

Accepted Manuscript

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PII: S1464-343X(18)30113-4

DOI: [10.1016/j.jafrearsci.2018.04.015](https://doi.org/10.1016/j.jafrearsci.2018.04.015)

Reference: AES 3199

To appear in: *Journal of African Earth Sciences*

Received Date: 4 August 2016

Revised Date: 19 April 2018

Accepted Date: 20 April 2018

Please cite this article as: Birhanu, Y., Williams, S., Bendick, R., Fisseha, S., Time dependence of noise characteristics in continuous GPS observations from East Africa, *Journal of African Earth Sciences* (2018), doi: 10.1016/j.jafrearsci.2018.04.015.

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Time dependence of noise characteristics in continuous GPS observations from East Africa

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Abstract

A noise model for the regional continuous GPS (cGPS) timeseries in East Africa (Ethiopia and Eritrea) was computed using the maximum likelihood estimation (MLE) method. Using this method and assigning different noise models for each cGPS site and each component (north, east and vertical) may bias the noise level of the velocity solutions due to the non-uniformity of the length of the timeseries. Within the whole regional network, the length of the timeseries varies from one to seven years. We

23 compute a preferred regional noise model for the whole data sets using a stacked
 24 maximum likelihood values for the different power – law indexes (between -2 and 0 with
 25 a time step of 0.1), presuming that there is only one noise model that exists in the
 26 regional cGPS timeseries. Therefore, a single power – law index (flicker plus white
 27 noise) was assigned for the whole regional network irrespective of the length of the
 28 timeseries. This approach is more robust and “realistic” to determine the noise
 29 characteristics of the regional GPS network.

30
 31 **Keyword:** timeseries, maximum likelihood, noise model, power – law index

32 33 **Introduction**

34 The noise characteristics of cGPS coordinate timeseries have been studied by
 35 various authors [e.g. Agnew, 1992; Mao et al., 1999; Williams et al., 2004; Hackl et al.,
 36 2011] in different parts of the world, using different processing methods, from regional
 37 networks through to global solutions. It is a well-understood phenomenon that assigning
 38 white noise model for the cGPS timeseries will underestimate the noise uncertainty level
 39 in the final velocity solution [Mao et al., 1999] since it does not account for the
 40 contribution of time varying noises. Work done by these authors [Agnew, 1992; Zhang et
 41 al., 1997; Mao et al., 1999; Williams, 2003a; Williams et al., 2004] showed that the cGPS
 42 timeseries noise characteristics consists of a combination of white noise and other time-
 43 correlated noise models. We used 16 cGPS sites in the northeast part of the East African
 44 Rift System (EARS) (Figure 1) and its surroundings, especially Ethiopia and Eritrea, to

45 study the noise characteristics of the whole regional network. The cGPS sites have
46 timeseries that span from early 2007 to 2014 as shown in figure 2. However, the
47 timeseries are non-uniform over that period because of their deployment for various
48 studies and contain discontinuities and gaps in the timeseries caused by various
49 instrumental or logistical problems.

50 In this study we discuss the noise level of the regional cGPS timeseries using the
51 maximum likelihood estimation (MLE) technique and the time dependence of the noise
52 characteristics as a function of the length of timeseries. We compute a regional noise
53 model for the entire network and compare velocity uncertainties derived from CATS and
54 GLOBK which are based on different noise models.

55

56 **Methods**

57 All the cGPS data were processed using GAMIT/GLOBK software [Herring et al.,
58 2010]. 15 International GNSS Service (IGS) sites closest to the study area were used as
59 reference sites with one IGS site (ADIS) located within the study area (Figure 1), and
60 included in this study. Using GAMIT, we applied double differencing on each daily
61 phase observation in order to estimate station coordinates, phase ambiguities and satellite
62 state vectors. At every station seven tropospheric delay and two tropospheric gradient
63 parameters per day were estimated [Reilinger et al., 2006]. After applying double
64 differencing to each daily phase observation, the daily solutions (h-files) of the cGPS data
65 were combined with the daily global solutions (H-files) obtained from MIT, using the
66 global kalman filter (GLOBK). The resulting daily timeseries were closely inspected for

quality check in order to remove outliers (above two-sigma threshold), and discontinuities caused by antenna changes, receiver changes or earthquakes, together with other time-dependent changes [Reilinger et al., 2006]. Before assigning a Gauss Markov noise model for our cGPS timeseries, the usual routine in GLOBK, we tested different stochastic noise models for each cGPS timeseries. The noise characteristics of the north, east, and up components were computed individually using the CATS GPS coordinate timeseries analysis software [Williams, 2005] that applies maximum likelihood to estimate the noise parameters.

The observed GPS motion at each site is a superposition of secular trend, a periodic component (mostly annual and semi-annual signal), offsets (discontinuities caused by tectonic and non-tectonic processes) and a noise component as shown in equation (1) [Williams, 2003b; Yuan et al., 2007].

$$y_i = a + bt_i + \sum_{j=1}^l a_j H(t_j - T_j) + \sum_{n=1}^m A_n \sin(\omega_n t + \varphi_n) + n_i \quad (1)$$

Where the first term (a) is the site coordinate, the second term (bt_i) is the linear rate (Figure S1), the third term is made up of the Heaviside step function [$H(t_j - T_j)$] and an offset amplitude (a_j where $t_j = T_j$) mostly caused by earthquakes, various tectonic processes and other phenomena like antenna or receiver change, and the fourth term consists of periodic components, mainly the annual and semi-annual signals (Figure S2)

86 where $\omega_n = \frac{2\pi}{n}$ *rad/year*. The final term is the noise component of the GPS timeseries.

87 In this study we are interested to study the noise term.

88 From equation (1) above, in order to study the various noise models of the cGPS
89 timeseries the other terms have to be removed from the data. The data has to be
90 detrended, offsets and seasonal variations must also be removed from the data (Figure
91 S3). The correlation in cGPS timeseries may be approximated by a power law process
92 [Agnew, 1992; Mao et al., 1999; Williams, 2003a and b; Williams et al, 2004]:

93

$$94 \quad P(f) = P_o \left(\frac{f}{f_o}\right)^{-\alpha} \quad (2)$$

95

96 Where α is the spectral index (white noise has $\alpha = 0$, flicker noise has $\alpha = 1$ and random
97 walk noise has a spectral index of 2), P_o is a constant, f_o is a constant frequency and f is
98 the frequency.

99 We used the method described in CATS [Williams, 2005] in order to compute a
100 “realistic” noise model for the GPS timeseries. The following steps were implemented in
101 order to characterize the noise in the regional network.

- 102 **1.** All the GPS data were detrended and the linear terms in the timeseries were
103 removed using the weighted least squares fit.
- 104 **2.** The detrended data were closely inspected and offsets caused by any tectonic or
105 non-tectonic processes were removed using two-sigma significance level. Tsview
106 software [Herring and McClusky, 2009] was used to inspect and remove the

outliers and offsets. Although some of the cGPS sites are located at a closer proximity to the EARS, we did not see any offsets which is caused by the local earthquakes.

3. The annual and semi-annual terms in the data were removed using the fourth term in equation (1). Before removing the annual and semi-annual signals from the data a spectral analysis of the timeseries were computed using the Lomb-Scargle Algorithm [Press et al., 2001] as shown in (Figure S4). This method can help in identifying any dominant periodic signal. We used the R function 'lsp' [R Core Team, 2013] that computes the Lomb-Scargle periodogram [Ruf, 1999; Press et al., 2001]. This algorithm is different from the Fast Fourier Transform because it can handle data sets that are not evenly spaced and does not require data that are a factor of 2 in length (or padded). Since some of our data sets have gaps due to receiver and/or antenna malfunctions or other issues the Lomb-Scargle method is appropriate for these datasets.
4. The residual timeseries (called here the unfiltered timeseries), after the trend, offsets and seasonal signals were removed are then summed on a daily basis to produce mean values that are termed as a Common Mode Error (CME). The CME was then removed from each GPS timeseries to produce a filtered timeseries. However since the trends are ultimately derived from the unfiltered series we use these for the MLE estimation of the noise rather than the filtered series.
5. The CME is useful in explaining how spatially coherent the timeseries are. We use MLE estimation to compare different candidate noise models for each series

individually. We also use a method to produce a combined regional estimate of the noise in the timeseries.

6. Finally, Principal Component Analysis (PCA) was applied as an alternate to the CME approach to the three individual components of the unfiltered timeseries in order to gauge the usefulness of each technique.

Results

In order to ascertain which noise model is the most realistic for the regional datasets, we look at the sum of the natural logarithm of the ML values for each noise model and each component for the whole data set. After that, the summed noise models were differenced from the ML estimate of the white noise model (the white noise model is used as a null hypothesis), and values greater than the white noise ML indicate that this model is more likely. The top rows of bar graphs in Figure 3 show the difference between the seven stochastic models (flicker noise (F), random walk noise (R), power-law (P), Gauss Markov noise (G), flicker plus white noise (F+W), random walk plus white noise (R+W) and Gauss Markov plus white noise (G+W)) with the null hypothesis (white noise (W)). All the models that include time correlated noise and white noise have larger collective MLs. In the plot the time-correlated noise (R and G) models have values generally smaller than the null hypothesis. In the second row bar plots (Figure 3) instead of taking white noise as the null hypothesis, flicker plus white noise was used as a null hypothesis, and similar computation has been done as before. The result shows that power law, first order Markov noise plus white noise model have almost zero MLE

151 differences with flicker plus white noise, except the north component of the first order
 152 G+W noise model which has a higher MLE (Figure 3). The other noise models show a
 153 negative value with the null hypothesis (F+W) and therefore less likely to be the noise
 154 model candidates

155 Unlike the above procedure i.e. computing the ML values for the known noise models
 156 (flicker, random walk, power-law, etc.) the spectral index of the cGPS timeseries was
 157 estimated from each components of the whole regional network. Dmetrieva et al., [2015]
 158 used a combination of Kalman filter and MLE to estimate a single noise model for a
 159 network of 15 sites in Eastern North America, instead of assigning an individual noise
 160 model for each site and component separately. We use a similar approach here and solve
 161 for the spectral index that maximizes the likelihood of the whole regional network. If we
 162 assume that the individual timeseries are not spatially correlated (or at least that the time
 163 and spatial correlations are orthogonal to one another) then the maximum likelihood of
 164 the network wide estimation of the spectral index would be the sum of the individual site
 165 MLs. In order to estimate a regional noise model for the network the following
 166 procedures were implemented. First the ML values for a set of fixed spectral indices for
 167 each site and each component (north, east, and up) were computed for the unfiltered
 168 timeseries. The computed ML values for each spectral index (between -2 and 0 and time
 169 steps of 0.1) were summed in order to compute the ML for the whole network. Figure 6
 170 shows the network-wide ML as a function of spectral index for each component. Instead
 171 of using a maximization algorithm to find the spectral index that gives the largest
 172 network-wide ML, which could potentially be CPU intensive, we took the three largest

ML values for each component (based on the fixed spectral index values) and fitted a polynomial of degree 2 to those values. Using only the three largest values ensures we can find a maximum to the polynomial. We then solve the polynomial to find the spectral index which maximizes the fit. In Figure 4, the computed spectral index for each cGPS site and each component are compared against the length of the timeseries. The plot shows that, as the length of the timeseries increases, the estimated spectral index tends towards -1 (flicker noise) and when the length of the timeseries is short the estimated spectral index tends to zero (white noise). This is primarily a result of lack of data, biasing the results [Santamaria-Gomez et al., 2011].

Principal Component Analysis (PCA) [Björnsson and Venegas, 1997] attempts to split time series up into several significant modes that are orthogonal to each other. The first mode is often very similar to the CME unless it has been prior to the PCA. The first three Eigenvectors for each component are plotted in Figure 5. In all components, there is a distinct change in the values for sites that started in 2012 onwards. The Eigenvectors are a lot more consistent with more variation in the older sites. This probably reflects that many of the major modes are reflecting the dual subspaces of the dataset, that is those sites containing data mainly prior to 2012 and those sites with data after 2012. In terms of the amount of variance explained the first mode accounts for 68%, 57% and 43% for the north, east and vertical components respectively. The second and third modes explain around 19% and 9%.

For the whole network approach the result showed that the three components have a power-law dependence with estimated spectral index between $(-1 < \text{spectral index} < -$

0.85) for the north, east and up components. We then assign the index that maximizes the fit as the preferred noise model for the whole network. This computation is preferred over assigning different noise models for each component, and each site, as this uses the entire cGPS data sets as compared to using the individual timeseries.

Discussion

The preferred noise model was computed based on the above two methods: using the difference of the natural logarithm MLE values and the stacked MLE values and fitting a polynomial of degree 2 that maximizes the spectral index. The regional noise analysis that uses the stacked MLE values is a more appropriate way to assign a “realistic” noise model for the regional timeseries computation. For the entire cGPS network, we expected to have one type of noise in the timeseries and not different power-law noises for each site and each component. Since our data sets have different timeseries length, assigning a different noise model for each component of the timeseries may bias the noise estimates (Figure 4). The regional noise model analysis uses the whole data sets and assign a single spectral index for the entire network. This regional analysis is a more robust way of assigning a noise model.

Conclusion

The GLOBK realistic sigma estimation, that uses first order Gauss Markov noise, shows coherent sigma estimates with the CATS estimates, that we assign flicker plus white noise as a preferred noise model. As shown in Figure S5 GLOBK and CATS

217 velocity noise analysis are most correlated when the timeseries are longer and least
 218 correlated when the timeseries are shorter. Shorter timeseries bias the noise estimates
 219 (Figure 4) and affect the velocity error estimates. In order to assign the optimal noise
 220 model for the entire cGPS network, for a non-uniform length of a cGPS timeseries, care
 221 has to be taken. Therefore, in this study, we have assigned flicker plus white noise model
 222 as a preferred noise model for the whole regional network regardless of the length of the
 223 timeseries.

224

225 **Acknowledgements**

226 This work is supported by NSF EAR-1119209 and NERC funding through RiftVolc
 227 NE/L013932/1. We would like to thank Institute of Geophysics, Space Sciences and
 228 Astronomy, Addis Ababa University and Eritrean Institute of Technology. Data are
 229 archived in the UNAVCO archive.

230

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290

291 **Figure captions**

292 **Figure 1.** Location of the cGPS sites overlain on the DEM map of the study area. Each
293 triangle shows the location of the GPS sites with site names next to each triangle.

294

295 **Figure 2.** The detrended cGPS displacement timeseries for each site and for the north,
296 east and vertical components respectively. The y-values are scaled for plotting purpose
297 only.

298

299 **Figure 3.** Histogram of the log maximum likelihood values for the various noise models.
300 The labels in the horizontal axis, W- white noise, F – flicker noise, R- random walk
301 noise, P – power-law noise, G – first order Gauss Markov noise, F+W – flicker plus
302 white noise, R+W – random walk plus white noise, and G+W – first order Gauss Markov

plus white noise model. The first row, three bar graphs are for the north, east and vertical components respectively. The first row three bar plots show white noise (W) as a null hypothesis and compared against the mean of the difference of the natural logarithm ML of white noise with other noise models. In the second row, the bar plots (for the north, east and vertical), the similar procedure was done and in this case flicker plus white noise (F+W) was taken as a null hypothesis.

Figure 4. Time dependence of the estimated spectral index of the cGPS timeseries for the north south, east west and up components. The estimated spectral index is computed for the different duration of the timeseries (from one to seven years).

Figure 5. The Principal Component Analysis (PCA) for the north, east and up components. The three colors (purple, blue and red) indicates the first, second and third principal component (pc) respectively for the north, east and vertical components. The cGPS site names are labeled below the horizontal axis.

Figure 6. The first three plots show, the stacked ML values for a selected spectral index between -2 and 0 with an interval of 0.1. In the second three plots the three maximum-stacked MLE values are selected and the associated polynomial fit of degree 2 was fitted to the values that maximizes the spectral index. The broken red vertical lines indicate the maximum of the fit (-0.89, -0.92 and -0.98) for the north, east and vertical respectively.

325 **Figure S1** GPS time series of ADIS station. In this plot the linear trend, annual and
 326 semiannual terms are not removed. The dotted red line is the weighted least square fit for
 327 the north, east and up components. The north and east components show significant
 328 linear trend.

329

330 **Figure S2** The linear trend of the time series were removed and the red line is the annual
 331 and semiannual fit. The up component shows significant annual and semiannual
 332 component.

333

334 **Figure S3** Residual GPS time series where the linear, annual and semiannual terms were
 335 removed from the time series.

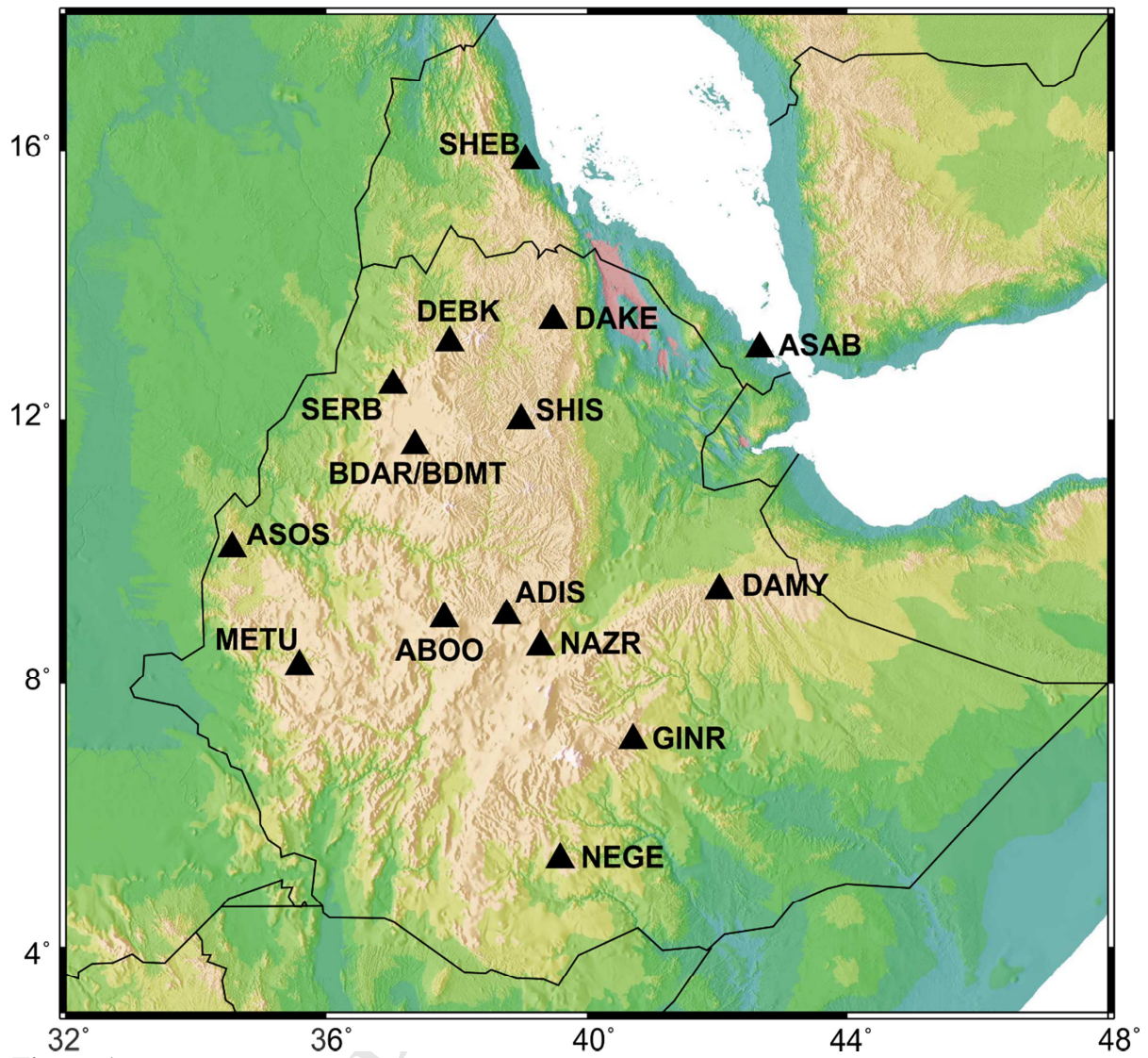
336

337 **Figure S4** Lomb – Scargle periodogram of the cGPS timeseries for north, east and up
 338 components with the corresponding site names on the left bottom corners of the plots.
 339 These are the selected periodogram that have longer timeseries and based on the
 340 geographic distribution. The blue plots are for the north, purple for the east and red for
 341 vertical components. The broken black lines in the plots show the annual frequency. The
 342 label in the bottom right corner indicates, slope of north component (slope_n), slope of
 343 east component (slope_e) and slope of up/vertical component (slope_u).

344

345 **Figure S5** Comparison of the north-south and east-west CATS velocity uncertainties
 346 (purple color), which uses flicker plus white noise as a preferred noise model, in this

347 study, and the GLOBK velocity uncertainties (blue color) based on Gauss Markov noise
348 model. The cGPS site names are labeled below the horizontal axis and the velocity
349 uncertainties in the vertical axis.



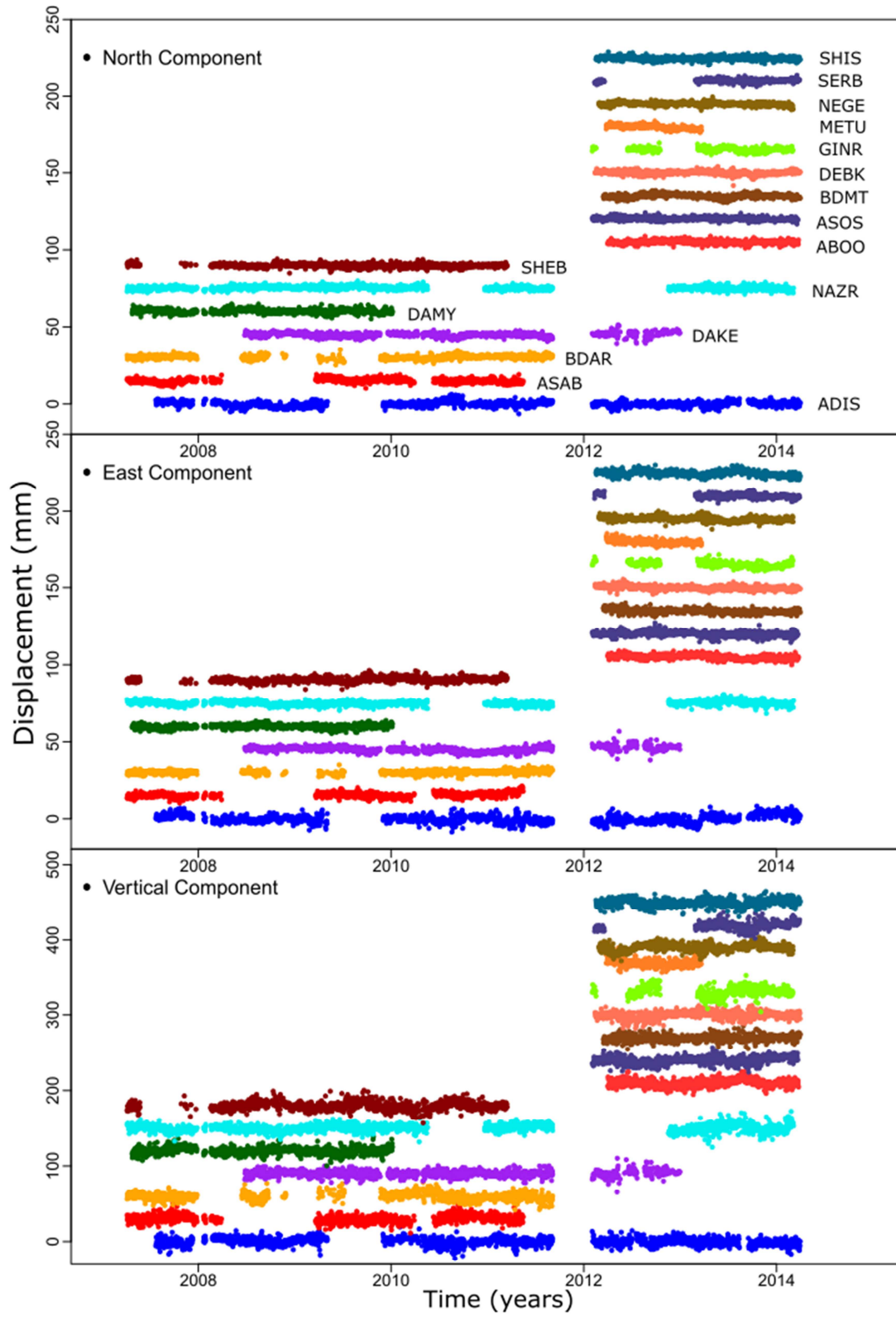


Figure 2

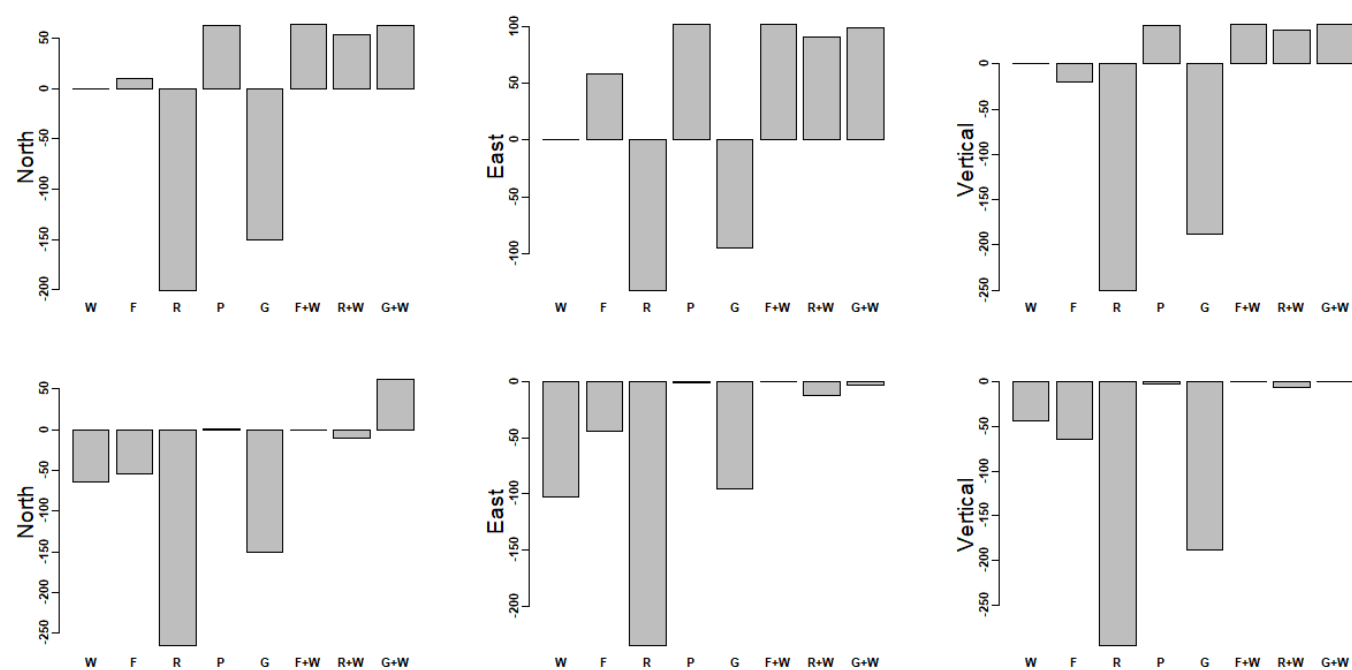


Figure 3

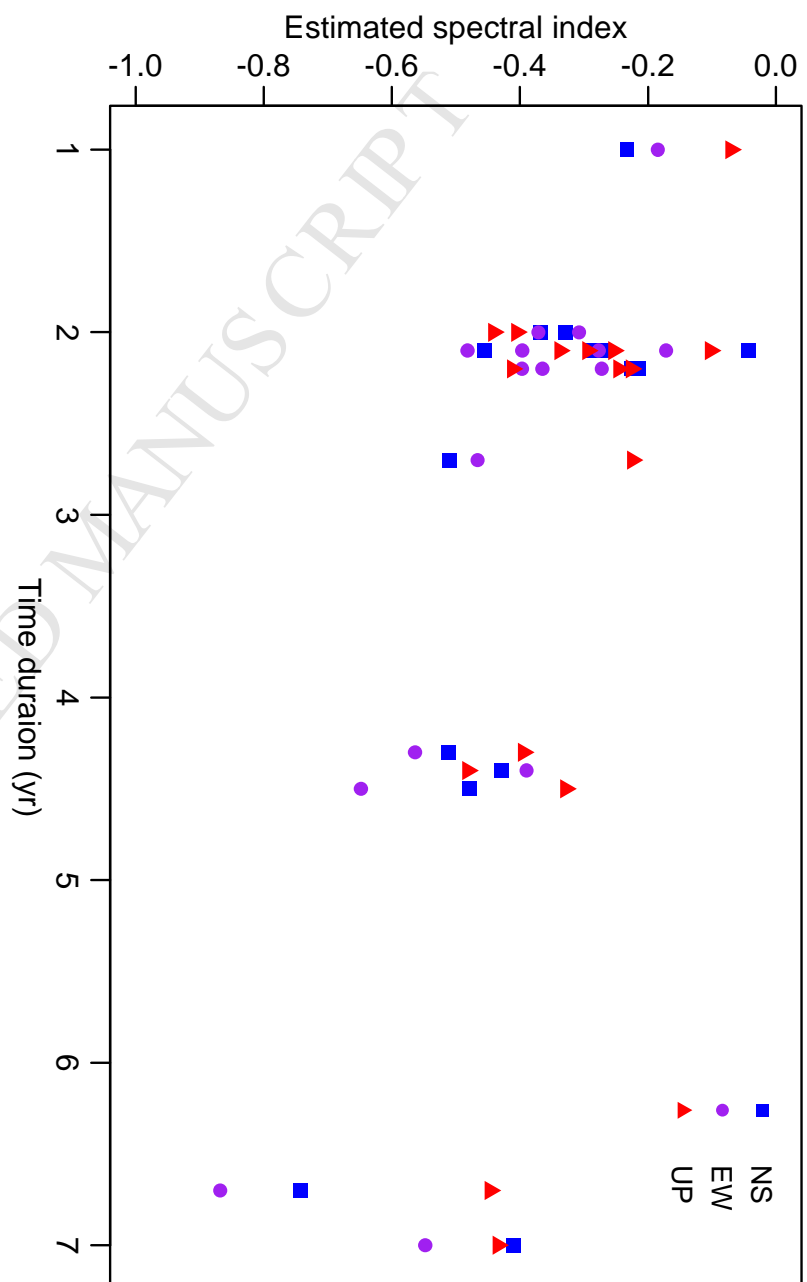


Figure 4

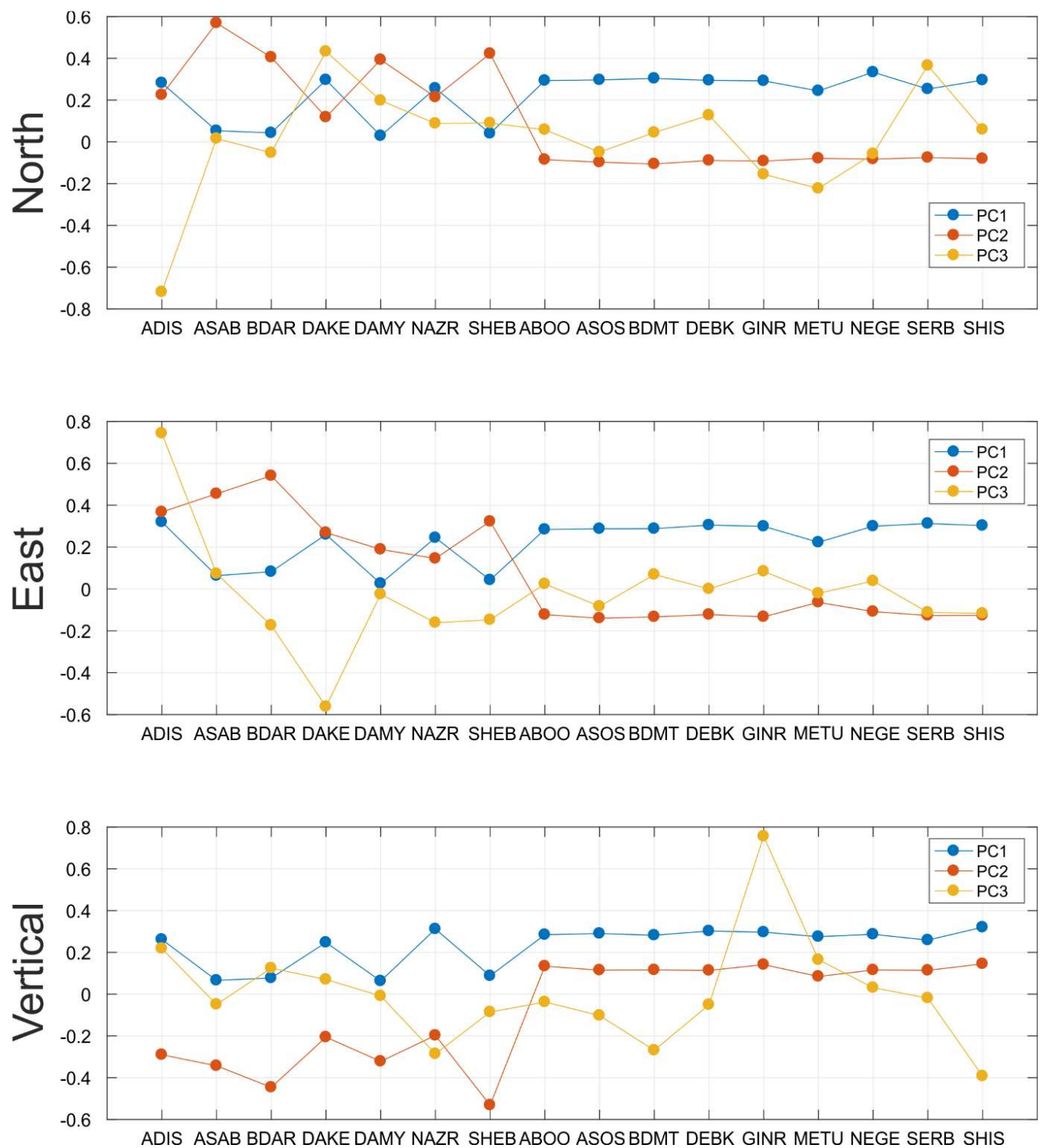


Figure 5

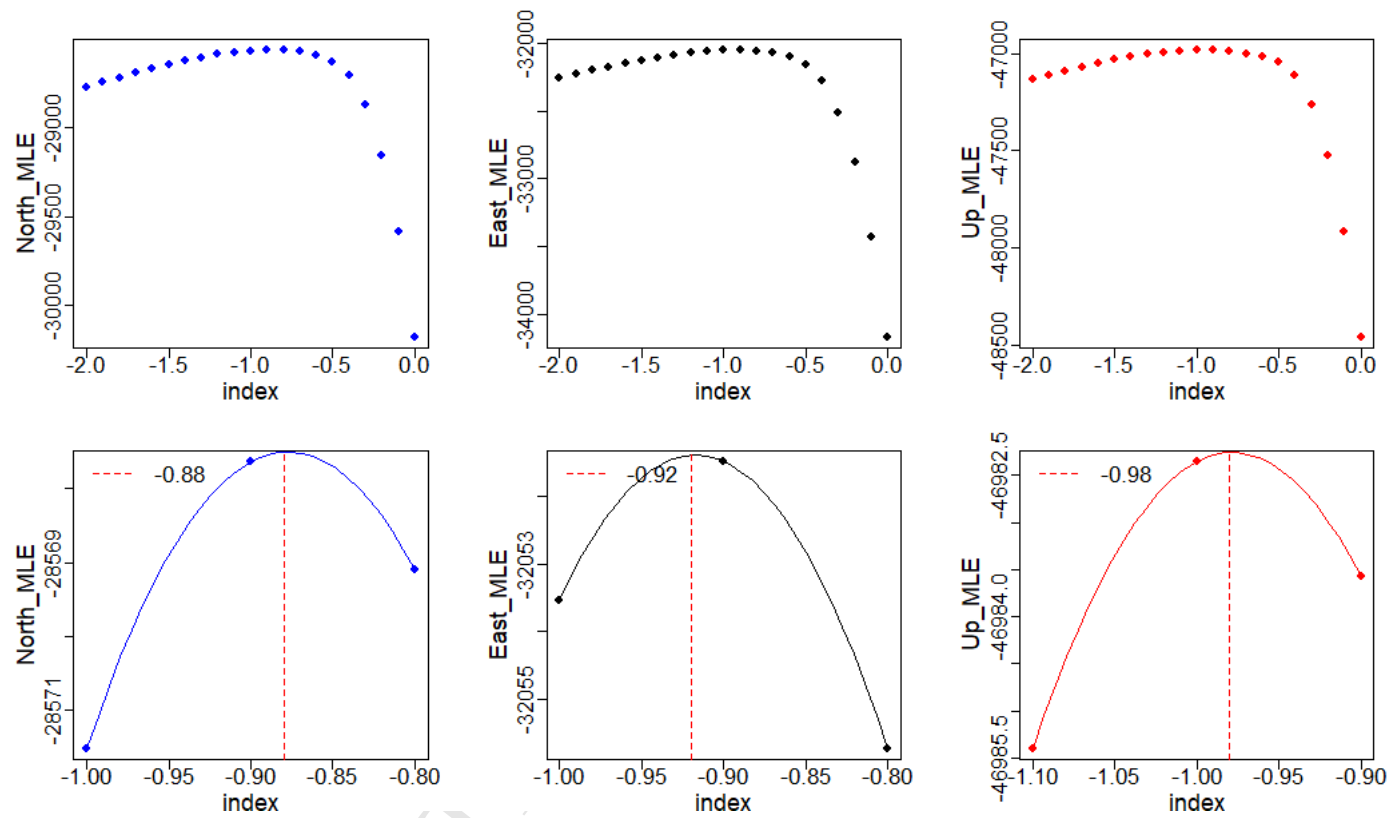


Figure 6

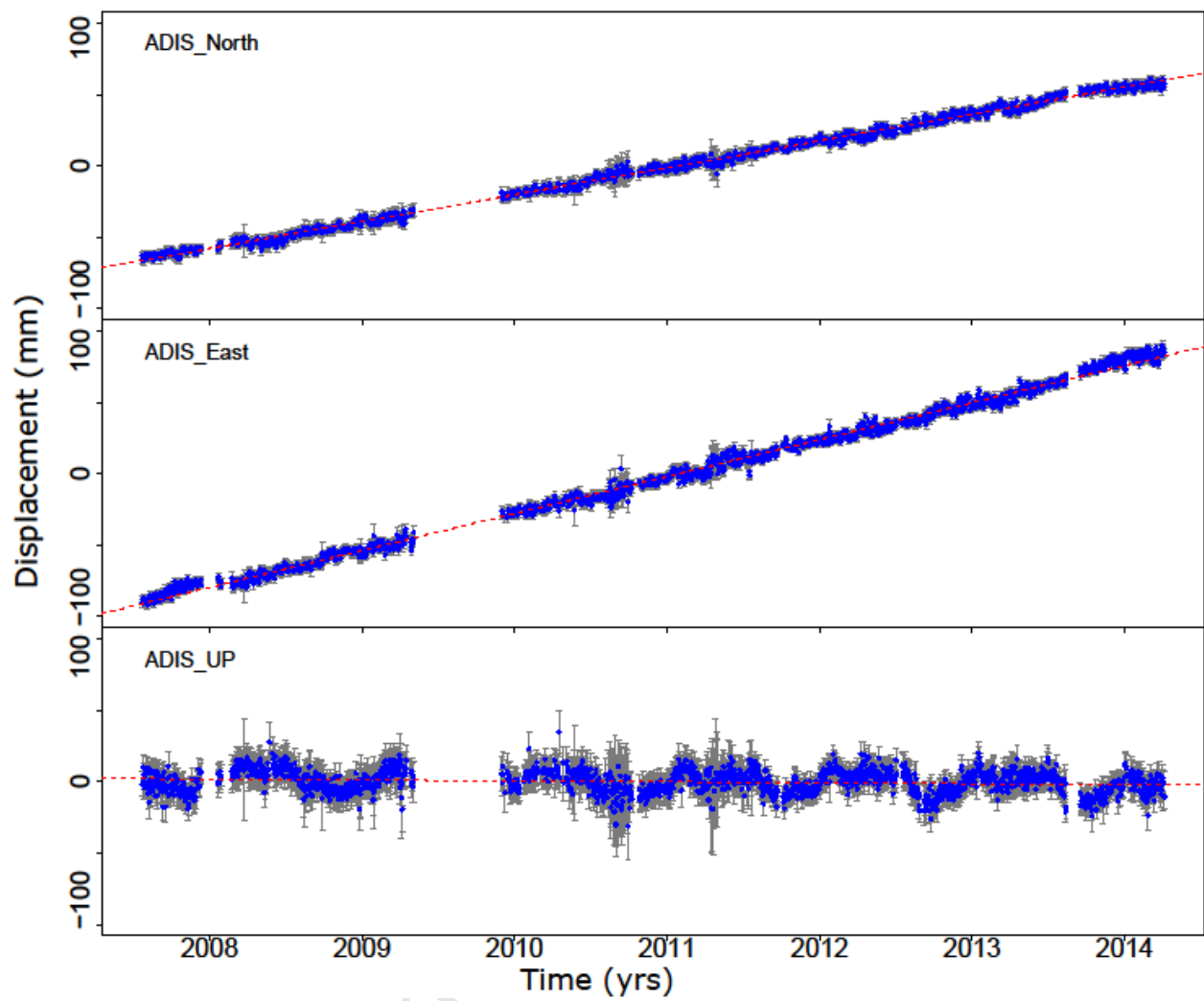


Figure S1

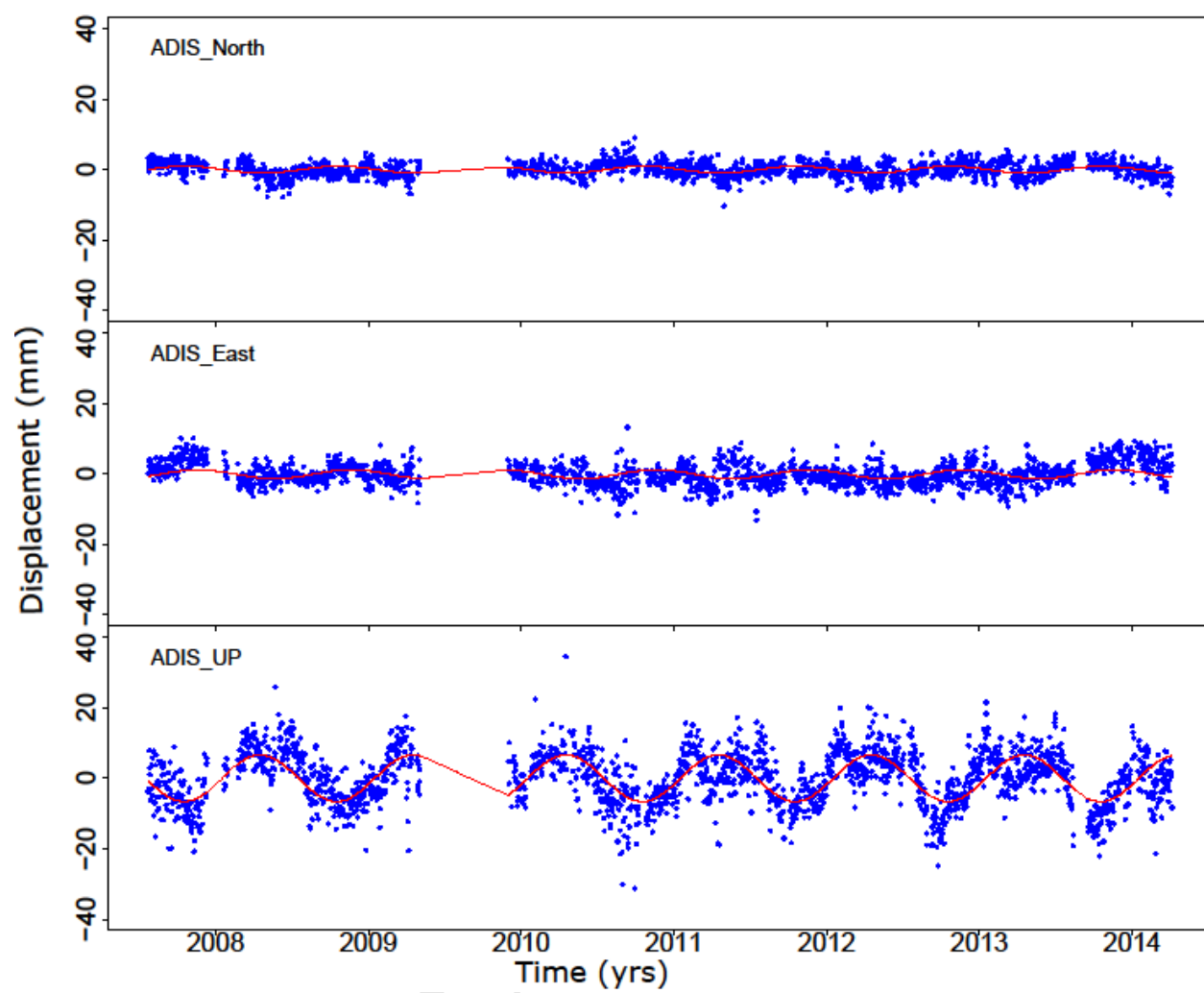


Figure S2

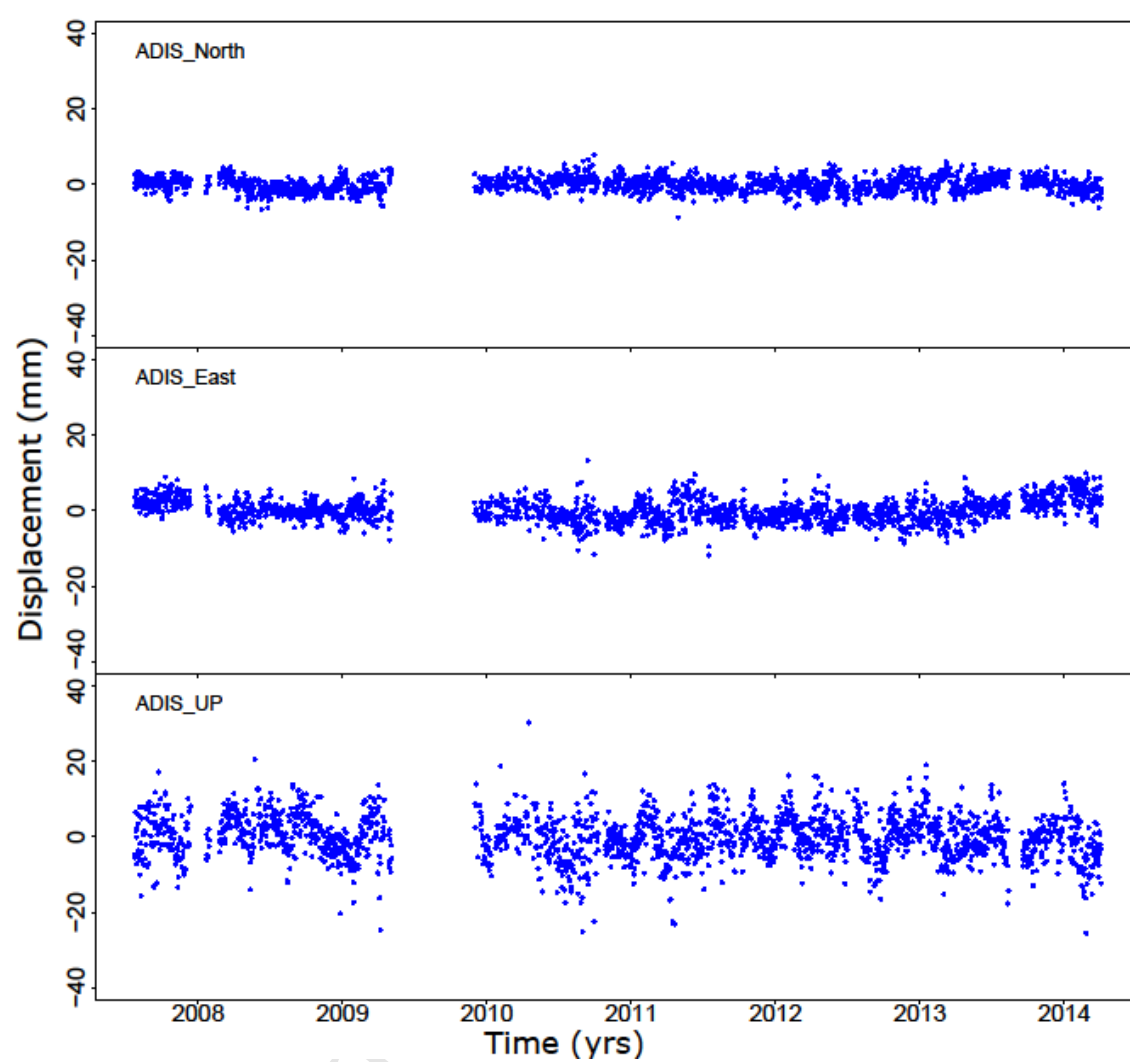


Figure S3

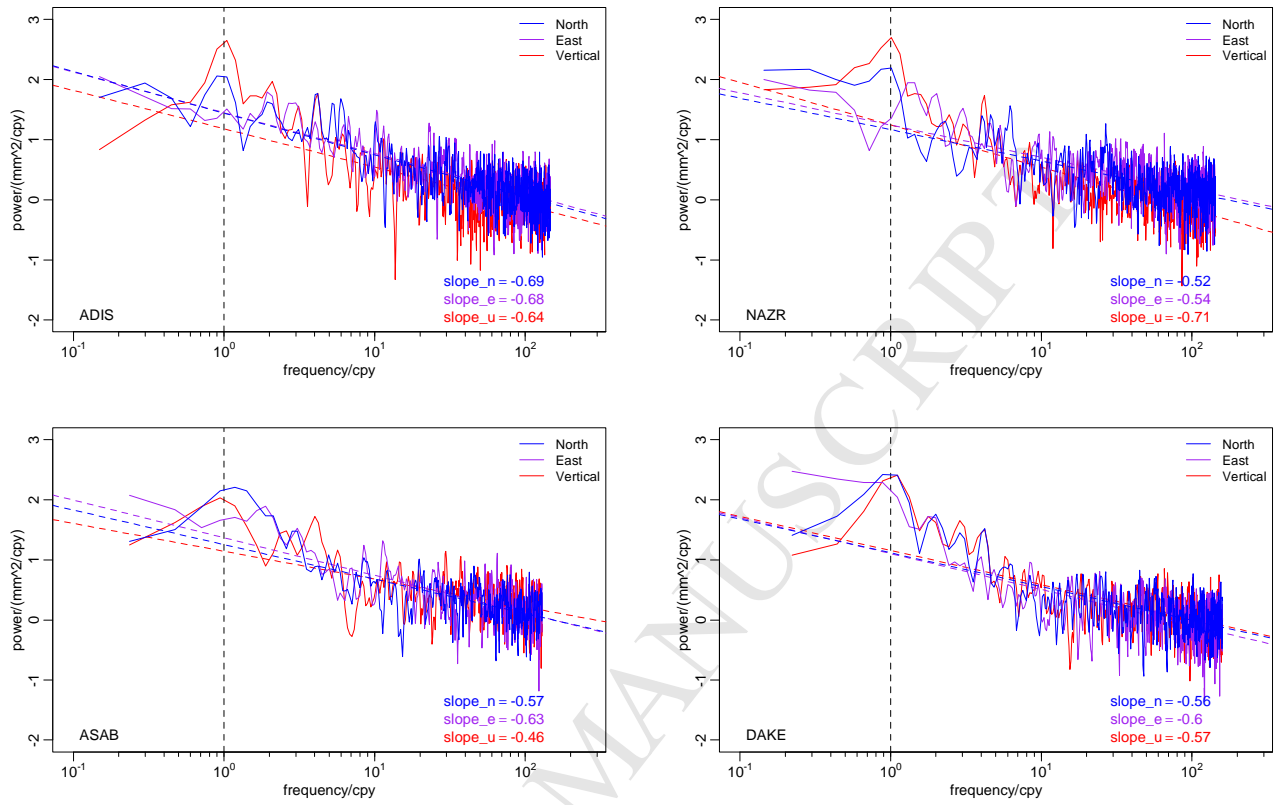
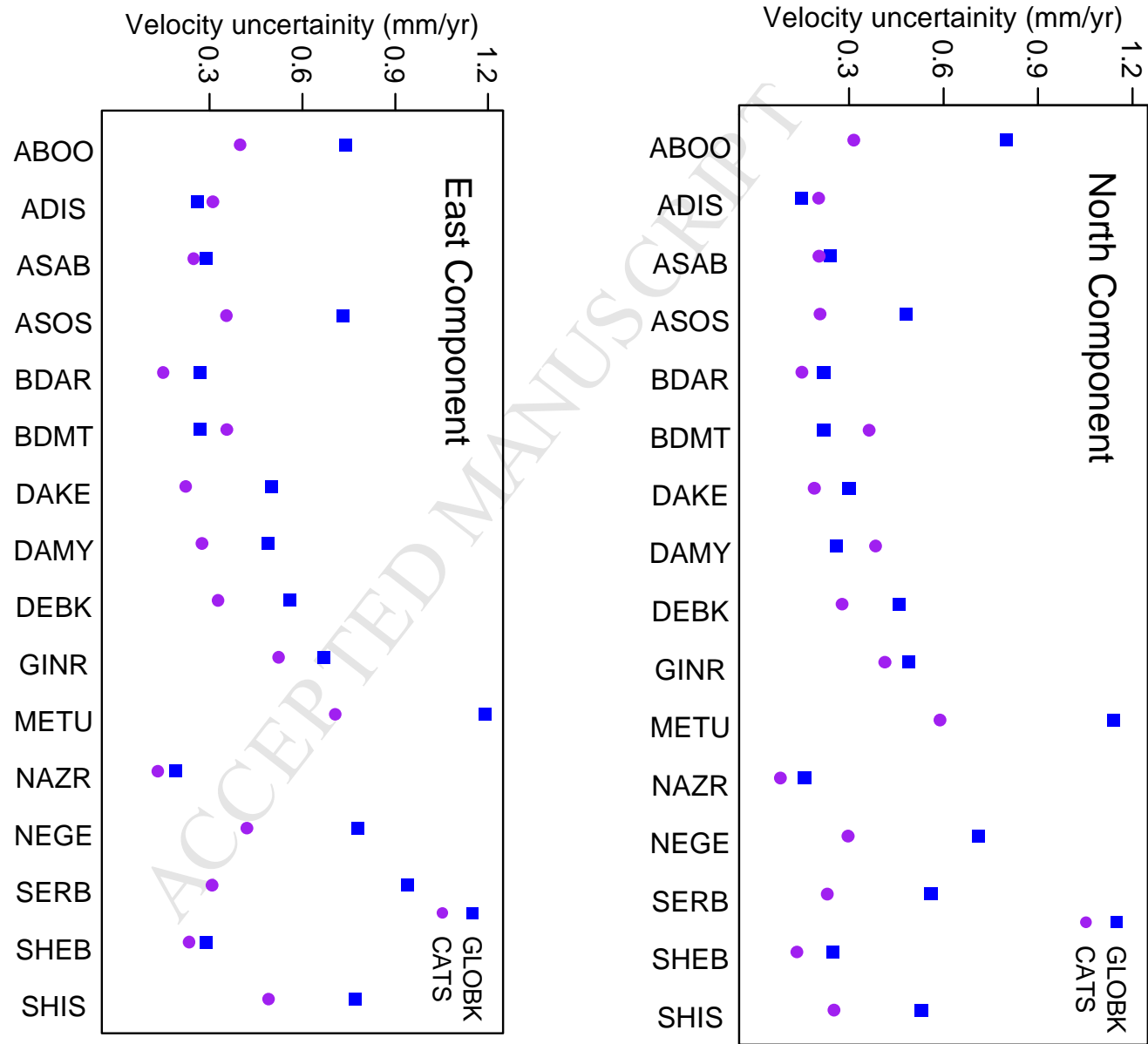


Figure S4

Figure S5



- 16 continuous GPS sites in Ethiopia and Eritrea have been used to compute the noise characteristics of the GPS velocity uncertainties.
- The network wide regional noise analysis shows flicker plus white noise model is the “robust” noise model for the regional GPS velocity uncertainties.
- The uncertainties of the GPS velocity estimates are biased due to the length of the GPS timeseries (we have used 1 to 7 years of GPS data).