

1 **The rate of sedimentation from turbulent suspension: an experimental model with application**
2 **to pyroclastic density currents and discussion on the grain-size dependence of flow runout**

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8
9 **Abstract**

10 Large-scale experiments generating ground-hugging multiphase flows were carried out with the aim
11 of modelling the rate of sedimentation S_r of pyroclastic density currents. The current was initiated by
12 the impact on the ground of a dense gas-particle fountain issuing from a vertical conduit. On impact,
13 a thick massive deposit was formed. The grain size of the massive deposit is almost identical to that
14 of the mixture feeding the fountain, suggesting that similar layers formed at the impact of a natural
15 volcanic fountain should be representative of the parent grain-size distribution of the eruption. The
16 flow evolved laterally into a turbulent suspension current that sedimented a thin, tractive layer. A
17 good correlation was found between the ratio transported/sedimented load and the normalized Rouse
18 number P_n^* of the turbulent current. A model of the sedimentation rate was developed, which shows
19 a relationship between grain size and flow runout. A current fed with coarser particles have a higher
20 sedimentation rate, a larger grain-size selectivity and runs shorter than a current fed with finer
21 particles. Application of the model to pyroclastic deposits of Vesuvius and Campi Flegrei of Southern
22 Italy resulted in sedimentation rates falling inside the range of experiments and allowed defining the
23 duration of pyroclastic density currents τ_{dep} , which add important information on the hazard of such
24 dangerous flows. The model could be possibly extended, in the future, to other Geological density
25 currents as, for example, turbidity currents.

26
27 **Keywords:** Pyroclastic density currents, sedimentation rate, turbulent suspension, experiments

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30 **1. Introduction**

31 The formation of sedimentary deposits is in many cases regulated by the passage of turbulent currents
32 carrying a substantial particle load. Common examples are rivers, turbidity currents and pyroclastic
33 density currents (Gladstone et al., 1998; Kneller and Buckee, 2000). The flow carries solid particles
34 by three main modes (Rouse, 1939; Middleton and Southard, 1984): bedload, suspension and wash
35 load. Bedload concerns the material moved by traction on the ground; wash load concerns particles
36 so fine that are intimately coupled to fluid turbulence and are carried away by the current. Suspension
37 refers to particles that are sustained by fluid turbulence and settle when their terminal velocity is
38 lower than the current shear velocity. The suspension load is thought to represent about 90-95% of
39 the total particulate material in the current. It is the rate of sedimentation from turbulent suspension
40 that feeds the sediment layer, allows aggradation of deposit thickness and regulates the current runout.
41 In the time-space evolution of a flow, particles, after transportation in suspension, eventually settle
42 to the ground and form the bedload that, upon some tractional movement, comes to rest and forms
43 the final deposit (Branney and Kokelaar, 2002; Sulpizio and Dellino, 2008; Sulpizio et al., 2014;
44 Dufek, 2016). The structural configuration of the layer is acquired by the last movement of the
45 sediment in the bedload, which leads to the formation of asymmetrical structures such as ripples, sand
46 waves and cross lamination.

47 The flow of pyroclastic density currents moves in contact with the ground due to its higher density
48 with respect to the surrounding atmosphere. The loss of particles from sedimentation, combined with
49 fluid entrainment from the atmosphere, results in a reduction of concentration, with a consequent
50 lowering of fluid density. It is related to the fact that the flow is composed of gas and particles and
51 behaves as a “pseudofluid” which density is given by $\rho_f = \rho_s C + \rho_g (C - 1)$, where ρ_s is particle
52 density, ρ_f is fluid density, ρ_g is gas density and C is particle volumetric concentration (all symbols
53 are defined in Table 1). It is known that pyroclastic density currents, as results from the density
54 reduction due to sedimentation, at some point stop moving laterally and start lofting from the ground
55 in the form of a buoyant phoenix cloud (Neri and Macedonio, 1996; Sparks et al., 1997; Branney and

56 Kokelaar, 2002; Andrews and Manga, 2011, 2012), aided by buoyancy resulting from entrainment
57 and thermal expansion of atmosphere.

58 The ability of a turbulent current to transport a particle in suspension is a function of the particle
59 Rouse number $P_n = w_t/ku^*$ (Rouse, 1939), where u^* is the current shear velocity, which is related to
60 the turbulent shear stress (Pope, 2000; Schlichting and Gersten, 2000), k is Von Karman constant =
61 0.4 and w_t is particle terminal velocity

$$62 \quad w_t = \sqrt{\frac{4Dg(\rho_s - \rho_f)}{3C_d\rho_f}} \quad (1)$$

63 where D is particle size, g is gravity acceleration, C_d is particle drag coefficient. Particles with P_n
64 lower than 2.5 are carried in suspension by turbulence, meaning that they are suspended until u^*
65 doesn't drop to values lower than w_t (Middleton and Southard, 1984; Valentine, 1987; Branney and
66 Kokelaar, 2002; Dellino et al., 2008).

67 Equation (1) gives a good estimation of particle settling velocity if particle volumetric concentration
68 does not exceed a few percent, which is the case of the suspension currents considered in this paper.
69 The particle volumetric concentration (hence density) is stratified within the current. The maximum
70 volumetric concentration of particles that can be transported in suspension, i.e. the maximum current
71 capacity, is a function of the Rouse number of the particulate mixture taken in suspension P_{nsusp} . It
72 is regulated by the Rouse concentration profile over current height y

$$73 \quad C_{tot} = C_0 \frac{1}{H_T - y_0} \int_{y_0}^{H_T} \left(\frac{H_T - y_0}{y_0} \frac{y}{H_T - y} \right)^{P_{nsusp}} dy \quad (2)$$

74 where C_{tot} is the total concentration of the current, H_T is current total thickness and C_0 is a value of
75 known concentration at a specific height y_0 , which is assumed to be the value of concentration at
76 maximum packing in contact with the ground (0.75 in this paper). From (2) it can be inferred that a
77 current carrying a finer mixture (lower P_{nsusp}) can transport, at maximum capacity, a higher
78 concentration than one having the same shear velocity, but carrying coarser particles (higher P_{nsusp}).
79 The solid load constituting a suspension current, especially in the case of pyroclastic density currents,
80 is made up of a mixture of different components (lithics, glassy fragments and crystals) with different

81 size, density and shape, thus different terminal velocity. The Rouse number of the solid material in
 82 the current must be expressed as the average of the particulate mixture,

$$83 \quad P_{n_{avg}} = \sum_{i=1}^n P_{ni} C_i / C_{tot} \quad (3)$$

84 with the subscript i referring to the i th particle-size class and n being the number of size classes.

85 The ratio between $P_{n_{avg}}$ of the material in the current and P_{nsusp} is here defined as the normalized
 86 Rouse number P_n^* of the current. When it is higher than 1, a current has a particle volumetric
 87 concentration in excess of its maximum capacity, e.g. it is over-saturated of particles, which favours
 88 sedimentation. When it is lower than 1, a current has a particle volumetric concentration lower than
 89 its maximum capacity, e.g. it is under-saturated, and could potentially include additional sediment
 90 that is being eroded from the substrate. Very coarse particles, namely those with P_n higher than 5,
 91 settle from suspension without being much influenced in their trajectory by turbulence.

92 Particles in a pyroclastic density current often come from the fountaining of an eruption column and
 93 generally are over-saturated with particles. In fact, pyroclastic density currents leave continuous
 94 deposits on the ground, meaning that during most of the runout they are in sedimentation mode. The
 95 sedimentation rate $S_r = w_t \rho_s \gamma$ is a measure of the mass of particles sedimenting with time per unit
 96 area, where γ is the proportion of particles settling from suspension. It is convenient to express the
 97 sedimentation rate by means of the sum of the contribution of each size class in the mixture $S_r =$
 98 $\sum_{i=1}^n S_{ri}$ where $S_{ri} = w_{ti} \rho_{si} \gamma_i$, where S_{ri} is the sedimentation rate of the i th size class, w_{ti} is the
 99 terminal velocity of the i th size class, ρ_{si} is the density of the i th size class and γ_i is the proportion of
 100 particles of the i th size class settling from suspension.

101 The grain-size distribution of a deposit is generally represented by a histogram expressed in ϕ units
 102 $\phi = -\log_2 d$, with d particle diameter in millimetres. It represents the distribution of the weight
 103 fraction ϕ_i of each size class in the deposit, with $\sum_{i=1}^n \phi_i$ summing to 1 (or 100%). In the case of
 104 deposits formed by sedimentation from turbulent suspension, it is here assumed that the grain-size
 105 distribution represents also the proportion of the sedimentation rate of each size class. Thanks to this

106 assumption, the values of γ_i can be easily calculated once the total sedimentation rate S_r , the terminal
107 velocity w_{ti} and density of each size class ρ_{si} are known.

108 The growth of deposit thickness with time at a location, i.e. the layer aggradation rate, is given by
109 $A_r = S_r / \rho_{dep}$ where A_r is the aggradation rate and ρ_{dep} is deposit density, measured in the field as
110 $0.6\rho_s$ in this study. Depositional time τ_{dep} is given by $\tau_{dep} = \frac{H_{dep}}{A_r}$ where H_{dep} is deposit thickness.

111 If deposit density and thickness are measured in the field, and the rate of sedimentation can be
112 modelled, it is possible to reconstruct the depositional time, which to a good approximation represents
113 the time it took for the current to pass that particular location. The depositional time is an important
114 indicator of the potential impact that a pyroclastic density current can have on human health, since it
115 quantifies the residence time of hot volcanic ash that can be inhaled by people potentially exposed to
116 these dangerous flows (Horwell and Baxter, 2006). Even a very low volumetric concentration of ash
117 in suspension is unbreathable, and is one of the main causes of mortality of pyroclastic density
118 currents. A model of the sedimentation rate from suspension could greatly help assessing the hazard
119 of pyroclastic density currents. Unfortunately, up to now, no such model exists. In fact, in the
120 computational fluid dynamic simulations of pyroclastic density currents, the effect of sedimentation
121 is generally not included.

122 This paper describes experiments carried out for developing a model of the sedimentation rate based
123 on data derived from deposits. The model highlights the grain-size dependence of flow runout.
124 Application to natural deposits leads to calculation of the depositional time of natural currents, which
125 helps assessing the hazard to human health.

126

127 **2. Experiments and laboratory investigation**

128 The experiments were carried out with the apparatus described in detail in Dellino et al., 2007; 2010a;
129 and 2010b, which allowed the reproduction of various regimes of explosive eruptions (Dellino et al.,
130 2014). In this paper, only the results of experiments generating substantial density currents are

131 considered (Dellino et al., 2010b). The particulate material used in the experiments comes from
132 deposits of Vesuvius and Campi Flegrei volcanoes in Southern Italy, and covers an ample range of
133 size, density and particle shape. For each run, up to 350 kg of particles were used. The grain size of
134 two compositions, representing the coarse (from Vesuvius) and fine (from Campi Flegrei) end
135 members, are shown on fig.1. The coarser composition, ranging from lapilli to fine ash (fig. 1a), is
136 made of dense lithic, vesicular glass and crystal components, while the finer one, mostly fine ash (fig.
137 1b), is made almost exclusively of glass fragments.

138 Experiments were carried out at various temperatures, from ambient up to 300 °C. The effect of high
139 temperature was that of reducing the density of the carrier fluid and forming a buoyant phoenix cloud
140 at the end of runout (Dellino et al., 2010b). Additional details that emerged, by the experiments, on
141 the effect of temperature on various regimes are deferred to Dellino et al. (2014). The experimental
142 design (fig. 2) consists of 2 interconnected packs of 16 pressurized-gas bottles (the gas storage
143 compartment); a high-pressure section consisting of 18 steel-reinforced rubber hoses each 30 m long;
144 a rapid-compression section consisting of 18 steel-reinforced rubber hoses each 1.5 m long; and a
145 low-pressure section consisting of a 3.2-m-long stainless-steel conduit with a 0.6-m internal diameter,
146 mounted on a massive base plate. The gas bottles are coupled to the high-pressure section via two
147 valves and a hub, in line with manometers that control the reservoir pressure and the pressure in the
148 high-pressure section. High-speed solenoid valves connect the high-pressure section via a second hub
149 to the rapid compression section. The short hoses are connected to eighteen blow nozzles in the base
150 plate of the low-pressure section. The pyroclastic material is placed into the conduit and rests directly
151 on the base plate. The experiment starts by opening the valves that connect the gas-storage
152 compartment to the high-pressure section until the desired pressure is reached in the 30 m long hoses.
153 The computer controlled opening of the solenoid valves connects the high-pressure section to the
154 low-pressure section (via the rapid compression section) and allows a fast coupling of the pressurized
155 gas with the pyroclastic material filling the conduit, which while mixing with the expanding gas,
156 accelerates along the conduit. The two-phase mixture is finally expelled from the conduit in the form

157 of a dense gas-particle fountain, reaching a maximum height over 10 m (fig. 3a). On hitting the
158 ground, the fountain resembled the collapse of an eruptive column similar to that generating a natural
159 pyroclastic density current.

160 Upon the impact of the fountain on the ground (fig. 3b), the normal stress of the fluid was transformed
161 into tangential stress, which led to a flow that evolved laterally into a fully turbulent, gas-particle
162 shear current, a few meters thick, moving at several m/s (fig. 3c). Deposits of measurable thickness
163 (at least a few millimetres) formed on the ground upon the passage of the current. At the impact zone,
164 where the lateral flow was not fully developed yet, a tens of centimetres thick, massive deposit, was
165 formed (fig. 3d). It represents the excess of solid material that could not be transported into the lateral
166 flow, and was emplaced “en masse” (Sulpizio et al., 2014; Roche, 2015). It is important to note that
167 the grain-size distribution of the massive layer is very similar to that of the original particle load in
168 the conduit. This happens both with the coarse and fine particle mixtures used in the experiments (fig.
169 4a and b). It reveals that there is not an effective grain-size selection between the particulate mixture
170 issuing from the conduit and the material emplaced en masse at the impact on the ground. It means
171 that the grain size of massive layers formed by similar collapses of dense volcanic fountains should
172 be considered as representative of the parent particle population of natural eruptions. This deposit
173 facies, which resembles a massive pyroclastic flow (Branney and Kokelaar, 2002), makes transition,
174 laterally, into a thin structured layer, similarly to what is observed and documented in certain
175 ignimbrites (Brown and Branney, 2013). The thin layer shows sedimentary structures such as ripples
176 (fig. 3e), which are characterized by an asymmetrical distribution of particles. The finer load occurs
177 at the foreset and the coarser load at the backset, suggesting a selective transportation of the bedload,
178 which is typical of tractive processes occurring at the base of natural currents. These features are
179 common among deposits formed by pyroclastic density currents, with the difference that natural
180 layers have a much higher thickness and represent the aggradation of multiple tractional structures
181 formed during the time integrated passage of the flow, which is much longer than that of experiments.
182 The deposit thins out with increasing distance from the impact zone and has a fan shaped distribution

183 covering, with a thickness ranging from a few millimetres to a few centimeters, an area of up to about
 184 2000 m² (fig. 3f). The shear current was continuously fed from the fountain for several seconds. In
 185 that time period the deposit was formed by steady sedimentation of particles from suspension, and
 186 final bedload traction. When the fountain stopped feeding the current, the flow rapidly decelerated
 187 and only the finest particulate material of the upper part of the current continued moving as a wash
 188 load, for a long time (Supporting video). The wash load was spread well over the deposit fan-shaped
 189 area and formed a very thin, submillimetric, veil of ash.

190 The current runout was recorded by a network of pressure sensors and multiple high-definition digital
 191 video cameras (Dellino et al., 2007, 2010a, 2010b, 2014). For each experiment, thickness and speed
 192 of the current were recorded at multiple stations along runout, starting from the impact point and up
 193 to about 20 m of distance. The distance between successive stations was set at 1 m for runs spreading
 194 on a smaller area and at 2 meters for larger ones. Sediment samples were collected from each station.
 195 Sampling was done by collecting the sediment from a rectangular area on the ground, about 1 m²,
 196 which allowed calculation of the mass per unit area of sediment deposited at each station. A total of
 197 18 samples representing the total number of locations out of 6 experimental runs is considered in this
 198 paper, on which grain-size, density and particle shape analyses were carried on.

199 For each of the 18 locations, by combining the processing of sensors and video camera recordings
 200 and laboratory analyses of the sediment samples, the following parameters were obtained:

- 201 • the shear velocity of the current u^* ;
- 202 • the particle volumetric concentration of each size class C_i and the total particle volumetric
 203 concentration C_{tot} ;
- 204 • the Rouse number of each size class P_{ni} and the average Rouse number $P_{n_{avg}}$;
- 205 • the Rouse number at maximum suspension capacity P_{nsusp} and the normalized Rouse number
 206 P_n^* ;
- 207 • the proportion of particles of each size settling from suspension γ_i ;

- 208 • the rate of sedimentation of each size class S_{ri} and the total sedimentation rate S_r ;
- 209 • the settling velocity of each size class w_{ii} ;
- 210 • the density of each size class ρ_{si} ;
- 211 • the flow density ρ_f ;
- 212 • the multicomponent grain-size distribution including the shape of particles;
- 213 • the particle mass flow rate PMFR.

214 In particular, the particulate mass flow rate at the impact zone was measured as the product of the
 215 area of impact, velocity of impact and density of the flow. The PMFR of each size class in the current
 216 was calculated, at each station, by subtracting the total mass of sediment deposited at all previous
 217 stations from the particulate mass flow rate at the impact, and using as the area crossed by the flow,
 218 the value measured by image analysis upon flow front passing from each station. The current density
 219 due to the particle load of each size class was calculated by dividing the mass flow rate of each size
 220 class by the average velocity of the current. The total mass and grain-size distribution of each
 221 component in the particulate mixture was measured in the laboratory before each run, as it was done
 222 also for the samples taken at each station. The mass of material transported in the current at successive
 223 stations was calculated by subtracting the mass of sediment deposited at previous stations from the
 224 total mass. The particle volumetric concentration of each size class of the current C_i was calculated
 225 by dividing the bulk density of the current by the particle density of each size class. The total
 226 sedimentation rate was calculated at each station by the mass of sediment per unit area divided by the
 227 time of sedimentation. The time of sedimentation was measured at each station starting from the
 228 arrival of the flow front and ending by the passage of the wash load. The total sedimentation rate was
 229 partitioned among size fractions according to the partitioning of the grain-size distribution. The
 230 proportion of particles settling from suspension γ_i of each size class was calculated by dividing the
 231 sedimentation rate of each size class by the settling velocity and particle density. For more details on
 232 the experimental methods, techniques and uncertainties, see the Supporting file A. The experiments
 233 covered an ample range of flow parameters. In particular, the sedimentation rate S_r was between 0.009

234 / 1.17 kgm⁻²s. It is in the same order of magnitude of the sedimentation rate obtained by means of
 235 lagrangian multiphase numerical simulations (Valentine et al., 2011; Doronzo et al., 2017). For the
 236 range of other experimental parameters see the summary Table 2.

237

238 **3. The experimental model**

239 The ratio between the particle volumetric concentration of each size class C_i and the proportion of
 240 particles of each size class settling from suspension γ_i of each experiment is well correlated with the
 241 Rouse number of each size class P_{ni} , as it should be expected in a turbulent suspension current where
 242 the attitude of particles to be transported (or sedimented) is a function of the balance between terminal
 243 velocity and shear velocity. On fig. 5, the different slopes in the regression equation of a current
 244 carrying coarser particles compared with one carrying finer ones demonstrates that flows having a
 245 different normalized Rouse number P_n^* have also a different attitude toward sedimentation (or
 246 transportation), which depends on the excess of particle load (oversaturation) with respect to
 247 maximum current capacity. To take into account this factor, the ratio C_i/γ_i of all particle sizes and
 248 components of all samples was plotted against P_{ni}/P_n^* . The equation of the regression line:

$$249 \quad \frac{C_i}{\gamma_i} = \frac{P_{ni}}{P_n^*} 10.065 + 0.1579 \quad (4)$$

250 well approximates data of all experiments (fig. 6).

251 The regression line of equation (4) can be used either to predict the proportion of each size class of
 252 particles settling from suspension γ_i if the Rouse number of each size class P_{ni} , the normalized Rouse
 253 number P_n^* and the particle volumetric concentration of each size class C_i are known or to obtain C_i
 254 if P_{ni} , P_n^* and γ_i are known. Unfortunately, it is difficult to estimate the values of C_i or γ_i of natural
 255 pyroclastic density currents. The particle parent population that issues from the volcanic conduit and
 256 feeds pyroclastic density currents is generally unknown. In fact, there is a strong geological evidence
 257 that it changes from volcano to volcano and from eruption to eruption, depending mostly on magma
 258 fragmentation processes. It is to expect that the relative proportions of the size fractions in the

259 transported material should be different from the proportions of the material settling on the ground.
 260 In fact, the grain-size distribution will evolve as particles selectively sediment as a function of grain
 261 size and density. While it is reasonable to hypothesize a substantial grain-size difference, along
 262 runout, between the material transported at a certain location and the material deposited far away, the
 263 difference between the grain size transported at some point and that deposited at the same point should
 264 be smaller. Following this line, we tested whether the difference in grain size between the sediment
 265 sampled at a station and that of the particulate mixture transported in the current at the same station
 266 was small enough as to permit the use of the sediment size as a “first guess” of the transported material
 267 in equation (4) for reconstructing the sedimentation rate of the experiments. The relative proportions
 268 of the size fractions in the transported material at a station were obtained, as described in the previous
 269 paragraph, by subtracting the total mass of sediment deposited at all previous stations from the total
 270 mass. We made the test by using a goodness-of-fit formula based on the chi-square statistics: $100\% -$
 271 $\left(\sum_{i=1}^n \left(\frac{O_i - E_i}{E_i} \right)^2 \right)$. O_i is the weight % of the transported material of the i_{th} size class (Observed value
 272 in Statistics), E_i is the weight % of the deposited material of the i_{th} size class (Expected value in
 273 Statistics). The components were summed together as to obtain, for each class, a weight not smaller
 274 than 5%, as it is suggested in Statistics when using percent data in the chi-square test (Davis, 2002).
 275 We obtained a fitting typically better than 90% (see Supporting file C and D for examples), which
 276 ensures that the grain-size distribution of the deposit can be used as a “first guess” of the grain-size
 277 distribution of the transported material, without too much error. The concentration of each component
 278 of each size class of the transported material was, then, reconstructed by means of the grain-size
 279 distribution of the deposit at each station by $c_{trans_i} = \frac{\phi_i / \rho_{s_i}}{\sum_{i=1}^n \phi_i / \rho_{s_i}} * C_{tot}$, where c_{trans_i} is the particle
 280 volumetric concentration of the i_{th} size class, $\frac{\phi_i / \rho_{s_i}}{\sum_{i=1}^n \phi_i / \rho_{s_i}}$ is the volume fraction occupied by the i_{th} size
 281 class, and C_{tot} is the total particle volumetric concentration of the current.

282 By means of the values of c_{trans_i} , and rearranging equation (4), the contribution of each particle
 283 size of each component in the sedimentation rate $\gamma_{proxy_i} = \frac{c_{trans_i}}{\left(\left(10.065 * P_{n_i}^* / P_n^*\right) + 0.1579\right)}$ was obtained.

284 By means of γ_{proxy_i} , and using the values of settling velocity and particle density of each size
 285 class, the sedimentation rates were calculated and compared with the experimental values. The plot
 286 on fig. 7 shows the regression line approximating data points. Judging from the correlation
 287 coefficient, while some scatter is visible, the fitting is good. The slope of the regression line is,
 288 however, a little smaller than 1, suggesting that the calculated values are a little underestimated with
 289 respect to the experimental ones, which can be attributed to the approximation that was made by
 290 using the grain-size distribution of the deposit as a "first guess" of the grain-size distribution of the
 291 transported material. The underestimation suggests that the grain size of the sediment must be a
 292 little coarser than that of the transported material, as it is expected from a current that settles,
 293 selectively, more of the coarser than of the finer particle load. We looked for correcting the
 294 underestimation and found the grain-size shift necessary to adjust the γ_{proxy_i} values. Details of the
 295 method are shown in the Supporting file B. By means of the application of the grain-size shift, the
 296 corrected proportions of the sedimentation rate of each size class are recalculated as: $\gamma_{true_i} =$
 297 $\gamma_{proxy_i} * 0.7 + \gamma_{proxy_{i+1}} * 0.3$, where γ_{true_i} is the correct value.

298 By means of the values of γ_{true_i} the sedimentation rates were recalculated and compared with the
 299 experimental ones, resulting in the regression of fig. 8, by which the final model equation of the
 300 sedimentation rate is obtained

$$301 \quad S_r = \left(\sum_i^n \rho_{s_i} w_{t_i} \left(\frac{c_{trans_i}}{\left(\left(10.065 * P_{n_i}^* / P_n^*\right) + 0.1579\right)} * 0,7 + \gamma_{proxy_{i+1}} * 0.3 \right) \right) - 0,01 \quad (5)$$

302 By means of equation (5), the final fitting of fig. 9 is obtained, which shows, to a good
 303 approximation, a 1 to 1 ratio between measured and calculated sedimentation rates.

304 By rearranging terms, it is also possible to reconstruct the particle volumetric concentration of each
 305 size and component transported in the current, starting from the proportion in the deposit, by

$$306 \quad c_{transtrue_i} = \left(\gamma_{proxy_i} * 0.7 \right) + \left(\gamma_{proxy_{i+1}} * 0.3 \right) * \left(\left(10.065 P_{n_i}^* \right) + 0.1579 \right) \quad (6)$$

307 where $c_{transtrue_i}$ is the corrected concentration of the i_{th} size class in the current. By normalizing to
 308 1 the sum of the values of $c_{transtrue_i}$ of all the size classes, the grain-size distribution of the particle
 309 mixture in the current is calculated.

310 On fig. 10, examples of the comparison between the grain-size distribution of the transported material
 311 and that of the sediment material are shown. The sediment particulate mixture is a little bit coarser
 312 than the particle load transported in the current, as it is expected in a current that settles selectively
 313 more of the coarser than of the finer particle load. This happens when the coarse composition of
 314 Vesuvius is used in experiments (fig. 10a), where the weight % of the coarser size classes is constantly
 315 higher in the sedimented than in the transported material down to a size of 3ϕ , then the behaviour is
 316 inverted for the finest class sizes. The difference is less obvious when the particulate mixture is
 317 composed of fine material (fig. 10b), as observed in the experiment with the composition of Campi
 318 Flegrei. In this case, in fact, the weight % of the coarser size classes is higher in the transported
 319 material down to 2.5ϕ , then it is higher in the sedimented material from 2.5 and 5ϕ , and finally it is
 320 again higher in the transported material, indicating a lack of a particular selectivity of grain size. An
 321 additional proof of the difference in selectivity between “coarse” and “fine” currents is shown on
 322 Fig.11 where a comparison between the grain-size evolution of deposits as a function of distance
 323 from the impact zone is shown for two experiments. In the “coarse” run a small but significant
 324 variation of grain size as a function of increasing distance is noticeable, while for the “fine” run the
 325 grain size is more or less the same at various distance. While the experiments do not represent the
 326 real scale of distance travelled by natural currents, the clear difference between currents carrying
 327 coarser vs fine pyroclasts suggests that pyroclastic density currents transporting mostly fine ash
 328 should show less grain-size variations along runout. The comparison of the sedimentation rate at two
 329 locations, which represent experiments fed with the coarse and fine end members, reveals some
 330 additional aspects of the grain-size dependence of runout of pyroclastic density currents. The relative

331 data sets are included in the Excel worksheets of the Supporting file C and D for a detail analysis.
332 Here we just discuss the fundamental results. While the flow characteristics, i.e. current velocity, are
333 not much different, a big difference, between the coarse and fine cases, is in the ratio between the
334 particle mass flow rate and the sedimentation rate. It is much smaller in the case of the experiment
335 with the coarser material compared to the finer one (four times smaller, see Supporting file C and D).
336 The coarser current has a much higher sedimentation rate than the finer one (tens of times, see
337 Supporting file C and D). Summing up, finer currents can suspend a higher amount of particles
338 (because of the lower Pn), emplace less material along runout, maintain a significant density
339 difference with respect to the atmosphere, hence a higher mobility and a longer runout compared to
340 coarser currents. As a consequence, fine pyroclastic density currents can run faster, longer and leave
341 thin, widely spread deposits. This conclusion seems to be corroborated by the fact that some of the
342 most widespread historical pyroclastic density currents known up to date, for example the 1.8 ka
343 Taupo ignimbrite of New Zealand (Wilson et al., 1995; Dade and Huppert, 1996; Manville et al.,
344 2009), show thin, fine-grained deposits, which grain size doesn't change much with travel distance.

345

346 **4. Model application to natural pyroclastic deposits and scaling to experiments**

347 The experimental model developed in this paper was applied to the deposits of pyroclastic density
348 currents of the Mercato Plinian eruption at Vesuvius and of the Agnano Monte Spina Plinian eruption
349 of Campi Flegrei in Southern Italy. Details on the eruptions and stratigraphy of deposits can be found
350 in de Vita et al., 1999; Dellino et al., 2004; Mele et al., 2011, 2015. The layers considered in the
351 present study were formed by the passage of density currents fed by the collapse of an eruption
352 column and show, at the localities sampled in this study, a facies characterized by tractional structures
353 (fig. 12a) and inclined lamination (fig. 12b), suggesting that transportation and sedimentation were
354 from flows carrying a particulate load by turbulent suspension, and final tractional movement at the
355 bedload.

356 The layers are 0.5 and 0.2 m thick for Mercato and Agnano Monte Spina, respectively. Deposit
357 density is 1476 and 1295 kg/m³ for Mercato and Agnano Monte Spina, respectively. They are
358 composed of vesicular glass, dense lithics and crystals, which multicomponent grain size is shown
359 on fig. 13 a and b, respectively. The density, shape and settling velocity of each size class of each
360 component of the deposits were calculated using the same techniques of the experimental samples.
361 The flow parameters needed for the application of our sedimentation rate model were calculated by
362 means of the software Pyflow (Dioguardi and Dellino, 2014), which is based on the models of Dellino
363 et al. (2008) and Dioguardi and Mele (2015). The calculation used in the present paper utilizes the
364 concept of hydraulic equivalence. If two components with different median size, density and shape,
365 settle together, they are hydraulically equivalent and have the same settling velocity. By this
366 assumption, the software equates the settling velocity of the two components and solves for the
367 current shear velocity u_* , total concentration over flow height C_{tot} and Rouse number at maximum
368 suspension capacity Pn_{susp} . The software finds a range of solutions that considers the variation of
369 deposit particle characteristics. For the sake of simplicity, we restrict our analysis to the average
370 solution, and give the uncertainty in terms of \pm one standard deviation around the average.
371 By combining the particles data and flow parameters obtained by the software Pyflow: C_i , P_{ni} , $P_{n_{avg}}$
372 and P_n^* were calculated, and by means of the combined use of eq. (4), (5) and (6) the sedimentation
373 rate was obtained. In the Supporting file E and F, an Excel worksheet contains all the input data and
374 results of the average solution, as to allow following step by step the calculations. The sedimentation
375 rate is about $0.59_{-0.22}^{+0.19}$ and $0.38_{-0.17}^{+0.08}$ kg/m²s for Mercato and Agnano Monte Spina, respectively. It
376 falls inside the range of experiments, suggesting that the application of the model to natural deposits
377 doesn't imply an unwarranted extrapolation of results outside the experimental range. By comparing
378 data of file E and F and table 2 it is possible to judge how other important parameters scale between
379 experiments and natural pyroclastic density currents. The shear velocity of the Mercato and Agnano
380 Monte Spina pyroclastic density currents, while in the same order of magnitude of experiments, is
381 about threefold. The thickness of natural currents is much larger than that of the experiments.

382 Combining shear velocity and flow thickness and recalling that the velocity profile is a function of
 383 the shear velocity (see Supporting material A), it results that natural currents typically reach, with
 384 increasing height, a speed of tens of m/s, while in the experiments the maximum speed was a little
 385 bit lower than 10 m/s. The grain size of natural deposits is in the same range of experiments as it is
 386 also the particle volumetric concentration and the normalized Rouse number, P_n^* . Summing up, while
 387 velocity and thickness of natural currents are larger than experiments, the experiments well preserve
 388 the scale of natural flows in their basal part, where sedimentation occurs, justifying the fact that the
 389 sedimentation rate of natural deposits is well inside the range of experiments. The tractional features
 390 of natural deposits are similar to the experimental ones, whereas the thickness of deposits is much
 391 larger in the natural case. Since the growth of deposit thickness with time is a function of the
 392 aggradation of the material sedimented from turbulent suspension, the larger thickness of natural
 393 deposits means a longer duration of the passage of natural pyroclastic density currents with respect
 394 to the experiments. For approximating the duration of the passage of the natural currents, first the
 395 aggradation rate A_r and finally the deposition time τ_{dep} were calculated, by recalling the definition
 396 given in the introduction section. A_r is $4_{-1.5}^{+1.3} \times 10^{-4}$ and $2.9_{-0.13}^{+0.07} \times 10^{-4}$ m/s for Mercato and Agnano
 397 Monte Spina, respectively, and τ_{dep} is 1240_{-307}^{+765} and 681_{-125}^{+554} s. It means that the passage of the
 398 currents, at the location where the deposits were sampled, lasted around 20 minutes in the case of
 399 Mercato and around 11 minutes in the case of Agnano Monte Spina. This is consistent with the
 400 observation of historical eruptions, where the flow lasted for several minutes to hours (e.g. Lube et
 401 al., 2007). During that time period the territory was engulfed with thick, expanded, fast and hazardous
 402 currents, loaded with unbreathable hot ash (Horwell and Baxter, 2006). It is important to take note of
 403 such information, when projecting for emergency plans and risk-reduction measures.

404

405 **5. Discussion and future perspective**

406 By means of large-scale experiments, a novel model of the sedimentation rate from turbulent
 407 suspension (Equation 5) was obtained. The sedimentation rate strongly influences the runout of

408 pyroclastic density currents, depending on the grain-size of the particulate mixture. If the grain size
409 of the current is coarser (coarse ash to lapilli), the flow sediments selectively the particulate load,
410 making the particle mixture gradually finer along the runout. When, instead, the particulate mixture
411 is finer (fine ash) there is less selective transportation, hence deposition. In this case, particles have a
412 smaller Rouse number, which is the exponent of the concentration profile, resulting in an almost
413 evenly distributed concentration of the sediment along flow height. In principle, fine particles should
414 be transported in continuous suspension, but if the current is oversaturated ($P_n^* > 1$), a sedimentation
415 rate must be anyway allowed, although it is very small (tens of times lower than the case with coarser
416 particles, see Supporting file C and D). The settled fine ash remains attached to the ground and cannot
417 be re-eroded from the substrate (Gladstone et al., 1998). In summary: fine-grained pyroclastic density
418 currents, while leaving thin layers on the ground, travel further and possess a higher capacity of
419 impact over the territory. The impact potential is related to the presence of unbreathable hot ash
420 (Horwell and Baxter, 2006) and to the dynamic pressure of the flow $\frac{1}{2} \rho_f u^2$, which in extreme cases
421 is able to destroy buildings (Valentine, 1998; Baxter et al., 2005; Neri et al., 2015). The distribution
422 of these impact parameters along flow runout is strongly influenced by the sedimentation rate. We
423 believe that the inclusion of the sedimentation rate in the numerical multiphase simulation codes, by
424 means of equations (5) and (6), would improve the ability to predict the hazard of pyroclastic density
425 currents on active volcanoes.

426 In order to effectively use our new model, it is important to have precise data on the physical
427 characteristics of the particles present in a current. Unfortunately, there are no theoretical models
428 giving a priori insights into the grain size, density and shape of the particulate mixture. Pyroclastic
429 deposits are the only record of the passage of pyroclastic density currents, and a prerequisite work in
430 the field is needed for getting information of the real particle population that feeds the flow. After a
431 detailed facies analysis that includes measurements of thickness and density of deposits, samples
432 collected from representative layers need to be processed in the laboratory for multicomponent grain-
433 size analysis, including density and shape. On this regard, a word of caution must be spent on the

434 conditions that permit a proper application. Since our model is based on the concepts of sedimentation
435 from turbulent suspension and final traction at the bedload, a careful study of the deposit facies
436 architecture is needed in order to ascertain that such conditions are met. The occurrence of
437 asymmetrical bedforms, such as ripples and sand waves and of internal inclined lamination (see fig.
438 12) are features indicative of tractional processes at the bedload, which guarantee the application of
439 the model presented in this paper. However, pyroclastic density currents do not always behave as
440 turbulent suspensions, as it is the case of massive pyroclastic flows (Branney and Kokelaar, 2002),
441 or also the case of massive deposits from pyroclastic density currents found in proximal locations,
442 i.e. at the impact zone of the collapsing fountain (Sulpizio and Dellino, 2008; Sulpizio et al., 2014;
443 Dufek, 2016). An example of a metric thick, massive, structureless deposit formed by the impact of
444 the eruptive fountain feeding pyroclastic density currents of the Mercato eruption is shown, as an
445 example, on Fig. 14. The layer does not show any feature suggesting a particle selective transportation
446 in suspension or traction at the bedload. The model of sedimentation rate by turbulent suspension is
447 not applicable to this layer. In fact, at the impact, the particle volumetric concentration was so high
448 that particle-particle interaction played a stronger role than turbulent suspension upon deposition.
449 Judging from the experiments described in the present paper, it seems that the particulate mixture
450 issuing from the conduit and feeding the dense fountaining column did not undergo a grain size
451 selection upon the impact on the ground from where the “en mass” deposition of the massive layer
452 occurred. The grain size selection, in fact, started only after the development of the lateral turbulent
453 suspension current. It implies that, if thick, massive, structureless layers are formed by the collapse
454 of a natural eruptive column in the same way as in the experiments, their grain size, not having
455 underwent a selective process, can be taken as representative of the parent grain-size population
456 feeding the eruption. This outcome has important implication on the modelling of explosive
457 volcanism, since the parent grain-size population is one of the main parameters used for initializing
458 eruption simulations (Neri and Macedonio, 1996). Furthermore, concerning the modelling of grain
459 size of pyroclastic density currents, equation (6) allows the reconstruction of the grain size of the

460 material transported by turbulent suspension from the deposit or, vice-versa, the reconstruction of the
461 sedimented material grain size starting from that of the transported material. This information can be
462 used, in models, to predict the granulometric evolution of the particulate mixture during runout.

463 The sedimentation rate calculated for some pyroclastic density currents of Vesuvius and Campi
464 Flegrei falls within the range of experimental data, which is a guarantee of good scalability of the
465 model. Furthermore, the duration of the natural currents, as calculated by the model, is compatible
466 with the observations of historical eruptions, making us confident that the model allows a reasonable
467 reconstruction of the behaviour of natural currents. We expect that higher values of the sedimentation
468 rate will result when the model is applied to more powerful eruptions than those studied in this paper,
469 and a systematic investigation will allow acknowledging the true range that can be reached by
470 pyroclastic density currents. Probably there is an upper limit over which massive deposition takes
471 over suspension-sedimentation plus bedload-traction. Future investigations are required to determine
472 this limit.

473 As a conclusive remark, we suggest that the model proposed in this paper, as it is based on the general
474 laws that regulate the sedimentation of particulate material from turbulent suspension, can have
475 applicability beyond the study of pyroclastic density currents, for example to other geological density
476 currents such as turbidites.

477

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628 **Caption of figures**

629 Fig. 1. Grain-size distribution of the material used for experiments. A: multicomponent grain-size
 630 distribution of the coarse material coming from the Veusvius composition. The relative fractions of
 631 components are shown. The xx symbol means crystals. B: grain-size distribution of the fine glassy
 632 material coming from the Campi Flegrei composition.

633

634 Fig. 2. Skech design of the experimental apparatus with description of the main parts. Modified
 635 after Dellino et al., 2017.

636

637 Fig. 3. Display mount showing phases of the experiment and associated deposits. A: formation of
 638 the dense gas-particle fountain at the conduit exit. B: Collapse of the fountain and impact on the
 639 ground. C: development of the fully turbulent current. D: Massive deposit formed at the impact area
 640 where the flow is not established yet. Deposit with tractional features of the type of ripples formed
 641 by the fully turbulent suspension current. E: Fan shape distribution of the deposits.

642

643 Fig. 4. Grain-size distribution of the massive layer formed at the impact of the experimental
 644 fountain. A: grain size of a “coarse” run. B: grain size of a “fine” run.

645

646

647 Fig. 5. Graph showing the correlation between the ratio of the particle volumetric concentration of
 648 the transported material, C_i , and the proportion of the material sedimented from turbulent
 649 suspension, γ_i of the particle size classes as a function of the Rouse number of the size class P_{ni} . For
 650 comparison, data from a coarse run and data from a fine run are represented, together with the
 651 respective correlation coefficient, regression equation and normalized Rouse number Pn^* .

652

653 Fig. 6. Graph showing the correlation between the ratio of the particle volumetric concentration of
 654 the transported material, C_i , and the proportion of the material sedimented from turbulent
 655 suspension, γ_i of the particle size classes as a function of the ratio of Particle Rouse number of the
 656 size fraction P_{ni} and normalized Rouse number Pn^* . Data of all the components and size classes of
 657 all experiments are included. The correlation coefficient and regression equation are inset.

658

659 Fig. 7. Graph showing the correlation between the measured sedimentation rate and the
 660 sedimentation rate calculated by means of γ_{proxy_i} . The regression equation and correlation
 661 coefficient are inset.

662

663 Fig. 8. Graph showing the correlation between the measured sedimentation rate and the
 664 sedimentation rate calculated by means of γ_{true_i} . The regression equation and correlation
 665 coefficient are inset.

666

667 Fig. 9. Graph showing the correlation between the measured sedimentation rate and the
 668 sedimentation rate calculated by means of equation (5). The regression equation and correlation
 669 coefficient are inset.

670

671 Fig. 10. Comparison between the grain-size distribution of the transported material and that of the
 672 sedimented material. A: Coarse composition coming from Vesuvius. B: fine composition coming
 673 from Campi Flegrei.

674

675 Fig. 11. Graph showing the variation of the median size of the grain-size distribution of samples as
 676 a function of distance from the impact location for an experimental run fed with coarse material
 677 (dots) and one with fine material (triangles). D is maximum distance, d is distance from the impact.

678

679 Fig. 12. Photos showing the facies of deposits used for the application of the model of the
 680 sedimentation rate. A: layer of the Mercato eruption at Vesuvius showing tractional structures. B:
 681 layer of the Agnano Monte Spina eruption showing inclined laminae.

682

683 Fig. 13. Multicomponent grain-size distribution of layers used for the application of the model of
 684 the sedimentation rate. A: multicomponent grain-size distribution of the layer from the Mercato
 685 eruption at Vesuvius. B: multicomponent grain-size distribution of the layer from the Agnano
 686 Monte Spina eruption at Campi Flegrei.

687

688 Fig. 14. Photo showing a massive, structureless layer of the Mercato eruption formed at the impact
 689 of a collapsing eruptive fountain.

690