

How does the Southern Ocean palaeoenvironment during Marine Isotope Stage 5e compare to the modern?

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Abstract

Marine Isotope Stage (MIS) 5e (130-116 ka) represents an important ‘process analogue’ for understanding the climatic feedbacks and responses likely active under future anthropogenic warming. Reconstructing the Southern Ocean (SO) palaeoenvironment during MIS 5e and comparing it to the present day provides insights into the different responses of the SO sectors to a warmer climate. This study presents new records from seven marine sediment cores for MIS 5e together with their surface sediment records; all cores are located south of 55 °S. We investigate changes in diatom species assemblage and the accompanying variations in sea surface temperatures, winter sea-ice extent (WSIE) and glacial meltwater flux. All records show warmer conditions and a reduced WSIE during MIS 5e relative to the surface sediments. While the Pacific and Indian Sector records present very stable conditions throughout MIS 5e, the Atlantic Sector records display much more changeable conditions, particularly with respect to the WSIE. These variable conditions are attributed to higher iceberg and glacial meltwater flux in the Weddell Sea. This evidence for increased iceberg and glacial meltwater flux in the Weddell Sea during MIS 5e may have significant implications for understanding the stability of the West Antarctic Ice Sheet, both during MIS 5e and under future warming.

Keywords

MIS 5e; Palaeoenvironment; Diatom; Southern Ocean; Marine Sediment Core

1. Introduction

The Antarctic continent and Southern Ocean (SO) play a critical role in the global climate system through the albedo-radiation feedbacks induced by the vast extent of the Antarctic ice sheets and SO sea ice. Sea-ice cover also regulates heat and gas exchange between the SO and the atmosphere and, through changes in sea surface temperatures (SSTs) and salinity, affects Antarctic Bottom Water production and thereby impacts upon global ocean circulation (Abernathey et al. 2016).

33 Rising greenhouse gas concentrations are driving current global warming, with polar regions warming
34 twice as fast as the global average (IPCC 2019). High latitudes have a greater sensitivity to radiative
35 forcing and therefore tend to amplify the effects of rising temperatures through ocean and cryosphere
36 feedbacks (Vaughan et al. 2013). This greater sensitivity makes higher latitudes particularly important
37 regions for studying and understanding climate dynamics. However, the very short length of
38 observational records limits our understanding of the underlying processes. Studying past warm
39 periods, when the extent of land ice and sea ice were reduced, can help guide our understanding of
40 the impact of predicted future climate change in these key regions.

41 Marine Isotope Stage (MIS) 5e (130 – 116 ka) was the last period when the Antarctic region was
42 substantially warmer. SSTs and global mean annual atmospheric temperatures peaked at around 0.8
43 °C warmer during MIS 5e than present (Otto-Bliesner et al. 2013, Capron et al. 2014, Fischer et al.
44 2018) and global sea levels were possibly 5-9 m higher than now (Kopp et al. 2009). Proxy
45 reconstructions of mean annual SSTs in middle to low latitudes (between 51 °N and 51 °S) peaked at
46 just 0.5 ± 0.3 °C warmer than preindustrial during MIS 5e (Hoffman et al. 2017) whereas model results
47 suggest that summer SSTs in the SO peaked at 1.8 ± 0.8 °C warmer than preindustrial (Capron et al.
48 2017), indicating strong polar amplification during MIS 5e. Unlike future anthropogenic warming, MIS
49 5e peak temperatures were orbitally forced rather than primarily through rising greenhouse gas
50 concentrations, making MIS 5e an important ‘process analogue’ for understanding the climate
51 mechanisms and natural feedbacks that will be active under future warmer conditions (Stone et al.
52 2016).

53 In the modern SO there is substantial spatial heterogeneity in the observed sea-ice trends, with sea-
54 ice reductions in the Bellingshausen and Amundsen seas concurrent with increases in the (outer)
55 Weddell Sea and Ross Sea Sectors (Stammerjohn et al. 2008a, Hobbs et al. 2016, Parkinson 2019).
56 Trends in modern SO surface, deep and bottom water temperatures display similar heterogeneity
57 (Maheshwari et al. 2013, Schmidtko et al. 2014). Model simulations are unable to replicate the
58 observations of recent sea-ice change without reducing the regional warming trends (Rosenblum &
59 Eisenman 2017). These difficulties are indicative of the complexities of the climate dynamics, which
60 drive SST and sea-ice change in the SO today (Stammerjohn et al. 2008b, King 2014, Hobbs et al. 2016,
61 Purich et al. 2016). Model deficiencies in the SO region suggest that the large signal-to-noise ratio in
62 SO temperature and sea-ice conditions during MIS 5e will be useful for improving climate model
63 simulations of this region for warmer climates.

64 The diatom assemblages preserved in SO marine sediments are a valuable tool for reconstructing past
65 oceanographic conditions. Diatoms are phototrophic algae, which are prevalent in the SO euphotic

66 zone, and their species distribution patterns are closely related to the environmental conditions in the
67 surface waters, principally the sea ice cover and SSTs (Zielinski & Gersonde 1997, Gersonde & Zielinski
68 2000, Armand et al. 2005, Crosta et al. 2005, Romero et al. 2005, Esper et al. 2010). Several previous
69 studies have used marine sediment core records to reconstruct the palaeoceanographic conditions
70 during MIS 5e, often alongside model simulations (Bianchi & Gersonde 2002, Turney & Jones 2010,
71 Otto-Bliesner et al. 2013, Capron et al. 2014, Hoffman et al. 2017, Turney et al. 2020). However, these
72 studies contain few, or no, marine records from south of 55 °S and therefore are unable to capture
73 the MIS 5e environmental conditions in the Antarctic Zone (south of the Polar Front). Due to the
74 uncertainties in the chronologies of SO proxy records (Govin et al. 2015), previous studies have often
75 either averaged SSTs across MIS 5e (Cortese et al. 2013, Turney et al. 2020) or assumed peak SSTs
76 occur synchronously throughout the SO and are coincident with peak atmospheric temperatures in
77 Antarctica (Otto-Bliesner et al. 2013, Capron et al. 2014, Hoffman et al. 2017).

78 This study aims to compare the environmental conditions in the SO between MIS 5e and the modern
79 using diatom assemblage data preserved in marine sediments. New MIS 5e assemblage data from
80 seven sediment cores located south of 55 °S (Figure 1) are compared to the surface sediment
81 assemblages to determine whether all three sectors of the SO had warmer SSTs and a reduced winter
82 sea-ice extent (WSIE) during MIS 5e, relative to the modern, and whether changes in SSTs and WSIE
83 during MIS 5e occurred in a uniform pattern across all SO sectors.

84 2. Modern oceanography

85 The SO comprises the southern-most basins of the Atlantic, Indian and Pacific oceans and acts as a
86 linkage between the water masses and oceanic circulation within these basins. The dominant
87 oceanographic feature in the modern SO is the clockwise flowing Antarctic Circumpolar Current (ACC),
88 which forms a continuous belt separating subtropical waters to the north from Antarctic waters to the
89 south (Orsi et al. 1995). The ACC is characterised by five fronts, marked by steep horizontal density
90 gradients associated with specific SSTs and salinities (Orsi et al. 1995, Dong et al. 2006, Sokolov &
91 Rintoul 2009). The most southerly fronts are the Southern Boundary of the ACC and the Southern ACC
92 Front (Orsi et al. 1995), which do not separate distinct surface water masses and thus will not be
93 considered in this study. The Subtropical Front marks the northern boundary of the ACC, separating
94 Subantarctic and Subtropical surface waters, and is located too far north (~40 °S) to influence the core
95 sites used in this study. The Subantarctic Front is the most northerly of the remaining two fronts (white
96 line in Figure 1) and is marked by SSTs ≥ 8 °C to the north and SSTs ≤ 7 °C to the south (Meinen et al.
97 2003). The final ACC front is the Polar Front (black line in Figure 1) which is marked by the subsurface
98 (200 m) 2 °C isotherm (Orsi et al. 1995) and generally corresponds to SSTs of ~2-3 °C (Dong et al. 2006).

99 The Subantarctic Front separates the Subantarctic Zone (to the north) from the Polar Front Zone (to
 100 the south), which is in turn separated from the more southerly Antarctic Zone by the Polar Front (Orsi
 101 et al. 1995).

102 The flow of deep and bottom water masses and the locations of the ACC fronts are strongly influenced
 103 by the SO bathymetry. Bathymetric highs, such as the Kerguelen Plateau and Pacific-Antarctic Ridge,
 104 and oceanic gateways, such as the Drake Passage and Tasman Gateway, act to ‘pin’ and constrain ACC
 105 fronts and restrict their latitudinal migration (Dong et al. 2006).

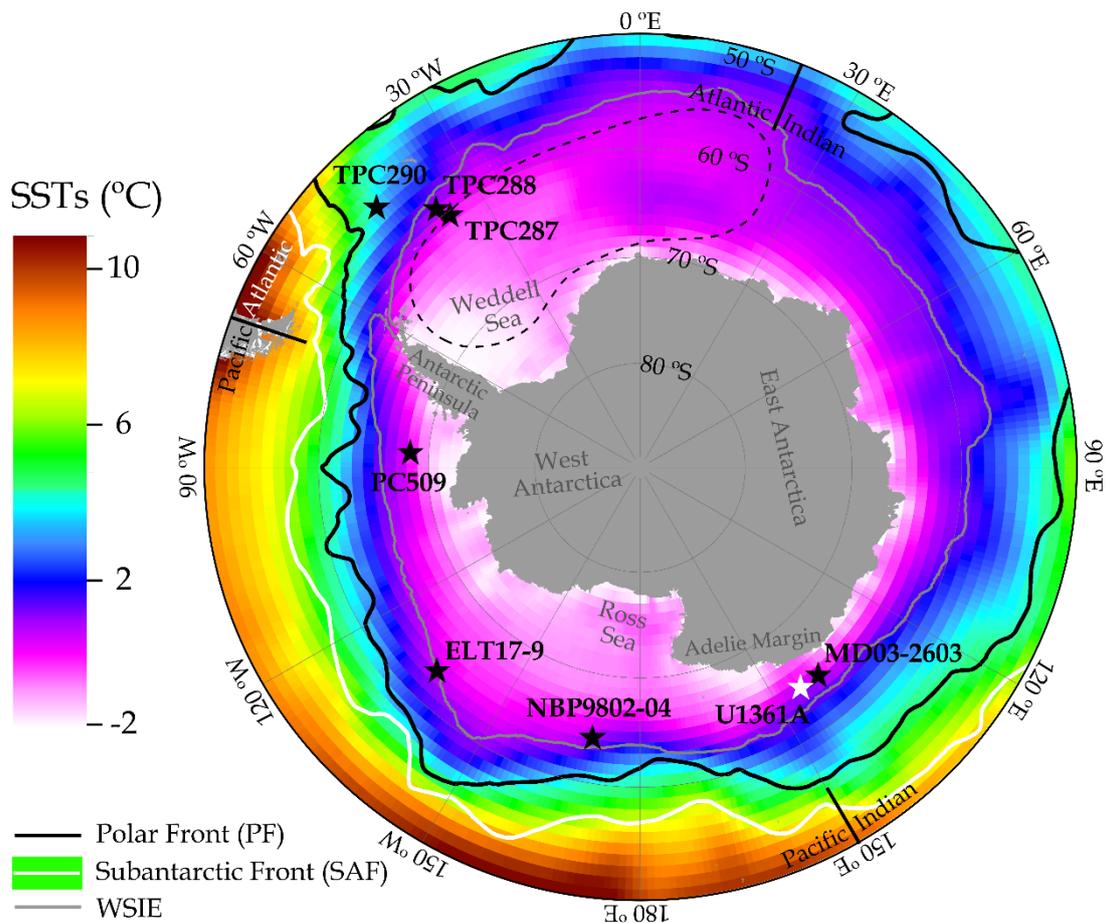


Figure 1: Map of core locations (black stars) with modern SSTs, locations of SO fronts and WSIE. The white star indicates a surface sediment sample (U1361A) assumed to contain a modern diatom assemblage characteristic for the Indian Sector. The white line marks the modern Subantarctic Front (position from Orsi et al. (1995)), the solid black line marks the modern Polar Front (position from Trathan et al. (2000)), the dashed black line marks the extent of the Weddell Gyre (Vernet et al. 2019) and the grey line marks the mean September sea-ice extent from 1981-2010 (data from Fetterer et al. (2017)). The background shadings display the mean annual SSTs from 1981-2010 using the COBE-SST2 dataset provided by the NOAA PSL, Boulder, Colorado, USA (<https://psl.noaa.gov/>). The boundaries between the three SO sectors (Atlantic, Indian and Pacific) are highlighted by straight black lines.

106 **3. Materials and methods**

107 **3.1. Core sites**

108 This study presents new MIS 5e diatom assemblage data from seven marine sediment cores (Table 1
109 & Figure 1) in all sectors of the SO – three in the Atlantic Sector (70 °W – 20 °E), one in the Indian
110 Sector (20 °E – 150 °E) and three in the Pacific Sector (150 °E – 70 °W). These cores were selected as
111 they contain >20 cm thick intervals of diatom rich MIS 5e sediments and are located further south
112 than almost all cores with existing MIS 5e diatom records (Chadwick et al. 2020). Seafloor surface
113 sediment from an additional Indian Sector core, International Ocean Discovery Program (IODP)
114 Expedition 318 Hole U1361A, is also included.

Core	Latitude, Longitude	Water depth (m)	Cruise, Year	Ship	Core length (cm)
TPC290	55.55 °S, 45.02 °W	3826	JR48, 2000	<i>RRS James Clark Ross</i>	1179*
TPC288	59.14 °S, 37.96 °W	2864	JR48, 2000	<i>RRS James Clark Ross</i>	940*
TPC287	60.31 °S, 36.65 °W	1998	JR48, 2000	<i>RRS James Clark Ross</i>	615*
ELT17-9	63.08 °S, 135.12 °W	4935	ELT17, 1965	<i>R/V Eltanin</i>	2018
NBP9802-04	64.20 °S, 170.08 °W	2696	PA9802, 1998	<i>RV/IB Nathaniel B. Palmer</i>	740
MD03-2603	64.28 °S, 139.38 °E	3320	MD130, 2003	<i>R/V Marion DuFresne II</i>	3033
U1361A	64.41 °S, 143.89 °E	3459	IODP Exp. 318, 2010	<i>JOIDES Resolution</i>	38800
PC509	68.31 °S, 86.03 °W	3559	JR179, 2008	<i>RRS James Clark Ross</i>	989

Table 1: Details of the location and recovery information for the eight sediment cores analysed in this study. Cores are ordered by latitude. For IODP Site U1361, only the core top surface sample from Hole U1361A was analysed. *For the three TPC cores (TPC290, TPC288 and TPC287), a trigger core and piston core were taken at the same location, and the records were spliced together to form a composite trigger-piston core record.

115 **3.2. Diatom counts**

116 For the diatom assemblage data, microscope slides were produced using a method adapted from
117 Scherer (1994). Samples of 7-20 mg were exposed to 10% hydrochloric acid to remove any carbonate,
118 30% hydrogen peroxide to break down organic material and a 4% sodium hexametaphosphate
119 solution to promote disaggregation during their placement in a warm water bath for a minimum of 12
120 hours. The material was homogenised into a ~10 cm water column and allowed to settle randomly
121 onto coverslips over a minimum of 4 hours. The water was drained away and coverslips were mounted
122 on microscope slides with Norland Optical Adhesive (NOA 61). Slides were investigated with a light
123 microscope (Olympus BH-2 at x1000 magnification) and a minimum of 300 diatom valves were
124 counted for each sample.

Diatom species/genera	Modern summer SST (°C)	Modern sea-ice duration (months/yr)
<i>Actinocyclus actinochilus</i>	-0.5 – 0.5 ^a	7.5 - 9 ^a
<i>Azpeitia tabularis</i>	1 – 22.5 ^b	0 – 3.5 ^b
<i>Chaetoceros</i> rs.	-1.3 – 3.5 ^a	0 – 10.5 ^a
<i>Eucampia antarctica</i>	-2 – 9.5 ^c	–
<i>Fragilariopsis curta</i>	-1.3 – 2.5 ^a	5 – 10.5 ^a
<i>Fragilariopsis cylindrus</i>	-1.3 – 1 ^a	7.5 – 10.5 ^a
<i>Fragilariopsis kerguelensis</i>	-1 – 22 ^d	0 – 9 ^d
<i>Rhizosolenia antennata</i> f. <i>semispina</i>	0.5 – 2 ^e	1 – 3.5 ^d

Table 2: Modern summer SSTs and sea-ice duration ranges for diatom species and genera that are presented in this study. The SST and sea-ice duration ranges for the present day are based on surface samples where the listed species/genera is >2 % of the assemblage. *Eucampia antarctica* does not show any clear association with modern sea-ice duration. ^a Armand et al. (2005), ^b Romero et al. (2005), ^c Zielinski & Gersonde (1997), ^d Crosta et al. (2005), ^e Armand & Zielinski (2001).

125 Diatom relative abundances for each sediment core are reported for species or groups with well-
126 constrained present-day ecologies/habitats (Table 2). *Actinocyclus actinochilus* is a cold-water species
127 (Table 2) generally found in low abundances (<3 %) in SO seafloor surface sediment samples within
128 the maximum WSIE (Armand et al. 2005, Esper & Gersonde 2014b). Increasing relative abundances of
129 this species in southern high latitude sediments suggest colder SSTs and more severe sea-ice cover.
130 *Azpeitia tabularis* is a warm-water species (Table 2), reaching up to 20 % of the total diatom
131 assemblages in the Subantarctic Zone and presenting relative abundances <5 % in surface sediments
132 south of the Polar Front (Romero et al. 2005, Esper & Gersonde 2014b). This species additionally has
133 a southerly occurrence restricted by the maximum WSIE (Zielinski & Gersonde 1997). Increasing
134 abundances of this group in southern high latitude sediments therefore suggest warmer SSTs and ice-
135 free conditions.

136 The abundance of *Chaetoceros* rs. in SO surface sediments is dominantly influenced by meltwater
137 surface stratification and nutrient availability (Armand et al. 2005), with high surface stratification and
138 nutrient availability resulting in *Chaetoceros* rs. dominated assemblages (>60 %) in coastal Antarctic
139 systems (Leventer & Dunbar 1988, Leventer 1991, 1992, Crosta et al. 1997). High *Chaetoceros* rs.
140 abundances are also associated with moderately consolidated winter sea ice (sea-ice duration = 3-9
141 months/yr), but this relationship is still poorly understood (Armand et al. 2005). The *Chaetoceros* rs.
142 abundances in this study include a small number (up to 1 %) of *Chaetoceros (Hyalochaete)* vegetative
143 cells in some samples. Other than *Chaetoceros* rs., *Fragilariopsis kerguelensis* is the dominant diatom

144 species/group in SO surface sediments, with the greatest abundances found in locations with year-
145 round open ocean conditions (Crosta et al. 2005, Cefarelli et al. 2010, Esper et al. 2010) and, as a
146 result, changes in *F. kerguelensis* abundance are often negatively correlated with changes in
147 *Chaetoceros* rs. abundance. Taken together, an increase in the relative abundance of *F. kerguelensis*,
148 and concurrent decrease in the relative abundance of *Chaetoceros* rs. in our core records indicate a
149 shift from conditions with moderate sea-ice cover and stratified surface waters to open ocean
150 conditions with low or no winter sea-ice cover.

151 *Fragilariopsis curta* and *F. cylindrus*, composing the FCC group, are sea-ice associated species (Kang &
152 Fryxell 1992, Beans et al. 2008), presenting their maximum abundances in modern sediments at winter
153 sea-ice concentrations >70 % and SSTs <1 °C (Armand et al. 2005, Esper & Gersonde 2014b, a). The
154 FCC group is used as an indicator of winter sea-ice presence (Gersonde & Zielinski 2000), with
155 abundances >3 % associated with locations south of the mean WSIE, abundances 1-3 % found between
156 the mean and maximum WSIE and abundances <1 % being indicative of conditions north of the
157 maximum WSIE (Gersonde & Zielinski 2000, Gersonde et al. 2005). Increasing relative abundances of
158 the FCC group in our cores therefore infer heavy sea-ice conditions and cold SSTs.

159 The abundance of *Eucampia antarctica* in SO surface sediments does not show a clear pattern relative
160 to SSTs or sea-ice extent (Zielinski & Gersonde 1997), probably because its two varieties have usually
161 been combined in abundance counts. The cold variety of *E. antarctica* has, however, been related to
162 iceberg flux, with high iceberg flux promoting high *E. antarctica* abundances through meltwater-
163 induced buoyancy and high iron availability (Burckle 1984, Fryxell & Prasad 1990, Allen 2014). Based
164 on restricted modern studies, high relative abundances of *E. antarctica* cold variety encountered
165 downcore have been used as an indicator of iceberg or marine-terminating glaciers melting (Barbara
166 et al. 2016). The *E. antarctica* relative abundances reported in this study only include valves from the
167 cold variety. *Rhizosolenia antennata* f. *semispina* reaches its maximum abundance in SO surface
168 sediments located within, and just north of, the mean WSIE (Crosta et al. 2005) and is also an indicator
169 of high meltwater flux and surface stratification (Allen et al. 2005).

170 Surface sediment samples from four of the core sites (TPC290, TPC288, TPC287 and PC509) are used
171 to obtain modern diatom assemblages for those sites. For core MD03-2603, the surface sediment
172 sample from the nearby Site U1361 (Table 1 & Figure 1) is used for the modern diatom assemblage.
173 There was no surface material available for the central Pacific Sector cores (ELT17-9 and NBP9802-04),
174 so the MIS 5e diatom assemblages in these cores are not compared against any modern assemblages.
175 Using surface sediment samples to represent the modern surface water conditions is consistent with
176 previous studies (Zielinski & Gersonde 1997, Crosta et al. 1998, Armand et al. 2005, Crosta et al. 2005,

177 Romero et al. 2005, Esper & Gersonde 2014b, a), but we note that the assemblage preserved in surface
 178 sediments is likely an integrated signal of up to 500 years (Miklasz & Denny 2010).

179 **4. Age models**

180 *4.1. Published chronologies*

181 Of the seven sediment cores, for which MIS 5e data are presented in this study, five utilise previously
 182 published age models (Table 3). The chronology for cores TPC290 and TPC288 is given in Pugh et al.
 183 (2009) and utilises the correlation between the magnetic susceptibility (MS) record in marine
 184 sediment cores from the Scotia Sea and the dust record in the EPICA Dome C (EDC) ice core over past
 185 glacial-interglacial cycles (Pugh et al. 2009, Weber et al. 2012). In both cores this chronology is
 186 combined with the abundance stratigraphy of the radiolarian species *Cycladophora davisiana*, with
 187 the e₃ low abundance event indicating MIS 5e (Brathauer et al. 2001). The Termination II tiepoint in
 188 TPC290 was adjusted from the 7.11 metres below seafloor (mbsf) given in Pugh et al. (2009) to 7.23

Core	Latitude, Longitude	SO sector	Chronology for MIS 5e	MIS 5e sampling interval (ka)
TPC290	55.55 °S, 45.02 °W	Atlantic	Correlating MS from TPC290 to EDC ice core dust record combined with <i>C. davisiana</i> abundances (Pugh et al. 2009)*	0.6
TPC288	59.14 °S, 37.96 °W	Atlantic	Correlating MS from TPC288 to EDC ice core dust record combined with <i>C. davisiana</i> abundances (Pugh et al. 2009)	0.7 - 1.1
TPC287	60.31 °S, 36.65 °W	Atlantic	Correlating MS from TPC287 to MS from TPC288 (this study; Figure 2)	0.5 - 1.2
ELT17-9	63.08 °S, 135.12 °W	Pacific	Combined abundance stratigraphies of <i>E. antarctica</i> and <i>C. davisiana</i> on SPECMAP age scale (Chase et al. 2003)	1.2 - 1.3
NBP9802-04	64.20 °S, 170.08 °W	Pacific	Correlating MS from NBP9802-04 to EDC ice core dust record combined with Last Occurrence Datum of <i>H. karstenii</i> (Williams 2018)	1.4
MD03-2603	64.28 °S, 139.38 °E	Indian	Correlating Ba/Al and Ba/Ti ratios from MD03-2603 to LR04 benthic oxygen isotope stack combined with diatom biostratigraphy (Presti et al. 2011)	0.4 - 0.9
PC509	68.32 °S, 86.03 °W	Pacific	Correlating wet bulk density (=proxy mirroring biogenic opal content) from PC509 to the LR04 benthic oxygen isotope stack (this study; Figure 3)	0.6 - 1.3

Table 3: Summary of the locations and chronologies for the seven sediment cores analysed in this study. Cores are ordered by latitude. *For core TPC290 the chronology was adjusted from the age model previously published in Pugh et al. (2009) by shifting the Termination II tiepoint to improve alignment of its MS signal with the EDC dust record.

189 mbsf to improve the alignment of the MS signal of the sediments with the EDC dust record. The
190 chronology for core NBP9802-04 also utilises the correlation between sediment MS and EDC dust
191 (Pugh et al. 2009), alongside the presence/absence of the diatom species *Hemidiscus karstenii*
192 (Williams 2018), which is a biostratigraphic marker for the MIS 6/7 boundary (Burckle et al. 1978). All
193 three of these cores have chronologies tied to the EDC3 time scale (Parrenin et al. 2007).

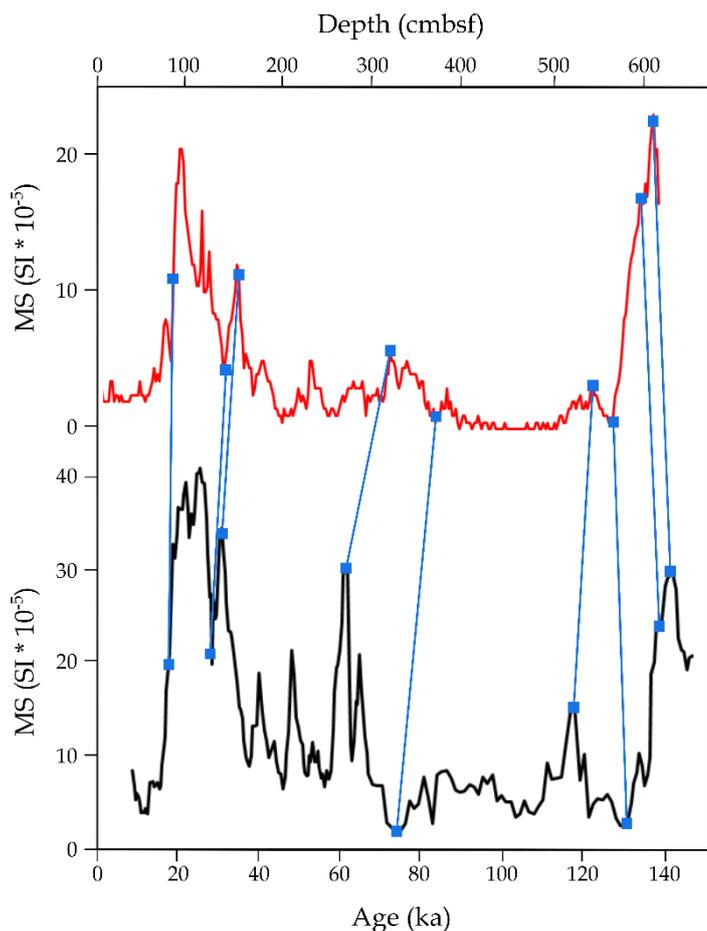
194 Cores MD03-2603 and ELT17-9 have published chronologies tied to the LR04 and SPECMAP age scales,
195 respectively (Table 3). For core MD03-2603, Presti et al. (2011) correlated the downcore records of
196 Ba/Al and Ba/Ti ratios, which are palaeo-productivity proxies, to the LR04 benthic foraminifera $\delta^{18}\text{O}$
197 stack (Lisiecki & Raymo 2005). The chronology for core ELT17-9 was published by Chase et al. (2003)
198 and uses abundances for *C. davisiana* (Hays J. unpublished data) and *E. antarctica* (Burckle L.H.
199 unpublished data), which provide well established abundance stratigraphies (Burckle & Burak 1995,
200 Brathauer et al. 2001). To allow for consistent comparison of timings between cores, both MD03-2603
201 and ELT17-9 are translated across onto the EDC3 chronology using the conversion tables published by
202 Lisiecki & Raymo (2005) and Parrenin et al. (2013). During MIS 5e, EDC3 ages are ~ 1 ka older than LR04
203 ages, with this offset due to the LR04 chronology being based upon a benthic record that incorporates
204 both sea-level and temperature components (Parrenin et al. 2007). A chronology tuned to surface
205 water rather than deep water changes was chosen because we are investigating environmental
206 changes in the surface ocean.

207 4.2. TPC287 chronology

208 The chronology for core TPC287 was constructed by aligning the downcore MS records in cores
209 TPC287 and TPC288 (Figure 2). TPC287 is located approximately 150 km southeast of TPC288 (Figure
210 1), and thus the MS variations in both cores are expected to occur synchronously across glacial and
211 interglacial cycles. The two MS records were graphically aligned by eye using the AnalySeries software
212 (Paillard et al. 1996) by choosing prominent features as tiepoints (Figure 2 & Table 4).

213 4.3. PC509 chronology

214 The chronology for core PC509 was constructed by visually aligning the wet bulk density, a proxy
215 mirroring biogenic opal content (Busch 1991, Weber et al. 1997, Hillenbrand et al. 2009), to the LR04
216 benthic foraminifera $\delta^{18}\text{O}$ stack (Lisiecki & Raymo 2005) using the AnalySeries software (Paillard et al.
217 1996). Tiepoints were selected in the wet bulk density record at MIS stage and sub-stage boundaries
218 (Table 5 & Figure 3). The MIS 5 sub-stages use the age assignments from Govin et al. (2009), and the
219 ages are translated across from the LR04 chronology to the EDC3 chronology using the conversion
220 table published by Parrenin et al. (2013).



← **Figure 2:** Alignment between the MS downcore records from cores TPC287 (red) and TPC288 (black) using the AnalySeries software (Paillard et al. 1996). The blue squares and connecting lines mark tiepoints between the records. The age model for core TPC288 was published by Pugh et al. (2009).

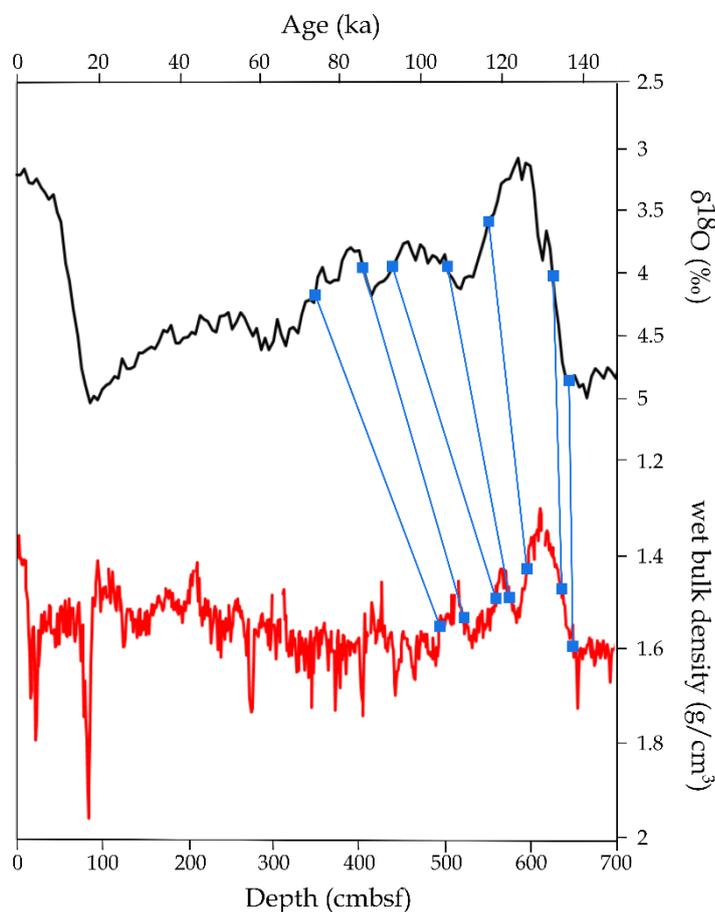
TPC287 depth (mbsf)	TPC288 age (ka)
0.84	17.5
1.39	28.5
1.53	31
3.21	61
3.71	74
5.43	118
5.65	131
5.95	138.5
6.09	142

Table 4: Tiepoints for the TPC287 chronology with depths in core TPC287 being tied to the EDC3 ages published by Pugh et al. (2009) for core TPC288.

PC509 depth (mbsf)	LR04 age (ka)	MIS stage/sub-stage boundary
5.00	73	4-5a
5.20	84	5a-5b
5.64	91	5b-5c
5.78	106	5c-5d
5.97	116	5d-5e
6.38	130	5e-6
6.48	136	-

Table 5: Tiepoints for the PC509 chronology. The wet bulk density record for PC509 was aligned with the LR04 benthic stack using the AnalySeries software (Paillard et al. 1996). The ages for MIS stage/sub-stage boundaries used as tiepoints are listed.

Figure 3: Alignment between the downcore wet bulk density record of core PC509 (red) and the LR04 benthic $\delta^{18}O$ stack (black) using the AnalySeries software (Paillard et al. 1996). The tiepoints are marked by blue squares and connecting lines.



222 **5. Results and discussion**

223 **5.1. MIS 5e diatom assemblages**

224 Relative diatom abundances in the sediments deposited during the time interval 132-120 ka are
225 presented for all seven core sites (Figures 4-6) in order to capture the palaeoenvironmental signal
226 both from the Termination II deglaciation and 'peak' MIS 5e. The *Azpeitia tabularis* and *Actinocyclus*
227 *actinochilus* abundances are low in all seven cores (0.3 ± 0.4 % and 0.5 ± 0.6 %), with cores TPC288,
228 ELT17-9 and PC509 recording only negligible contributions (0.3 ± 0.3 %) of either species (Figures 4-6).
229 Core TPC287 has the largest 'cold signal' (Table 2), with an *A. actinochilus* peak of 2 % at 120 ka (Figure
230 4), and core TPC290 has the greatest 'warm signal' (Table 2) with high *A. tabularis* abundances of 1.8
231 ± 0.7 % and 2.0 ± 0.4 % from 126-124 ka and 121-120 ka, respectively (Figure 4). Core NBP9802-04
232 also has a strong 'warm signal' with *A. tabularis* being present almost throughout MIS 5e (Figure 6).

233 The highest MIS 5e *Chaetoceros* rs. abundances occur in core PC509 (78 ± 4 %). This is likely due to
234 high input of meltwater and nutrients, like in the vicinity of the Antarctic Peninsula, where high
235 meltwater stratification and nutrient availability promote extensive *Chaetoceros* blooms (Crosta et al.
236 1997). The three Atlantic Sector cores (TPC290, TPC288 and TPC287) are dominated by both
237 *Chaetoceros* rs. (38 ± 10 %, 40 ± 8 % and 28 ± 11 %, respectively) and *F. kerguelensis* (34 ± 8 %, 27 ± 6
238 % and 33 ± 12 %, respectively) throughout MIS 5e with peaks in one group coinciding with troughs in
239 the other (Figure 4). This alternation is particularly evident in core TPC287, where *Chaetoceros* rs.
240 abundance declined by ~ 40 % after 131 ka, concurrent with an equivalent increase in *F. kerguelensis*
241 abundance (Figure 4). Both TPC288 and TPC287 have similar *F. kerguelensis* abundance profiles with
242 higher values of 34 ± 4 % and 53 ± 6 %, respectively, between 130-127 ka and 126-124 ka (Figure 4).
243 This contrasts with core TPC290, where the *F. kerguelensis* abundance is lowest (27 ± 4 %) from 128-
244 122 ka (Figure 4). Consistent with the modern distribution pattern published by Crosta et al. (1997),
245 the Indian and central Pacific Sector cores (ELT17-9, NBP9802-04 and MD03-2603) have low
246 *Chaetoceros* rs. abundances (10 ± 2 %, 5 ± 2 % and 17 ± 5 %, respectively) during MIS 5e and are
247 dominated instead by *F. kerguelensis* (63 ± 4 %, 74 ± 4 % and 59 ± 8 %, respectively) (Figures 5 & 6).

248 The FCC abundances in cores TPC288 and TPC287 are very similar with minima (0.3 % and 1 %) early
249 in the 132-120 ka interval, followed by an increase to maxima of ~ 9 % and ~ 15 %, respectively, at ~ 127 -
250 126 ka before decreasing to largely steady abundances of ~ 3 % and ~ 6 %, respectively, between 125
251 and 120 ka (Figure 4). In contrast, the FCC abundances in core TPC290 remain largely consistent at 2.3
252 ± 0.7 % from 130 to 124 ka before gradually declining to a minimum of ~ 0.6 % at 120 ka (Figure 4).

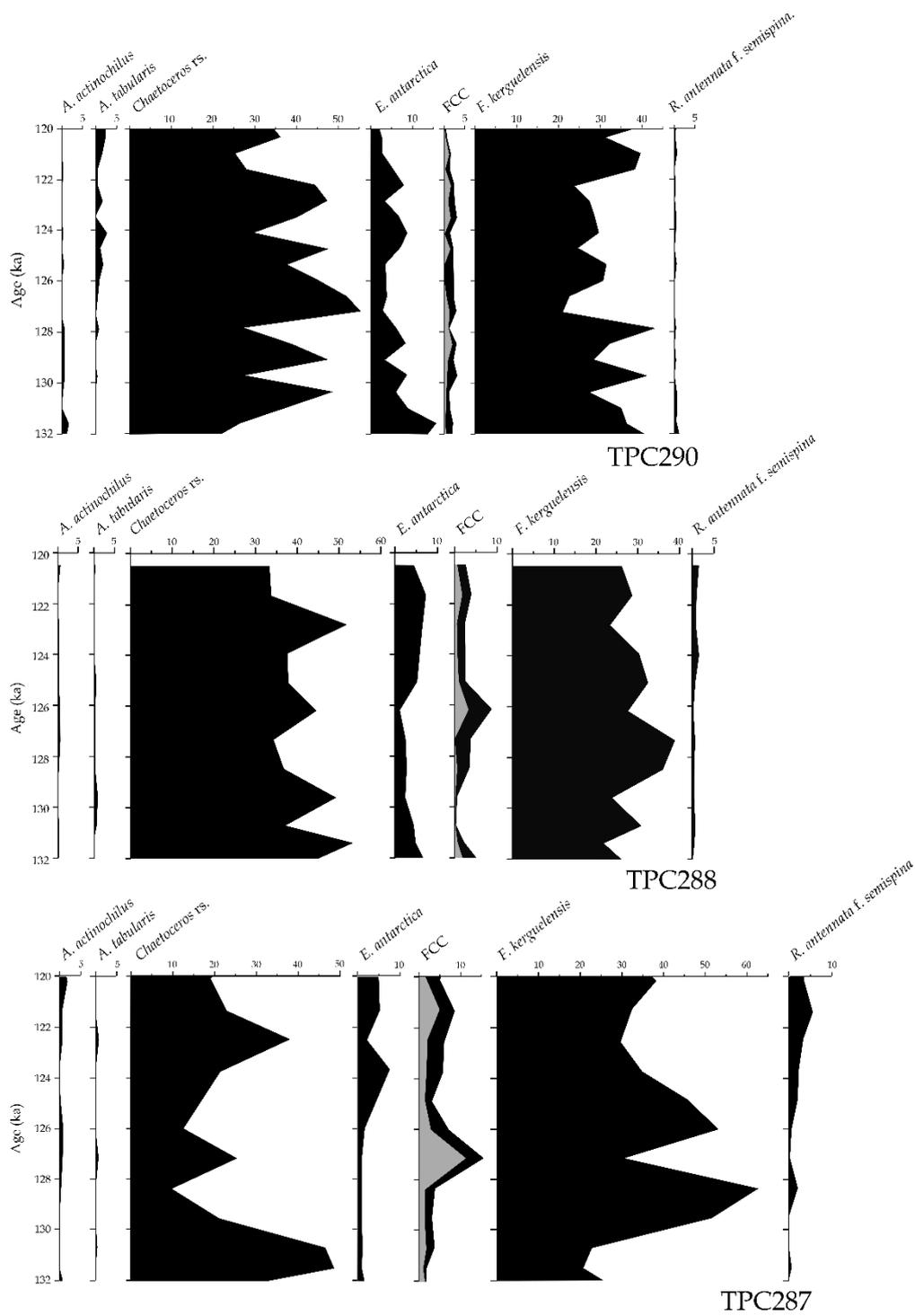


Figure 4: Downcore relative diatom abundances for the three Atlantic Sector cores (TPC290, TPC288 and TPC287) covering the 132-120 ka period. For the FCC group, the grey shading shows the *F. cylindrus* abundance and the black shading shows the *F. curta* abundance.

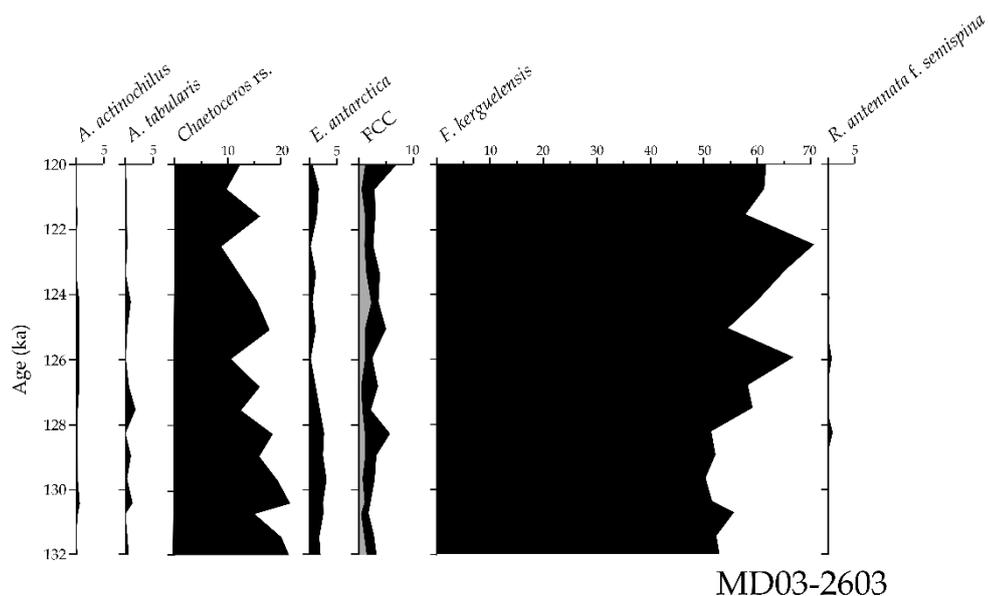


Figure 5: Downcore relative diatom abundances for the Indian Sector core (MD03-2603) covering the 132-120 ka period. For the FCC group, the grey shading shows the *F. cylindrus* abundance and the black shading shows the *F. curta* abundance.

254 In the Pacific Sector cores (ELT17-9, NBP9802-04 and PC509), the FCC abundances are largely uniform
 255 throughout MIS 5e, at ~2 %, ~3.5 % and ~6 % respectively (Figure 6). Cores ELT17-9 and NBP9802-04
 256 reach a FCC abundance minimum of ~0.9 % at ~124 ka, whereas the minimum (4.5 %) in PC509 occurs
 257 earlier at ~129 ka, consistent with the more southerly Atlantic Sector cores (TPC288 and TPC287)
 258 (Figures 4 & 6). Core MD03-2603 also has a largely constant FCC abundance of ~3.5 % throughout MIS
 259 5e, although it reaches minima of 1.7 % at ~130.5 ka and 2.2 % at ~127.5 ka, both of which are
 260 concurrent with *A. tabularis* abundance peaks of >1 % (Figure 5).

261 The *Eucampia antarctica* (cold variety) and *Rhizosolenia antennata* f. *semispina* abundances in TPC288
 262 and TPC287 show coincident increases after 126 ka (Figure 4). This pattern is not seen in TPC290,
 263 where the *R. antennata* f. *semispina* abundance remains low (0.3 ± 0.4 %) throughout MIS 5e and the
 264 highest *E. antarctica* abundances (>10 %) are observed before 131 ka (Figure 4). In the four Indian and
 265 Pacific Sector cores both *R. antennata* f. *semispina* and *E. antarctica* have low abundances (<5 %)
 266 throughout the 132-120 ka interval (Figures 5 & 6).

267 5.2. Comparison between Termination II-MIS 5e and recent diatom assemblages

268 The diatom abundances in the surface sediments are compared with the average abundances in three
 269 Termination II-MIS 5e time slices – early (132-130 ka), mid (130-125 ka) and late (125-120 ka) (Table
 270 6). These MIS 5e time windows are chosen to reconstruct average palaeoenvironmental conditions
 271 for the end of the deglaciation, the peak of MIS 5e and the later stage of MIS 5e, respectively, thereby

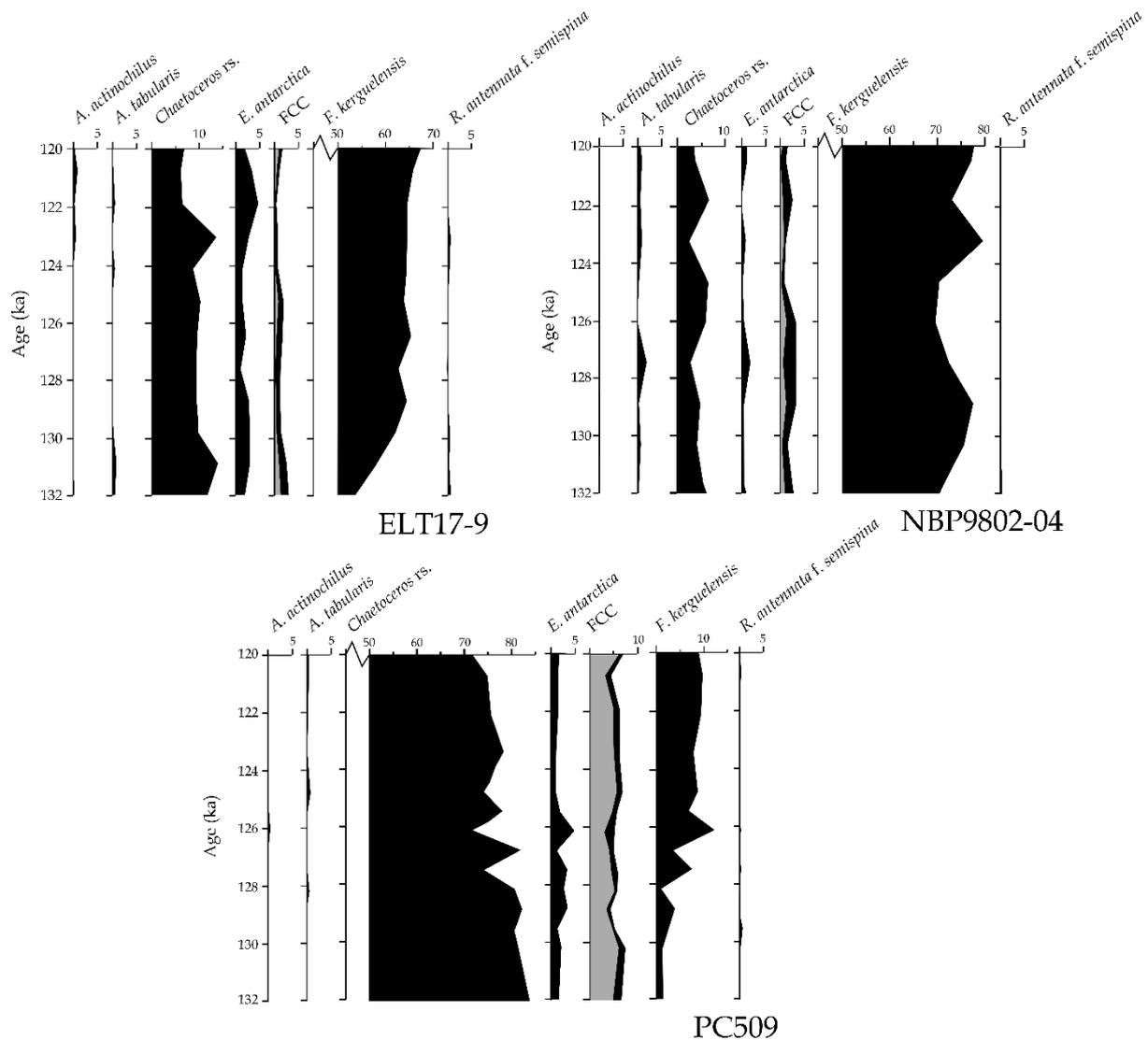


Figure 6: Downcore relative diatom abundances for the three Pacific Sector cores (ELT17-9, NBP9802-04 and PC509) covering the 132-120 ka period. For the FCC group, the grey shading shows the *F. cylindrus* abundance and the black shading shows the *F. curta* abundance. Note that the scales for species/groups with abundances >50 % of the assemblage throughout MIS 5e start at 50 %.

272 following the divisions suggested by Capron et al. (2014). To ensure there are data from at least 3
 273 samples for each time slice, 1 sample older than 132 ka had to be included in the 132-130 ka window
 274 for five cores (TPC288, TPC287, ELT17-9, NBP9802-04 and PC509). Figures 7 & 8 show the differences
 275 in spatial and temporal relative abundances for four of the key groups (*Chaetoceros rs.*, *E. antarctica*,
 276 FCC and *F. kerguelensis*) in the three Termination II-MIS 5e intervals and the surface sediments.
 277 Diatom assemblages in the surface sediments are consistent with the modern environmental setting.
 278 *A. actinochilus* and *A. tabularis* have similar, but opposing, offsets (<1 %) between the surface
 279 sediments and the three time slices in all the cores (Table 6). The higher abundances of *A. tabularis*
 280 and lower abundances of *A. actinochilus* in the MIS 5e sediments when compared to the surface

281 sediments supports the warmer than present conditions expected during MIS 5e (Capron et al. 2014).
282 Care should be taken when interpreting such small changes in species abundance. However, because
283 both *A. actinochilus* and *A. tabularis* occur in very low relative abundances (<3-5 %) throughout the
284 Antarctic Zone of the SO (Armand et al. 2005, Romero et al. 2005, Esper et al. 2010), even small
285 variations in their relative abundances, especially across multiple consecutive downcore sediment
286 samples, can indicate substantial environmental shifts.

287 Most of the cores have largely similar *Chaetoceros* rs. abundances across all three time slices and in
288 the surface sediments. Core TPC287 is a clear exception, with a decrease of >20 % between the early
289 (42.5 %) and the mid and late time slices (17.2 % and 23.4 % respectively) (Figure 8). The modern
290 *Chaetoceros* rs. abundance for TPC287 (33.3 %) is most similar to the early time slice (42.5 %) but is
291 ~10 % different from any of the time slices (Figure 8). The combined abundance of *F. kerguelensis* and
292 *Chaetoceros* rs. in TPC287 is very similar between the early and mid time slices (65.8 % and 66.7 %
293 respectively) indicating that the decrease in *Chaetoceros* rs. is almost exactly matched by an increase
294 in *F. kerguelensis* (Figure 4). This transition from a *Chaetoceros* rs. dominated assemblage during
295 Termination II to a *F. kerguelensis* dominated one during the mid and late time slices is likely related
296 to a change in the major oceanographic influence at this site. The stratified surface waters of the
297 Weddell Gyre overly core site TPC287 (Vernet et al. 2019), and, thus, clockwise lateral transport of
298 robust *Chaetoceros* rs. from the Antarctic Peninsula and Weddell Sea Embayment (Crosta et al. 1997)
299 may have caused the high *Chaetoceros* rs. abundances in this core. A poleward shift of the northern
300 boundary of the Weddell Gyre during MIS 5e, as indicated by multiple CMIP3 and CMIP5 models under
301 a warmer than present climate (Meijers et al. 2012, Wang 2013), and the accompanying southerly
302 displacement in surface water masses, would result in the replacement of high *Chaetoceros* rs.
303 abundances with high *F. kerguelensis* abundances, indicative of open ocean conditions (Hasle 1969,
304 Cefarelli et al. 2010). A reduction in the longitudinal extent of the Weddell Gyre during MIS 5e is
305 supported by other core records from the SW Indian Sector (Ghadi et al. 2020). The Indian and Pacific
306 Sector cores show similar *F. kerguelensis* abundance patterns with increasing abundances in the mid
307 and late time slices (Figure 8). All cores for which a surface sample is available have greater *F.*
308 *kerguelensis* abundances in the mid and late time slices than in the surface sediments (Figure 8),
309 indicating increased open ocean conditions during MIS 5e.

310 Consistent with the evidence of increased open ocean conditions during MIS 5e, the FCC abundances
311 in all cores indicate a reduced WSIE during MIS 5e relative to the surface sediments (Figure 8). All
312 three Atlantic cores share similar patterns in FCC abundances, with the lowest abundances occurring
313 during Termination II (average 2.1 %) followed by an increase during the mid time slice (average 4.5
314 %) and subsequent decrease during the late time slice (average 3.6 %) (Figure 8). The amplitude of

315 these abundance changes exhibits a N-S trend with the highest amplitude shifts at the most southerly
316 site (TPC287) and the least variation at the northernmost site (TPC290). The FCC abundances in core
317 PC509 are consistently >3 % from 132 ka to 120 ka which indicates that the site was located to the
318 south of the mean WSIE throughout this entire period (Figure 6). The two central Pacific Sector cores
319 (ELT17-9 and NBP9802-04) do not have surface sediment assemblages to compare with. However, the
320 FCC abundances (Figure 6) suggest that during Termination II and MIS 5e, site ELT17-9 (FCC ~1-2 %)
321 was located on the edge of the maximum WSIE and site NBP9802-04 (FCC ~3 %) was located near the
322 mean WSIE until 126 ka, when the winter sea-ice limit retreated. Compared to the modern September
323 sea-ice extent (Figures 1, 7 & 8), the FCC abundances (Figure 8) indicate a southward shift in sea-ice
324 cover for the central Pacific Sector during Termination II-MIS 5e. FCC abundances in core MD03-2603
325 show strong similarity between the mid and late time slices and the surface sediments (Figure 8). The
326 greater WSIE reduction in the Atlantic Sector compared to the Pacific and Indian Sectors supports the
327 pattern of the simulated MIS 5e WSIE minimum in Holloway et al. (2017).

328 The *E. antarctica* (cold variety) abundances are highest in cores TPC290 and TPC288 (5.5 % & 5.8 %
329 respectively), whilst the Indian and Pacific Sector cores have low abundances throughout Termination
330 II and MIS 5e (average 2.6 %) and the modern (average 1.5 %) (Figure 7). Higher *E. antarctica*
331 abundances in the Atlantic Sector cores when compared to the Pacific and Indian Sector cores (Figure
332 5) are likely linked to greater influence of iceberg flux from the Weddell Sea Sector than the other
333 Antarctic embayments (Death et al. 2014). The high *E. antarctica* abundances for the Termination II
334 interval in cores TPC290 and TPC288 when compared to surface sediments (Table 6 & Figure 7) suggest
335 a higher iceberg supply during the deglaciation, which is supported by high accumulation rates of
336 iceberg-rafted debris during this time recorded in Weddell Sea cores from the East Antarctic margin
337 (Diekmann et al. 2003). During the late time slice the high *E. antarctica* abundances (6.2 % and 5.1 %,
338 respectively) in cores TPC288 and TPC287 indicate a later period of substantial iceberg flux, which
339 could reflect a poleward migration and/or expansion of the iceberg tracks over the course of MIS 5e.
340 A poleward displacement of the iceberg tracks would support a contraction of the Weddell Gyre
341 (Tournadre et al. 2016) and suggest a southerly shift in the position of the ACC and wind fields
342 (Gladstone et al. 2001). Present day iceberg trajectories support the greater *E. antarctica* abundance
343 in the surface sediment assemblage of core TPC290 when compared with the *E. antarctica* abundance
344 in surface sediments of cores TPC288 and TPC287 (Silva et al. 2006).

345 5.3. Environmental heterogeneity during MIS 5e

346 Peaks in *A. tabularis* abundance (~2 %) at ~127.5 ka in cores MD03-2603 and NBP9802-04 (Figures 5
347 & 6) and the increased *A. tabularis* abundance between 126-124 ka and 121-120 ka in core TPC290

348 (Figure 4) could be related to higher SSTs and a more southerly Polar Front than the modern. Poleward
349 migration of the Polar Front in the Atlantic Sector during MIS 5e has been concluded from previously
350 published proxy reconstructions (Nürnberg et al. 1997, Bianchi & Gersonde 2002, Howe et al. 2002,
351 Kemp et al. 2010, Chadwick et al. 2020). Unlike NBP9802-04 and MD03-2603, the ELT17-9 and PC509
352 records have very low *A. tabularis* abundances throughout Termination II and MIS 5e (Figure 6) which
353 suggests that, if there was a southerly migration in the Polar Front, it did not occur uniformly across
354 the Pacific and Indian Sectors. Heterogeneous frontal migration during MIS 5e is supported by
355 Chadwick et al. (2020), and is also evident for the modern SO (Freeman et al. 2016). It could be caused
356 by the ‘pinning’ of fronts by bathymetric features in some regions, which will impede their migration
357 (Nghiem et al. 2016).

358 The higher *R. antennata* f. *semispina* abundances in the most southerly Atlantic cores (TPC288 and
359 TPC287, Figure 4) during the interval 126-120 ka indicate that the edge of the mean WSIE was closer
360 to both core sites during this interval than during the rest of MIS 5e (cf. Crosta et al. 2005) and also
361 indicate increased surface meltwater stratification (cf. Allen et al. 2005). This increased meltwater
362 stratification could have resulted either from the annual melting at the mean WSIE (Armand &
363 Leventer 2003) or melting associated with high iceberg flux indicated by the elevated *E. antarctica*
364 abundances (Figures 4 & 7), with high global sea level after 126 ka (Kopp et al. 2013) supporting a
365 large reduction in global ice volume at this time. The abundances of *E. antarctica*, FCC and *R.*
366 *antennata* f. *semispina* during Termination II and MIS 5e document clear environmental differences
367 between the largely stable conditions in the Pacific and Indian Sectors (Figures 5 & 6) and the more
368 variable conditions in the southerly Atlantic Sector (cores TPC288 and TPC287, Figure 4). FCC
369 abundances likely indicate an early (~130-129 ka) WSIE minimum at sites TPC288 and TPC287. This is
370 consistent with the FCC records for cores MD03-2603 and PC509 (Figures 5 & 6) but the substantial
371 re-expansion of WSIE at ~127-126 ka is only seen in the Atlantic Sector records (Figure 4). In the Indian
372 Sector, the similarity in FCC abundances between the Termination II-MIS 5e and surface samples in
373 core MD03-2603 (Figure 8) could be due to the influence of the Australian-Antarctic Basin gyre
374 regulating the position of the WSIE along the Adélie Land margin in East Antarctica (McCartney &
375 Donohue 2007, Carter et al. 2008).

	TPC290				TPC288				TPC287				ELT17-9			NBP9802-04			MD03-2603				PC509			
	132* – 130 ka	130– 125 ka	125 – 120 ka	Modern	132* – 130 ka	130– 125 ka	125 – 120 ka	Modern	132* – 130 ka	130– 125 ka	125 – 120 ka	Modern	132* – 130 ka	130– 125 ka	125 – 120 ka	132* – 130 ka	130– 125 ka	125 – 120 ka	Modern ^a	132* – 130 ka	130– 125 ka	125 – 120 ka	Modern			
<i>A. actinochilus</i>	0.7	0.4	0.1	0.0	0.3	0.4	0.3	0.6	0.2	0.5	0.7	1.0	0.1	0.0	0.5	0.1	0.1	0.0	0.3	0.5	0.2	0.7	0.0	0.1	0.0	0.0
<i>A. tabularis</i>	0.0	0.5	1.2	0.0	0.2	0.5	0.1	0.0	0.1	0.2	0.1	0.3	0.6	0.0	0.4	0.4	0.8	0.7	0.6	0.6	0.4	0.7	0.0	0.1	0.3	0.0
<i>Chaetoceros</i> rs.	38.6	42.5	38.1	46.8	45.2	40.9	39.4	39.0	42.5	17.2	23.4	33.3	12.0	9.7	8.8	6.2	4.6	4.9	19.1	16.1	12.6	12.5	82.9	78.6	76.0	73.5
<i>E. antarctica</i>	10.1	5.0	5.5	6.6	5.5	3.0	6.2	3.4	1.3	1.2	5.1	2.6	2.8	2.1	3.3	1.4	1.1	0.6	2.4	2.0	1.1	1.0	2.2	3.0	1.5	2.0
FCC	1.6	2.5	2.0	7.6	2.6	3.9	3.0	5.3	2.2	7.2	5.8	14.0	2.4	1.6	0.8	2.7	3.4	1.6	2.2	3.5	3.2	4.8	7.0	5.4	6.1	7.8
<i>F. kerguelensis</i>	33.4	31.9	30.7	20.8	23.8	30.2	24.7	23.9	23.3	49.5	37.0	16.6	57.9	64.1	64.9	71.8	73.5	75.3	55.1	58.4	66.3	56.1	1.5	5.5	9.1	5.1
<i>R. antennata</i> f. <i>semispina</i>	0.5	0.2	0.2	0.7	0.9	0.8	2.4	7.8	0.6	0.9	3.6	6.2	0.2	0.3	0.3	0.1	0.0	0.0	0.2	0.2	0.1	0.3	0.3	0.2	0.3	0.0

Table 6: Mean abundances for the three Termination II-MIS 5e time slices (early = 132*-130 ka; mid = 130-125 ka; late = 125-120 ka) at each core site and modern abundances in surface sediments. The 132*-130 ka time slice includes samples older than 132 ka in five of the cores (TPC288, TPC287, ELT17-9, NBP9802-04 and PC509).

^a modern abundances for MD03-2603 are from the surface sediments at IODP Site U1361.

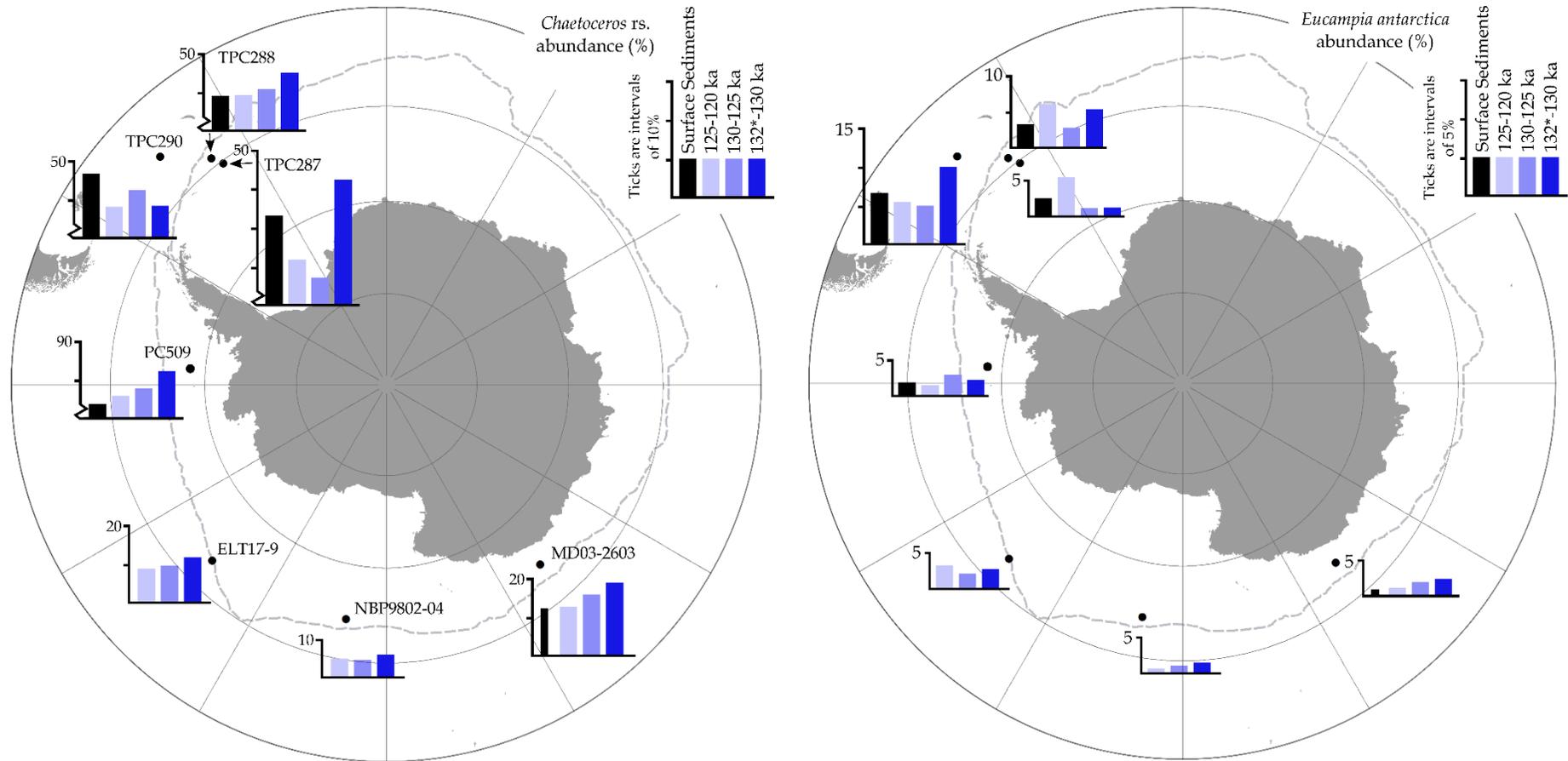


Figure 7: Maps of *Chaetoceros rs.* (LHS) and *Eucampia antarctica* (RHS) abundances in seven marine sediment cores. Modern diatom abundances are marked by black bars and mean abundances during three Termination II-MIS 5e time slices (early = 132*-130 ka; mid = 130-125 ka; late = 125-120 ka) are indicated by bars with different shades of blue. Thin black bars on the MD03-2603 graphs indicate diatom abundances in the surface sediment samples from the nearby IODP Site U1361. The black dots mark the core locations and the grey dashed line is the median modern (1981-2010) September sea-ice extent from Fetterer et al. (2017). *: the 132-130 ka time slice includes samples older than 132 ka to ensure that at least three samples are included in the average.

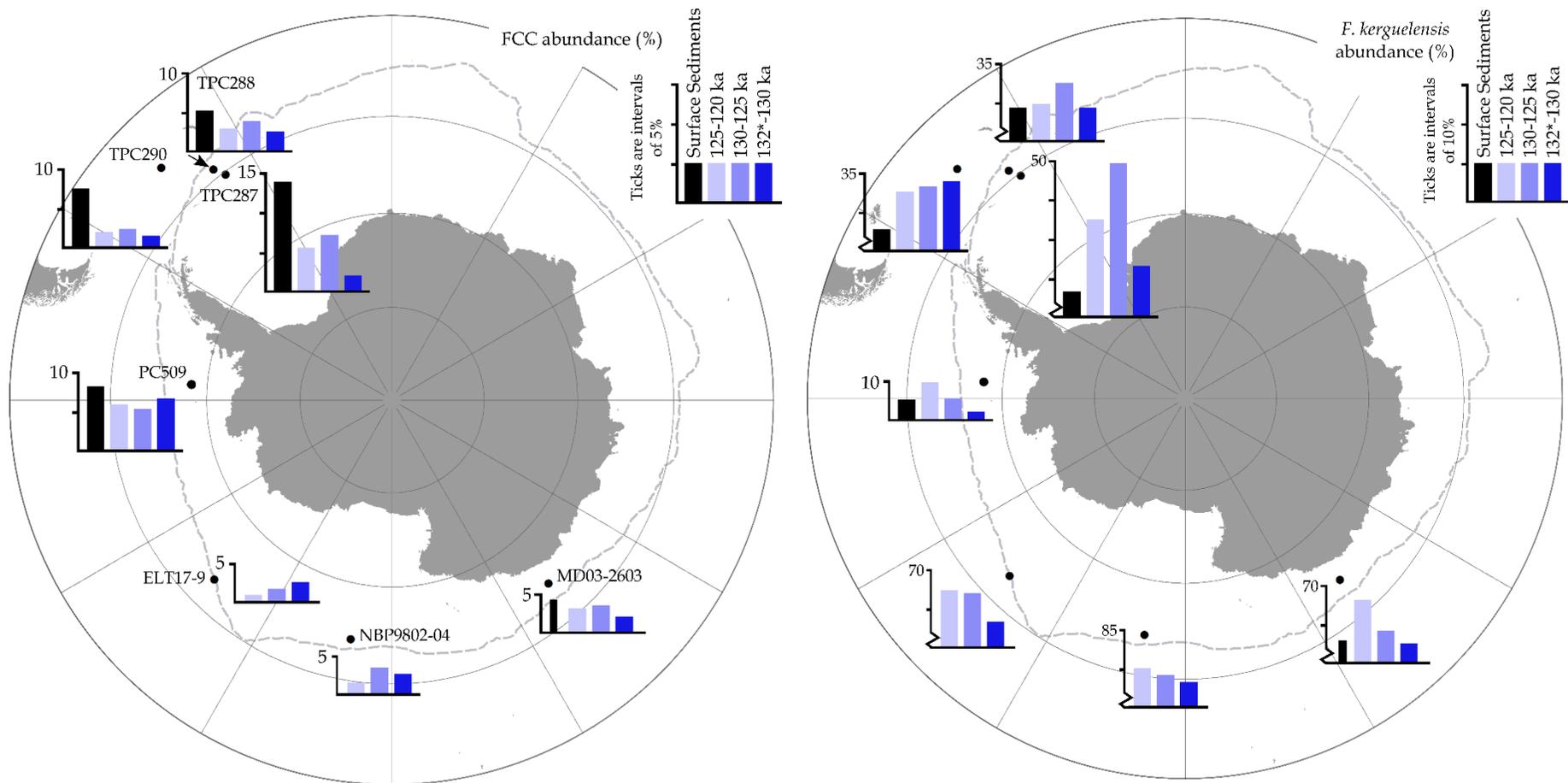


Figure 8: Maps of FCC (LHS) and *F. kerguelensis* (RHS) abundances in seven marine sediment cores. Modern diatom abundances are marked by black bars and mean abundances during three Termination II-MIS 5e time slices (early = 132*-130 ka; mid = 130-125 ka; late = 125-120 ka) are indicated by bars with different shades of blue. Thin black bars on the MD03-2603 graphs indicate diatom abundances in the surface sediment sample from the nearby IODP Site U1361. The black dots mark the core locations and the grey dashed line is the median modern (1981-2010) September sea-ice extent from Fetterer et al. (2017). *: the 132-130 ka time slice includes samples older than 132 ka to ensure that at least three samples are included in the average.

379 **6. Conclusions and wider implications**

380 The early (~130 ka) reduction in WSIE and increase in SSTs at the start of MIS 5e for the two most
381 southerly Atlantic Sector cores (TPC288 and TPC287, Figure 4), coupled with the indication of a
382 poleward contraction of the Weddell Gyre, have large implications for the region and further afield.
383 The downwelling of dense water masses in the Weddell Sea is a key component in the formation of
384 Antarctic Bottom Water (Orsi et al. 2002), which helps drive global ocean overturning circulation (Brix
385 & Gerdes 2003). Reduced WSIE, with a winter sea-ice edge located further south than its modern
386 position, resulted from less sea ice formation in coastal polynyas and from less sea ice advection to
387 the north by winds and subsequent Ekman transport. This, in turn, suggests less brine rejection which
388 may lead to a subsequent decrease in the rates of deep and bottom water mass production along the
389 Antarctic coast as well as a warming of the abyssal waters (Bouttes et al. 2010, Ferrari et al. 2014,
390 Marzocchi & Jansen 2019). Warmer surface and abyssal waters in the Weddell Sea would imply
391 accelerated basal melting of ice shelves and increased grounding line retreat of marine terminating
392 ice streams in the region, which in turn would induce substantial mass loss from the Antarctic ice
393 sheets (Pollard & DeConto 2009, DeConto & Pollard 2016). This increased melting and ice sheet loss
394 may account for the sea ice resurgence (Merino et al. 2018) and increased iceberg discharge (Liu et al.
395 2015) inferred from cores TPC288 and TPC287 after 127 ka (Figure 4). Surface water freshening from
396 glacial meltwater input has also been linked to a reduction in the formation rates of Antarctic Bottom
397 Water, causing further warming of the abyssal ocean (Fogwill et al. 2015, Lago & England 2019).

398 The possible southerly shifts in the Polar Front near sites NBP9802-04 and MD03-2603 and the
399 poleward contraction of the Weddell Gyre south of site TPC287 suggest a poleward migration of the
400 ACC during MIS 5e. A more southerly ACC causes increased advection of relatively warm ACC water
401 masses, such as Circumpolar Deep Water, onto the Antarctic continental shelf (Fogwill et al. 2014,
402 Spence et al. 2017). These warm upwelling ACC water masses contribute to the melting of glacial ice
403 (Hellmer et al. 2012), similar to what is observed today in the Amundsen-Bellingshausen Sea sectors
404 of West Antarctica (Rignot et al. 2019). A southern shift of the ACC would also have caused a poleward
405 movement of the precipitation field and storm tracks (Liu & Curry 2010). A poleward migration of the
406 precipitation field near site MD03-2603 would result in drier conditions across Southern Australia
407 (Saunders et al. 2012), as seen today (CSIRO 2018).

408 Warming and reduced WSIE in the Weddell Sea could also have a substantial impact on the SO
409 biosphere. At present, the Weddell Sea has the highest area-normalised primary productivity rates in
410 the SO (Vernet et al. 2019). Reduced sea-ice extent and increased glacial meltwater supply during MIS
411 5e likely promoted greater primary productivity, as observed today (de Jong et al. 2012, Kahru et al.

412 2016). In contrast, Antarctic krill (*Euphausia superba*), a key trophic intermediary in the modern SO
413 (Knox 2006), prefers lower water temperatures (Siegel & Watkins 2016, Atkinson et al. 2017). A repeat
414 of warmer MIS 5e-like conditions in the future will therefore likely cause a substantial reduction in the
415 habitat and abundances of Antarctic krill in the SO and impact the populations of megafauna that rely
416 on them (Hill et al. 2013), as can be seen today in the rapidly warming northern region of the West
417 Antarctic Peninsula (Montes-Hugo et al. 2009). A WSIE reduction in the Weddell Sea like during MIS
418 5e would also have damaging impacts for modern day sea-ice obligate species, e.g. Emperor Penguins
419 (Jenouvrier et al. 2005).

420 The largely stable environmental conditions at the Pacific Sector core sites, especially ELT17-9, during
421 Termination II-MIS 5e (Figure 6) suggest that this region may be more resilient to future changes. This
422 is consistent with the observations that the modern WSIE in the central and western Pacific Sector is
423 strongly influenced by the topography and bathymetry, through the pinning of fronts and currents
424 (Nghiem et al. 2016). Greater stability of the WSIE in the Pacific Sector would have resulted in
425 protection of ice shelves, such as the Ross Ice Shelf, and maintained their buttressing effect for
426 grounded ice, similar to what is seen in present day Greenland (Walter et al. 2012). This could also
427 have substantial implications for the stability of the West Antarctic Ice Sheet during MIS 5e, with no
428 evidence in the PC509 record (Figure 6) of high iceberg flux originating from glaciers draining the
429 Bellingshausen Sea Sector of the West Antarctic Ice Sheet (Gardner et al. 2018).

430 It is clear that, similar to today (Hobbs et al. 2016, Parkinson 2019), changes to the SO during MIS 5e
431 were not spatially and temporally homogeneous. Some of the climatic variability during Termination
432 II and MIS 5e is due to the difference in the topographic characteristics and oceanographic conditions,
433 with the stability in the Pacific and East Indian Sectors likely due to bathymetry pinning and the
434 stability of the Australian-Antarctic Basin gyre, respectively. This variation in the controls on and
435 magnitude of changes in the Antarctic and SO climate system are important factors to be included
436 into model simulations of future warming. MIS 5e is a valuable 'laboratory' for understanding how the
437 Antarctic and SO region responds to a warmer climate, especially for regions like the Weddell Sea,
438 where climatic trends during MIS 5e diverge from what is observed today (Purich et al. 2016, Parkinson
439 2019).

440 **Data availability**

441 Termination II and MIS 5e diatom assemblage data for all samples are available from the Polar Data
442 Centre (Chadwick & Allen 2021a, b, c, d, e, f, g) and surface sediment diatom assemblages can be
443 found at <http://dx.doi.org/10.17632/2tnxcww6c8.1> an open-source online data repository hosted at
444 Mendeley Data (Chadwick 2020).

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