1 Crustal structure of the Mid Black Sea High from wide-angle seismic

- 2 data
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14 Abstract

15 The Mid Black Sea High comprises two en-echelon basement ridges, the

16 Archangelsky and Andrusov Ridges that separate the western and eastern Black Sea

17 basins. The sediment coverage above these ridges has extensive seismic reflection

18 coverage, but the crustal structure beneath is poorly known. We present results from

19 a densely sampled wide-angle seismic profile, coincident with a pre-existing seismic

20 reflection profile, which elucidates the crustal structure. We show that the basement

21 ridges are covered by c. 1-2 km of pre-rift sedimentary rocks. The Archangelsky

22 Ridge has higher pre-rift sedimentary velocities and higher velocities at the top of

23 basement (~6 km/s). The Andrusov Ridge has lower pre-rift sedimentary velocities

- 24 and velocities less than 5 km/s at the top of the basement. Both ridges are underlain
- 25 by c. 20-km thick crust with velocities reaching c. 7.2 km/s at their base, interpreted
- 26 as thinned continental crust. These high velocities are consistent with the geology of

the Pontides, which is formed of accreted island arcs, oceanic plateaux and

28 accretionary complexes. The crustal thickness implies crustal thinning factors of ~1.5-

29 2. The differences between the ridges reflect different sedimentary and tectonic

30 histories.

31

32 Introduction

33 Several episodes of extension and shortening have shaped the Black Sea region since 34 Permian times [e.g., Nikishin et al., 2003; Robertson et al., 2004; Yilmaz et al., 1997], 35 which led to the addition of a series of volcanic arcs, oceanic plateaux and 36 accretionary complexes to the Eurasian margin [e.g., Okay et al., 2013]. The basin is 37 thought to have formed in a back-arc extensional environment because of its close 38 spatial association with the subduction of both the Paleo- and Neo-Tethys Oceans 39 [e.g., Letouzey et al., 1977], but the timing and style of this opening history remain 40 controversial, partly because the thick sediment coverage means that the oldest 41 sedimentary fill has not been drilled [Banks et al., 1997; Nikishin et al., 2015a; Okay 42 et al., 1994; Zonenshain and Le Pichon, 1986; Okay et al., this volume]. The Black 43 Sea is commonly subdivided into eastern and western basins; these sub-basins are 44 separated by the Mid Black Sea High (MBSH), a system of buried basement ridges 45 that runs SW-NE [Fig. 1; e.g., Nikishin et al., 2015b; Okay et al., 1994].

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The opening of the western basin may be estimated from the ages of arc volcanic rocks in the western Pontides and from associated plate reconstructions; this evidence suggests a Middle to Upper Cretaceous age [*Görür*, 1988; *Okay et al.*, 1994; *Okay et al.*, 1994; *Okay et al.*, this volume]. Based on seismic refraction and gravity data, the crust in the centre of the basin is 7-8 km thick and has velocities consistent with the presence of oceanic

- crust, suggesting that rifting culminated in seafloor spreading [*Belousov et al.*, 1988; *Letouzey et al.*, 1977; *Starostenko et al.*, 2004].
- 54

55	The age and nature of the eastern basin are more controversial. The basin is thought to
56	have formed by rotation of the Shatsky Ridge relative to the Mid Black Sea High
57	[Figs. 1 and 2; Nikishin et al., 2003; Okay et al., 1994]. The main phase of opening
58	has been interpreted as Jurassic, Cretaceous [Nikishin et al., 2003; Nikishin et al.,
59	2015a; Okay et al., 1994; Zonenshain and Le Pichon, 1986], Early Eocene/Paleocene
60	[Banks et al., 1997; Robinson et al., 1995; Shillington et al., 2008], or Eocene
61	[Kazmin et al., 2000; Vincent et al., 2005]. Based on gravity and early seismic data,
62	the crust in the centre of this basin was inferred to have a thickness of \sim 10-11 km and
63	seismic velocities are lower than those of typical oceanic crust, suggesting the
64	presence of thinned continental crust [Belousov et al., 1988; Starostenko et al., 2004].
65	However, results from a wide-angle seismic experiment in 2005 suggest that the
66	crustal structure varies along the basin, with the western part floored by thinned
67	continental crust (7-9 km thick), and thicker, higher velocity crust below the eastern
68	part that is attributed to magmatically robust early seafloor spreading resulting in
69	early oceanic crust that is thicker and has higher velocities than average oceanic crust
70	[Shillington et al., 2009].

The Mid Black Sea High itself is divided into the en-echelon Archangelsky and
Andrusov ridges, which have different sediment thicknesses and are inferred to have
different structure and origin [*Nikishin et al.*, 2015b; *Robinson et al.*, 1996] (Fig. 1b).
These ridges are poorly explored compared to the basins either side. The Andrusov

76 Ridge is inferred to have formed during early opening of the eastern basin [Nikishin et 77 al., 2015a; Okay et al., 1994; Robinson et al., 1996]. This rifting event is inferred to 78 have been amagmatic in this part of the basin [Shillington et al., 2009]. Alternatively, 79 the Andrusov Ridge is interpreted as a marginal ridge associated with the opening of 80 the western basin along the West Crimean transform fault [*Tari et al.*, 2015]. The 81 Archangelsky Ridge was formed by the opening of the Sinop Trough, which is linked 82 to the western basin and is interpreted to have opened in Cretaceous to Palaeocene 83 times [Espurt et al., 2014; Robinson et al., 1996], with ongoing extension into the 84 Miocene [Espurt et al., 2014; Rangin et al., 2002]. An Upper Cretaceous sedimentary 85 sequence and lower Cretaceous platform carbonate rocks have been dredged where 86 the pre-rift sequences outcrops on the flank of Archangelsky Ridge, providing an 87 upper limit on its age of formation [Rudat and Macgregor, 1993; Robinson et al. 88 1996]. 89 After their formation, both ridges have also experienced compressional deformation 90 [Espurt et al., 2014; Rangin et al., 2002]. This region has likely experienced multiple

91 episodes of compression, continuing to the present; apatite fission track data and

92 paleostress measurements onshore show that inversion of rifting structure onshore

93 occurred as early as 55 Ma [Saintot & Angelier, 2002; Espurt et al., 2014] following

94 extension leading to opening of eastern Black Sea. Active compression continues

95 around margins of easternmost Black Sea today based on seismicity and onshore

96 geology, particularly in the Caucaus [Saintot & Angelier, 2002; Gobarenko et al.

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2016].

99 Published constraints on crustal structure beneath the ridges are sparse. Seismic
100 refraction data acquired in the 1960s were recently re-analysed using more modern

101	ray-tracing techniques [Yegorova and Gobarenko, 2010]. This analysis suggests a
102	crustal thickness of c. 20 km beneath both ridges and crustal velocities in the range
103	6.0-7.0 km/s, interpreted as representing thinned continental crust. A more modern
104	profile crossing the southern part of Archangelsky Ridge suggests that here, crustal
105	thickness reaches c. 25 km [Shillington et al., 2009]. In this paper we present results
106	from a modern, densely sampled wide-angle seismic profile that crosses the Andrusov
107	Ridge close to its southern tip and the Archangelsky Ridge at its northern tip (Fig. 1).
108	
109	Wide-Angle Seismic Data
110	
111	An onshore-offshore wide-angle seismic dataset was collected in 2005 using the R/V
112	Iskatel to determine the deep structure of the eastern basin and Mid Black Sea High.
113	Seventeen four-component short-period ocean-bottom seismometers (OBSs) from
114	GeoPro were deployed on Profile 4 across the Andrusov Ridge (Fig. 1; Table 1), and
115	they recorded seismic shots generated from an airgun array with a total volume of
116	3140 in ³ that was triggered every 90 s (shot spacing c. 150 m). Profile 4 was co-
117	located with existing industry seismic reflection data, 91-106 (Figs. 1 and 2).
118	
119	Data Analysis
120	
121	Data processing
122	Water-wave arrivals were used to relocate OBS positions on the seafloor, using a
123	seafloor depth determined by echosounder at the position of each deployment and a
124	water velocity of 1.47 km/s. Relocated positions were typically less than 75 m from
125	deployment positions, but three OBS have relocated positions that differ by 200-300

m from deployment positions. We applied a minimum phase band-phase filter with
corners at 3, 5, 15, 20 Hz to suppress noise, and applied offset dependent gains and a
reduction velocity of 8 km/s.

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130 Phase identification

131 We identified refractions and wide-angle reflections from the pre-rift sedimentary 132 section, the crust and the upper mantle that could be consistently identified on a 133 majority of the receiver gathers. Phase interpretations and velocity models of the 134 overlying syn- and post-rift sedimentary section have been presented elsewhere [Scott 135 et al., 2009]. Travel-time picks were made manually of the following phases: 136 reflections off the base of an interpreted pre-rift sedimentary layer (PprP), crustal 137 refractions (Pg), reflections from the base of the crust (PmP), and upper mantle 138 refractions (Pn) (Fig. 3; Table 2). Reflections from the base of the interpreted pre-rift 139 sedimentary section are observed from near-vertical incidence to offsets up to ~30 km 140 and have picking uncertainties of 30-50 ms. Crustal refractions are observed as first 141 arrivals at offsets from \sim 12-100 km and have picking uncertainties of 30 to 75 ms. 142 Reflections from the base of the crust are observed at offsets between ~35-100 km; 143 the offsets where PmP reflections are observed vary significantly over the line, 144 indicating variations in crustal thickness. Likewise, the amplitude and character of 145 PmP reflections is also highly variable and thus picks of this phase have relatively 146 high uncertainties of 125 ms. We observed limited and relatively low amplitude 147 refractions interpreted to arise from the upper mantle in some receiver gathers; these 148 refractions are weak and variable, and have a picking uncertainty of 125 ms. Figure 3 149 shows examples of OBS data, phase identifications, and associated ray paths.

Wide-angle reflections interpreted to originate from the base of the interpreted pre-rift sedimentary layer can be linked to a coincident industry seismic reflection profile (BP91-106, Fig. 2). Picks of this interface were thus also made on the reflection profile (Fig. 2, red dotted line) and included in the inversion. We assigned an uncertainty of 100 ms to these picks to account for uncertainties in associating MCS and wide-angle reflections and for small-scale variations in interface geometry that cannot be recovered by inversion.

158

159 Velocity modelling

160 The travel-time picks described above were used to invert for velocities of the pre-rift

sedimentary section, crust and upper mantle. We used JIVE3D, a regularized

162 tomographic inversion code [Hobro et al., 2003], which solves for a minimum

163 structure layer-interface model that fits the data within its uncertainties. Velocities

164 within each layer and interface depths are defined by splines and vary smoothly;

165 interfaces represent velocity discontinuities. The forward problem involves tracing a

166 fan of rays from each OBS position through specified layers in the model to generate

167 predicted travel times (i.e., ray shooting); the ray that arrives within a distance

tolerance of the target with the minimum travel time is used. Inversion involves a

169 sequence of linear steps to reduce the difference between observed and predicted

travel times (e.g., Figs. 4d and 5d) and satisfy other smoothing criteria. In each step,

171 smoothing is reduced and structure is allowed to develop to improve data fit.

172 Smoothing is implemented during inversion by minimizing a function of data misfit

and model roughness.

174

175 We employed a layer stripping approach for this line. The previously determined 176 velocity structure of the post- and syn-rift sediment from Scott et al. [2009] was held 177 fixed. We first inverted for the interpreted pre-rift sediment layer using picks of wide-178 angle reflections from OBS data and vertically-incident reflections from the 179 coincident seismic reflection profile (Fig. 2). This layer was then held fixed during 180 the inversion for crustal and upper mantle structure. The inversion converged more 181 quickly and stably for both the pre-rift sedimentary section and for the crustal/mantle 182 sections when we inverted for them separately. However, inverting for all layers 183 simultaneously yielded the same overall velocity structure. We also performed two 184 different inversions for crust/mantle structure. The first inversion used only first 185 arriving refractions from the crust and mantle. The second inversion included 186 interpreted wide-angle reflections from the base of the crust (PmP) in addition to the 187 first arrivals. The purpose of performing two inversions for the crust and upper 188 mantle structure was to assess which features in the model arise from the inclusion of 189 wide-angle reflections from the base of the crust; identifying PmP is associated with 190 more uncertainty and subjectivity than first arrivals. We are most confident of 191 features that are present in both the first-arrival and reflection/refraction tomographic 192 inversions, and more cautious of features that are primarily constrained by the PmP 193 reflections.

194

We used a grid spacing of 1 km x 0.5 km in the pre-rift interval, and 1x1 km in the crust and upper mantle. For both inversions, we applied twice as much horizontal smoothing than vertical smoothing and allowed more interface roughness than velocity roughness. A simple 1D velocity model and constant interfaces were used for the starting models in both inversions.

The inversion for the pre-rift layer used 825 picks from the OBS data and 129 picks from the MCS data. The final model has a chi-squared misfit of 1.29 and RMS residual of 72 ms if only the OBS picks are included. Larger misfits are associated with the MCS picks since they include smaller scale variations in interface geometry than can be recovered by the inversion. If these are included, the overall chi-squared misfit is 1.65, and the RMS residual is 90 ms.



209 The final model has a chi-squared value of 0.96 and an RMS residual of 76 ms. The

210 reflection/refraction inversion used 7085 picks. The final model has a chi-squared

211 value of 2.23 and an RMS residual of 127 ms.

212

213 Based on ray coverage, data fit and testing of different inversion parameterizations, 214 we discuss the confidence that should be placed in different features of our final 215 models. The upper crustal structure is very well sampled by ray coverage associated 216 with our travel-time picks, and refractions from this part of the model have relatively 217 low misfits (Figs. 3-5). Similar features are apparent in both the reflection/refraction 218 tomography and the first-arrival tomography. Thus, we consider the variations in 219 upper crustal velocity structure between the Andrusov and Archangelsky Ridge to be 220 a robust result (Figs. 4b and 5b). The lowermost crustal sections beneath the 221 Andrusov and Archangelsky ridges are only constrained by sparse turning wave 222 coverage and relatively sparse reflections from the base of the crust (Figs. 3 and 5). 223 Because the uppermost part of the lower crust is sampled by reversed refracted 224 arrivals, we are confident that high velocities are required. However, we cannot

225 constrain the velocity gradient of the lowermost crust or absolute velocity at the very 226 base of the lower crust, and there are thus tradeoffs between velocities in the 227 lowermost crust and depth to the base of the crust. Both wide-angle reflections and 228 vertically incident reflections constrain the interpreted pre-rift sedimentary layer on 229 top of the MBSH. We find relatively high data misfits for phases defining this layer 230 (Table 2), which we attribute to substantial lateral variability that cannot be accounted 231 for in the analysis of OBS spaced at ~15 km. However, we think that the large-scale 232 patterns of thickness and velocity are well constrained.

233

Although we obtained an excellent misfit for the first-arrival tomography model (chisquared value of 0.96), our favored model from reflection/refraction tomography has a higher chi-squared value of 2.23. We relaxed the data misfit criteria to obtain a relatively smooth model; models with better data fit were substantially rougher. We feel this choice is justified by the likely three-dimensionality of velocity structure beneath these complex ridges and the complexity of sedimentary, crustal and upper mantle phases observed on OBS.

241

- 242 **Results and Discussion**
- 243

The final velocity models across the Mid-Black Sea High provide constraints on the

245 deep sedimentary and crustal structure of this composite ridge.

246

247 Sedimentary rocks overlying the Mid Black Sea High

248 The flat-lying post-rift sedimentary rocks exhibit a low-velocity zone in the Miocene

249 Maikop formation (Figs. 5 and 6) that extends across the eastern basin and also

appears to be present above parts of the MBSH and in the Sinop Trough [Fig. 6; *Scott et al.*, 2009]. The low-velocity zone is attributed to fluid overpressure, and fluid
pressures close to lithostatic have been inferred [*Scott et al.*, 2009], though application
of a more sophisticated approach in the eastern basin [*Marin-Moreno et al.*, 2013a; b]
suggests that fluid pressures are lower than those derived from the empirical
approaches of *Scott et al.* [2009].

256

257 Wide-angle reflections in the OBS data (Fig. 3) and reflections in the reflection 258 profile (Fig. 6) define a distinct layer with a thickness of 1-2 km and velocities of 3.0-259 4.75 km/s on top of the Andrusov and Archangelsky Ridges (Fig. 5). Based on the 260 character of this layer in the reflection profile, dredging on the Archangelsky Ridge 261 and drilling of the Andrusov Ridge, we interpret this layer to represent a sequence of 262 prerift Upper Cretaceous sedimentary rocks [Rudat & McGregor, 1993; Avdemir & 263 Demirer, 2013]. This layer is characterized by brightly reflective layering in the 264 reflection profile, which is consistent with a sedimentary origin (Figs. 2, 6). Drilling 265 on Andrusov Ridge at Sinop-1 recovered a relatively thin layer of Upper Cretaceous 266 carbonate rocks [Aydemir & Demirer, 2013]. Aydemir & Demirer [2013] suggest that 267 the thickness of this interval would be strongly controlled by basement topography at 268 the time of deposition and thus be highly variable, which may explain why we appear 269 to observe a thicker Upper Cretaceous layer on Profile 4. A similar sequence overlies 270 the Shatsky Ridge to the north [Fig. 1; Nikishin et al., 2015b; Robinson et al., 1996]. 271

272 The base of this layer is marked by a bright, continuous reflection in the reflection

273 profile (Fig. 6), which has been interpreted to mark the top of Lower Cretaceous

274 platform carbonate rocks [Rudat & McGregor, 1993; Robinson et al., 1996]. Based on

275 dredging results on the shallow part of the Archangelsky ridge, we interpret the 276 uppermost basement beneath this reflection as being composed of Lower Cretaceous 277 platform carbonate rocks and other older prerift sedimentary rocks. Platform 278 carbonate rocks are expected to have similar P-wave velocities to upper crystalline 279 crust [Christensen & Mooney, 1995], so it is not possible for us to definitely identify 280 carbonate rocks or quantify their thickness, but the nearby dredging results suggest 281 prerift sedimentary rocks are likely present in the uppermost basement here. The 282 uppermost basement beneath the prominent reflection described above reaches 6-6.25 283 km/s beneath the top of the Archangelsky Ridge, and drops to c. 4.5 km/s beneath the 284 Andrusov Ridge. The overlying layer interpreted to represent Upper Cretaceous 285 prerift sedimentary rocks also has significantly higher velocities beneath 286 Archangelsky Ridge than beneath Andrusov Ridge. These differences may be 287 attributed to several factors. First, although Archangelsky Ridge is generally a 288 shallower feature (Fig. 1), at the location of Profile 4 it is more deeply buried, so the 289 pre-rift sedimentary rocks may have undergone greater compaction and diagenesis. 290 Secondly, seismic reflection data suggest that the Andrusov Ridge is disrupted by 291 more faults than the Archangelsky Ridge [Robinson et al., 1996], and fracturing 292 associated with these faults may reduce the velocity by creating zones of higher 293 porosity and/or causing an elongation of pores, which have a bigger impact on elastic 294 properties [*Töksöz et al.*, 1976]. Thirdly, other differences in lithology may 295 contribute to observed variations in velocity. Finally, the low-velocity layer in the 296 post-rift directly abuts the Andrusov Ridge, but is separated from Archangelsky Ridge 297 by a layer of higher-velocity material. Therefore it is possible that fluid overpressure 298 is transmitted into pre-rift sedimentary rocks on the Andrusov Ridge but not on the 299 Archangelsky Ridge.

301 Crustal structure and Implications for Tectonic Evolution

302 The Andrusov and Archangelsky Ridges exhibit distinctly different crustal velocity 303 structures. As described in the previous section, the Archangelsky Ridge has higher 304 velocities in the uppermost basement (6-6.25 km/s) and a relatively low velocity 305 gradient (~0.075 km/s/km). In contrast, the Andrusov ridge has velocities in the 306 shallow basement as low as 4.5 km/s and a high velocity gradient in the upper 10 km 307 of 0.25 km/s/km. These differences might be associated with different degrees of 308 fracturing of platform carbonate rocks (see previous section) or of crystalline rocks, or 309 might arise because the prerift sedimentary sequence within the basement is thicker 310 beneath Andrusov Ridge, as perhaps suggested by seismic reflection data (Fig. 2). 311 312 Beneath both ridges, the velocity gradient is reduced in the lower crust, and velocities 313 reach a maximum of 7.2-7.3 km/s at the base of the crust (Fig. 5). These velocities 314 are somewhat higher than those observed beneath Archangelsky Ridge on Profile 3 315 (~6.75-7 km/s) [Shillington et al., 2009] (Fig. 1), and may indicate the presence of a 316 more mafic pre-rift crust [e.g., Christensen and Mooney, 1995]. Rifting to form the 317 eastern Black Sea occurred in a series of terranes accreted to the Euroasian margin, 318 which include volcanic arcs and oceanic plateaux, both of which are typified by high-319 velocity lower crust in modern analogues [Calvert, 2011; Kodaira et al., 2007; 320 Shillington et al., 2004].

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322 These velocities are also only slightly lower than lower-crustal velocities observed in

323 crust within the centre of the eastern part of the Eastern Basin [Shillington et al.,

324 2009], which were interpreted as evidence for new magmatic crust formed during

325 magma rich rifting and early spreading. However, the relationship between lower crustal velocity and crustal thickness suggests that synrift magmatism is not 326 327 responsible for the high lower crustal velocities beneath the MBSH. In the eastern 328 part of the Eastern Basin [Shillington et al., 2009] and at other volcanic rifted margins 329 worldwide [e.g., Holbrook and Kelemen, 1993; White et al., 2008], high-velocity 330 lower crust (~7.4-7.5 km/s) interpreted to represent mafic synrift intrusions is most 331 prominent in the area of crustal thinning. In contrast, the highest velocities observed 332 beneath the MBSH occur in the thickest crust and do not increase towards the thinned 333 margins of the ridge. Consequently, we propose that high lower crustal velocities 334 beneath the MBSH represent high velocities associated with accreted volcanic arcs 335 and oceanic plateaux in the pre-rift crust. Hence our observations from Profile 4 is 336 consistent the view that extension in the western part of the eastern Black Sea Basin 337 was largely amagmatic [Shillington et al., 2009].

338

339 The crustal layer, that may include platform carbonate rocks and possibly other pre-340 rift sedimentary rocks, thickens beneath both ridges to reach a maximum of 20-23 km 341 (Fig. 5). Between the two ridges, it decreases to c. 16 km, providing evidence that the 342 modest increase in sediment thickness between the two ridges (Fig. 1) is associated 343 with crustal-scale extension. Although the Archangelsky Ridge is deeply buried at 344 the location of Profile 4 (Fig. 1), it clearly remains a major crustal feature at this 345 location. Uppermost mantle velocities are a little below 8 km/s. Based on teleseismic 346 receiver functions, gravity data and limited wide-angle seismic constraints, the crustal 347 thickness onshore Turkey in the vicinity of Archangelsky Ridge is c. 35 km [Ozacar 348 et al., 2010; Yegorova et al., 2013], with thicker crust farther east where it is affected 349 more by compressional deformation. Therefore the crust along Profile 4 has been

350	thinned by a factor of 1.5-2. The degree of thinning is somewhat lower than inferred
351	by Shillington et al. [2008] based on the relationship between sediment thickness and
352	thinning factor on a well-constrained profile; this relationship gives a thinning factor
353	of 2-2.5 along most of Profile 4 (Fig. 7). One possible explanation for this difference
354	is that the "crust" of the Mid Black Sea High may include sections of pre-rift
355	sedimentary rocks that are not a part of the unthinned crustal section onshore [Okay et
356	<i>al</i> , this volume].

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358	Conclusions
220	Conclusions

359 From our analysis of data from a wide-angle seismic profile across the Mid Black Sea 360 High, comprising the en echelon Archangelsky and Andrusov ridges, we conclude 361 that:

362	1.	The basement highs are covered by at least 1-2 km layer of pre-rift
363		sedimentary rocks overlying a higher-velocity basement that may include pre-
364		rift sedimentary rocks, including platform carbonates that cannot be readily
365		distinguished from underlying crystalline crust.
366	2.	The pre-rift sedimentary rocks and upper basement have higher velocities on

- the Archangelsky Ridge and lower velocities on the Andrusov Ridge. These 367 368 differences could be explained by different amounts of faulting or changes in the abundance and/or composition of prerift sedimentary rocks. 369
- 370 3. The lower crust has a low velocity gradient and velocities exceed 7.0 km/s at 371 its base; the velocity structure is consistent with the presence of a mafic prerift crust with little magmatic addition during rifting. 372

- 373
 4. The crust is 20-23 km thick beneath the ridges and c. 16 km thick between
 374 them, representing thinning factors of 1.5-2.0 compared to adjacent crust in
 375 northeastern Turkey.
- 376

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387 Figure Captions

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389 Figure 1. a. Elevation/bathymetry of Black Sea region from GEBCO showing the 390 location of 2005 onshore/offshore seismic refraction experiment. Shot lines are 391 indicated with white lines, OBS are shown with white circles and seismometers 392 deployed onshore shown with white triangles. OBS from Line 4, which are used in 393 this study, are indicated with solid circles. Major tectonic elements indicated with 394 dashed yellow lines [Zonenshain and Le Pichon, 1986]. Black box indicates area 395 shown in Figure 1b. b. Close-up of Mid Black Sea High showing sediment thickness 396 [Shillington et al., 2008] and OBS locations from 2005 experiment. Note that Mid-397 Black Sea High separates the Western and Eastern basins of the Black Sea and 398 comprises two ridges: the Archangelsky Ridge and the Andrusov Ridge. Seismic 399 reflection profile 91-106 (Fig. 2) is shown with thick white line. It is coincident with 400 Profile 4 but shorter; it extends southwest to between OBS 3 and 4. 401 402 Figure 2. a. Seismic reflection profile 91-106 across the Mid-Black Sea High, which 403 is coincident with the Line 4 OBS profile (courtesy of BP and TPAO) (see Fig. 1 for 404 location). b. Seismic reflection profile with interfaces used in seismic inversion. The 405 blue, green and orange dotted lines show interpreted horizons used to invert for post-

406 and syn-rift sedimentary structure by Scott et al. [2009]. The red dotted line shows

407 the interpreted pre-rift sedimentary horizon used in the inversions presented here.

408

409 **Figure 3**. Receiver gather without picks (top panel). Data with observed picks and

410 picking errors (closed circles and bars) and predicted picks (solid, lighter colored

411 circles) (middle panel). Orange – PprP; Blue – Pg; Green – PmP; Red – Pn. Ray paths

412 through final model from reflection/refraction tomography model. a. OBS 2, b.
413 OBS9, c. OBS13, d. OBS15.

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Figure 4. a. Result of inversion for pre-rift sedimentary reflections (PprP) and first-416 417 arriving refractions from crust and upper mantle (Pg and Pn). Velocities contoured at 418 0.25 km/s. Velocity model is masked by density of ray coverage. **b.** Density of ray 419 coverage over the velocity model in a. c. Observed and predicted travel-time picks. 420 Uncertainty of observed picks indicated with bars. d. Travel-time residuals for picks. 421 422 Figure 5. a. Result of inversion for pre-rift sedimentary reflections (PprP), first-423 arriving refractions from crust and upper mantle (Pg and Pn), and reflections from the 424 base of the crust (PmP). Velocities contoured at 0.25 km/s. Velocity model masked by 425 density of ray coverage. b. Density of ray coverage over the velocity model in a. c. 426 Observed and predicted travel-time picks. Uncertainty of observed picks indicated 427 with bars. d. Travel-time residuals for picks. 428 429 Figure 6: Overlay of reflection profile 91-106 on final velocity model from 430 reflection/refraction tomography (Fig. 5), which was converted to two-way travel 431 time. 432 433 Figure 7: Comparison of crustal thinning factor (beta = initial thickness/rifted 434 thickness) along Line 4 from subsidence analysis based on sediment thickness 435 [Shillington et al., 2008] and from this study assuming an initial crustal thickness of 436 35 km.

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580	

Table 1: Relocated OBS positions

OBS	Latitude (°N)	Longitude (°E)
1	42.511005	35.212699
2	42.549179	35.322692
3	42.589855	35.432331
4	42.625923	35.543609
5	42.663829	35.654865
6	42.701	35.766201
7	42.738536	35.876911
8	42.777019	35.987457
9	42.813636	36.09951
10	42.851139	36.21101
11	42.887451	36.323356
12	42.925537	36.434604
13	42.960388	36.540924
14	42.995098	36.645301
15	43.035087	36.765087
16	43.073787	36.882984
17	43.10337	36.974617

Table 2

Phase	Number Picks	Chi Squared	RMS Misfit (s)
PprP	866	3.442881645	0.129567735
Pg	5334	2.038537344	0.106868266
PmP	1502	2.182268919	0.182210481
Pn	249	2.604642144	0.200007259



Figure 2



Figure 3







Figure 4



Figure 5







Figure 7

