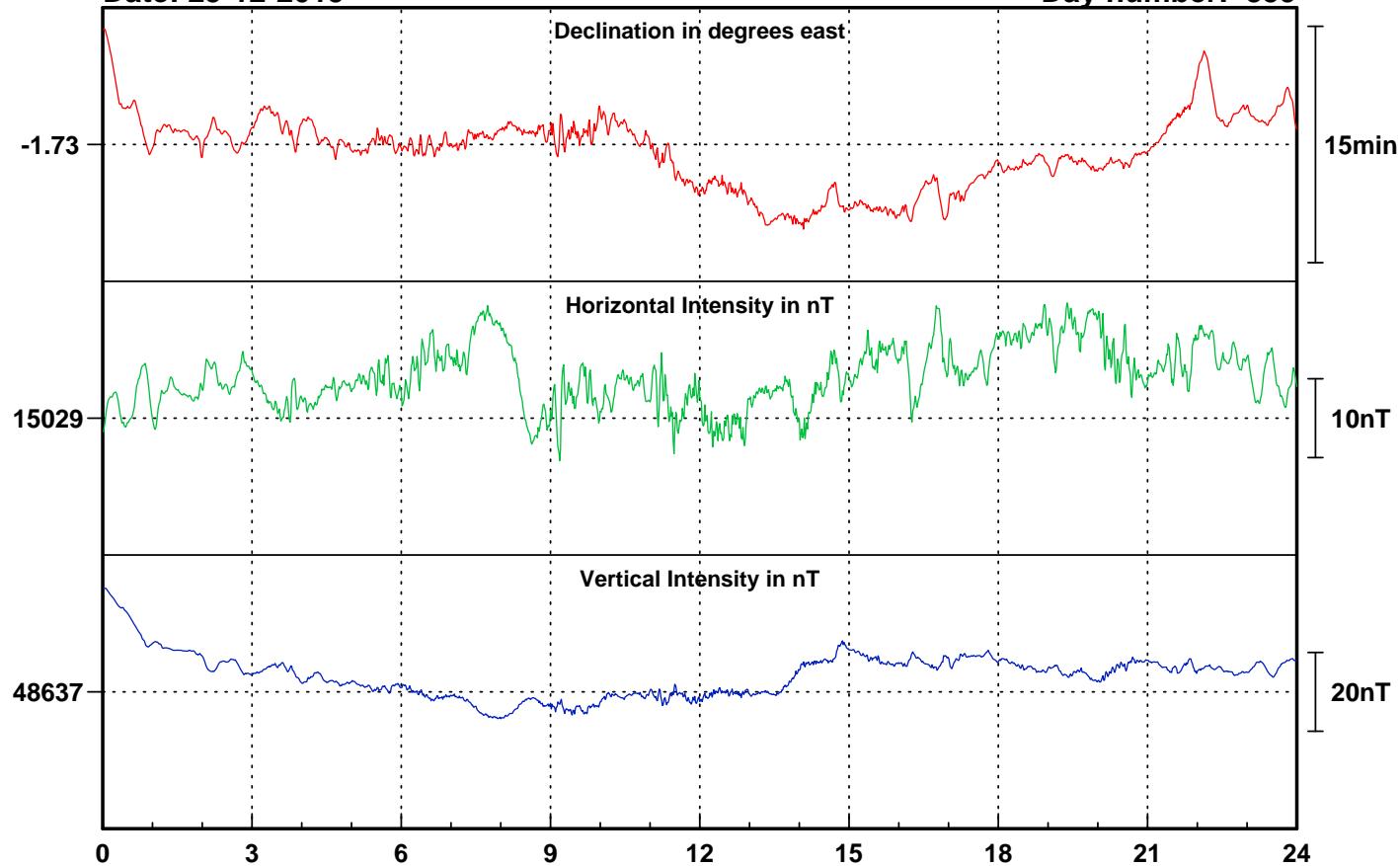
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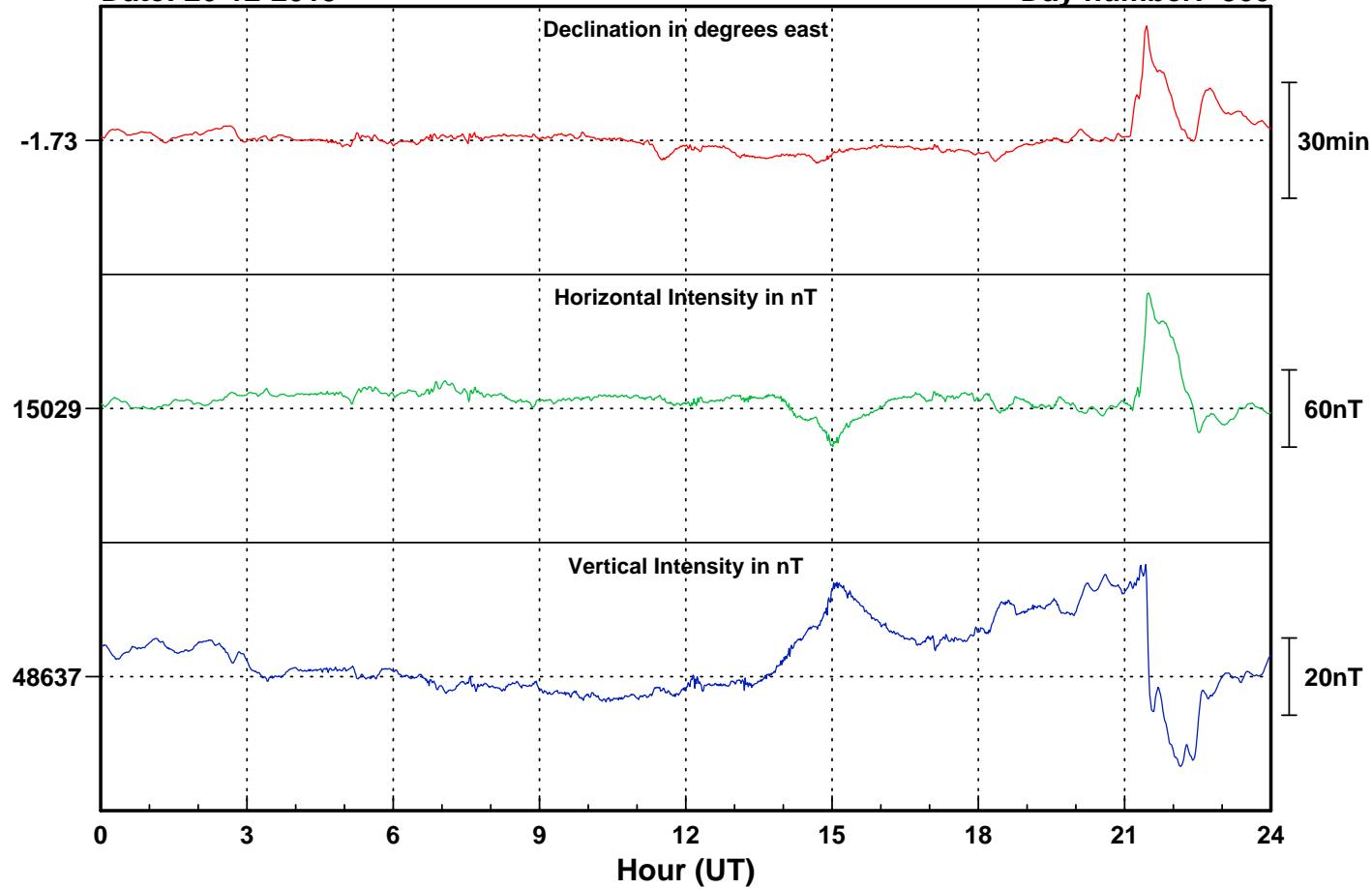
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Day number: 359



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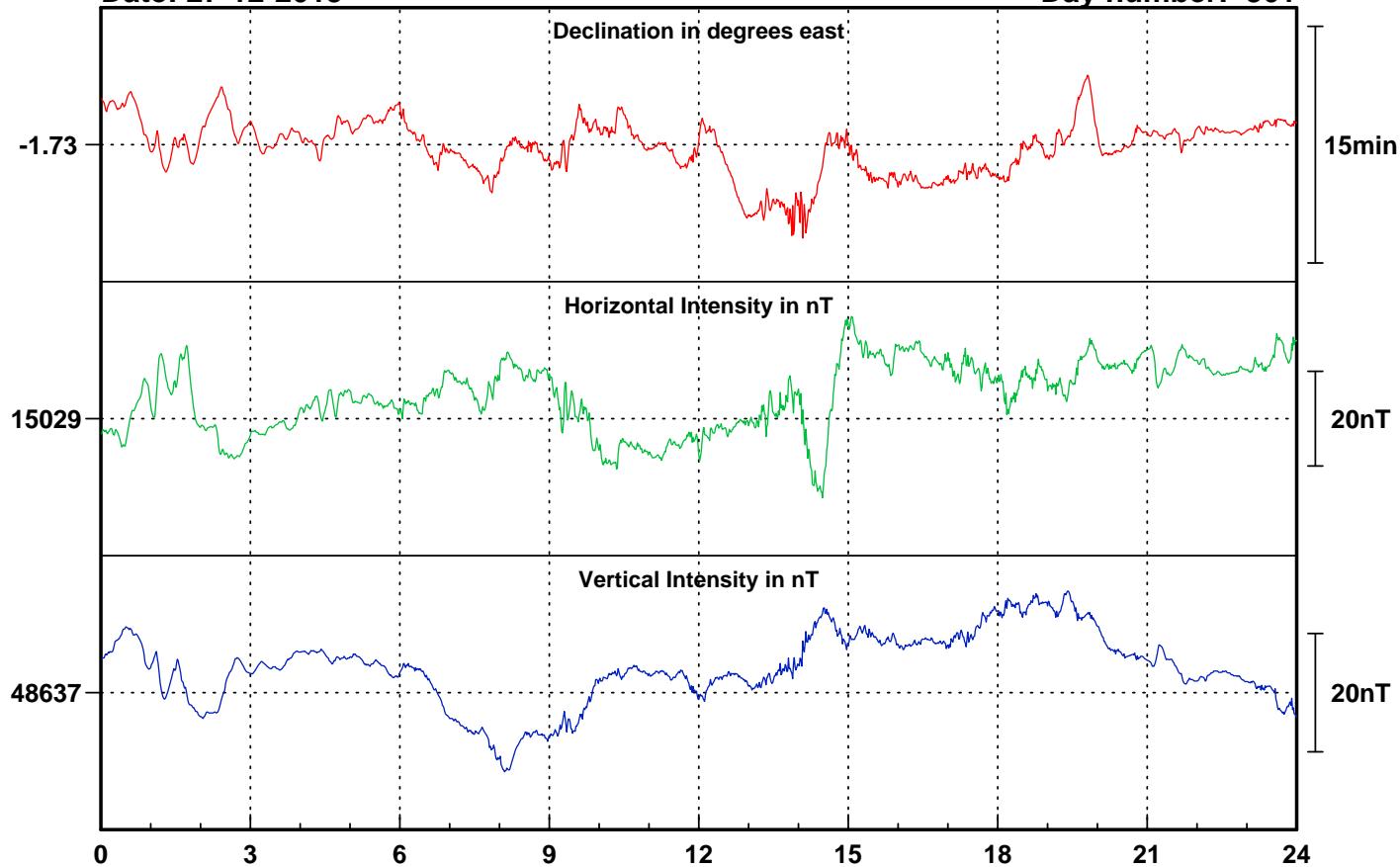
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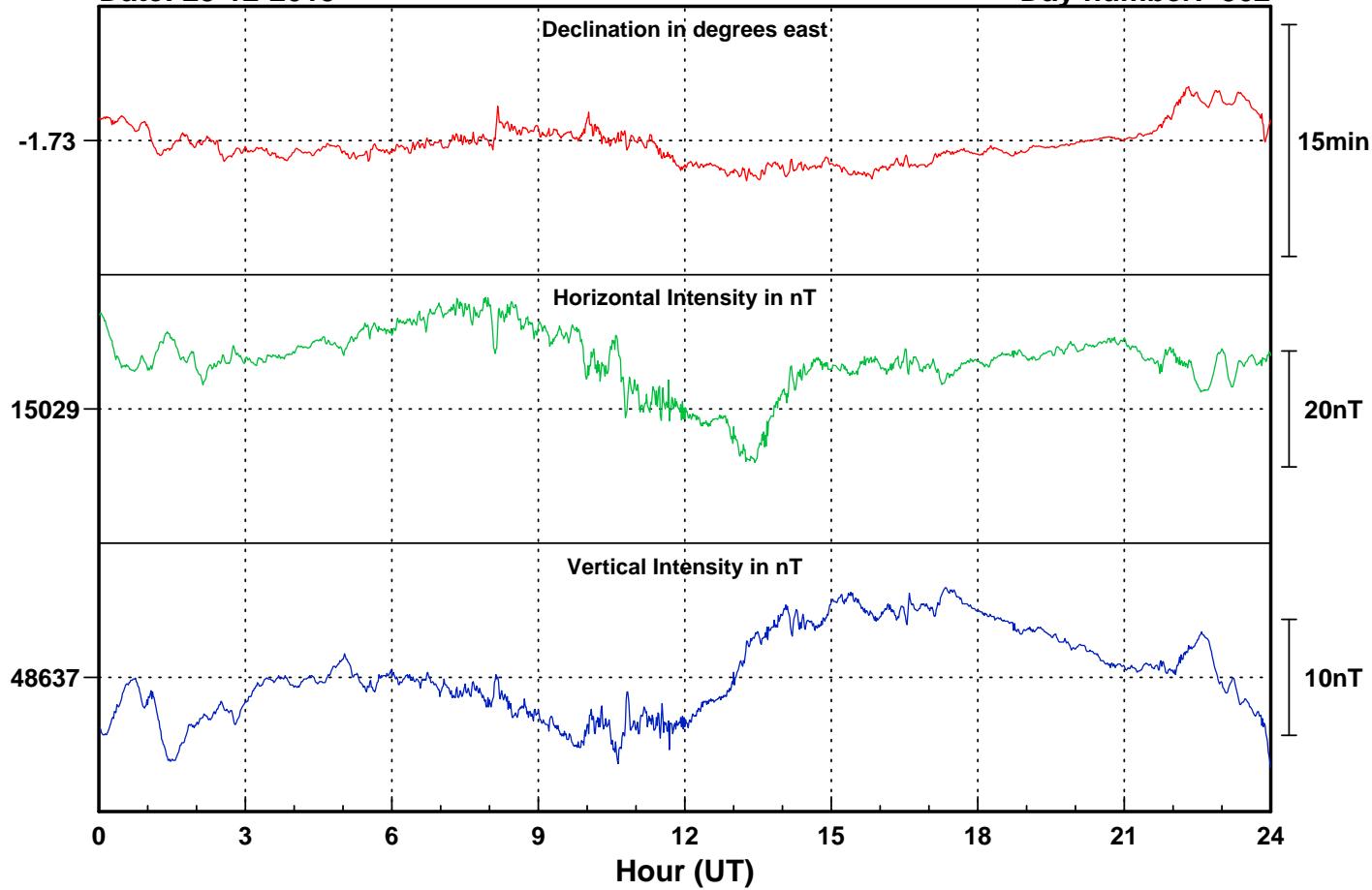
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Day number: 361



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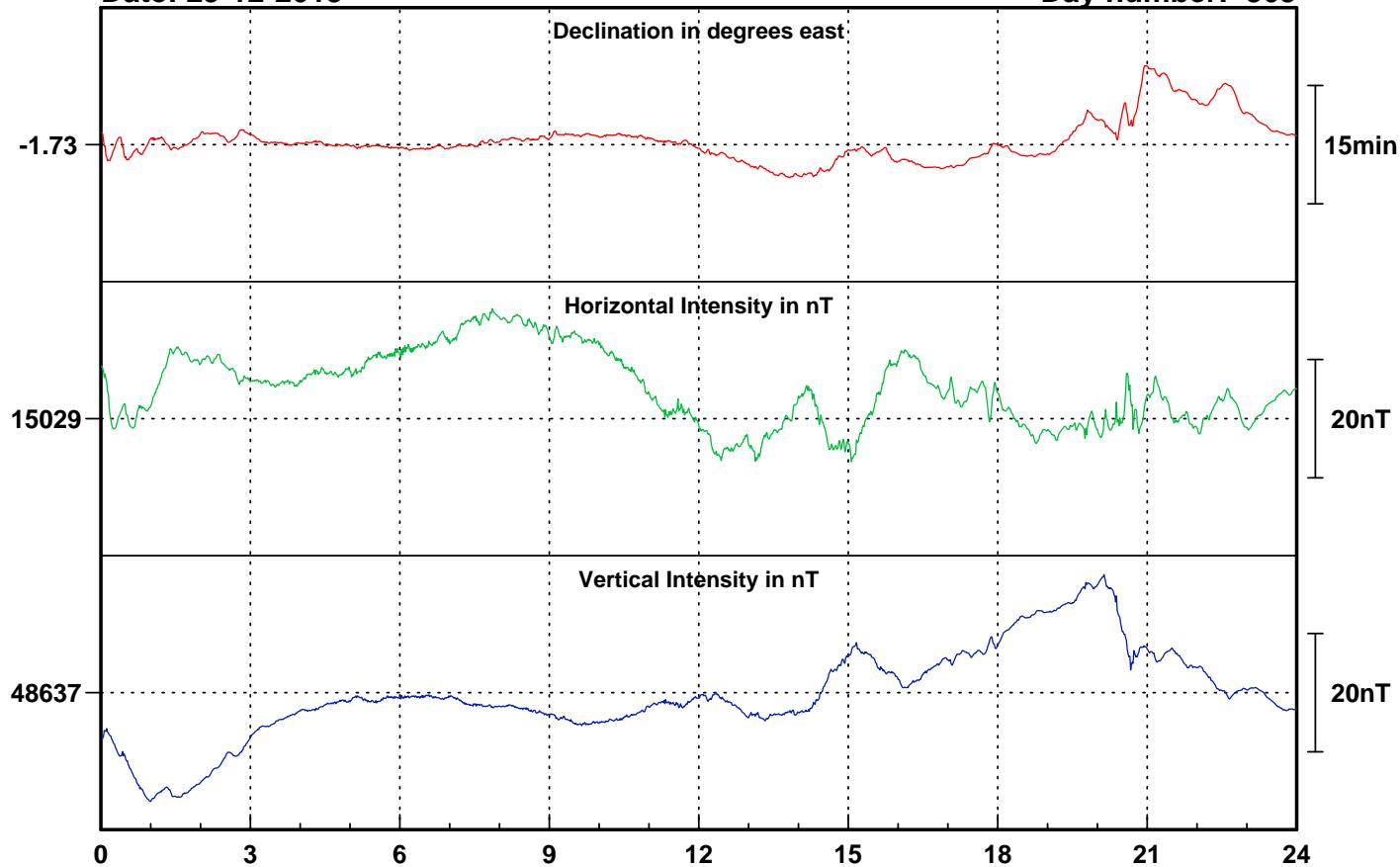
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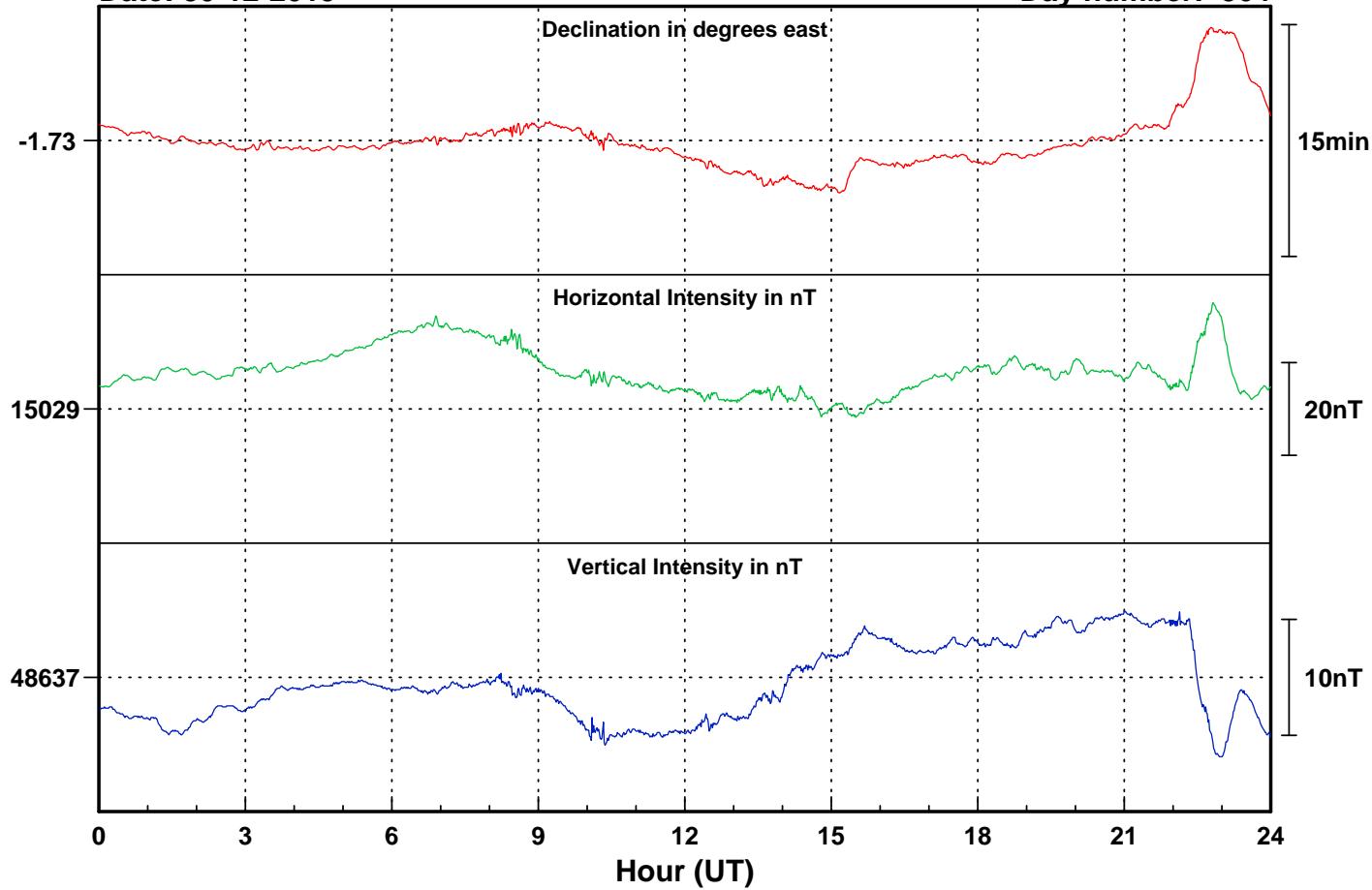
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Day number: 363



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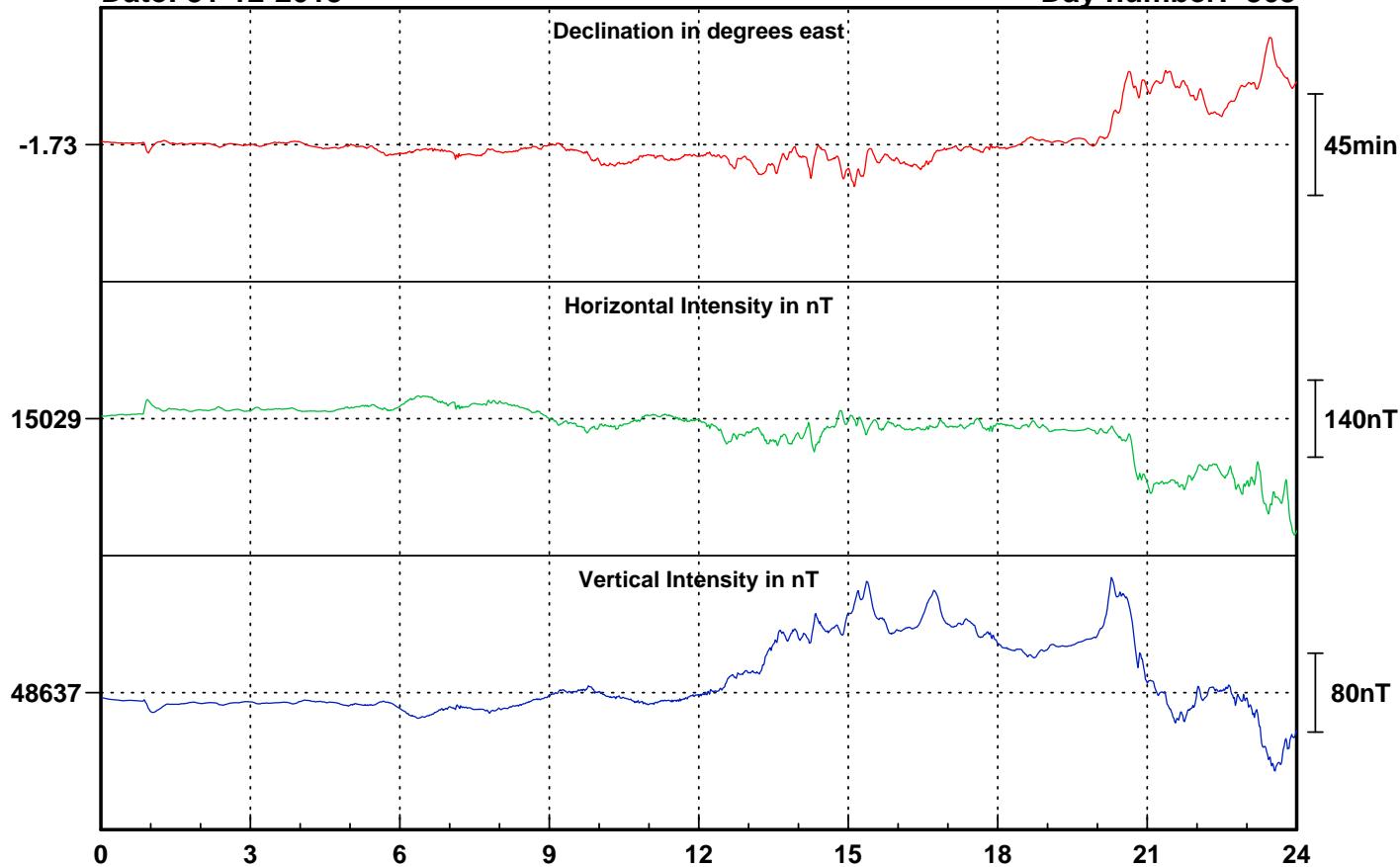
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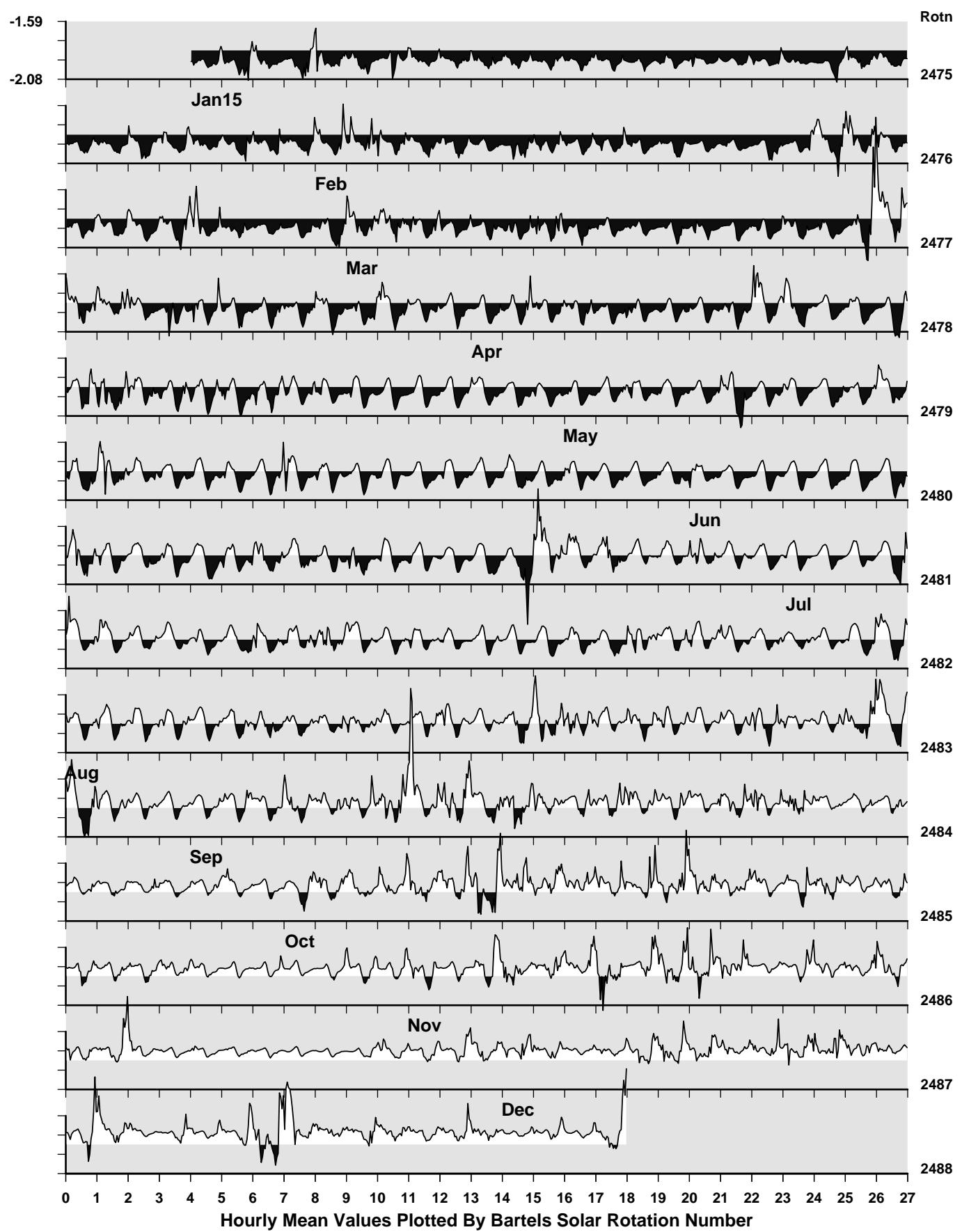
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Lerwick

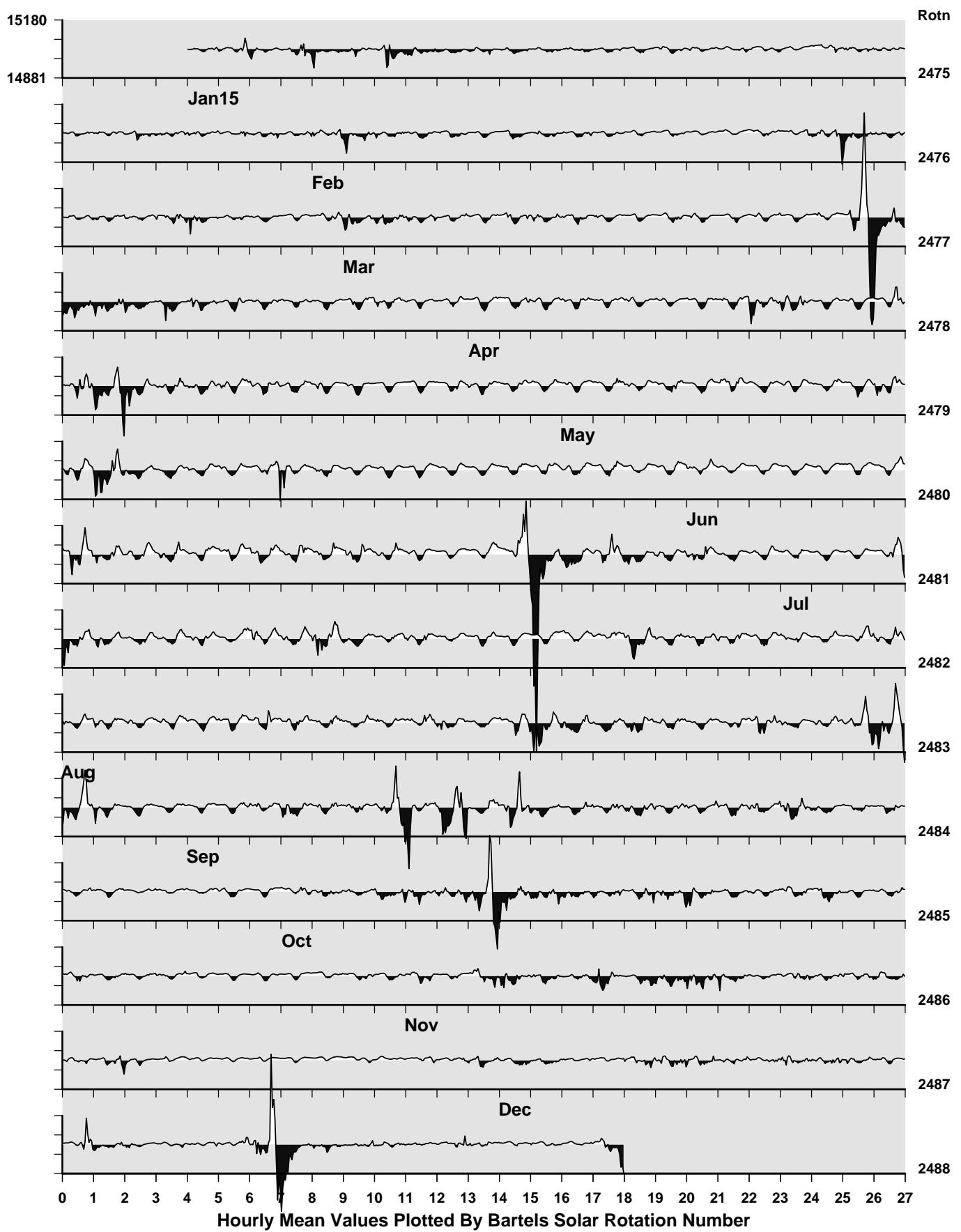
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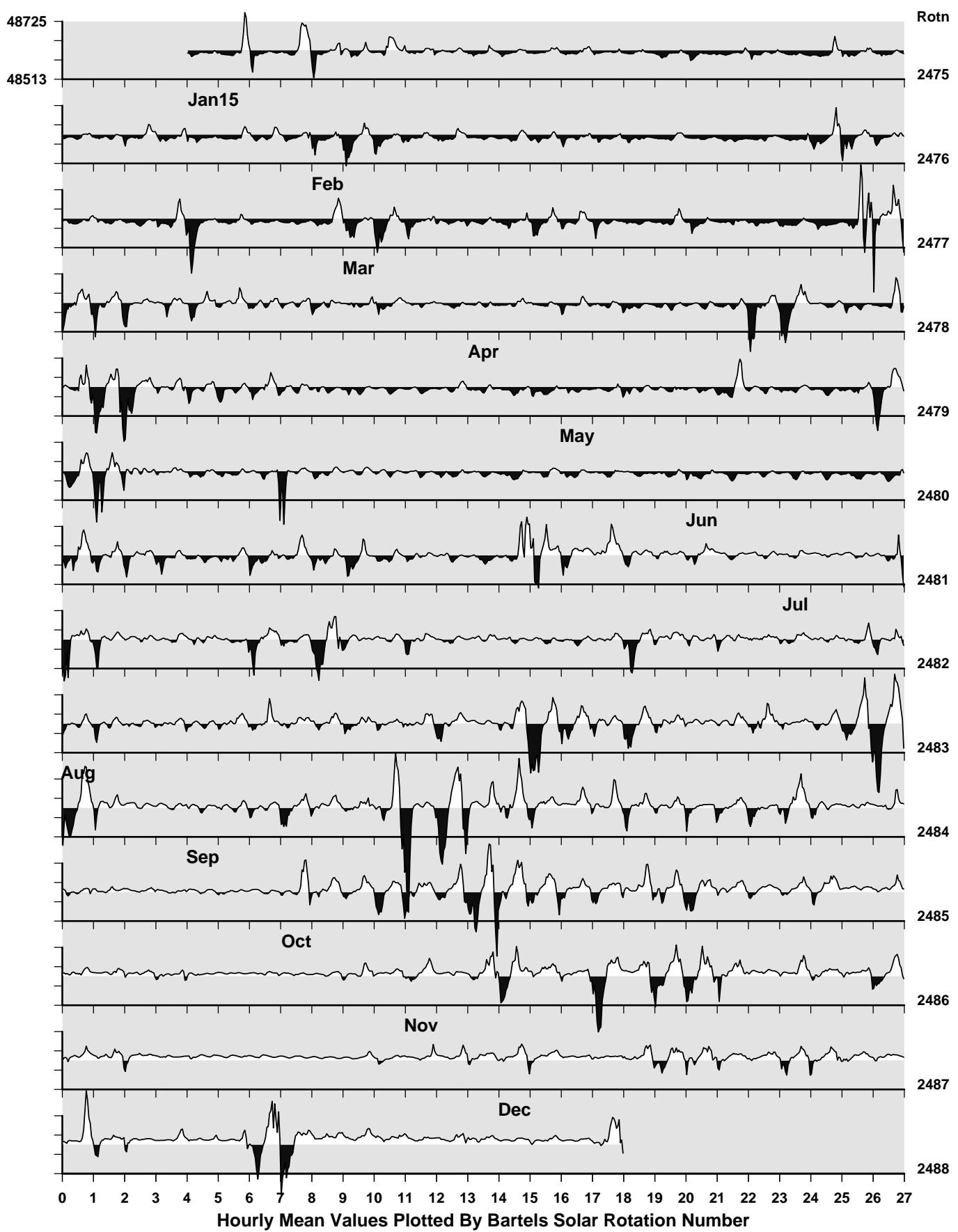
Lerwick Observatory: Declination (degrees)



Lerwick Observatory: Horizontal Intensity fhT_L

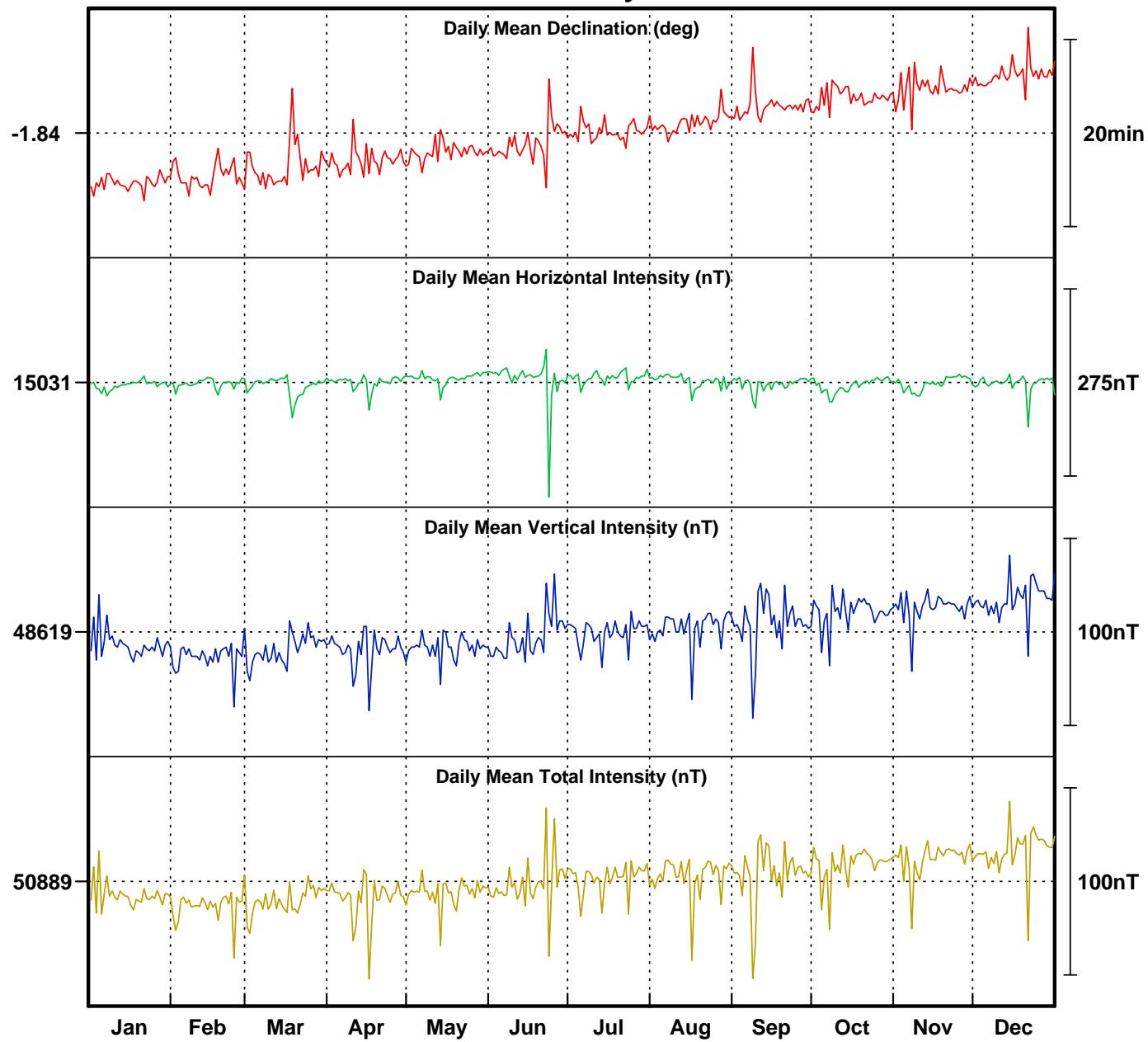


Lerwick Observatory: Vertical Intensity (nT)



Lerwick Observatory

2015



Monthly Mean Values for Lerwick Observatory 2015

Month	<i>D</i>	<i>H</i>	<i>I</i>	<i>X</i>	<i>Y</i>	<i>Z</i>	<i>F</i>
January	-1° 55.7'	15027 nT	72° 49.4'	15018 nT	-506 nT	48612 nT	50882 nT
February	-1° 55.0'	15029 nT	72° 49.1'	15020 nT	-503 nT	48606 nT	50876 nT
March	-1° 53.9'	15025 nT	72° 49.4'	15017 nT	-498 nT	48610 nT	50879 nT
April	-1° 53.3'	15031 nT	72° 49.1'	15023 nT	-495 nT	48609 nT	50880 nT
May	-1° 52.4'	15037 nT	72° 48.7'	15029 nT	-491 nT	48611 nT	50884 nT
June	-1° 51.4'	15035 nT	72° 48.9'	15027 nT	-487 nT	48617 nT	50888 nT
July	-1° 50.1'	15037 nT	72° 48.8'	15030 nT	-482 nT	48618 nT	50891 nT
August	-1° 49.3'	15033 nT	72° 49.2'	15025 nT	-478 nT	48622 nT	50893 nT
September	-1° 47.5'	15028 nT	72° 49.5'	15021 nT	-470 nT	48624 nT	50893 nT
October	-1° 46.5'	15026 nT	72° 49.8'	15019 nT	-465 nT	48630 nT	50898 nT
November	-1° 45.6'	15029 nT	72° 49.6'	15022 nT	-461 nT	48632 nT	50901 nT
December	-1° 44.1'	15029 nT	72° 49.7'	15022 nT	-455 nT	48637 nT	50906 nT

Note

- i. The values shown here are provisional.

LERWICK RAPID VARIATIONS

SIs and SSCs

Date	Time (UT)	Type	Quality	H (nT)	D (min)	Z (nT)
14-12-15	13 21	SSC*	C	10.6	-2.06	-3.3
19-12-15	16 16	SSC*	A	28.2	5.54	9.9
31-12-15	00 50	SSC*	B	26.3	-4.80	-12.5

Notes:

An asterisk (*) indicates that the principal impulse was preceded by a smaller reversed impulse.

The quality of the event is classified as follows:

A = very distinct

B = fair, ordinary, but unmistakable

C = doubtful

The amplitudes given are for the first chief movement of the event.

SFEs

Date	Universal Time			H (nT)	D (min)	Z (nT)
	Start	Maximum	End			
None						

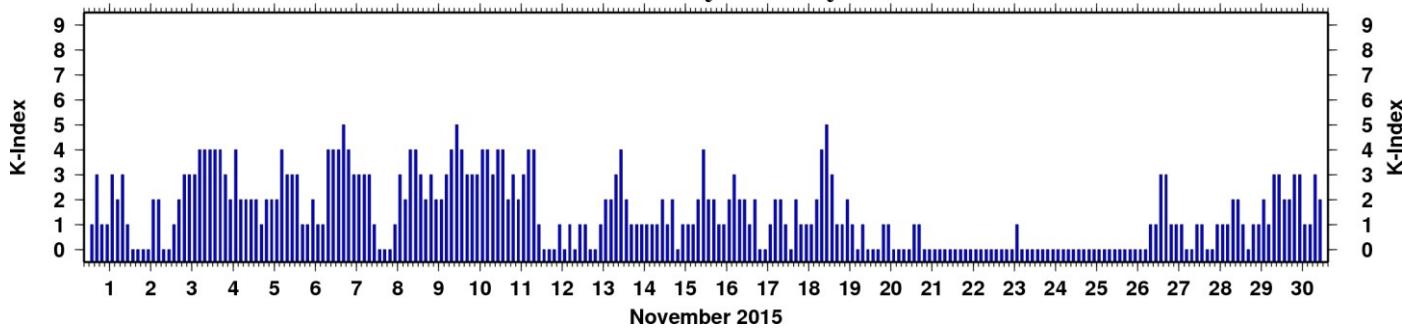
Note:

The amplitudes given are for the first chief movement of the event.

INDICES OF GEOMAGNETIC ACTIVITY

Day	K - INDICES FOR THREE-HOUR INTERVAL							
	00-03	03-06	06-09	09-12	12-15	15-18	18-21	21-24
1	2	2	0	1	2	3	3	4
2	2	1	1	0	2	2	2	2
3	0	0	0	0	0	0	0	1
4	1	1	1	0	0	0	0	2
5	1	0	2	3	1	2	3	2
6	2	3	2	2	3	3	4	3
7	3	3	2	2	2	4	4	2
8	2	2	1	2	2	1	1	2
9	2	1	0	1	1	2	4	2
10	3	4	2	2	2	4	4	3
11	3	2	1	2	2	4	4	3
12	2	2	1	1	1	2	2	2
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15	4	2	2	1	2	2	1	3
16	3	2	0	0	0	0	0	0
17	0	0	0	0	1	2	4	1
18	1	0	1	1	0	0	0	3
19	0	0	0	0	0	3	3	4
20	3	5	4	3	4	7	7	7
21	6	5	4	2	2	1	2	2
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24	2	1	2	1	2	1	3	2
25	2	1	1	1	1	1	1	2
26	2	1	1	2	2	3	2	4
27	2	1	2	1	2	2	2	0
28	1	0	1	1	1	0	0	1
29	1	0	0	0	1	1	2	2
30	0	0	0	0	0	0	0	2
31	2	2	2	2	3	4	5	5

Lerwick Observatory 3-hourly K-Indices



The *aa* Index

Date	Day	3-hourly <i>aa</i> -indices									<i>Aa_{am}</i>	<i>Aa_{pm}</i>	<i>Aa</i>
01-12-15	335	24	24	12	20	46	45	24	37	20.0	38.1	29.1	
02-12-15	336	12	12	24	12	24	24	12	12	15.0	18.0	16.5	
03-12-15	337	8	8	5	5	5	5	8	12	6.5	7.5	7.0	
04-12-15	338	12	16	8	5	5	8	2	12	10.2	6.8	8.5	
05-12-15	339	8	12	32	46	24	59	37	37	24.4	39.4	31.9	
06-12-15	340	16	32	24	59	45	45	58	45	32.8	48.6	40.7	
07-12-15	341	32	32	32	32	46	45	45	16	32.1	38.1	35.1	
08-12-15	342	32	16	12	24	32	12	16	24	21.1	21.0	21.1	
09-12-15	343	12	16	5	8	12	24	37	12	10.2	21.3	15.8	
10-12-15	344	32	80	32	32	24	37	45	45	44.1	38.1	41.1	
11-12-15	345	37	24	24	32	24	67	37	37	29.5	41.3	35.4	
12-12-15	346	16	24	24	12	16	32	12	24	19.1	21.0	20.1	
13-12-15	347	12	16	12	16	16	8	2	8	14.0	8.5	11.3	
14-12-15	348	20	12	8	12	32	80	59	115	12.9	71.4	42.2	
15-12-15	349	59	32	16	16	32	46	16	24	30.8	29.5	30.1	
16-12-15	350	45	16	5	2	2	2	5	5	17.3	3.8	10.5	
17-12-15	351	5	5	2	12	12	12	45	16	6.2	21.3	13.8	
18-12-15	352	12	8	8	5	5	5	8	20	8.2	9.5	8.8	
19-12-15	353	2	5	5	8	8	46	59	59	5.2	42.8	24.0	
20-12-15	354	45	102	80	67	116	102	136	136	73.7	122.5	98.1	
21-12-15	355	101	67	45	46	38	24	12	24	64.7	24.5	44.6	
22-12-15	356	24	16	46	59	24	8	20	12	36.2	15.9	26.1	
23-12-15	357	16	16	12	16	16	16	24	37	15.1	23.4	19.2	
24-12-15	358	16	16	32	16	38	16	20	16	20.1	22.4	21.3	
25-12-15	359	16	12	16	12	20	16	16	16	14.0	17.1	15.5	
26-12-15	360	16	16	16	24	45	45	24	67	18.1	45.4	31.7	
27-12-15	361	24	16	16	32	32	16	12	16	22.1	19.0	20.6	
28-12-15	362	12	12	12	16	24	8	2	12	13.0	11.6	12.3	
29-12-15	363	12	5	2	8	24	24	20	20	6.8	22.0	14.4	
30-12-15	364	2	2	5	5	12	12	5	20	3.8	12.2	8.0	
31-12-15	365	32	24	32	46	137	59	80	102	33.5	94.5	64.0	
Monthly Mean Value												26.4	

Notes

- i. The units of the *aa* index are nT.
- ii. The 3-hour *aa* values are rounded to the nearest integer. Where $aa = \lceil * .5 \rceil$, *aa* is rounded down.
- iii. Daily values (*Aa_{am}*, *Aa_{pm}* and *Aa*) are computed from *aa* values of original resolution.
- iv. The monthly mean value is computed from the daily mean values, *Aa*.
- v. Definitive *aa* indices are derived and published by the International Service for Geomagnetic Indices.

