

Climate and geology – a Phanerozoic perspective

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Abstract

The Phanerozoic is over 540 million years long, and, with its defining accompaniment of abundant complex life, provides us with a unique perspective on the extremes of climate change. Understanding these extremes is particularly important if we are to anticipate the possible effects of global warming. The broad sweep of climate change through the Phanerozoic began with relatively cool global temperatures and recovery from late Proterozoic glaciation. This was followed by a mid Cambrian to Ordovician episode of relatively warm global climate, after which global climate cooled, culminating in the major glaciations of the Carboniferous and Permian periods. The Triassic and Early Jurassic were warm. The Late Jurassic–Early Cretaceous period was cool, although without full global glaciation. Global temperatures peaked in the mid-Cretaceous. Since then global climates have cooled, culminating in Neogene glaciation. These ~100 million year trends in overall climate show short intense excursions of contrasting climate, many of which have been associated with the mass extinction of life, and with major volcanic and tectonic events. This chapter argues that, through the Phanerozoic, two overlapping stable climate regimes appear to have dominated: a high-CO₂ (>1000 ppmv), largely warm climate regime, punctuated by many short-lived episodes of glaciation; and a low-CO₂ (<1000 ppmv), largely cool regime, marked by protracted episodes of superglaciation.

As Hay *et al.* (1997) pointed out, "the climate of the Holocene is not well-suited to be the baseline climate of the planet". This is particularly important in light of predicted changes to the global climate anticipated as a result of global warming (Hay *et al.* 1997; IPCC 2001). To understand what may happen we must examine and try to simulate the warm and cool climates of the more distant past, ideally of the whole Phanerozoic (Fig. 1). The most

complete summary of Phanerozoic warm and cool climates continues to be Frakes *et al.* (1992), based on geological evidence from the rock record and oxygen isotopes. They subdivided Phanerozoic time into palaeoclimate "modes" (Table 1; Fig. 1a), independent of geological timescale boundaries, improving on the chronostratigraphic system by system treatment by Frakes (1979). These "modes" are still valid and useful today, although Barron (1993) pointed out that they may represent oversimplifications. For example, Barron (1993) argued that Frakes *et al.* (1992) placed the Pliocene warm interval, currently the focus of much study to understand the possible effects of global warming (e.g. Haywood *et al.* 2005; Barreiro *et al.* 2006), in one of their cold modes. Barron (1993) also contended that they grouped quite dissimilar Late Cretaceous and Early Eocene warm intervals together in the same warm mode. In this chapter, the warm and cool climate modes of Frakes *et al.* (1992) will be reviewed in time-sequential order, and only developments or changes since their synthesis will be covered in any detail.

Between 1992 and the advent of this book there have been several reviews of aspects of the Phanerozoic palaeoclimate record, although none on the scale of Frakes *et al.* (1992). The volume edited by Walliser (1995) reviewed global event stratigraphy for the Phanerozoic, which includes many aspects of climate change. Crowell (1999) examined the implications of pre-Mesozoic ice ages on our understanding of the climate system, and Huber *et al.* (2000) specifically treated former warm climates. Boucot & Gray (2001) assessed Phanerozoic climate models based on atmospheric CO₂, and reviewed the many CO₂ proxies, with a particular emphasis on soil and soil-forming organisms. Wallmann (2001) has come closest to the treatment by Frakes *et al.* (1992) and used the record of marine $\delta^{18}\text{O}$ values to subdivide the Phanerozoic. He identified three combined icehouse-greenhouse cycles with durations of approximately 127 My between the Cambrian and the Triassic, followed by a

longer cycle spanning the Jurassic to Cenozoic, which largely correspond to the climate modes proposed by Frakes *et al.* (1992) and discussed in this chapter. Subsequently, Wallmann (2004) used a multi-proxy approach to model Phanerozoic climates and identify warm and cool climate epochs based on CO₂ levels, proposing a link with the intensity of the galactic cosmic ray flux.

Palaeoclimate proxies

We cannot directly observe the climates of the geological past in the way that we have been able to observe climate parameters since the 17th century (e.g. Plaut *et al.* 1995). The key variables, which are both physical and chemical, and include temperature, atmospheric composition and seawater salinity, cannot be measured directly. We depend on chemical and biological systems that have responded, hopefully systematically, to changes in climate, and which have left a record in sediments and ice, which if not pristine, has at least been altered in a systematic way. These so-called proxy data, proxies for short, can be interpreted to yield numerical and qualitative data for the key variables of former climates in the geological past (e.g. Wefer *et al.* 1999). Proxies must be evaluated carefully and the level of confidence in a proxy can change as the factors affecting its reliability are better understood (e.g. Elderfield 2002; Williams *et al.* 2005). Geological proxies, i.e. those based on sedimentary facies such as coal or evaporites, can appear relatively straightforward to interpret but may introduce a bias in the signal through incompleteness of the stratigraphic record. There are also parallel factors, such that often only extreme events get preserved, and that preservation of ordinary events often requires peculiarities of tectonic setting (for a discussion of some of these caveats see Sellwood *et al.* 1993). It is useful to think of the tides in this respect and when or where today the constant shifting of sand on an average beach might be preserved to record a

shoreline. Even in tectonically active areas, such as the western margin of the Americas, earthquakes that shift the base level, such as the 1960 Chile earthquake (e.g. Cisternas *et al.* 2005), or the 1964 Alaska earthquake (e.g. Hamilton *et al.* 2005), happen only once every few hundred years along any particular section of coastline. Therefore, of the order of 150,000 tidal events may go unrecorded, with only the last tides preceding a major earthquake having a chance of preservation. By comparison, chemical and palaeontological methods of estimating Phanerozoic palaeoclimate are complex to interpret, but, and certainly in the case of oceanic records, they can offer better guarantees of completeness, at least back into the Jurassic.

In the sections below, proxies are reviewed (summarized in Table 2) and organised according to the way that proxies are grouped within particular subdisciplines. Chemistry-based proxies are the most abstract and tend to focus on the chemical or isotopic system applied. These are grouped by parameter estimated. Lithological proxies are the most directly related to the parameter estimated and are grouped by rock type, mineralogy or facies interpreted. Palaeontological proxies are subdivided by whether they use taxonomic methods or focus on some morphological aspect of a group of organisms. For clarity and ease of reference, all the proxies reviewed are grouped by parameter estimated in Table 2.

Chemistry-based proxies.

Chemistry-based palaeoclimate proxy studies were made possible by the stable isotope work of Harold Urey and others in the 1940's and 1950's (Urey 1947; Epstein *et al.* 1953). Emiliani (1955) published the first quantitative Pleistocene palaeotemperature curve based on the new $\delta^{18}\text{O}$ palaeothermometer. This work was substantially built upon by Shackleton and others (Shackleton 1965; Shackleton & Opdyke 1973; Shackleton & Kennett 1975;

Shackleton 1977), culminating in the enormous success of the CLIMAP project in the 1970's (CLIMAP 1981). Table 2 summarizes chemical proxies based on isotope-, element- and biogeochemistry-based systems.

Henderson (2002) has provided a useful framework for the treatment of chemistry-based oceanic proxies, which will form the basis for this section, though here applied to both oceanic and continental chemistry-based proxies. Henderson (2002) distinguished proxies that can be used to reconstruct the physical environment, such as palaeotemperature and mass flows, and those that can be used to reconstruct aspects of the carbon cycle, such as oceanic pH or atmospheric CO₂ concentrations.

Palaeotemperature. Surface temperature is a reflection of time-averaged net solar energy input. This is energy that is not immediately reradiated back to space or involved in photosynthesis, but which is trapped for a time, either in the atmosphere via, for example, the greenhouse effect and hydrological cycle, e.g. evaporation and ice melting, or in the solid surface, lakes, seas and oceans. Surface temperature is the key variable for the Earth system. In the oceans, it drives the weather and winds, which ultimately power circulation in the atmosphere and oceans. On land and sea, it controls evaporation, which is a major influence on the water cycle and weathering.

As outlined above, $\delta^{18}\text{O}$ in carbonate has been used as a surface palaeotemperature proxy since the 1950's (e.g. Emiliani 1955). $\delta^{18}\text{O}$ values are determined from skeletal and non-skeletal carbonates (Marshall 1992), with varying degrees of success, depending on whether the carbonate analysed is pristine or secondary. In marine settings $\delta^{18}\text{O}$ ratios are derived from calcareous plankton, such as foraminifera or coccolithophores, as well as other marine

invertebrates, such as brachiopods (Korte *et al.* 2005), corals, and various molluscs (e.g. Gazdzicki *et al.* 1992; Elorza *et al.* 1996; Bailey *et al.* 2003), and also bulk sediments (Tobin & Walker 1997) and cements (Hays & Grossman 1991). Marine records extend back to the Proterozoic (e.g. Samuelsson 1998). In terrestrial settings, speleothems (Winograd *et al.* 1992) provide a detailed record, as do lake-dwelling ostracods (Schwalb 2003) and gastropods (Grimes *et al.* 2003), although only extending back to the Pleistocene. Pedogenic carbonate provides a longer time record (e.g. Mack & Cole 2005), although with significant gaps and problems of interpretation. Overall, interpretation of the signal from carbonate is complicated by, for example, the presence of cryptic species, such as in the foraminifera (e.g. Kucera & Darling 2002), biological isotope fractionation during development, so called "vital effects" (e.g. Friedrich *et al.* 2006), and diagenesis (e.g. Wenzel *et al.* 2000; Williams *et al.* 2005). $\delta^{18}\text{O}$ ratios can also be derived from non-calcareous sources, such as vertebrate tooth enamel (e.g. Bentaleb *et al.* 2006) and bone (e.g. Barrick *et al.* 1999), as well as fish scales and otoliths (e.g. Grimes *et al.* 2003), and diatom silica (e.g. Leng & Barker 2006).

Other carbonate-based palaeotemperature proxies use the ratios of trace elements, such as foraminiferal Mg/Ca (e.g. Elderfield 2002; Elderfield *et al.* 2002), coralline Sr/Ca (e.g. Devilliers *et al.* 1994; Wei *et al.* 2004) and Li/Ca (Marriott *et al.* 2004), or Ca isotope ratios such as foraminiferal $\delta^{44}\text{Ca}$ (e.g. Gussone *et al.* 2003). However, Ca isotope ratios have complex species-related variations in sensitivity (e.g. Sime *et al.* 2005) that complicate their usefulness. Mg/Ca and Sr/Ca from belemnites appears to be effective from the Late Cretaceous back to the Early Jurassic (Rosales *et al.* 2004).

More recent non-carbonate palaeotemperature proxies include the index of unsaturation of alkenones from the membranes of some Prymnesiophyceae algae (coccolithophores) such as

Emiliana huxleyi (e.g. Prahl *et al.* 1988), the so-called UK'37 proxy. This proxy has proved to be very successful (e.g. Bard 2001), is effective back to the Miocene for marine sediments (e.g. Haywood *et al.* 2005), and has also been applied to lakes (Li *et al.* 1996). Oxidation effects needs to be considered for the older record (e.g. Hoefs *et al.* 1998). Amino acid racemization epimerization in molluscan calcite is another terrestrial palaeotemperature proxy (e.g. Murray-Wallace *et al.* 1988), which initially showed promise back to the Pliocene (e.g. Kaufman & Brigham-grette 1993). However, recent studies have shown its application to be limited to the Late Pleistocene (e.g. Andersson *et al.* 2000). Similar to UK37', TEX86 is a palaeotemperature proxy based on the tetraether lipids of crenarcheota (Schouten *et al.* 2002). Initial results suggest that the proxy is robust (e.g. Wuchter *et al.* 2004). It appears to be resistant to post-depositional oxidation (e.g. Schouten *et al.* 2004) and is capable of recording palaeotemperatures above the ~29°C limit of UK37' (e.g. Pelejero & Calvo 2003; Schouten *et al.* 2003). TEX86 is also applicable to lakes (e.g. Powers *et al.* 2004), and it has an effective time range extending back to the Mesozoic (e.g. Schouten *et al.* 2003).

Palaeosalinity. Henderson (2002) highlighted palaeosalinity as a major environmental parameter for which we have no direct marine proxy. Calculations based on $\delta^{18}\text{O}$ are prone to error (Schmidt 1999; Rohling 2000), but have been applied back to the Early Triassic (Korte *et al.* 2005). Recent developments suggest that U/Ca ratios in corals (Ourbak *et al.* 2006) and alkenone unsaturation ratios (Sikes & Sicre 2002; Blanz *et al.* 2005; Schouten *et al.* 2006) show some promise.

Palaeocirculation of oceans. Palaeocirculation proxies either trace the movement of water masses or provide estimates of deep water flow (Henderson 2002). Of the former, $\delta^{13}\text{C}$ (e.g. Lynch-Stieglitz & Fairbanks 1994), Zn/Ca (e.g. Marchitto *et al.* 2000) and Cd/Ca (Boyle 1988) ratios in foraminiferal carbonate and $^9\text{Be}/^{10}\text{Be}$ ratios in manganese crusts (Kusakabe *et*

al. 1987; vonBlanckenburg *et al.* 1996) trace water masses by mimicking nutrients. More recent tracers of water masses include isotopes of Nd and Pb (Ling *et al.* 1997), and Hf (David *et al.* 2001) mainly from ferromagnesian crusts on manganese nodules, but also from foraminifera (e.g. Vance & Burton 1999), and Ag/Si ratios (Zhang *et al.* 2004). For deep water flow, radioactive proxies include ^{14}C , and the U-series daughter products ^{231}Pa and ^{230}Th (Henderson 2002). These have half-lives of less than 100 ky (e.g. Edwards *et al.* 1997) and are really only applicable to the Pleistocene.

Sea-ice. IP25 (Belt *et al.* in press) is a proxy for palaeo-sea-ice distribution based on the distribution of C25 highly branched isoprenoid alkenes biosynthesised by diatoms living in sea-ice, notably genera of *Haska* and *Rhizosolenia* (Belt *et al.* 2006). The technique has so far been demonstrated in the Canadian Arctic for the Holocene (Belt *et al.* in press).

The carbon cycle. Understanding and quantifying the carbon cycle is key to determining the controls on past atmospheric CO_2 (Henderson 2002). For long timescales, i.e. longer than a million years, the main control on atmospheric CO_2 is the balance between volcanogenic gas output and net consumption by silicate weathering and carbonate deposition on the seafloor (Walker *et al.* 1981). For shorter timescales, the oceans (e.g. Henderson 2002), terrestrial vegetation (e.g. Schimel 1995) and soils (e.g. Schlesinger & Andrews 2000) are likely to be the main controls. For example the oceans contain 50 to 60 times as much CO_2 as the atmosphere (e.g. Henderson 2002; Barker *et al.* 2003). Quantifying, for example, the influence of the oceanic carbon cycle on atmospheric CO_2 must take into account the four species of dissolved inorganic carbon, i.e. dissolved CO_2 , undissociated carbonic acid, bicarbonate ion and carbonate ion, by assessing the factors controlling the concentration of these species, in particular, productivity, nutrient utilization, carbonate alkalinity/weathering fluxes, and pH (Henderson 2002). Even in the modern era, the role of soils and terrestrial

vegetation remains uncertain (House *et al.* 2003), this chapter will deal with the role of soils and vegetation in the section on lithological proxies below.

Palaeoproductivity. Biological productivity has the effect of removing carbon to depth from the surface oceans and the atmosphere. Biogenic barite has traditionally been used as a palaeoproductivity proxy because it sinks through the water column and enters the sedimentary record without much dissolution (e.g. Gingele & Dahmke 1994). New palaeoproductivity proxies for the Pleistocene include the difference in solubility of Th, Pa, and Be (e.g. Chase *et al.* 2002), and sedimentary U concentration (e.g. Francois *et al.* 1997). Isotopic fractionation of transition metal isotopes by biological activity appears to select for lighter values (Zhu *et al.* 2002). The $\delta^{66}\text{Zn}$ ratio of manganese crusts shows increasingly positive values with increasing bioproductivity (e.g. Maréchal *et al.* 2000). Fe isotopes (e.g. Beard *et al.* 1999) have proved so far ambiguous (e.g. Zhu *et al.* 2000). Zn/Si ratios in deep sea hexactinellid sponges increase with increasing rain of carbon as particulate matter from shallower oceanic levels, reflecting increased bioproductivity (Ellwood *et al.* 2004). Mo/Al ratios in black shales provide a palaeoproductivity proxy with a particularly long time range (e.g. Wilde *et al.* 2004). Rare earth element distribution coefficients in planktonic foraminifera also show promise as palaeoproductivity proxies (e.g. Haley *et al.* 2005). However, a recent intercomparison of palaeoproductivity proxies, particularly Ba, has shown some significant interpretation problems (Averyt & Paytan 2004).

Nutrient utilization. Water upwelling from the deep ocean carries with it nutrients and dissolved inorganic carbon. In areas of lower bioproductivity, some of this carbon is returned as CO₂ to the upper oceans and atmosphere. This has been particularly true in the Southern Ocean (e.g. Sigman 2000). The three major biolimiting nutrients are phosphate, nitrate and

silicate (Henderson 2002), with an increasing role being recognised for iron (e.g. Kumar *et al.* 1995; Noiri *et al.* 2005). Phosphate is not well-preserved in the sedimentary record, but its behaviour is tracked well by Cd, which substitutes into calcite (Henderson 2002). Cd/Ca ratio in foraminifera allows reconstruction of phosphate utilization (e.g. Rickaby & Elderfield 1999). A recent new proxy for phosphate utilization is Cd/P ratio in phytoplankton particulate matter (Elderfield & Rickaby 2000; Cullen & Sherrell 2005). Nitrate utilization can be tracked by $\delta^{15}\text{N}$, which increases in organic matter with increasing nitrogen uptake (Sigman *et al.* 1997; Robinson *et al.* 2005). This proxy has been demonstrated also to work for the Mesozoic (Jenkyns *et al.* 2001). For silicate, biogenic opal (e.g. diatom skeletons) preferentially incorporates the light isotope of silicon, ^{28}Si , resulting in increasingly positive $\delta^{30}\text{Si}$ values with increasing silicic acid utilization (De La Rocha *et al.* 1998; Reynolds *et al.* 2006). Rare earth element distribution coefficients in benthic foraminifera also show promise as nutrient utilization proxies (e.g. Haley *et al.* 2005)

Carbonate alkalinity/weathering fluxes. Carbonate alkalinity (the average concentration and distribution of carbonate and bicarbonate ions) determines the chemical speciation of carbon in the upper ocean and therefore the amount of CO_2 that can be drawn into the surface ocean from the atmosphere (Henderson 2002). Rivers constitute the major delivery system for alkalinity (generally bicarbonate ion) and weathering fluxes (most importantly Ca^{2+}) to the oceans (e.g. Raymond & Cole 2003). Foraminiferal Ba/Ca ratios have been used as a proxy for past alkalinity distributions (Lea 1993) although the relationship between the Ba and alkalinity cycles is complex (Rubin *et al.* 2003). $^{87}\text{Sr}/^{86}\text{Sr}$ in carbonate (e.g. Dessert *et al.* 2001) and $^{187}\text{Os}/^{186}\text{Os}$ ratios in clastic sediments (e.g. Cohen *et al.* 2004) measure weathering flux from continental sources, but mainly pick up periods of high-rate continental weathering (e.g. Ravizza *et al.* 2001). From marine sediments, the $^{87}\text{Sr}/^{86}\text{Sr}$ curve is particularly useful

as a weathering proxy, and, with biostratigraphical context, a powerful estimator of numerical age (Francois & Walker 1992; McArthur *et al.* 2001). Ge/Si ratios in diatom silica (opal) also show some promise as continental weathering proxies (e.g. Munhoven & Francois 1996; Jones *et al.* 2002), although temperature-dependent effects in the oceans need to be accounted for (e.g. Hammond *et al.* 2004). Hf and Nd isotope ratio time trajectories in manganese crusts show relative deviations during glacial periods that have been interpreted as a mechanical weathering proxy (van de Flierdt *et al.* 2002). Overall, good proxies for global oceanic alkalinity and weathering are lacking (Henderson 2002).

pH. Oceanic pH provides insights into how the carbonate chemistry of the oceans, including depth to lysocline, has changed through time (e.g. Sanyal *et al.* 1995). With increasing oceanic pH, concentration of the chemical species $B(OH)_3$ reduces and $B(OH)_4^-$ increases (e.g. Vengosh *et al.* 1991; Zeebe 2005). Isotopic fractionation between these two species results in an up to ~20 ‰ isotopic shift in $\delta^{11}B$ which is recorded by foraminiferal carbonate (e.g. Vengosh *et al.* 1991; Zeebe 2005). This proxy has been applied to other carbonate reservoirs (e.g. Honisch *et al.* 2004) and has been applied over a long time span, even back to the Silurian (e.g. Joachimski *et al.* 2005), although reliable estimates only extend back to the Palaeocene (Pearson & Palmer 2000). However, several parameters of the boron/borate/boric acid system need to be better known before long-term estimates can be used with confidence (e.g. Pagani *et al.* 2005a). Palaeo-pH over Phanerozoic timescales has also been estimated by Royer *et al.* (2004) based on the calcium-ion concentration of seawater and modelled atmospheric CO_2 concentrations.

Atmospheric CO_2 . Atmospheric CO_2 concentrations that pre-date the ~800 ky direct record provided by ice cores (e.g. Siegenthaler *et al.* 2005) are of extreme importance for palaeoenvironmental reconstruction. Of oceanic proxies, $\delta^{13}C$ of organic materials has been

particularly successful, and has recently been refined with studies that focus on $\delta^{13}\text{C}$ derived from a single group of organisms, for example, alkenones (e.g. Pagani 2002), rather than total marine organic material (Henderson 2002). The alkenones record extends back to the Eocene–Oligocene boundary (e.g. Pagani *et al.* 2005b). Atmospheric CO_2 from $\delta^{13}\text{C}$ of other organic materials, including terrestrial materials such as fossil wood (Hesselbo *et al.* 2000) has been demonstrated back to the Jurassic. $\delta^{13}\text{C}$ of carbonate, including both marine (Buggisch 1991), freshwater (Yemane & Kelts 1996) and pedogenic (Royer *et al.* 2001), has also been used as a proxy for atmospheric and oceanic carbon source and for rates of carbon burial (e.g. Schouten *et al.* 2000; Strauss & Peters-Kottig 2003). Oceanic pH has also been used as a proxy for atmospheric CO_2 concentration, although this requires assumptions to be made about past dissolved inorganic carbon, Ca flux to the oceans and alkalinity (Henderson 2002; Pearson & Palmer 2002). By this method, Pearson & Palmer (2000) and Demicco *et al.* (2003) have reconstructed atmospheric CO_2 concentration back to 60 m.y. ago. Recently, the technique has been demonstrated for Neoproterozoic carbonates (Kasemann *et al.* 2005).

Lithological proxies.

Sedimentary rocks are the most venerable palaeoclimate proxies in terms of when they were first used (e.g. Lyell 1830; Croll 1867) and they offer palaeoclimatic information from some of the earliest periods of Earth history (e.g. Moore *et al.* 2001). Köppen & Wegener (1924) made the earliest attempt at quantitative use of sedimentary facies for palaeoclimate reconstruction. The 1960's saw a period of critical appraisal of the value of sedimentary facies as palaeoclimate indicators (e.g. Nairn 1961; Van Houten 1961; Smith 1963). Since the early 1970's, the focus has been increasingly on chemical or isotopic components of the facies rather than interpretation of the facies themselves, most particularly for carbonate sediments, although they continue to be important because of the time range over which they

are useful. Sellwood *et al.* (1993) reviewed the sedimentary facies that are key palaeoclimatic indicators.

Carbonate rocks. Marine carbonate facies are some of the most important palaeoclimate indicators. Sellwood *et al.* (1993) distinguished between temperature-related facies changes, i.e. related to cool or warm water, and carbonate structures, such as reefs. Modern-day warm-water shelves carbonates form between 30° S and 30° N and include reef-building corals, codiacian algae, and ooids, aggregates and pellets. Temperate water carbonates, which extended as far as 50° N during the Cretaceous, are characterized by benthic foraminiferans, molluscs, bryozoans, barnacles and calcareous red algae, along with ahermatypic (non-reef-forming) corals. The latitudinal distribution of shelf carbonates has varied through the Phanerozoic, although total area appears to have been relatively invariant (Kiessling *et al.* 2003). Reefs and carbonate build-ups are interpreted as indicating the presence of former warm seaways (Sellwood *et al.* 1993), although the communities of organisms responsible have varied through the Phanerozoic (e.g. Braga & Aguirre 2001; Edinger *et al.* 2002; Leinfelder *et al.* 2005). Sellwood *et al.* (1993) gave the poleward limit of reefs as greater than 30° away from the equator. In the interim, it has been recognised that deep-water, cold-temperate reefs can form at latitudes of up to 65° N (e.g. Freiwald *et al.* 1997; Freiwald & Wilson 1998).

Evaporites. Evaporites are the second major facies treated by Sellwood *et al.* (1993) who defined them as forming anywhere on the Earth's surface where evaporation exceeds precipitation and/or rate of water inflow. Ziegler (2003) added that this must occur under the descending limbs of Hadley cells, today centred between 10° and 40° north or south. However, it is worth noting that hypersaline brines form in Antarctic dry valleys today

(Doran *et al.* 2003), at high latitudes and very low temperatures. The earliest evaporites are Archaean in age (e.g. Zharkov 2005).

Storm deposits. Storm deposits are another important palaeoclimatic indicator with high preservation potential (Sellwood *et al.* 1993). Tempestites (e.g. Myrow & Southard 1996; Mohseni & Al-Aasm 2004), other storm deposits (e.g. Hentschke & Milkert 1996; Chaudhuri 2005), and windblown dust (e.g. Dodonov & Baiguzina 1995; Clemens 1998; Qin *et al.* 2005) can indicate changing storminess, and, with modelling, even show the geographical distribution of storm belts through time (Agustsdottir *et al.* 1999). Palaeowind directions are also derivable in some cases (e.g. Allen 1996; Pochat *et al.* 2005).

Glacial sediments. Glacial sediments are archetypal indicators of cold palaeoclimates (Agassiz 1828; Darwin 1842). The direct products of glaciation, deposited in the immediate glacial environment, are generally unsorted diamictons (or, when lithified, diamictite), consisting of rounded granule to large boulder clasts, matrix supported in clay-silt. These are often referred to as boulder clay or till/tillite (Sellwood *et al.* 1993) and are in many cases intercalated with clast-supported conglomerate lenses. In the older record tillites/glacial diamictites need careful interpretation to avoid confusion with diamictites resulting from meteorite impact (e.g. Mory *et al.* 2000), subaerial debris flows (e.g. Blair 1999), or non-glacially derived sub-aqueous gravity flows (e.g. Eyles 1993), but reliable records extend back to the Archaean (e.g. Young *et al.* 1998). Fluvial systems driven by glacial meltwater produce outwash facies (e.g. Visser 1997; O'Brien *et al.* 1998), which can feed into glaciomarine systems including submarine fans. Glaciers feeding directly into the sea, with or without intervening ice shelves also produce distinctive glaciomarine deposits, commonly with dropstones from floating ice or icebergs (e.g. Bennett & Doyle 1996; Visser 1997; Price

1999). More distally, marine sequences can preserve layers of ice-rafted debris, shed from more far-travelled floating ice (e.g. Keany *et al.* 1976; Isbell *et al.* 2001). For periglacial environments, evidence pre-Late Cenozoic is sparse, but there are records from the Neoproterozoic (Williams 1998), and Mesozoic (Constantine *et al.* 1998).

Glendonites. The questions raised by Sellwood *et al.* (1993) over the palaeoclimatic significance of glendonite carbonate nodules have largely been resolved and their status as indicators of cold subaqueous environments confirmed (e.g. Swainson & Hammond 2001). Glendonite carbonate nodules are distinctive pseudomorphs after ikaite, a low-temperature, hydrous form of calcite that may be associated with methane hydrates (Greinert & Derkachev 2004). The precursor ikaite forms under near-freezing conditions at moderately elevated hydrostatic pressure (Swainson & Hammond 2001). Glendonites are characteristic of glaciomarine sediments throughout the Neoproterozoic and Phanerozoic (e.g. Price 1999; McLachlan *et al.* 2001; Alley & Frakes 2003; James *et al.* 2005).

Sub-glacial volcanic deposits. The deposits of sub-glacial volcanic eruptions, which include lavas overlying or intercalated with glacial diamicton, and intercalated or laterally associated with hyaloclastite breccia deltas (e.g. Smellie *et al.* 2006) indicate the presence of former ice caps and ice sheets (Smellie 2000). These have been described from Antarctica (e.g. Smellie & Skilling 1994; Smellie *et al.* 2006) and Iceland (e.g. Schopka *et al.* 2006). They constitute a new quantitative proxy for the presence or absence of ice caps (e.g. Schopka *et al.* 2006) and palaeo ice sheet thickness (e.g. Smellie *et al.* 2006), complementing nunatak and trimline studies (e.g. Rae *et al.* 2004; Paus *et al.* 2006). Zeolite compositions from subglacially erupted lavas can also be used to distinguish between marine and freshwater/glacial eruptive environments (Johnson & Smellie in press).

Coal, lignite and peat. Coal, lignite and peat deposits are well-recognised indicators of terrestrial humidity (Sellwood *et al.* 1993), i.e. areas where precipitation exceeds evaporation (e.g. Parrish *et al.* 1982; Hallam 1985). A combination of tectonic (e.g. McCabe 1991), biological (e.g. McCabe & Shanley 1992), eustatic (e.g. Staub 2002) and climatic (e.g. McCabe & Parrish 1992) factors is required for the preservation of coal- or lignite-forming mires. There are two main types (Moore 1995): rain-fed mires, which form in humid maritime climates at high-mid latitudes today; and flow-fed mires, which currently form at low latitudes. The latitudes of coal formation may, however, have been different in the past (Crowley & North 1991), although, for the Palaeozoic, for example, both low latitude flow-fed mires (e.g. Edwards 1998) and high latitude, rain- and flow-fed mires (e.g. Michaelsen & Henderson 2000) have been recognised.

Fusain. Wildfires (Harris 1981; Sellwood *et al.* 1993; Glasspool 2000) are represented in the geological record by the presence of fossil charcoal, predominantly fusain (Scott 1989; Scott

et al. 2000), in sedimentary rocks, and are indicators of a rapid growing season punctuated by periods of drought (Finkelstein *et al.* 2005) terminated with thunderstorms (Edwards 1984). The lack of a more complete treatment of fire through the Phanerozoic, as highlighted by Sellwood *et al.* (1993), has since been addressed by the excellent review by Scott (2000). Late Carboniferous wildfires in tropical lowland peats (Scott 2000) may reflect elevated levels of atmospheric O₂ at that time (Berner *et al.* 2003).

Palaeosols. Palaeosols provide important palaeoclimatic information, in particular palaeoprecipitation (e.g. Sellwood *et al.* 1993; Retallack 2001). Palaeosols represent intervals or areas of low or no sedimentation. Their mineral and chemical composition reflects the interaction between their source terrigenous clastic sediments and the processes of weathering, which can be physical, chemical and biological. Sellwood *et al.* (1993) identified five different, major, climatically significant palaeosol types. Laterites and bauxites they linked with humid tropical climate with a long dry season. They distinguished between pedogenic and groundwater laterites, the latter of which occur in low land and coastal settings, deriving their iron from groundwater. Sellwood *et al.* (1993) pointed out that although most bauxites occur with laterites, some may form in depressions in karst. Price *et al.* (1997a) modelled the occurrence of bauxites and showed that they require mean annual temperatures greater than 22 C° or 23 C° and high precipitation for at least 6 months of the year. Clay-rich vertisols, with characteristic internal features and micro structures, were linked by Sellwood *et al.* (1993) with exclusively warm temperate to tropical climates with four to eight dry months each year, in some cases associated with playas in otherwise arid or semi-arid regions. However, recent data suggests that they can also form in humid climates (e.g. Nordt *et al.* 2004). Chemical analyses of vertisol profiles provide palaeoprecipitation proxies (e.g. Stiles *et al.* 2001; Driese *et al.* 2005).

Calcretes and dolocretes, calcium or magnesium carbonate-rich soils respectively, are also a marker for arid or semi-arid areas, both cold (e.g. Lauriol & Clark 1999; Rowe & Maher 2000) and warm (e.g. Jimenez-Espinosa & Jimenez-Millan 2003), and are often found in association with zones of gypsum precipitation (gypcretes). Alonso-Zarza (2003) has pointed out that there is a continuum between calcretes and palustrine (swamp or marsh) carbonates, and that the latter can form rapidly and under much more humid conditions than calcretes, requiring care when making palaeoenvironmental interpretations. This is particularly important when determining palaeoprecipitation via transfer functions derived from the depth to the nodular, pedogenic carbonate horizon (Retallack 2005).

Cementation of soil, or the underlying saprolite, by secondary silica forms silcretes (e.g. Webb & Golding 1998). Since the work by Sellwood *et al.* (1993) silcretes have been identified forming today in central Australia. Although Webb & Golding (1998) suggested an association with long-term aridity and seasonally high evaporation, allowing silica to precipitate from highly saline ground waters, they argue that silcretes still have no clear climatic association. Recent work by Ulliyott & Nash (2006) and Alexandre *et al.* (2004) suggested that silcretes may also form in cool and wet climates.

The final category of palaeosol identified by Sellwood *et al.* (1993) are podzols (spodosols of Mack *et al.* (1993)). These form in cool wet environments where humus accumulates, promoting acidic conditions (e.g. Bonifacio *et al.* 2006). Sellwood *et al.* (1993) pointed out, however, that podzols may form under warmer climatic regimes in well-drained siliceous substrates (e.g. Van Niekerk *et al.* 1999).

Mack & James (1994) proposed a simpler scheme using the classification scheme of Mack *et al.* (1993), and based on a theoretical model of four palaeoclimatic palaeosol associations. They defined: 1) a wet equatorial zone, characterized by oxisols (laterites and bauxites), with secondary clay-rich argillisols and humic histosols (including coal); 2) a subtropical dry zone containing calcisols (calcretes and dolocrete) with subsidiary gypsisols (gypcretes) and vertisols; 3) a moist mid-latitude zone with argillisols, spodosols (rich in organic matter and iron/aluminium oxides; common in coniferous forests today and equivalent to podzols), gleysols (waterlogged soils) and histosols; and 4) a polar zone with gleysols and protosols (regolithic soils).

Clay mineralogy. Of parallel importance to soils, the distribution in oceanic sediments of clay minerals will reflect weathering processes and soil development in adjacent continents, providing an indirect record of terrestrial climate (e.g. Chamley 1989; Sellwood *et al.* 1993; Chamley 1998; Thiry 2000). A recent review by Thiry (2000) highlights the difficulties in extracting a climate signal, in particular the time lag introduced by the time it takes for soils to form, and biases in the signal that result from, for example, the longevity of kaolinite deposits (indicative of tropical wet climates). Differential settling of clay particles between proximal and distal settings may also introduce a bias in the signal (e.g. Simkevicius *et al.* 2003). Nevertheless, for example, the declining proportions of crystallized smectite and chlorite, and increasing illite in marine sediments have been used to trace the transition from humid to subpolar and polar conditions in the high southern latitudes (e.g. Ehrmann *et al.* 2005), and characteristic proportions of kaolinite, smectite, chlorite and illite indicate warm equable Late Triassic and humid Early Jurassic climates at mid-northern latitudes (Ahlberg *et al.* 2003). In terrestrial settings, Ballantyne (1994) suggested that low gibbsite concentration

of soils in formerly glaciated is a proxy for glacial erosion, which has recently been confirmed by cosmogenic isotope dating (Ballantyne *et al.* 2006).

Sortable silt. This is a relatively new proxy for palaeocurrent speed and rates of deep-water flow (e.g. McCave *et al.* 1995). It uses the 10–63 µm fraction of marine sediment because this fraction is most sensitive to sorting in response to hydrodynamic processes and its properties can be used to infer current speed (McCave *et al.* 1995). The proxy works well for the Cenozoic (e.g. Pfuhl & McCave 2005), where contourite drifts have been identified from sea-floor topography and geographic association, but should, in theory be applicable to the entire interrogateable sea-floor record.

Aeolianites. The final class of palaeoclimate indicator facies treated by Sellwood *et al.* (1993) was aeolianites (wind-blown sediments). As Sellwood *et al.* (1993) pointed out, the simple association between these deposits and hot arid climates has been seriously questioned and their main palaeoclimatic significance is as palaeowind indicators (e.g. Allen 1993; Adams 2003; Segalen *et al.* 2004; Le Guern & Davaud 2005). Fine-grained aeolianites, such as loess deposits and dust (for review see Kohfeld & Harrison 2001) are important in that they provide the only proxy for atmosphere palaeocirculation (e.g. Henderson 2002). Palaeocirculation information can be derived from interpretation of temporal variations in source region from the chemical and mineralogical composition of dust (e.g. Nakai *et al.* 1993). In addition to providing palaeowind and palaeocirculation data (e.g. Sun *et al.* 2004), loess deposits also indicate cold, arid periglacial climates (Lagroix & Banerjee 2002; Sun & An 2005).

Palaeontological proxies.

The empirical use of fossils and fossil assemblages as palaeoclimate indicators can be traced back at least as far as Lyell (1830) and their value was even recognised in classical times (Imbrie & Newell 1964). Quantifying former climates using biological data is based on three key assumptions (Woodcock 1992): 1) that climatic factors limit the occurrences or associations of taxa; 2) that taxa have characteristics (such as leaf shape, or stomatal density) that respond to climate; and 3) that climate causes variance in relative frequency of taxa. It is useful to subdivide studies that use biological data into taxonomic and morphological types (Wing & Greenwood 1993) based on whether or not they address assumption "2" above. Quantitative, proxy-based palaeoclimate reconstruction based on fossil data began for terrestrial environments in the 1950's, based on pollen analysis (Faegri & Iversen 1950), and for the marine environment in the early 1970's with the development of micropalaeontology-based transfer functions by Imbrie & Kipp (1971). Quantitative techniques relate biological parameters to changes in environment variables such as temperature, pH or salinity. These parameters can either be measured of individuals (morphological techniques), or of a faunal assemblage (taxonomic techniques).

Taxonomic methods. For the terrestrial realm the primary taxonomic methods for palaeoclimate reconstruction use assemblages of pollen (e.g. Birks 1981) or other palynomorphs (e.g. Stanley 1970; Loboziak *et al.* 1989). Pollen analysis, using climatic amplitude (e.g. Jimenez-Moreno *et al.* 2005) and mutual climatic range (e.g. Klotz *et al.* 2006) techniques, provides data on palaeotemperature and palaeoprecipitation (e.g. Fauquette & Bertini 2003) as far back as the Eocene (e.g. Harrington 2004). Taxonomic methods based on plant macrofossils (e.g. Baghai & Jorstad 1995), vertebrates (Markwick 1998; Friedman *et al.* 2003), molluscs (e.g. Moine & Rousseau 2002) and arthropods (e.g. Pilny *et al.* 1987; Williams & Eyles 1995) provide palaeotemperature data, and have been demonstrated for the

Cretaceous. Taxon analysis of chironomid midges provides lake palaeosalinity data back to the Late Pleistocene (e.g. Walker 1991) and insect assemblages have been proposed as a marker for low salinity for the Early Cretaceous (Coram & Jarzembowski 2002). More sophisticated palaeoclimatic treatments of terrestrial faunal assemblages use herbivorous mammals and their feeding habits. These include ecological diversity spectra (e.g. Andrews *et al.* 1979; Fernandez & Pelaez-Campomanes 2005), which quantify a range of parameters including temperature, humidity/aridity and annual precipitation potentially back to the Miocene, and cenograms (e.g. Legendre 1986; Rodriguez 1999; Legendre *et al.* 2005), which provide information on temperature and humidity/aridity back to the Eocene. The nearest living relative method (e.g. MacGinitie 1941), which blurs the boundary between taxonomic and morphological techniques, makes palaeoclimatic inferences based on fossil groups and assemblages by using the ranges and climatic responses of their modern descendants or sister groups (Wing & Greenwood 1993), and is particularly good for the Cenozoic. The nearest living relative method provides palaeotemperature data from both plants (Poole *et al.* 2005; Wang *et al.* 2005), back to the Jurassic, and animals (e.g. Hutchison 1982; Moe & Smith 2005), back to the Palaeocene.

For the oceanic realm, taxonomic methods of deriving palaeoclimate parameters, such as temperature and salinity, are most commonly derived via transfer functions, which were first developed by Imbrie & Kipp (1971). These can be calculated from marine (e.g. Sejrup *et al.* 2004) and lacustrine (e.g. Hausmann & Kienast 2006) microfossil assemblages, and are commonly applied to Pliocene or younger assemblages (e.g. Andersson 1997), although for palaeoproductivity they extend back to the Palaeocene (Siesser 1995). Transfer functions have evolved since the principal components regression work of Imbrie & Kipp (1971) and now include weighted averaging and weighted averaging least squares methods (Sejrup *et al.*

2004; Hausmann & Kienast 2006), as well as artificial neural networks (Malmgren & Nordlund 1997; Peyron & De Vernal 2001) and maximum likelihood and Bayesian techniques (e.g. Robertson *et al.* 1999). The applications of transfer function techniques have broadened away from the planktic foraminifera and radiolaria of Imbrie & Kipp (1971). They now include, in the oceans, benthic foraminifera (e.g. Sejrup *et al.* 2004), diatoms (e.g. Zielinski *et al.* 1998), ostracods (e.g. Brouwers *et al.* 1991), dinoflagellates (e.g. Peyron & De Vernal 2001), and in lakes, chironomid midges (e.g. Korhola *et al.* 2001) and diatoms (e.g. Roberts & McMinn 1999). Parameters estimated include, oceanic palaeoproductivity (e.g. Siesser 1995), ocean bottom water palaeotemperatures (e.g. Brouwers *et al.* 1991), palaeosalinity (e.g. Sejrup *et al.* 2004), sea-surface palaeotemperatures (e.g. Andersson 1997; Malmgren & Nordlund 1997; Zielinski *et al.* 1998), and seasonal extent of former sea-ice cover (e.g. Peyron & De Vernal 2001). In lakes, palaeonutrient utilization (e.g. Hausmann & Kienast 2006), surface palaeotemperature (Korhola *et al.* 2001) and palaeosalinity (Roberts & McMinn 1999) have been estimated. Non-transfer-function-based techniques on diatoms have used faunal assemblages (Gersonde & Zielinski 2000) and modern analogue techniques (Crosta *et al.* 1998) as proxies for sea-ice extent back to the Pleistocene. Similarity maximum modern-analogue techniques on foraminiferal assemblages from core tops have been also been used as a proxy for sea-ice extent (Sarnthein *et al.* 2003), again back to the Late Pleistocene. For the older, pre-Cenozoic, record, other non-transfer-function-based methods use faunal assemblages of marine molluscan macrofossils to estimate palaeotemperature (e.g. Kafanov & Volvenko 1997) and palynomorphs, such as acritarchs, for palaeoproductivity (e.g. Vecoli & Le Herisse 2004) with applications back to Ordovician times.

Morphological methods. Morphological methods generally centre around leaf margin analysis, beginning with the work of Bailey & Sinnott (1916) and culminating in the work of Wolfe (1993) on the climate-leaf analysis multivariate program, or CLAMP. This is widely used as a palaeotemperature proxy (e.g. Gregory-Wodzicki 2000; Kennedy 2003; Spicer *et al.* 2005) and has been extended back to the Permian (Glasspool *et al.* 2004). The width of growth rings in fossil wood is also an important morphological technique for palaeotemperature and palaeoprecipitation (e.g. Fritts 1976; Creber & Chaloner 1985; Francis 1986; Poole *et al.* 2005). Although results need to be assessed in terms of taxonomic analysis (Brison *et al.* 2001) the technique has been demonstrated to be effective from Carboniferous times to Recent (e.g. Falcon-Lang 1999; Morgans 1999; Francis & Poole 2002). Measuring maximum latewood density of tree rings gives the ability to measure hemispheric variations in temperature (Briffa *et al.* 2004), although this has only been applied to historically recent material. The density of stomata in leaf cuticle has been used as a proxy for palaeo-CO₂ concentration (Woodward 1987; Woodward & Bazzaz 1988). This has been shown to be effective back to the Carboniferous (e.g. Otto-Bliesner & Upchurch 1997; McElwain *et al.* 1999; Retallack 2002; Thorn & DeConto 2006). For the oceans, palaeo-carbonate ion concentration, which determines the water depth at which all calcite has dissolved from the sediment or lysocline (e.g. Ridgwell 2005), has been determined by estimating the degree of dissolution of individual foraminifera of a particular size (e.g. Lohmann 1995). However, the situation is more complicated in glacial periods (Broecker & Clark 2001) and dissolution effects at the sediment-water interface cannot be discounted (e.g. de Villiers 2005).

Modelling

Palaeoclimate modelling falls into two main categories: The main, and most significant, category is by the generation and interpretation of climate analogues. These can take the form

of conceptual models (Parrish 1993), or be carried out using numerical, computer-based global circulation models (GCMs) or Earth models of intermediate complexity (EMICs) (for a recent review see McGuffie & Henderson-Sellers 2001). These attempt to generate analogues for climate over a range of spatial and temporal scales up to and including the complete globe. The second category is what are called geochemical (Royer *et al.* 2001), mass-balance (e.g. Beerling 1999), or trend models (e.g. Richards 1998). These use palaeoclimate proxies or model outputs to reconstruct the evolution of a particular atmospheric or oceanic chemical species, such as oxygen or carbon dioxide, over a range of geological time, up to and including the entire geological record.

Climate analogue modelling. It was recognised in the early 19th Century that former global climates differed from today's (Agassiz 1828; Darwin 1842). Following the work of Tyndall (1861) and Arrhenius (1896) a key role for CO₂ was suspected from the beginning (Chamberlin 1897, 1899). The first empirical palaeoclimatic reconstructions for the Phanerozoic were carried out by Köppen & Wegener (1924), who used a simple zonal climate model and the distribution of palaeoclimatic indicator facies such as coal or glacial sediments in a continental drift framework to refine the fit of the continents. These applications of palaeoclimate studies based on simple zonal schemes continued into the 1960's (e.g. Nairn 1961). It was only with the advent of palaeomagnetism in the 1950's and plate tectonics in the 1960's (summarized in Irving 2005) that continental palaeolatitudes could be independently fixed. This freed the geological record to provide ground truth for increasingly empirical, conceptual (Parrish 1993) and numerical (Schneider & Dickinson 1974) palaeoclimate models. The earliest numerical models were relatively simple analogues of Late Pleistocene glacial climates (Alyea 1972; Gates 1974). These quickly moved to modelling of palaeoclimates with global circulation models (GCMs) initially of Late

Pleistocene glacial climates (Gates 1976), and at the end of the decade the CLIMAP project saw the first attempts to provide model boundary conditions and to evaluate climate models with proxy data (CLIMAP 1981). From the 1980's, palaeoclimate modelling was extended first to the Mesozoic (Barron *et al.* 1981; Kutzbach & Gallimore 1989; Chandler *et al.* 1992; Valdes & Sellwood 1992) and then the Palaeozoic (Crowley & Baum 1992; Kutzbach & Ziegler 1993; Otto-Bliesner 1995; Gibbs *et al.* 1997).

Although the first coupled ocean-atmosphere GCM was realised in the 1960's (Manabe & Bryan 1969), it was not until the 1990's that these were applied to palaeoclimate studies (Stocker *et al.* 1992; Cubasch *et al.* 1997; Schiller *et al.* 1997) initially for the early Holocene and the last interglacial (Eemian) (Texier *et al.* 1997; Montoya *et al.* 1998). Since 2000, these GCMs have been increasingly applied to the older Cenozoic and Mesozoic (Otto-Bliesner *et al.* 2002; Haywood & Valdes 2004; Haywood *et al.* 2004) (Fig. 3). In the 1990's, multiple or ensemble runs (e.g. Cubasch *et al.* 1994; Hoar *et al.* 2004) and asynchronous coupling (e.g. Liu *et al.* 1999; Dutton & Barron 2000), where the outputs of one EMIC or GCM is used to force climate parameters in a second EMIC or GCM, were used to model climate behaviour.

Since the late 90's, the emphasis has been on model "validation" or "evaluation" (the second term is preferred because it does not confer any sense of "approval" (Kohfeld & Harrison 2000)). This uses multi-proxy data sets (e.g. Price *et al.* 1995; Price *et al.* 1997b; Sellwood & Valdes 1997; Sellwood *et al.* 2000; Felzer & Thompson 2001; Haywood *et al.* 2004) to minimize the uncertainties inherent in numerical models (Dickinson 1989). A second trend is that models have become more sophisticated with coupling of ice-sheet (e.g. Ridley *et al.*

2005) and biome (e.g. Haywood *et al.* 2002a; Snyder *et al.* 2004; Brentnall *et al.* 2005) models to the coupled atmosphere–ocean GCMs.

Geochemical or mass-balance modelling. Geochemical or mass balance modelling uses calculations of the transfer of material between the lithosphere, hydrosphere, and atmosphere (Li 1972; Garrels & Lerman 1984; Lasaga & Berner 1998), to assess the long term evolution of atmospheric and oceanic composition. It is by necessity generalized and the geographical patchiness and temporal incompleteness of the terrestrial and particularly the pre-Jurassic marine record must qualify its interpretation. Modelling of the long-term carbon cycle is one of the keys to understanding Phanerozoic climate change (e.g. Berner & Barron 1984; Gifford 1994; Berner 1998; Berner & Kothavala 2001). Changes in the carbon cycle can be measured directly, with varying degrees of precision, from organic carbon in sedimentary rocks or carbonate in limestone (this latter may take the form of sedimentary rock or carbonate cements in, for example, soils) (e.g. Berner 1998).

Of secondary importance is the long-term oxygen cycle (e.g. Berner & Canfield 1989; Berner 1999; Berner *et al.* 2000; Berner 2001). Some authors argue that oxygen concentration changed substantially during the Phanerozoic, possibly peaking in the late Carboniferous (e.g. Berner & Canfield 1989; Beerling *et al.* 2002) although there are strong counterarguments that it was relatively unvarying (e.g. Lenton & Watson 2000). The long-term oxygen cycle is indirectly modelled from the carbon cycle and the sulphur cycle (e.g. Berner 2005) (the sulphur cycle is measured directly from evaporitic rocks and iron pyrites in sedimentary rocks (e.g. Francois & Gerard 1986; Railsback 1992)).

Three other cycles that are important for understanding atmospheric evolution are the nitrogen cycle (e.g. Falkowski 1997), the phosphorus cycle (Guidry & Mackenzie 2000; Lenton 2001) and the silicon cycle in the oceans (Ragueneau *et al.* 2000; Matsumoto *et al.* 2002). All three are implicated in the drawdown of carbon and its incorporation in marine

sediments (e.g. Falkowski 1997; Lenton 2001; Matsumoto *et al.* 2002). One of the most spectacular successes of mass-balance modelling is in providing support for the extreme atmospheric changes associated with the Neoproterozoic 'Snowball Earth' model (Hoffman *et al.* 1998). Recent mass-balance models are sophisticated, incorporating multiple elements and both major oceanic and atmospheric cycles (e.g. Hansen & Wallmann 2003; Bergman *et al.* 2004)

Phanerozoic climate modes

Frakes *et al.* (1992) looked at climates back to 600 million years ago and identified what they termed "climate modes" with alternating episodes of cool or warm climates (Table 1; Fig. 1a). They defined "cool modes" as "times of global refrigeration during which ice was present on earth". Their "warm modes" they defined as times "when climates were globally warm, as indicated by the abundance of evaporites, geochemical data, faunal distributions, etc, and with little or no polar ice". Frakes *et al.* (1992) argued that warm-cool climate mode pairs spanned intervals of approximately 150 million years: about half a galactic year. Superimposed upon these climate modes were brief intervals of contrasting climates (Fig. 1b, Table 1), which they did not consider sufficiently long to merit climate modes of their own. These are important, however, and the geological background to some will be dealt with in slightly more detail in this chapter, as they give insights into the extremes of climate change. All numerical ages quoted for chronostratigraphic units and boundaries are according to Gradstein *et al.* (2004).

Frakes *et al.* (1992) proposed eight climate modes for the Phanerozoic (Table 1; Fig. 1a). They grouped these into four warm: Cambrian to Middle Ordovician; middle Silurian to

Early Carboniferous; latest Permian to early Jurassic; late Cretaceous to early Palaeogene, inclusive; and four cool: late Ordovician to early Silurian; early Carboniferous to late Permian; late Jurassic to early Cretaceous; mid-Palaeogene to Recent, inclusive. These are treated alternately in time-sequential order below.

Climate modes

Earliest Cambrian to Middle Ordovician warm mode. Frakes *et al.* (1992) placed the beginning of this warm mode at the end of the latest Neoproterozoic glaciation, which they assumed to be at the Precambrian–Cambrian boundary. Since then, the latest Precambrian, Ediacaran Period has been added (Knoll *et al.* 2006) post-dating the last Neoproterozoic glacial deposits, this period starting between 635 and 600 million years ago (Knoll *et al.* 2006). The Ediacaran terminates at the earliest Cambrian, now 540 million years ago (Gradstein *et al.* 2004), 30 million years later than the dates used by Frakes *et al.* (1992) from the Decade of North American Geology (DNAG) timescale of Palmer (1983). Ediacaran stratigraphic (Knoll *et al.* 2006) and stable isotope (Le Guerroue *et al.* 2006) data suggest that it post-dates major Neoproterozoic glaciation and should be included in the warm mode. Frakes *et al.* (1992) suggested that this warm mode ended prior to the beginning of the Ordovician Katian stage which is now dated at about 461 million years ago, an interpretation endorsed by Page *et al.* (in press) in this volume.

This warm mode is marked by substantial deposition of evaporites and moderately high rates of carbonate formation by Phanerozoic standards (Frakes *et al.* 1992). One notable biogeochemical event during this warm mode is the deposition of substantial phosphorite deposits around the time of the Ediacaran–Cambrian boundary (Cook 1992). This accumulation of phosphorite differs from other Phanerozoic accumulations in that it does not

appear to be associated with oceanic upwelling and may be related to a change in ocean chemistry associated with a rise in atmospheric oxygen and the formation of the first hard skeletons.

At least three brief intervals of probable cooler climate punctuate this warm mode. A brief glacial episode is recognised in the Ediacaran at about 580 Ma (c1, Fig. 1b) from $\delta^{13}\text{C}$ values in marine carbonates (Le Guerroue *et al.* 2006). Another short-lived episode of cool, possibly glacial climate is identified in the Steptoean of the Late Cambrian (e.g. Saltzman *et al.* 2000), at ~494 Ma (c2, Fig. 1b), again from a positive excursion of up to 5 ‰ in brachiopod carbonate (Cowan *et al.* 2005) immediately followed by a major global sea-level drop (Glumac & Spivak-Bimdorf 2002). Although Frakes *et al.* (1992) observed that unequivocal Cambrian or Early Ordovician glacial sediments are rare, Early Ordovician, Tremadocian, ~480 Ma (c3, Fig. 1b), glacial deposits have been identified in the Meguma Zone of Nova Scotia (Schenk 1995).

Late Ordovician to early Silurian cool mode. Suggestions for the duration of this cool mode have varied since Frakes *et al.* (1992). From a review of Phanerozoic glacial deposits, Evans (2003) estimated that it lasted of the order of 17 million years, from the latest Ordovician (Hirnantian), ~445 Ma until the Llandovery–Wenlock boundary at the end of the Early Silurian, 428 Ma. Royer *et al.* (2004), using pH-corrected Phanerozoic $\delta^{18}\text{O}$ trends, estimate the duration at less than 10 million years and Brenchley *et al.* (2003) estimate that widespread ice sheets lasted less than 1 m.y. However, Page *et al.* (in press) in this volume presents evidence that support the assessment of Frakes *et al.* (1992) with a beginning at ~461 Ma.

Frakes *et al.* (1992) documented widespread glacial sediments over a large area of the African and South American parts of Gondwana, and in terranes probably sourced from there. Page *et al.* (in press) in this volume present sea-level data and lithological evidence from West Africa, Canada and France for two middle Caradoc and one early Ashgill cool or glacial episodes in the Late Ordovician (Katian stage) (c4-c6, Table 1 and Fig. 1b) and there is some evidence in Baltica for cool phases at these times (Kaljo *et al.* 2003). These do not appear to represent glaciation on the scale of Hirnantian and Early Silurian episodes and may indicate a gradual transition from the preceding warm mode. Data from Baltica restricts full glaciation to the latest Ordovician and Early Silurian (Jeppsson & Calner 2003; Kaljo *et al.* 2003; Legrand 2003). Four glacial episodes are recognised (Kaljo *et al.* 2003), including the latest Ordovician Hirnantian episode, at ~445 Ma (c7, Fig. 1b), and Silurian glacial episodes in the Aeronian, ~439 Ma (c8, Fig. 1b), Telychian, ~436 Ma (c9, Fig. 1b), and earliest Wenlock (Sheinwoodian), ~428 Ma (c10, Fig. 1b)). The Early Silurian is marked by a series of global oceanic changes, termed Primo and Secundo states by Aldridge *et al.* (1993), inferred to have been determined by whether deeper oceanic circulation is driven by cold, or saline, density currents (e.g. Jeppsson & Aldridge 2000). Given the relative brevity of this cool mode it is not marked by any notable deviations from cool climate, although Secundo states in the Llandovery to earliest Wenlock interval are marked by warm high latitudes (Aldridge *et al.* 1993).

Mid Silurian to Early Carboniferous warm mode. The latest clear evidence for Silurian glaciation comes from isotopic and sedimentological evidence in the Early Wenlock (Azmy *et al.* 1998; Kaljo *et al.* 2003), placing the start of this warm mode slightly earlier than the end-Wenlock suggested by Frakes *et al.* (1992). The end of this warm mode is open to interpretation. Frakes *et al.* (1992) placed it at the Viséan–Namurian boundary (the Namurian is now called the Serpukhovian (Gradstein *et al.* 2004)), ~326 m.y. ago according

to Gradstein *et al.* (2004). However, Evans (2003) included the high-latitude Late Devonian, Famennian and Early Carboniferous, Tournaisian/Visean glaciations in the following cool mode, which would terminate this warm mode at ~375 Ma.

Based on the distribution of carbonates, evaporites, and coal deposits, Frakes *et al.* (1992) argued that this warm mode demonstrated a progressive global warming to Late Devonian times and tentatively placed its end at the beginning of the Serpukhovian. Carbonates show a progressive expansion to higher and higher latitudes during this period. However, within this warm mode, there is evidence for Late Devonian glaciation at the Frasnian–Famennian boundary (Evans 2003), ~375 Ma (c11, Fig. 1b), a Gondwana "mini-glaciation" at the Devonian–Carboniferous boundary (Caplan & Bustin 1999), ~360 Ma (c12, Fig. 1b) and glaciation in the Early Carboniferous (Tournaisian–Visean boundary), ~345 Ma (c13, Fig. 1b). Frakes *et al.* (1992) explained these glaciations by the passage of parts of Gondwana over the South Pole, largely confirmed by Evans (2003), i.e., high latitude glaciations not requiring any major perturbation of the Earth System. However, the pH-corrected CO₂ and palaeotemperature curve of Royer *et al.* (2004) suggest that following a palaeotemperature peak at ~380 Ma (early Late Devonian), there is a steady decline into the Carboniferous, with notable steeper drops between around 375 Ma and 360 Ma, which coincide with the Famennian (Streel *et al.* 2000), Devonian–Carboniferous boundary (Caplan & Bustin 1999) and Early Carboniferous (Evans 2003) glaciations.

Early Carboniferous to late Permian cool mode. Frakes *et al.* (1992) placed the beginning of this cool mode in the early Serpukhovian, at ~ 325 Ma. Evidence from South Africa and Australia supports an early glacial phase in this cool mode between 325 and 315 Ma (Evans 2003). The end of this cool mode was estimated by Frakes *et al.* (1992) to be in the Late

Permian at what was the Kazanian–Tatarian boundary (now the Wuchiapingian–Changhsingian) at ~254 Ma. The most recent data (Eyles *et al.* 2002; Eyles *et al.* 2006) place the latest glacially-influenced sedimentation in the latest Early Permian, Kungurian, at ~271 Ma.

The latitudinal distribution of glacial sediments summarised by Frakes *et al.* (1992) suggests that low-latitude glaciation (35–40° from the equator) during this cool mode occurred in two main phases, the first in the late Serpukhovian to end-Moscovian (previously called late Namurian and Westphalian) and the second in the Asselian and Sakmarian of the earliest Permian (Frakes *et al.* 1992). Evans (2003) argues that this second phase was at its most intense in a brief period around 297 Ma. In the intervening Kasimovian/Gzelian (Stephanian) glacial deposits extended no farther towards the equator than 50° (Frakes *et al.* 1992). Interpretation is made more complex because Gondwana was moving rapidly across the high southern latitudes during this cool mode (Frakes *et al.* 1992). Frakes *et al.* (1992) argued that post-Sakmarian glaciation was probably restricted to Antarctica, possibly providing a source of dropstone material to Australian and South African sedimentary basins until late in the Early Permian (e.g. Eyles *et al.* 2002; Evans 2003; Eyles *et al.* 2006). Carbonates and evaporites are restricted to low-latitudes during this cool mode (Frakes *et al.* 1992), with evaporites showing a retreat to more equatorial latitudes from the Visean onwards, with carbonates showing a similar change in the mid-Serpukhovian. Both carbonates and evaporites show a recovery from the mid Early Permian onwards (Frakes *et al.* 1992).

The short-lived climatic amelioration in the Kasimovian/Gzelian (Stephanian) (e.g. Frakes *et al.* 1992; Bruckschen *et al.* 1999) merits some discussion. Pennsylvanian cycles of terrestrial to fluvial to submarine sedimentation, called cyclothems and often associated with coal

formation (e.g. Heckel 1986), which are an expression of glacioeustatic sea-level changes (Veevers & Powell 1987), are linked to Milankovitch cyclicity, although there is some discussion over what periodicity may be the primary controlling one (e.g. Ross & Ross 1985; Heckel 1986; Crowley *et al.* 1993; Algeo *et al.* 2004; Driese & Ober 2005). In this scheme, the Kasimovian/Gzelian showed a shortening of duration of the major controlling cycles (Ross & Ross 1985). However, with a reduction in duration of the Pennsylvanian from 34 to 19 million years (Klein 1990; Gradstein *et al.* 2004), this simple explanation for Kasimovian/Gzelian warming, and a link to Milankovitch-paced cyclicity, was called into question. An examination of $\delta^{18}\text{O}$ records from Carboniferous brachiopods in western Europe and the former USSR (Bruckschen *et al.* 1999), coupled with the $^{87}\text{Sr}/^{86}\text{Sr}$ and $\delta^{13}\text{C}$ record shows drops in $\delta^{18}\text{O}$ and $^{87}\text{Sr}/^{86}\text{Sr}$ during the Kasimovian/Gzelian consistent with global warming and suggests that this was quite short-lived and restricted to the Kasimovian at ~307 Ma (w1, Fig. 1b).

Latest Permian to early Jurassic warm mode. Based on the evidence for the latest ice-rafting in Australian basins (Eyles *et al.* 2002; Eyles *et al.* 2006) the start of this warm mode fits best at the beginning of the Roadian at ~271 Ma, which is slightly earlier than the end-Wuchiapingian (Kazanian), ~254 Ma, proposed by Frakes *et al.* (1992). Frakes *et al.* (1992) suggested that this warm mode ended with the first evidence for ice-rafted debris in Jurassic sediments. These are recorded in the Middle Jurassic of northern Asia, at the Aalenian–Bajocian boundary (Chumakov & Frakes 1997), now dated at ~172 Ma (Fig. 1b), although Price (1999) summarised evidence for possible glendonites in the Early Jurassic in Pliensbachian times at ~190 Ma.

This warm mode is singular in that it contains the extinctions just before and at the Permian–Triassic boundary, the latter of which is the most severe in the Phanerozoic (e.g. Knoll *et al.*

1996; Hallam & Wignall 1999). These major changes in the biosphere have implications for lithology-based interpretation of early Triassic climates, for example, the "coal gap" (Retallack *et al.* 1996), when no coals formed anywhere on the globe even in appropriate sedimentary settings, and interpretations based on palaeosol formation, i.e. spodosols were absent (Retallack 1997). Based on relatively low-precision $\delta^{18}\text{O}$ data and a restricted set of lithological indicators (including palaeosols and reef limestones), Frakes *et al.* (1992) argued that this warm mode demonstrated an early rapid rise in temperatures in the latest Permian and Early Triassic, followed by a slow cooling trend through the Triassic and Early Jurassic. The more recent, pH-calibrated $\delta^{18}\text{O}$ -based palaeotemperature curve of Royer *et al.* (2004) indicates a very similar overall pattern, with a sharp rise in temperatures from late Early Permian times. Palaeotemperatures reached a peak in the early Middle Triassic, at ~245 Ma, followed by a slow decline towards the Mid Jurassic, only interrupted by a short-lived rise around the Pliensbachian–Toarcian boundary, at ~183 Ma (Hesselbo *et al.* 2000; Royer *et al.* 2004).

Cooler contrasting climates punctuated this warm mode, with, in some cases, combined evidence for glaciation from isotope excursions and rapid sea level changes. The earliest documented probable glacial event in this warm mode occurred across the Guadalupian–Lopingian boundary in the Late Permian (Isozaki *et al.* 2006), at ~260 Ma (c14 Fig. 1b). Although the Triassic appears to be devoid of cool intervals, stable isotope studies by Korte *et al.* (2005) indicated that short-lived, high-amplitude $\delta^{13}\text{O}$ excursions occurred in the Early Triassic, at ~250 Ma (c15, Fig. 1b), and a positive $\delta^{18}\text{O}$ excursion occurred at the Carnian–Norian boundary, at ~217 Ma (c16, Fig. 1b). In the Early Jurassic, rapid sea-level changes provide evidence of short-lived glaciation in the Late Sinemurian, ~192 Ma (c17, Fig. 1b), at the Sinemurian–Pliensbachian boundary, ~190 Ma (c18, Fig. 1b), in the late Pliensbachian,

~184 Ma (c19, Fig. 1b), and at the end Toarcian, ~176 Ma (c21, Fig. 1b), (e.g. Price 1999; Immenhauser 2005). This Early Jurassic interval is also marked by rapid temperature rise around the Pliensbachian–Toarcian boundary, ~183 Ma, (e.g. Pálffy & Smith 2000; Bailey *et al.* 2003), which appears to have been followed by a brief early Toarcian cool or possibly glacial episode, notionally at ~182 Ma (c20, Fig. 1b) (Morard *et al.* 2003; Wignall *et al.* 2005). The steep temperature rise deviates from the long-term cooling trend demonstrated by Royer *et al.* (2004). Belemnite Mg/Ca, Sr/Ca, and Na/Ca ratios increase by a factor of between 1.7 and 2, coincident with a 3 ‰ negative shift in $\delta^{18}\text{O}$ and indicating an abrupt warming in northwest Europe of 6–7° C and a more active hydrological cycle in this interval (Bailey *et al.* 2003).

Late Jurassic to Early Cretaceous cool mode. According to Frakes *et al.* (1992), the start of this cool mode is relatively clearly marked by the appearance of dropstones in high-latitude Mid Jurassic marine deposits from the Aalenian–Bajocian boundary (e.g. Chumakov & Frakes 1997), at ~172 Ma (c22, Fig. 1b). However, more recent evidence could place the beginning as early as the Pliensbachian (e.g. Price 1999; Immenhauser 2005). Compared with the evidence for its initiation, the termination of this cool mode is more difficult to define. Frakes *et al.* (1992) used the latest appearance of ice-rafted debris in Australian marine sequences of the Eromanga basin in the early Albian of the Late Cretaceous, which would approximately place it at ~110 Ma (Fig. 1b). However, a continuation of short-lived glacial episodes from the Mid Jurassic through the Late Cretaceous is supported by the pH-calibrated $\delta^{18}\text{O}$ -based palaeotemperature curve of Royer *et al.* (2004), who argue that the "Late Jurassic to Early Cretaceous" cool mode is fundamentally different from the "Early Carboniferous to late Permian" and "Mid-Palaeogene to Recent" cool modes and is more akin to the "Late Ordovician to early Silurian" one.

Frakes *et al.* (1992) argued that this mode was cooler than the preceding and following ones, based on the evidence for seasonal ice at high latitudes in both hemispheres, although not as cold as Palaeozoic cool modes. Based on a variety of indicators, including evidence of ice, palaeobotanical data, oxygen isotopes, clay minerals and the distribution of coals, evaporites and carbonates, they suggested that the intensity of cool climates varied through the mode. Frakes *et al.* (1992) proposed a cool Mid Jurassic, with seasonal ice in the Bajocian and Bathonian, ~167 Ma (c23, Fig. 1b), a conclusion supported by Price (1999). Frakes *et al.* (1992) suggested warming during the end-Mid Jurassic Callovian, although new oxygen isotope data from shark-tooth enamel and ammonite migration data suggest that the end-Mid Jurassic (late Callovian) was actually a time of glaciation, lasting ~2.6 million years (Dromart *et al.* 2003), ~164–161 Ma (c24, Fig. 1b). Glaciation at this time is also suggested by the pH-calibrated $\delta^{18}\text{O}$ -based palaeotemperature curve of Royer *et al.* (2004). For the Oxfordian, Kimmeridgian and Tithonian of the Late Jurassic, Frakes *et al.* (1992) suggested the warmest climates of the mode, although Price (1999) gave evidence for cold or sub-freezing polar climates in the Tithonian, at ~150 Ma (c25, Fig. 1b). The suggestion of a cooler and more humid earliest Cretaceous (Valanginian and Hauterivian) (Frakes *et al.* 1992) is consistent with the evidence for cool Valanginian climates, ~140 Ma (c26 Fig. 1b), presented by Price (1999). Frakes *et al.* (1992) proposed a gradual warming trend in the later Early Cretaceous following further evidence for ice in the late Early Aptian, ~120 Ma (c27, Fig. 1b), (e.g. Price 1999) towards a peak in temperatures in the Albian at the beginning of the following warm mode.

Given the variation in climates during this cool mode it is particularly difficult to pick out climates that deviate significantly from any chosen trend. However, the Late Callovian

cooling identified by Dromart *et al.* (2003) is particularly sharp (1–3 °C for lower to middle latitudes) and associated with increased drawdown of organic carbon and an abrupt fall in global sea levels.

Late Cretaceous to early Palaeogene warm mode. Frakes *et al.* (1992) place the beginning of this warm mode at the start of the Late Cretaceous in the mid-Albian at ~105 Ma. The end of this warm mode is defined by the onset of global cooling as documented in the $\delta^{18}\text{O}$ record of marine foraminifera from about ~55 Ma onwards (Frakes *et al.* 1992), which is the Palaeocene–Eocene boundary.

The status of this warm mode has been seriously questioned by Royer *et al.* (2004). Like the preceding cool mode, this warm mode shows an oscillating climate with shorter cool and warm periods, although the warm periods are substantially warmer than are those in the preceding cool mode. The mode is punctuated by a series of glacial episodes, with evidence of these from sea-level records (e.g. Miller *et al.* 2005), oxygen isotopes (e.g. Miller *et al.* 2003) and strontium (e.g. Stoll & Schrag 1996). Eustatic sea-level data (e.g. Immenhauser 2005), from the Russian and Arabian platforms, show sharp short-lived falls in the late Early and Late Albian, at ~110 and ~105 Ma (c28, c29, Fig. 1b), which may indicate high-latitude ice caps at that time. Similarly, eustatic sea-level data from offshore New Jersey, combined with the dates of positive excursions in the marine foraminiferal $\delta^{18}\text{O}$ record (e.g. Miller *et al.* 2005) indicates short-lived Antarctic ice-caps of the order of 40–50% Neogene volumes. The best evidence for these (e.g. Miller *et al.* 2005) is in the mid-Cenomanian, ~96 Ma (c30, Fig. 1b), mid-Turonian, 93–92 Ma (c31, Fig. 1b), and at the Campanian Maastrichtian boundary, 70.6 Ma (c32, Fig. 1b).

The temporal distribution of warm and cool or glacial episodes during this warm mode has changed substantially since the synthesis by Frakes *et al.* (1992). Frakes *et al.* (1992) suggested that the Albian was one of the warmest periods of this warm mode, along with the Coniacian and Campanian. New data suggests that the warmest period of this warm mode, and one of the warmest periods of the Phanerozoic, was during the Turonian (Wilson *et al.* 2002; Bice *et al.* 2006), and that the Albian was possibly quite cool with at least two short-lived high latitude glaciations (e.g. Immenhauser 2005). Additionally, the Turonian episode of maximum warmth is bracketed by short-lived glacial periods in the Cenomanian and Coniacian according to Royer *et al.* (2004), and punctuated by a glacial episode according to Miller *et al.* (2005).

Carbon dioxide levels modelled to generate sea-surface temperatures based on foraminiferal $\delta^{18}\text{O}$ and Mg/Ca data suggest variation by a factor of four (600–2400 ppmv) in the Albian to Turonian interval (Bice & Norris 2002; Bice *et al.* 2006). Benthic foraminifera assemblage data from Venezuela suggest that the following Santonian interval was warm, with some cooling towards the Campanian boundary (Rey *et al.* 2004). Frakes *et al.* (1992) suggested that the Campanian was also warm and this is supported by more recent calcareous nannofossil transfer function data from the Indian Ocean (Lees 2002).

As suggested by Frakes *et al.* (1992), the Maastrichtian was a time of cooler climate. Cooling began at the Campanian–Maastrichtian boundary. ~71 Ma, with evidence of cooler ocean temperatures from benthic foraminiferal and calcareous nannofossil assemblages (Friedrich *et al.* 2005), and eustatic sea-level drop consistent with substantial, short-lived glaciation at high latitudes (e.g. Miller *et al.* 2005). Indian Ocean calcareous nannofossil data suggest that the

late Maastrichtian was a period of warming (Lees 2002), which is also suggested by Canadian palaeosol $\delta^{13}\text{C}$ data (Nordt *et al.* 2002).

The very latest Maastrichtian to Early Danian (Early Palaeocene) saw a period of rapid changes in climate, associated with rapid changes in sea-level as indicated by planktonic foraminiferal assemblages (e.g. Keller *et al.* 2002; Keller 2004). Palaeobotanical data suggest that the Early Palaeocene, ~65 Ma (c33, Fig. 1b), was cool at high latitudes with warming by the end of the Danian (Poole *et al.* 2005). New data suggest that the later Palaeocene was subtropical at high northern latitudes, with a peak across the Palaeocene–Eocene boundary (Moran *et al.* 2006; Sluijs *et al.* 2006) at ~55 Ma, the so-called Palaeocene–Eocene thermal maximum (e.g. Zachos *et al.* 1993; Farley & Eltgroth 2003). Cooling occurred rapidly in the Early Eocene, ~55 Ma (c34, Fig. 1b), heralding the end of this warm mode, with freshening of Arctic Ocean surface waters in the latest Early Eocene, at ~49 Ma, and the first glacially-rafterd sediments evident in Arctic Ocean sequences in the early Mid Eocene at ~45 Ma (Moran *et al.* 2006).

Given the range of climate variability during this warm mode summarised above, it could be argued that the deviations from warm climates were more the norm. The short-lived Albian and Late Cretaceous glaciations are of particular interest. They are associated with sharp (< 1 m.y.) drops in sea-level of up to 80 m in the Albian (e.g. Immenhauser 2005) and up to 40 m in the Late Cretaceous (e.g. Miller *et al.* 2005). The Late Cretaceous examples are also associated with positive $\delta^{18}\text{O}$ excursions of the order of 1 ‰. As Miller *et al.* (2005) point out, these are comparable to pre-Pleistocene sea-level and $\delta^{18}\text{O}$ changes linked with substantial Antarctic ice caps. Given the recent evidence for large variations in Late Cretaceous atmospheric CO_2 concentrations (e.g. Bice *et al.* 2006), these brief glacial

excursions are less surprising than they would have been at the time of writing of Frakes *et al.* (1992). Another transient climatic episode in this warm period, the Palaeocene–Eocene thermal maximum (Zachos *et al.* 1993), at 55.8 Ma, may at face value seem an odd choice, given that it represents a period of extreme warmth. However, like the Pliensbachian–Toarcian event (e.g. Pálffy & Smith 2000; Bailey *et al.* 2003), it occurs towards the end of the warm mode (Zachos *et al.* 2001) and is followed by a short-lived episode of cooling, in the earliest Eocene (Wing *et al.* 2000). It is associated with a global strong positive excursion in $\delta^{13}\text{C}$ of 2.5 ‰ (Zachos *et al.* 1993; Dickens 2000) and sub-tropical Arctic Ocean temperatures (Sluijs *et al.* 2006). Extraterrestrial ^3He data suggest that it had a rapid onset (<few k.y.) and lasted ~120 k.y. with possible accelerated drawdown of CO_2 from the atmosphere and oceans (Farley & Eltgroth 2003).

Mid-Palaeogene to Recent cool mode.

This cool mode is perhaps the most intensively studied of all the climate modes described by Frakes *et al.* (1992), which is reflected by their devoting as many pages to it as all of the other climate modes combined. It differs from preceding climate modes in that continuous and high-resolution proxy records track its changes, e.g. $\delta^{18}\text{O}$ for palaeotemperature (e.g. Zachos *et al.* 2001; Royer *et al.* 2004), $\delta^{11}\text{B}$ for palaeo-pH and palaeo- CO_2 (Pearson & Palmer 2000), Sr/Ca from benthic foraminiferal calcite for weathering fluxes (Lear *et al.* 2003a), etc. Its beginning is marked by cooling in the early Middle Eocene (Moran *et al.* 2006), at ~49 Ma, following the Palaeocene–Eocene thermal maximum (Zachos *et al.* 1993). This is slightly later than the Early Eocene proposed by Frakes *et al.* (1992). In principle, we have not yet reached the end of this cool mode (e.g. Loutre & Berger 2000), but whether or not we have done so may be dependent on human activities in the near future (e.g. IPCC 2001).

The present cool mode is marked by an overall cooling trend, reflected in the $\delta^{18}\text{O}$ and fossil plant assemblages, punctuated by a series of sharp changes, culminating in Pleistocene bihemispheric glaciation (Frakes *et al.* 1992). The earliest evidence of cooling, that marks the beginning of this cool mode, comes from new marine core data from the Arctic Ocean that suggest substantial freshening of ocean waters close to the Early–Mid Eocene boundary (Moran *et al.* 2006), at ~49 Ma. These new data support the onset of Arctic ice-rafting from mid-Mid Eocene times onwards, at ~45 Ma, 35 million years earlier than previously assumed (Moran *et al.* 2006). This earlier evidence for Arctic ice is supported by recent sedimentary and foraminiferal geochemistry data from the Pacific, which suggest that bipolar glaciation was episodically present in the late Mid Eocene from ~42 Ma onwards (Tripathi *et al.* 2005).

The first sharp change in this cool mode, represented by major positive excursions of ~1.5 ‰ in $\delta^{18}\text{O}$ and ~1.4 ‰ in benthic $\delta^{13}\text{C}$, and sea-level drops of up to 125 m, occurred at the Eocene–Oligocene boundary (e.g. Tripathi *et al.* 2005), ~34 Ma. These changes are interpreted to mark the formation of the first large permanent ice sheets in Antarctica (e.g. Zachos & Kump 2005), although terrestrial palaeotemperature proxy data suggest that the positive excursion in $\delta^{18}\text{O}$ was more likely to be related to changes in Antarctic ice volume rather than a lowering of palaeotemperatures (e.g. Grimes *et al.* 2005). Alkenones data suggest that atmospheric CO_2 levels were still relatively high at this time, but that they continued to drop through the Oligocene, from 1000–1500 ppmv in the Late Eocene, reaching modern levels by latest Oligocene times (Pagani *et al.* 2005b).

The Early and early Late Oligocene is marked by cool climates with a series of short-lived positive excursions in the $\delta^{18}\text{O}$ curve (Miller *et al.* 1991), labelled Oi-1, 1a and 1b, at ~34

Ma, ~33 Ma and ~32 Ma (e.g. Zachos *et al.* 2001; Pekar *et al.* 2002), and Oi-2, 2*, 2a, 2b and 2c at ~30 Ma, ~29 Ma, ~28 Ma, ~27 Ma and ~26 Ma (Pekar *et al.* 2002; Wade & Palike 2004). These were associated with sea-level drops of up to 65 m (Wade & Palike 2004) and are interpreted to represent oscillating, large-scale Antarctic ice sheets (Zachos *et al.* 2001). The later Late Oligocene, ~25 Ma (w2, Fig. 1b), shows a brief return to warmer climates in both hemispheres (Barreda & Palamarczuk 2000; De Man & Van Simaeyns 2004; Villa & Persico 2006), followed by a return to glacial conditions at the Oligocene–Miocene boundary (e.g. Zachos *et al.* 2001).

Like the Oligocene, the Miocene is a time of cool climates with a series of short-lived positive excursions in the $\delta^{18}\text{O}$ curve associated with short-lived drops in sea-level (Miller *et al.* 1991). The first and most major of these, Mi-1 (Miller *et al.* 1991), is at the Oligocene–Miocene boundary, at ~23 Ma. This is followed by at least five further $\delta^{18}\text{O}$ stages, Mi-2 to Mi-6 (and substages), at ~16 Ma, ~15 Ma, ~13 Ma, ~12 Ma and ~10 Ma (Miller *et al.* 1991; Flower 1999). A significant overall positive step in $\delta^{18}\text{O}$ occurred between Mi-2 and Mi-4, between ~16 Ma and ~12 Ma (e.g. Flower 1999), the so-called "mid-Miocene cooling" event, associated with the establishment of a semi-permanent East Antarctic ice sheet and changes in oceanic circulation (e.g. Westerhold *et al.* 2005). From Mi-6 onwards, sedimentological evidence suggests the presence of glaciation outside Antarctica in both hemispheres (e.g. Denton & Armstrong 1969; Mercer & Sutter 1982; Duncan & Helgason 1998).

As suggested by Frakes *et al.* (1992), the early Tortonian of the Late Miocene, ~11 Ma (w3, Fig. 1b), shows a slight warming of global climates, from, for example, Mg/Ca ratios in benthic foraminifera (Lear *et al.* 2003b). The Late Miocene saw a resumption of significant cooling, from $\delta^{18}\text{O}$ in benthic foraminifera and accompanying sea-level falls, that resulted in

glaciation during the Messinian of the latest Miocene, between ~6.26 and ~5.50 Ma, although terminating prior to the Miocene–Pliocene boundary (e.g. Hodell *et al.* 2001), at ~5.3 Ma. Extreme desiccation of the Mediterranean basin occurred at this time (e.g. Hodell *et al.* 1986; Krijgsman *et al.* 1999). Although glacio-eustatic sea level falls and more arid global climate are contributory factors, the primary driver for isolation of the Mediterranean appears to be tectonic (e.g. Krijgsman *et al.* 1999; Hodell *et al.* 2001), particularly as similar tectonically-driven desiccation episodes affected basins at its western margin shortly before in the Tortonian (Krijgsman *et al.* 2000).

The very latest Miocene and Early Pliocene, ~5.3 Ma, show a return to milder climates (w4, Fig. 1b), as indicated by changes in the $\delta^{18}\text{O}$ of benthic and planktonic foraminifera from the Atlantic (Vidal *et al.* 2002; Reuning *et al.* 2006), thermohaline circulation modelling (Ravelo & Andreasen 2000), and Antarctic glacial sedimentology (Prentice & Krusic 2005). Climate warming peaked in the mid-Pliocene, between ~3.3 Ma and ~3 Ma (w5, Fig. 1b), and as described by Haywood *et al.* (2002b), "Earth experienced a significant sustained period of greater global warmth spanning a time frame longer than any interglacial of the Quaternary period". Evidence for this warmth comes from Australian ostracod faunal data (Warne 2005), alkenones palaeothermometry (Haywood *et al.* 2005), multiproxy analysis of Arctic Ocean ODP core data (Knies *et al.* 2002) and climate modelling (e.g. Haywood *et al.* 2000; Haywood *et al.* 2002a).

In the Late Pliocene, at some point around 2.7 Ma, probably at marine isotope stage G6 (~2.74 Ma) (Bartoli *et al.* 2005), global climate cooled. Ocean-bottom nitrogen isotope data (Sigman *et al.* 2004), alkenone unsaturation ratios and diatom oxygen isotope ratios (Haug *et al.* 2005), Ge/Si ratios (Lin & Chen 2002), and ostracod assemblages (Yamada *et al.* 2005),

suggest that this was the beginning of full-scale Northern Hemisphere glaciation (e.g. Klotz *et al.* 2006). It marks the onset of relatively continuous intensifying glacial-to-interglacial cycles to the present day (e.g. Bartoli *et al.* 2005). Late Pliocene and Early Pleistocene glacial–interglacial cycles show a 41 ky periodicity, probably controlled by the astronomical obliquity cycle (e.g. Park & Maasch 1993). Between ~1.2 and ~0.6 Ma, centred on ~0.95 Ma, the so-called mid-Pleistocene climate transition or climate revolution, the periodicity shifted to 100 ky (e.g. Tziperman & Gildor 2003). This is consistent, temporally, with control by the precessional cycle, although the variation in insolation associated with this cycle appears to be too small compared with the associated significant increase in the intensity of Northern Hemisphere glaciation and increase in amplitude of glacial–interglacial intervals (e.g. Winckler *et al.* 2004). Evidence for this change comes from positive excursions in oceanic $\delta^{13}\text{C}$ (Raymo *et al.* 1997; Wang *et al.* 2004), global eustatic sea-level fall (Kitamura & Kawagoe 2006), mass extinctions of benthic foraminifera (Hayward 2001; Kawagata *et al.* 2005; Kawagata *et al.* 2006), and the onset of major glaciation in the European Alps (Muttoni *et al.* 2003). The final stages of this transition are associated with unusual deposition of diatomaceous, carbonaceous and carbonate sediments in the South Atlantic and Mediterranean (Schmieder *et al.* 2000; Gingele & Schmieder 2001).

No particularly notable intervals of contrasting climate mark this cool mode. The Late Oligocene, ~25 Ma, mid-Miocene, ~11 Ma, and latest Miocene–Early Pliocene, ~5.3 Ma milder intervals (w2-4, Fig. 1b) are mainly of note because they briefly offset the overall cooling trend). The "mid-Miocene cooling" event between Mi-2 and Mi-4, ~16 Ma–12 Ma (e.g. Flower 1999) is interesting because the early stages coincide with eruption of the Columbia River basalts between ~16.1 and ~15 Ma (Hooper *et al.* 2002), which is the most recent large igneous province eruption known. The only interval comparable to the

Palaeozoic Kazimovian warm interval is, perhaps, the mid-Pliocene warm interval (w5, Fig. 1b), which had sea levels higher than today and East Antarctic climates similar to modern-day Chile (Haywood *et al.* 2002b).

Major Phanerozoic controls on climate

Long-term secular variation in the carbon cycle

The carbon cycle, with its links to biological and tectonomagmatic activity, is extremely important for the long-term evolution of climate in the Phanerozoic. The geochemically modelled atmospheric CO₂ curves of Berner (e.g. Berner & Kothavala 2001; Berner 2003a) (Fig. 3a) indicate that pCO₂^{atm} was high in the Early Palaeozoic, low during Permo-Carboniferous times, relatively high again in the Mesozoic, and that it has been in decline since then. Although the relationship between palaeotemperature and pCO₂^{atm} is less clear at the very highest concentrations seen during the Mesozoic (Bice *et al.* 2006), there is general agreement of a positive correlation between pCO₂^{atm} and global average temperature (e.g. Kump 2002; Retallack 2002; Royer *et al.* 2004; Siegenthaler *et al.* 2005). Examination of the curve of Berner (2003a), (Fig. 3a) and comparison with the palaeotemperature analysis by Royer *et al.* (2004) (Fig. 3b), suggests that Phanerozoic climate can be subdivided into long-term, fluctuating high CO₂ modes, interleaved with shorter-term low CO₂ modes (Fig 1b). Similarly, Raymond & Metz (2004) suggested that for Phanerozoic glaciation there are short-term, overall high CO₂ glacial modes, such as during the Late Ordovician and Cretaceous (Fig. 1b), and longer-term, low CO₂ glacial modes, such as in the Permo-Carboniferous and Late Cenozoic (Fig. 1b).

Of course, something that is not always expressed is what constitutes high or low pCO₂^{atm} levels and where the threshold lies. Royer *et al.* (2004) suggested that levels above 1000

ppmv should be considered as high and below that should be considered as low, with a transitional range between about 600 and 1000 ppmv. A brief examination of the geological record suggests that this is reasonable. The earliest Cenozoic Antarctic glaciation, Oi-1 at the beginning of the Oligocene (e.g. Tripathi *et al.* 2005), began when atmospheric CO₂ levels were at about 1000 parts per million and falling (Pagani *et al.* 2005b); recent data suggests that carbon dioxide levels in the cool Early Cretaceous were of the order of 560 to 1200 parts per million (Haworth *et al.* 2005), compared with up to 2400 ppmv for the warmer Late Cretaceous (Bice *et al.* 2006).

Tectonism

A key driver of long-term variability in atmospheric CO₂ is the amalgamation and fragmentation of continents, the so-called supercontinent cycle (e.g. Condie 2002; Murphy & Nance 2005). Major orogenic events (Fig. 4a), which exhume deep levels of the continental crust, have been proposed as drivers of major draw-down of atmospheric CO₂ (Raymo 1991; Hay 1996). These are associated with changes in palaeoelevation (e.g. Ruddiman & Kutzbach 1991), weathering rates (Raymo 1991), atmospheric circulation (Gunnell 1998) and atmospheric dustiness (e.g. Kohfeld & Harrison 2000). Silicate weathering products added to the ocean from the current orogenic belt that extends from Papua New Guinea to the Pyrenees, and includes the Himalayas, have been implicated in the overall decline of atmospheric CO₂ levels during the Late Palaeogene and Neogene and the onset of global glaciation (Raymo 1991), although modelling suggests that the net effect is not sufficient to explain the degree of cooling observed (Kerrick & Caldeira 1999).

The current phase of the supercontinent cycle is one of amalgamation (Fig. 4a), which is a likely time of major orogenesis. Similarly, in the Permo-Carboniferous, Pangaea was amalgamating (e.g. Veevers 2004), resulting in the Hercynian–Alleghenian–Uralian orogeny

(Fig. 4a), comparable in scale, duration, and rate of exhumation, to the modern-day collisional orogenic belt described above. Although the situation is slightly complicated by the rapid evolution of land plants in the Permo-Carboniferous, the addition of silicate weathering products to the Palaeo-Tethyan and Panthalassic oceans was likely to have been key in the large-scale decline of atmospheric CO₂ levels. Conversely, the rifting phase of Pangaea in the Jurassic was associated with elevated rates of CO₂ degassing (Kerrick 2001). Fig 3a and 3b suggest that there is a relatively good correlation between orogenesis that results in the amalgamation of supercontinents and low-CO₂ climate modes. Other orogenic events do not seem to correlate in time with cool episodes. The correlation with "aragonite" and "calcite" seas (Hardie 1996) (Fig. 4b) suggests that this is more closely linked to the supercontinental cycle with only very indirect influence on climate.

Another important aspect of tectonic changes over shorter timescales is the opening and closing of oceanic gateways (e.g. Smith & Pickering 2003) (Fig. 5) and change in the configuration of land and sea (e.g., Barron *et al.* 1980; Ramstein *et al.* 1997). The opening of the Drake Passage and Tasman gateways in the earliest Oligocene (Fig. 5) has been implicated in the onset of Antarctic isolation through initiation of the Antarctic Circumpolar Current (Livermore *et al.* 2004; Mackensen 2004). Very shortly afterwards, in the Early Oligocene, subsidence of the Greenland-Scotland ridge is suggested to be the trigger for the onset of North Atlantic deep water formation and the beginning of the modern thermohaline circulation (Via & Thomas 2006). In the middle Miocene, closure of the Panamanian gateway between North and South America appears to have triggered the major cooling step at that time (e.g. von der Heydt & Dijkstra 2006). Full-scale Northern Hemisphere glaciation may have followed as a result of advection of warmer waters northwards, with increased delivery of moisture to high latitudes (e.g. Lear *et al.* 2003b). Further back in time plate

configurations are much less certain. Nevertheless, opening and closing of gateways during the final amalgamation of Pangaea in the Permian have been linked to climate changes at that time (Saltzman 2003).

Magmatism

A major source of variation in the rate of addition of juvenile, mantle-derived CO₂ to the atmosphere and oceans is the rate of eruption of magmatic large igneous provinces (LIPs), volcanic eruption cycles with volumes greater than 100,000 km³ of magma, which appear to have a ~170 million year cyclicity over the past 1500 million years (Prokoph *et al.* 2004). McElwain *et al.* (2005) have suggested that LIP eruptions also add substantial CO₂ by intruding into coal seams. Enhanced rates of magmatism, particularly associated with superplume episodes, such as in the Cretaceous (Larson 1995), are implicated in intervals of high CO₂ (Caldeira & Rampino 1991). The LIP record for the Phanerozoic is biased towards the record of terrestrial LIP eruptions (Fig. 6). Consequently the record is much more complete for the period after 170 million years ago, when the presently interrogateable seafloor record began. The Jurassic–Cretaceous, Mesozoic peak in LIP emplacement (Fig. 6), the most recent peak in the superplume cycle, is well represented in the geological record and coincides with a relative peak in atmospheric CO₂ and some of the highest global temperatures in the Phanerozoic (Vaughan & Storey in press).

The record for LIP emplacement pre-170 million years ago is much more sketchy. Some Late Triassic and Early Jurassic oceanic LIPs have been recognised (e.g. Vaughan & Storey in press), but that is as far back as the record of oceanic LIPs extends, so far. Earlier LIPs are all terrestrial, and largely recognised from dyke swarms rather than basalt or rhyolite lavas

(Ernst *et al.* 2005). In Figure 6, LIP eruptions through the Phanerozoic are plotted by area (Fig. 6a) and by volume (Fig. 6b). Volume is one of the key parameters required to assess the impact of a LIP, but volumes are difficult to obtain for LIPs older than Early Jurassic. Dyke swarm extent gives a proxy for area for older LIPs, so for Phanerozoic-length timescales it is useful to examine LIPs by erupted area. For the Mesozoic, there is a striking time coincidence in some cases between LIP eruptions and cool or glacial episodes, e.g. the Siberian Traps, Karoo magmatism and Shatsky Rise LIP (Fig. 6a).

If basaltic LIP eruptions are possible drivers of elevated global temperature, silicic LIPs and super-eruptions appear to be associated with global cooling. For example, Briffa *et al.* (1998) showed a strong correspondence between large silicic eruptions over the past 600 years and severe Northern Hemisphere winters based on maximum latewood density in tree rings. Over longer timescales, Rampino & Self (1992) suggested that the Toba super-eruption 73,500 years ago accelerated the Oxygen Isotope Stage 5a to 4 interglacial to glacial transition, and Prueher & Rea (1998) suggested that a 10-fold increase in volcanicity in the North Pacific ~2.67 million years ago coincided with an intensification of Late Pliocene continental glaciation. Farther back in time, the Early and early Middle Jurassic period of Gondwana break-up silicic LIP eruptions (Fig. 6a) coincide with a period of four glacial episodes over that interval (Fig. 1b). Modelling by Jones *et al.* (2005) suggests that the silicic magmatism is insufficient on its own to trigger glaciation, although deviation from normal temperatures may last decades. The evidence suggests that cooling associated with silicic eruptions may act as an intensifier in the onset of glaciation where other conditions are favourable.

The rate of obduction of ophiolites (Fig. 7a), a slice of ocean floor added to continental margins rather than being subducted, has also been suggested as a proxy for peaks in the superplume cycle (Vaughan & Scarrow 2003). Again, as with LIP eruptions, from Fig. 7a, there is a striking positive time correlation between ophiolite obduction pulses and Palaeozoic and Mesozoic cool intervals, particularly for the Steptean in the Cambrian (c2, Table 1), the Tremadocian in the Early Ordovician (c3, Table 1), mid and Late Jurassic intervals, and the Late Cretaceous and Palaeogene.

The rate of eruption of kimberlites (Fig. 7b), also a significant source of juvenile CO₂, appears to correlate positively with the rate of basaltic LIP emplacement. At face value, kimberlites volumes appear too small to perturb the CO₂ balance of the atmosphere and oceans for more than a few years. However, there is a good general positive correlation, visible on Fig. 7b, between peaks in kimberlite eruption rate and high-CO₂ climate intervals and lulls in kimberlite eruption rates and low-CO₂ climate modes, which suggests that they may have a more significant and longer-term effect.

Overall, these proxies, in conjunction with the CO₂ and sea-floor record, suggest a major peak in oceanic LIP emplacement during the Cambro-Ordovician, even in the absence of surviving direct evidence. These LIPs were probably oceanic, as were most of the Cretaceous LIPs, and constitute the early Palaeozoic superplume event proposed by Larson (1991) and Barnes (2004). Significant LIP emplacement intervals, also probably representing superplume events, although on a smaller scale than the Mesozoic and Cambro-Ordovician events, have also been suggested for the Late Carboniferous–Early Permian (e.g. Larson 1991; Doblas *et al.* 1998) and Late Permian (e.g. Medvedev *et al.* 2003).

Biological evolution

The evolution of land plants with complex root systems caused a major change in atmospheric composition (e.g. Berner 1998) and the evolutionary rise in land plants has been strongly implicated in the severity of the long-term CO₂ reduction during the Permian–Carboniferous low–CO₂ interval and global glaciation (e.g. Berner 2003b; Igamberdiev & Lea 2006). Plants increased the rate of continental silicate weathering rates, enhancing the drawdown of CO₂ in the oceans as carbonate; organic acids generated by plant roots are significant agents of silicate weathering (Caldeira 2006). Aided by actively subsiding basins and many cycles of glacio-eustatic sea level rise and fall, or cycles of humidity–aridity, plants also directly contributed to the removal of CO₂ from the atmosphere by forming peat, which was subsequently turned to coal (e.g. Heckel 1996). Similar suggestions for the Cenozoic have been made with regard to the rise of the angiosperms (Igamberdiev & Lea 2006).

Another major change happened in the Late Triassic with the evolution of plankton capable of fixing carbonate to form hard skeletons, allowing the deep oceans to become a significant sink for carbonate (e.g. Ridgwell 2005; Erba 2006). This was initiated by the change from dominance in the oceans of phytoplankton with green algal plastids, e.g. acritarchs, to those with red algal plastids, e.g. coccolithophorids, at the Permian–Triassic boundary (Grzebyk *et al.* 2003) culminating in the Late Triassic (e.g. Ridgwell 2005). Prior to this, carbonate fixing was carried out in continental shelf areas by largely benthic and infaunal organisms (e.g. Ridgwell 2005). Major drops in sea level in pre-Mesozoic time, by diminishing shelf area, drastically reduced the capacity of the Earth system to sequester carbon as carbonate, a regulated and buffered process by comparison with the more rapid drawdown of CO₂ as organic carbon, saturating the oceans (Ridgwell 2005). Since the Late Triassic, only at times

of probable ocean acidification do we see a return to the Palaeozoic and Early Mesozoic mechanism, and, in addition to the swings in climate, sea-level, and isotopic systems described above, this is commonly associated with mass extinction of organisms, for example, at the Cretaceous–Palaeogene boundary (e.g. Hollander *et al.* 1993). However, the relationship between carbonate and organic carbon formation and atmospheric CO₂ is complex and still not well understood (see Barker *et al.* 2003 for a recent review)

Astronomical controls

Changes in insolation driven by the Earth's orbital parameters of precession, obliquity and eccentricity, have long been recognised as important drivers of climatic change, if only widely accepted in the last 30 years (e.g. Croll 1867; Milankovitch 1941; Hays *et al.* 1976; Zachos *et al.* 2001; Berger & Loutre 2004). These occur on 19 ky, 41 ky, and 100 ky frequencies, respectively. Additionally, there are so-called "grand" cycles related to the orbits of the major planets with periodicities at the present day of 400 ky, and 1.25, 2.35 and 4.6 My (e.g. Olsen 2001). Modelling based on Milankovitch theory has been applied with varying degrees of success to explanation of cyclical climate variations for much of the Neoproterozoic (e.g. Lewis *et al.* 2003) and Phanerozoic (e.g. Herrmann *et al.* 2003; Wade & Palike 2004; Holbourn *et al.* 2005; Huynh & Poulsen 2005). These studies suggest that insolation minima of the orbital cycles are the likely trigger for glaciation at low CO₂ levels at long and short timescales. The effects of long-term changes in the intensity of insolation at the peaks of the Milankovitch cycle are not known, but may be a significant factor in longer-term climate change.

Although the sun is recognised as a short-term variable (Solanki & Krivova 2004), the influence of variation in solar output at longer timescales, although potentially an extremely

important driver of climate change (e.g., Walker *et al.* 1981), is currently difficult to model or quantify. For example, Foukal *et al.* (2006) have shown that variations in the Sun's brightness (solar luminosity) caused by changes in sunspot area are insufficient to account for warming since the Little Ice Age; however, they point out that additional climate forcing by variations in solar ultraviolet output or the solar wind cannot be ruled out. Apart from the obvious effect of variation in insolation, solar activity, via accompanying changes in the strength of the Sun's magnetic field, may also inversely affect the galactic cosmic ray flux (Usoskin *et al.* 2005), with possible implications for low-level cloudiness (e.g. Yu 2002; Harrison & Stephenson 2006). For example, Ramirez *et al.* (2006) have suggested that cosmic ray modulated variation in low-level cloud cover may be responsible for temperature changes up to 50% that predicted for doubling of atmospheric CO₂. In terms of variations in solar output and insolation at the Earth's surface, a study of sunspot activity for the Holocene, using dendrochronologically-dated variations in ¹⁴C, suggests that current solar activity is at its highest for 11,000 years (Solanki *et al.* 2004).

Variations in the galactic cosmic ray flux as a function of the location of the solar system relative to the spiral arms of the galaxy has been implicated as a driver of long-term climate change (Wallmann 2004; Wendler 2004; Shaviv 2005). Figure 8 shows a summary of the warm and cool modes of Frakes *et al.* (1992) for the last 600 million years, overlain by blocks showing likely periods of enhanced galactic cosmic ray flux of Wallmann (2004). Interestingly, there appears to be a loose time correlation between peaks in the galactic cosmic ray flux and protracted periods of Phanerozoic glaciation (Wallmann 2004). However, this is less clear for high- and low-CO₂ climate modes (Fig. 8b). Mechanisms proposed to explain links between galactic cosmic ray flux and climate include the reflective and cooling effects of enhanced low-level cloudiness nucleated by cosmic rays penetrating

the atmosphere (Shaviv 2005). However, Royer *et al.* (2004) showed that a pH-corrected Phanerozoic palaeotemperature curve predicts proxy-based cool and warm epochs more accurately than predicted variation in galactic cosmic ray flux (Shaviv & Veizer 2003).

Figure 8 suggests that, stochastically, the periodicity of changes in the galactic cosmic ray flux and the more closely coupled supercontinent and superplume cycles have been partially in step, although not in any way coupled, during the Phanerozoic. Galactic cosmic ray flux mechanisms for global cooling (e.g. Shaviv & Veizer 2003) may have at times intensified major glaciations, but are unlikely to be directly responsible for them. Interestingly, Evans (2003) has pointed out that, taking into account preservational bias, sedimentological or isotopic evidence for major global glaciation is absent between the 2.5–2.2 Ga Palaeoproterozoic Huronian glaciation and the Neoproterozoic "Snowball Earth" glaciations of ~750 million years ago (Evans 2000). One outside possibility is that peaks in the galactic cosmic ray flux and periods of low-CO₂ (lows in the superplume cycle and times of supercontinental amalgamation) were out of phase during the later Palaeoproterozoic to early Neoproterozoic interval, moderating the severity of low-CO₂ mode glaciation.

Other mechanisms

Studies of the chemical composition of evaporite deposits and of fluid inclusions in halite (e.g. Hardie 1996; Lowenstein *et al.* 2001) suggest that the carbonate composition of seawater has varied systematically through the Phanerozoic, alternating between "aragonite seas", as at the present day, and "calcite seas" (e.g. Hardie 1996; Berner 2004) (Fig. 4c), such as in the Cretaceous. It has been argued that these changes in composition are variously related to rates of magmatic or hydrothermal activity at spreading ridges, changes in Mg/Ca ratio, and pCO₂ (summarized in Adabi 2004). As summarized by Erba (2006), a broad general relationship is evident with aragonite seas during periods of supercontinental

amalgamation and relatively low hydrothermal activity at spreading ridges, i.e. relatively low-CO₂ intervals, and calcite seas during periods of high hydrothermal activity at spreading ridges and at times of supercontinental dispersal, i.e. relatively high-CO₂ intervals.

Dissociation of seabed methane hydrate deposits (e.g. Nisbet 1992; Hesselbo *et al.* 2000), or high latitude permafrost-hosted methane hydrate deposits (e.g. Nisbet 2002), is implicated in generating peaks in atmospheric CO₂, with accompanying palaeotemperature rises, and paired negative–positive spikes in $\delta^{13}\text{C}$ (e.g. van de Schootbrugge *et al.* 2005). The proposal is that, destabilised by warming of ocean waters (e.g. Thomas *et al.* 2002a), or decompression caused by a drop in sea-level (e.g. Maslin *et al.* 2005), methane hydrate deposits in shelf sediments catastrophically decrepitate. This transfers large volumes of isotopically-light, up to -60 ‰ $\delta^{13}\text{C}$, carbon as methane to the atmosphere (e.g. Zachos *et al.* 2005). Methane is a powerful greenhouse gas in its own right (Badr *et al.* 1992), so this initial release may cause a pulse of global warming, but atmospheric chemistry models suggest that the methane rapidly oxidizes to CO₂ (< 10 years, Nisbet pers. comm.. 2006), prolonging the elevation of temperatures. The increase in atmospheric CO₂ triggers a rise in ocean productivity, initially resulting in a negative spike in $\delta^{13}\text{C}$, which progresses to more positive values as drawdown of CO₂ continues (Jenkyns 2003).

Methane hydrate dissociation has been proposed as the mechanism for events at the boundaries of the Permian–Triassic (e.g. Winguth & Maier-Reimer 2005), Pliensbachian–Toarcian (Hesselbo *et al.* 2000) and Palaeocene–Eocene (Dickens 2000). The Permian–Triassic event coincided with the eruption of the Siberian Traps (a major basaltic and silicic LIP) (Wignall 2001) and the Early Jurassic methane hydrate release was also coincident with

major LIP eruptions 183 Ma (Fig. 6a) associated with the onset of Gondwana break-up (Vaughan & Storey in press).

Discussion

Examination of the curves of Berner & Kothavala (2001), review of the geological data presented above, and comparison with the palaeotemperature analysis by Royer *et al.* (2004) suggests an alternative to the warm and cool modes of Frakes *et al.* (1992). As outlined in Table 1 and Fig. 1b, I propose that there are long-term high CO₂ warm modes to Phanerozoic climate, punctuated by short-lived contrasting cool or glacial intervals, and shorter term low CO₂ cool modes, such as the current Oligocene–Recent interval, with rarer contrasting short-lived warm intervals, that culminate in times of major global glaciation, or superglaciations.

Examination of the geological record of cool and warm climates, as summarized in Table 1, Fig. 1b, and Figs 3,4,6 & 7, would suggest that, high CO₂ modes, which have a predominance of warm climates, coincide with peaks in the superplume cycle, i.e. with high rates of emplacement of LIPs, and coincide with periods of continental dispersal in the supercontinent cycle. These are periods of high rates of magmatic activity and elevated hydrothermal activity at mid-ocean ridges. Although the correlation is not perfect, high CO₂ modes also loosely coincide with "calcite" seas. Overall, these modes are times of high CO₂ flux and relatively low rates of continental weathering.

Low CO₂ modes, which are generally of much shorter duration, appear to coincide with lulls in the superplume cycle, i.e. with very low rates of LIP emplacement (Fig. 6). They coincide with periods of amalgamation in the supercontinent cycle with very large-scale orogenesis and high rates of crustal exhumation (Fig. 4). These are periods of relatively low rates of

magmatic activity and low rates of hydrothermal activity at mid-ocean ridges. Again, although the correlation is not perfect, low carbon dioxide modes approximately with "aragonite" seas. Overall, these modes are times of relatively low CO₂ flux and elevated rates of continental weathering.

Of particular interest are the periods of contrasting climate (summarised in Table 1). High CO₂ modes are punctuated by frequent short-lived cool or glacial episodes. Low CO₂ modes, on the other hand, only rarely experience short-lived warming episodes, such as that in the mid-Pliocene. The short-lived cool or glacial episodes in high CO₂ modes have some common characteristics. They generally occur at times of eustatically high sea level, but are marked by a rapid drop in sea level, such as in the Late Callovian (Dromart *et al.* 2003). They are preceded by or overlap with episodes of deposition of carbonaceous sediments, such as black shales or sapropels, such as in the Turonian (Wilson *et al.* 2002; Bice *et al.* 2006). Where reliable isotope data can be extracted, such as for the Mesozoic and early Cenozoic, they are also associated with positive excursions in $\delta^{13}\text{C}$ and $\delta^{18}\text{O}$, again seen in the Late Callovian (Dromart *et al.* 2003).

A modified scheme of climate modes for the Phanerozoic

The Earliest Cambrian to mid-Ordovician, Late Ordovician to Early Silurian and Late Silurian to mid-Early Carboniferous climate modes of Frakes *et al.* (1992) (warm–cool–warm respectively) coincided with the early Palaeozoic high-CO₂ climate mode (Table 1, Fig. 1b), which was largely warm with CO₂ concentrations fluctuating on short timescales. As suggested above, this high-CO₂ mode can probably be extended back into the Ediacaran of the late Neoproterozoic. This is based on the analysis of Royer *et al.* (2004), and the CO₂ curve of Berner (2003a). It largely overlaps with the early Palaeozoic "calcite" sea (e.g.

Holland 2004), indicating a time of high hydrothermal activity at spreading ridges and continental dispersal. As discussed above, proxy data suggest that this is an interval of a peak in the superplume cycle with elevated rates of LIP eruptions.

Of the short-lived cool, possibly glacial, phases during this high-CO₂ mode, the Ediacaran short-lived glacial interval at ~580 Ma (c1, Fig. 1b; Table 1) coincides with the inception of the Middle Iapetus LIP of Ernst & Buchan (2001). Eruption of the Kalkarindji LIP in Australia (Glass & Phillips 2006) overlaps with the Middle Cambrian, Steptean positive $\delta^{13}\text{C}$ excursion, major sea-level drop and probable short-lived glacial episode (e.g. Saltzman *et al.* 2000) (c2, Fig. 1b; Table 1).

For the short-lived cool or glacial periods in the early Ordovician, late Ordovician, and Early Silurian (c3–c10, Fig. 1b; Table 1), there are no recorded episodes of basaltic LIP magmatism, although indirect proxy evidence suggests that they were likely. The late Ordovician Hirnantian glaciation (c7, Fig. 1b; Table 1) was probably the most severe and is associated, although not simultaneously, with black shale deposition and a mass extinction event (Marshall *et al.* 1997). The three Early Silurian glaciations (Kaljo *et al.* 2003) (c8–c10, Fig. 1b; Table 1) do coincide with an inferred period of elevated global silicic volcanism, mainly represented by thick bentonite horizons, representing individual explosive eruptions of up to 1000 cubic kilometres each (Fortey *et al.* 1996). These may have resulted in short periods of global cooling, as suggested for explosive eruptions during the past 600 years (e.g. Briffa *et al.* 1998), or have acted as a trigger for glaciation (e.g. Rampino & Self 1992; Jones *et al.* 2005).

During the latter part of this high-CO₂ mode, short-lived glaciation is associated with black shale deposition, sea level changes and mass extinction in the Late Devonian, the Kellwasser event of the Frasnian–Famennian (Buggisch 1991) (c11, Fig. 1b; Table 1), and at the Devonian–Carboniferous boundary, the Hangenberg event (Caplan & Bustin 1999) c12, Fig. 1b; Table 1). This latter event coincides with the eruption of the Yakutsk–Baltica LIP (Fig. 6a), one of the largest known LIPs of the Palaeozoic. A final phase of short-lived glaciation occurred in the Visean of the Early Carboniferous (e.g. Eyles 1993) (c13, Fig. 1b; Table 1).

The following late Palaeozoic cool, low-CO₂ mode (Table 1, Fig. 1b) coincides with the Permo-Carboniferous cool mode of Frakes *et al.* (1992) and was a period of superglaciation. It coincided with a peak of orogenesis in the supercontinent cycle, with activity along the Alleghenian-Hercynian-Uralian orogeny as Pangaea finished amalgamating (e.g. Veivers 2004). It coincided with the late Palaeozoic "aragonite" sea (e.g. Holland 2004), indicating a time of low hydrothermal activity at spreading ridges and continental amalgamation. This low-CO₂ mode occurs at a low in the superplume cycle, with only one LIP recognised in the Late Carboniferous and Early Permian (Kazimovian–Artinskian) (Doblas *et al.* 1998). This low-CO₂ cool mode is punctuated by only one short-lived warm interval, in the Kazimovian (Bruckschen *et al.* 1999) (w1, Fig. 1b; Table 1), which precedes the peak in Permo-Carboniferous superplume magmatism (Doblas *et al.* 1998) (Fig. 6a). Warming in the Kazimovian was associated with mass extinction of land plants (Edwards 1998).

The mid-Phanerozoic warm, high-CO₂ mode (Table 1, Fig. 1b) incorporates the Latest Permian to Early Jurassic, Middle Jurassic to Early Cretaceous, and Late Cretaceous to Palaeocene climate modes of Frakes *et al.* (1992) (warm-cool-warm, respectively). Like the early Palaeozoic high-CO₂ mode, this mode was largely warm with CO₂ concentrations

fluctuating on short timescales. This mode occurs during a period of continental dispersal following the break-up of Pangaea and Gondwana from Early Jurassic times onwards. It overlaps with the Mesozoic to Palaeogene "calcite" sea (e.g. Holland 2004), indicating a time of high hydrothermal activity at spreading ridges and continental dispersal. Direct magmatic evidence suggests that this is an interval of a peak in the superplume cycle with elevated rates of continental and oceanic LIP eruptions (Vaughan & Storey in press). At least 25 LIPs erupted between early Jurassic times and the Late Palaeocene, many of which coincided with cool or glacial intervals (Fig. 6a). The short-lived cooler episodes in this warm, high-CO₂ mode (Table 1, Fig. 1b) are associated with black shale deposition, oxygen and carbon isotope excursions, sea-level changes and mass extinctions.

Notable events occurred in the Late Permian (c14, Fig. 1b; Table 1), associated with the eruption of the Emeishan Traps and the Guadalupian–Lopingian mass extinction event (Wignall 2001), in the earliest Triassic (c15, Fig 1b, Table 1) associated with the eruption of the Siberian Traps and the largest mass extinction of the Phanerozoic (e.g. Reichow *et al.* 2002), and in the Late Triassic and Early Jurassic (c16–c21, Fig. 1b; Table 1), associated with emplacement of the Central Atlantic Magmatic Province and Gondwana Large Igneous Province and several mass extinction events (summarised in Vaughan & Livermore 2005). Further events occurred throughout the Jurassic and Cretaceous and in the Late Palaeocene–Early Eocene (c22–c34, Fig. 1b; Table 1). Mid-Cretaceous LIP eruptions are not particularly noted for mass extinction events, although these LIPs were largely submarine and oceanic, which may have reduced their impact on the biosphere (Wignall 2001).

The Cenozoic cool, low-CO₂ mode (Table 1, Fig. 1b) is equivalent to the Early Eocene to present cool mode of Frakes *et al.* (1992), and like the late Palaeozoic low-CO₂ mode is also

a time of superglaciation. It coincides with a peak of orogenesis in the supercontinent cycle, with Cenozoic activity along the orogenic belt that extends from the Pyrenees to Papua New Guinea, and includes the Himalayas, as a new supercontinent forms. Orogenesis is also active along the western margin of the Americas (Dickinson 2004; Sobolev & Babeyko 2005). This cool mode occurs during the current "aragonite" sea (e.g. Holland 2004), indicating a time of low hydrothermal activity at spreading ridges. This low-CO₂ mode occurs at a low in the superplume cycle, with only two small LIPs having erupted in the last 50 million years: the Ethiopian traps, which erupted ~30 million years ago (Hofmann *et al.* 1997), and the Columbia River basalts ~16 million years ago (Hooper *et al.* 2002). This low-CO₂ cool mode is punctuated by four short-lived warm intervals (w2–w5, Fig. 1b; Table 1), in the Late Oligocene, Late Miocene, latest Miocene–Early Pliocene and mid-Pliocene (mid-Neogene) (Table 1, Fig. 1b).

Conclusions

Emanuel (2002) observed that although climate appears to be sensitive to small orbitally driven changes in insolation, often with abrupt consequences, it is on another level stable, in that it has shown only relatively small variations despite an almost 30% increase in solar output since formation of the Earth 4,560 million years ago. Emanuel (2002) suggested that the climate may have a limited series of overlapping stable regimes that result in more than one equilibrium state for the same insolation conditions. This has also been modelled for the thermohaline circulation (Marotzke & Willebrand 1991), and was suggested by Chamberlin (1906) in the early 20th Century. A review of the palaeoclimatic evidence suggests that through the Phanerozoic two overlapping stable climate regimes appear to have dominated: a high-CO₂, largely warm climate regime, punctuated by short-lived episodes of glaciation

driven by negative-feedback processes in the carbon cycle; and a low-CO₂, largely cool regime, marked by protracted episodes of superglaciation. Other stable regimes are obviously possible, the most recent of which were seen during the low-latitude, "Snowball Earth" global glaciations of the Neoproterozoic (e.g. Hoffman *et al.* 1998).

The future

From a proxy point of view, a major gap in the arsenal of techniques available to scientists is a simple and reliable one for marine salinity, particularly given its importance for the density of ocean waters and the control on deep ocean circulation (Henderson 2002). Given the success of the crenarchaeotal-lipid-membrane-based Tex86 proxy for palaeotemperature (Schouten *et al.* 2002), perhaps the halophile euryarchaeota, which are widely present in terrestrial and marine environments (e.g. Elshahed *et al.* 2004; Herndl *et al.* 2005), may provide a palaeosalinity equivalent. A gap also exists for good proxies of global oceanic alkalinity and weathering (Henderson 2002), which exert a key control on CO₂ drawdown. From a modelling point of view, the way forward, given the steady increase in computer processing power, is almost certainly towards increasingly coupled GCMs. The most ambitious plans, at the time of writing, aim to couple biome and ice sheet models with coupled ocean-atmosphere GCMs (e.g. Haywood *et al.* 2002b). Model evaluation with the highest-quality proxy data from the geological record is the only way that we can increase our confidence in model outputs, an essential for anticipating future climates, as outlined by Kump (2002). From a geological point of view, the recent success of the Integrated Ocean Drilling Program Arctic Coring Expedition to the Lomonosov Ridge (Brinkhuis *et al.* 2006; Moran *et al.* 2006; Sluijs *et al.* 2006) illustrates the continued importance of coring the ocean

basins. Drilling programmes are planned in both polar regions and the importance of this cannot be overstated.

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Figure Captions

Figure 1: **(a)** Late Neoproterozoic and Phanerozoic climate modes of Frakes *et al.* (1992) from Table 1; **(b)** Intervals of glacial or cool (grey bars) and warm (white areas or bars) climates for the Late Neoproterozoic and Phanerozoic, modified from fig. 2b of Royer *et al.* (2004) including additional intervals as discussed in the text and listed in Table 1, with the Palaeozoic–Recent palaeotemperature curve of Royer *et al.* (2004) based on $\delta^{18}\text{O}$ calibrated with palaeo-pH based on CO_2 proxies and palaeo-concentrations of Ca in seawater; cool intervals labelled c1–c29 (Table 1); warm intervals labelled w1–w5 (Table 1). Brackets above Fig. 1b indicate the durations of high- and low- CO_2 modes for Phanerozoic climate as discussed in the text and listed in Table 1.

Figure 2: HadCM3 coupled ocean-atmosphere GCM prediction showing the difference in December, January and February (DJF) & June, July and August (JJA) surface temperatures ($^{\circ}\text{C}$) between a Mid Pliocene (~ 3 Ma BP) and pre-industrial experiment (redrawn from Haywood *et al.* in press).

Figure 3: **(a)** Palaeo- CO_2 atmospheric concentration as a multiple 'R' of pre-industrial values using the curve of Berner (2003a) superimposed on the climate modes of Frakes *et al.* (1992); **(b)** Intervals of glacial or cool (black bars) and warm (white areas) climates and durations of high- and low- CO_2 modes for the Late Neoproterozoic and Phanerozoic modified from Fig. 1b shown for reference.

Figure 4: Fig. 1a superimposed with **(a)** late Neoproterozoic and Phanerozoic tectonic events modified after Vaughan & Livermore (2005) fig. 3; except age of Rodinia break-up (Murphy *et al.* 2004), Pan-African events (Thomas *et al.* 2002b; Heilbron & Machado 2003; Meert

2003), Ross–Delamerian (Boger & Miller 2004), Grampian/Taconian (Pincivy *et al.* 2003; Dewey 2005), Scandian (Kirkland *et al.* 2006) and Acadian (Sherlock *et al.* 2003); **(b)** Intervals of glacial or cool (black bars) and warm (white areas) climates and durations of high- and low-CO₂ modes for the Late Neoproterozoic and Phanerozoic modified from Fig. 1b shown for reference; **(c)** Phanerozoic intervals of "calcite" and "aragonite" seas for the Phanerozoic as suggested by Hardie (1996).

Figure 5: Reconstruction of Cenozoic opening of the Drake Passage and Tasman oceanic gateways and onset of the circum-Antarctic current (Redrawn after Kennett 1977).

Figure 6: **(a)** Fig 1b intervals of glacial or cool (grey bars) and warm (white areas or bars) climates for the Late Neoproterozoic and Phanerozoic overlain by the timing, and approximate, estimated areas of large igneous province flood basalt events (including Gondwana rhyolite magmatism; duration from Riley *et al.* (2001)) for the Neoproterozoic to Neogene modified from Ernst & Buchan (2001) and on-line database at <http://www.largeigneousprovinces.org/downloads.html> (accessed 9/7/2006), except Ethiopia (Kieffer *et al.* 2004), North Atlantic volcanic province (Eldholm & Grue 1994), Bunbury Basalts (Ingle *et al.* 2004), CAMP (McHone 2002), Nilufer unit (Genc 2004), Brazil (Santos *et al.* 2002) and Kalkarindji LIP (Glass & Phillips 2006); Gondwana rhyolite magmatism shown as horizontal grey bar; **(b)** data as for (a) but showing large igneous province volumes instead. Names of continental flood basalts indicated in plain type; oceanic flood basalts indicated in italic type. Gradationally shaded area shows geological interval over which sea floor is preserved.

Figure 7: Fig 1b intervals of glacial or cool (grey bars) and warm (white areas or bars) climates for the Late Neoproterozoic and Phanerozoic overlain by **(a)** Cumulative frequency plot (Ludwig 1999) of Phanerozoic ophiolite obduction events modified from Vaughan & Scarrow (2003) **(b)** (Cumulative frequency plot (Ludwig 1999) of Phanerozoic kimberlite age data replotted from Griffin *et al.* (1999), Heaman & Kjarsgaard (2000), Belousova *et al.* (2001), Davis & Miller (2001), Heaman *et al.* (2003), Westerlund *et al.* (2004), and Jelsma *et al.* (2004).

Figure 8: **(a)** Fig. 1a superimposed with intervals of high galactic cosmic ray flux for the Phanerozoic from Wallmann (2004); **(b)** Intervals of glacial or cool (black bars) and warm (white areas) climates and durations of high- and low-CO₂ modes for the Late Neoproterozoic and Phanerozoic modified from Fig. 1b shown for reference.

Table 1: Climate modes and periods of contrasting climate for the Phanerozoic

Climate modes ¹ of Frakes <i>et al.</i> (1992)	Interval (inclusive)	Absolute age range ²	New CO ₂ modes ³ proposed in this paper	Overall climate	Interval (inclusive)	Absolute age range ²	Intervals of contrasting climate ⁴	Absolute age ²
Warm	Ediacaran to Late Ordovician	600–461 Ma	High	Warm, fluctuating	Ediacaran–Visean (late Early Carboniferous)	600–326.4 Ma	mid-Ediacaran (c1)	580 Ma
Cool	Late Ordovician to Early Silurian	461–428					Steptoan (mid-Late Cambrian) (c2) Tremadocian (Early Ordovician) (c3) early Katian stage, mid-Caradoc (Late Ordovician) (c4) mid-Katian stage, Late Caradoc (Late Ordovician) (c5) late Katian stage, Early Ashgill (Late Ordovician) (c6) Hirnantian (Late Ordovician) (c7) Aeronian (Early Silurian) (c8) Telychian (Early Silurian) (c9) Wenlock (late Early Silurian) (c10) Frasnian–Famennian boundary (Late Devonian) (c11)	494 Ma 480 Ma 455 Ma 452 Ma 449 Ma 445 Ma 439 Ma 436 Ma 428 Ma 374 Ma
Warm	mid-Silurian to mid-Early Carboniferous	428–326.4					Devonian–Carboniferous boundary (c12) Tournaisian–Visean boundary (Early Carboniferous) (c13)	360 Ma 345 Ma
Cool	Late Early Carboniferous to Late Permian	326.4–270.6	Low	Cool, glacial	Serpukhovian (early Late Carboniferous)–Kungurian (latest Early Permian)	326.4–270.6 Ma	Kazimovian (w1)	307 Ma
Warm	Latest Permian to Early Jurassic	270.6–172	High	Warm, fluctuating	Roadian (early mid-Permian)–Early Eocene (early mid-Palaeogene)	270.6–49 Ma	Guadalupian–Lopingian boundary (Late Permian) (c14) Early Triassic (c15) Carnian–Norian boundary (Late Triassic) (c16) late Sinemurian (Early Jurassic) (c17) Sinemurian–Pliensbachian boundary (Early Jurassic) (c18) late Pliensbachian (Early Jurassic) (c19) early Toarcian (Early Jurassic) (c20) end Toarcian (Early Jurassic) (c21) Aalenian–Bajocian boundary (mid-Jurassic) (c22)	260 Ma 250 Ma 217 Ma 192 Ma 190 Ma 184 Ma 182 Ma 176 Ma 172 Ma
Cool	Middle Jurassic to Early Cretaceous	172–105					Bathonian (mid-Jurassic) (c23) late Callovian (mid-Jurassic) (c24) Tithonian (Late Jurassic) (c25) Valanginian (Early Cretaceous) (c26) late Early Aptian (Early Cretaceous) (c27) Early Albian (Early Cretaceous) (c28) Late Albian (Early Cretaceous) (c29) mid-Cenomanian (Late Cretaceous) (c30) mid-Turonian (Late Cretaceous) (c31) Campanian–Maastrichtian boundary (c32) Early Palaeocene (Early Palaeogene) (c33) Early Eocene (mid-Palaeogene) (c34)	167 Ma 164–161 Ma 150 Ma 140 Ma 120 Ma 110 Ma 105 Ma 96 Ma 93–92 Ma 70.6 Ma 65 Ma 55 Ma
Warm	Late Cretaceous to Early Eocene	105–49					Late Oligocene (Late Palaeogene) (w2)	25 Ma
Cool	mid-Eocene to present	49–0	Low	Cool, glacial	mid-Eocene (mid-Palaeogene)–Recent (latest Neogene)	49–0 Ma	Late Miocene (Early Neogene) (w3) latest Miocene–Early Pliocene (mid Neogene) (w4) mid-Pliocene (mid-Neogene) (w5)	11 Ma 5.3 Ma 3.3–3 Ma

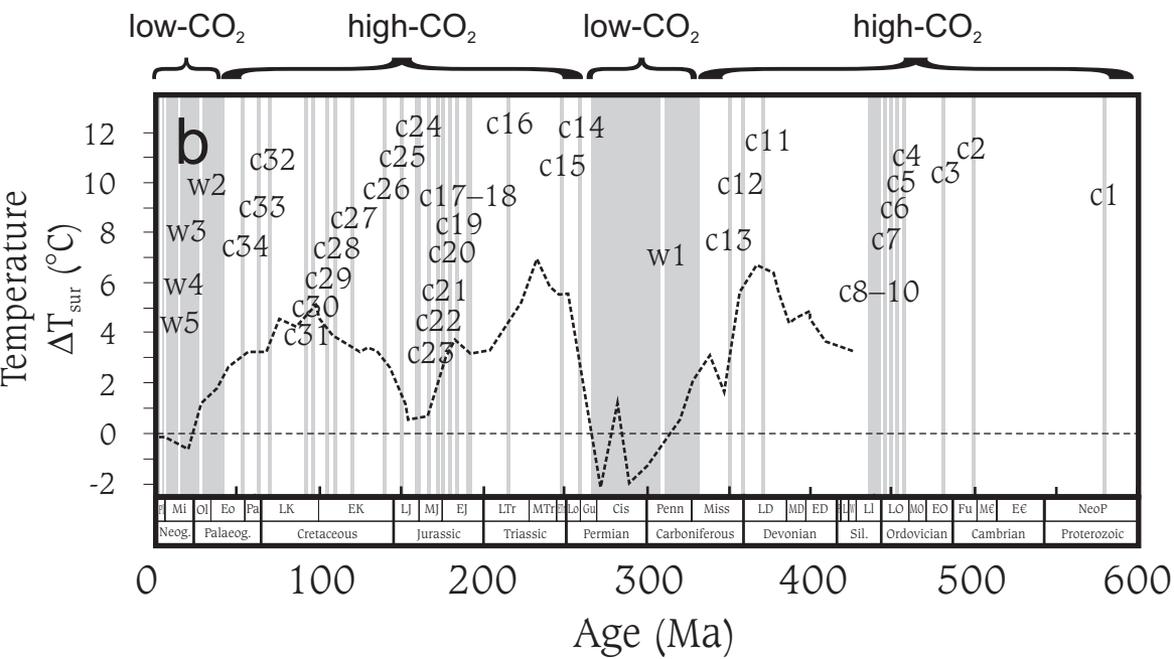
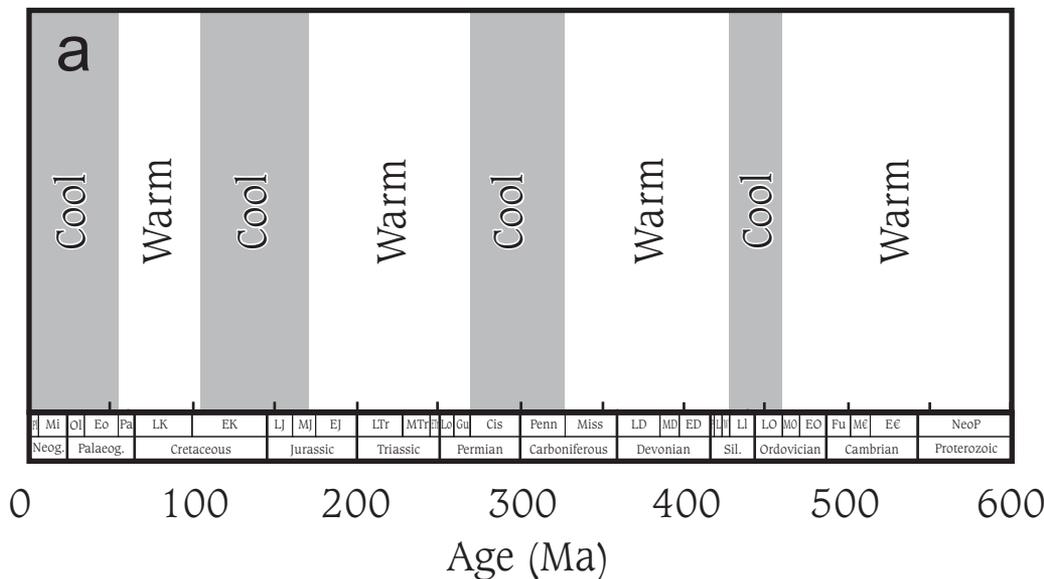
See text for explanation and bibliographic sources; ¹modified from Frakes *et al.* (1992); ²absolute ages from (Gradstein *et al.* 2004); ³as proposed here; ⁴codes in brackets indicate intervals as labelled on Fig. 1b.

Table 2: Palaeoclimatic proxies

Parameter	Proxy	Effective time range	Recent or key reference	
Temperature	$\delta^{18}\text{O}$	Archaean–Recent, but in practice only reliable from Late Cretaceous–Recent	Marshall (1992)	
	Foraminiferal Mg/Ca	Late Cretaceous–Recent	Elderfield <i>et al.</i> (2002)	
	Belemnite Mg/Ca	Jurassic–Cretaceous	Rosales <i>et al.</i> (2004)	
	Coralline Sr/Ca	Late Cretaceous–Recent	Wei <i>et al.</i> (2004)	
	Belemnite Sr/Ca	Jurassic–Cretaceous	Rosales <i>et al.</i> (2004)	
	Foraminiferal $\delta^{44}\text{Ca}$	Late Cretaceous–Recent	Gussone <i>et al.</i> (2003)	
	Alkenones unsaturation ratios UK'37	Miocene–Recent	Bard (2001)	
	Amino acid racemization epimerization in molluscs	Late Pleistocene–Recent	Murray-Wallace <i>et al.</i> (1988)	
	Crenarchaeotal tetraether lipids TEX86	Cretaceous–Recent	Schouten <i>et al.</i> (2002)	
	Carbonate facies analysis	Cambrian–Recent	Sellwood <i>et al.</i> (1993)	
	Pollen analysis using mutual climatic range and climatic amplitude techniques	Eocene–Recent	Jimenez-Moreno <i>et al.</i> (2005), Klotz <i>et al.</i> (2006)	
	Taxonomic methods based on plant macrofossils, vertebrates, molluscs and arthropods	Cretaceous–Recent	Baghai & Jorstad (1995), Friedman <i>et al.</i> (2003), Moine & Rousseau (2002), Williams & Eyles (1995)	
	Nearest living relative method on plant and animal taxa	Early Jurassic–Recent	Wang <i>et al.</i> (2005), Moe & Smith (2005)	
	Ecological diversity spectra of mammals	Pleistocene–Recent, but potentially applicable back to the Miocene	Fernandez & Pelaez-Campomanes (2005)	
	Mammal cenograms	Eocene–Recent	Legendre <i>et al.</i> (2005)	
	Transfer functions based on benthic ostracods, benthic and planktonic foraminifera, and diatom assemblages	Pliocene–Recent	Brouwers <i>et al.</i> (1991), Andersson (1997), Zielinski <i>et al.</i> (1998)	
	Molluscan faunal assemblage analysis	Palaeocene–Recent	Kafanov & Volvenko (1997)	
	Climate-leaf analysis multivariate program (CLAMP)	Permian–Recent	Glasspool <i>et al.</i> (2004)	
	Width of growth rings in fossil wood	Carboniferous–Recent	Poole <i>et al.</i> (2005)	
	Maximum latewood density of tree rings	Recent	Briffa <i>et al.</i> (2004)	
	Salinity	$\delta^{18}\text{O}$	Triassic–Recent	Korte <i>et al.</i> (2005)
		U/Ca ratios in corals	Pleistocene–Recent	Ourbak <i>et al.</i> (2006)
		Alkenones unsaturation ratios UK'37	Miocene–Recent	Schouten <i>et al.</i> (2006)
		Taxon analysis of chironomid midges	Pleistocene–Recent, applicable only to lakes	Walker (1991)
		Faunal assemblage analysis of insects	Cretaceous–Recent	Coram & Jarzembowski (2002)
		Transfer functions based on benthonic foraminifera	Pleistocene–Recent	Sejrup <i>et al.</i> (2004)
		Transfer functions based on diatom assemblages	Holocene, applicable only in lakes	Roberts & McMinn (1999)
Zeolite analysis of sub-glacial volcanic rocks		Miocene–Recent (potentially much further back)	Johnson & Smellie (in press)	
Circulation of oceans		$\delta^{13}\text{C}$	Pleistocene–Recent	Lynch-Stieglitz & Fairbanks (1994)
		Foraminiferal Zn/Ca	Pleistocene–Recent	Marchitto <i>et al.</i> (2000)
	Foraminiferal Cd/Ca	Pleistocene–Recent	Boyle (1988)	
	$^9\text{Be}/^{10}\text{Be}$ in manganese crusts	Holocene	vonBlankenbourg <i>et al.</i> (1996)	
	Nd and Pb isotopes in manganese crusts	Miocene–Recent	Ling <i>et al.</i> (1997)	
	Hf isotopes in manganese crusts	Miocene–Recent	David <i>et al.</i> (2001)	
	Nd isotopes in foraminifera	Pleistocene–Recent	Vance & Burton (1999)	
	Ag/Si ratios	Recent	Zhang <i>et al.</i> (2004)	
	^{14}C	Late Pleistocene–Recent	Henderson (2002)	
	^{231}Pa and ^{230}Th	Pleistocene–Recent	Henderson (2002)	
	Sortable silt	Palaeocene–Recent, but potentially back to the Jurassic	Pfuhl & McCave (2005)	
	Biogenic barite	Pleistocene–Recent	Gingele & Dahmke (1994)	
	Solubility of Th, Pa, and Be	Pleistocene–Recent	Chase <i>et al.</i> (2002)	
Productivity	Sedimentary U concentration	Pleistocene–Recent	Francois <i>et al.</i> (1997)	
	$\delta^{66}\text{Zn}$ ratio of manganese crusts	Holocene	Maréchal <i>et al.</i> (2000)	
	Zn/Si ratios in deep sea hexactinellid sponges	Late Pleistocene–Recent	Ellwood <i>et al.</i> (2004)	
	Mo/Al ratios in black shales	Early Ordovician–Recent	Wilde <i>et al.</i> (2004)	
	REE in foraminifera	Untested	Haley <i>et al.</i> (2005)	
	Transfer function based on carbonate mass accumulation rates of nanoplankton	Palaeocene–Recent	Siesser (1995)	
	Acritarch diversity analysis	Proterozoic–Permian	Vecoli & Le Herisse (2004)	
	Cd/Ca ratio in foraminifera	Pleistocene–Recent	Rickaby & Elderfield (1999)	
	Nutrient utilization			

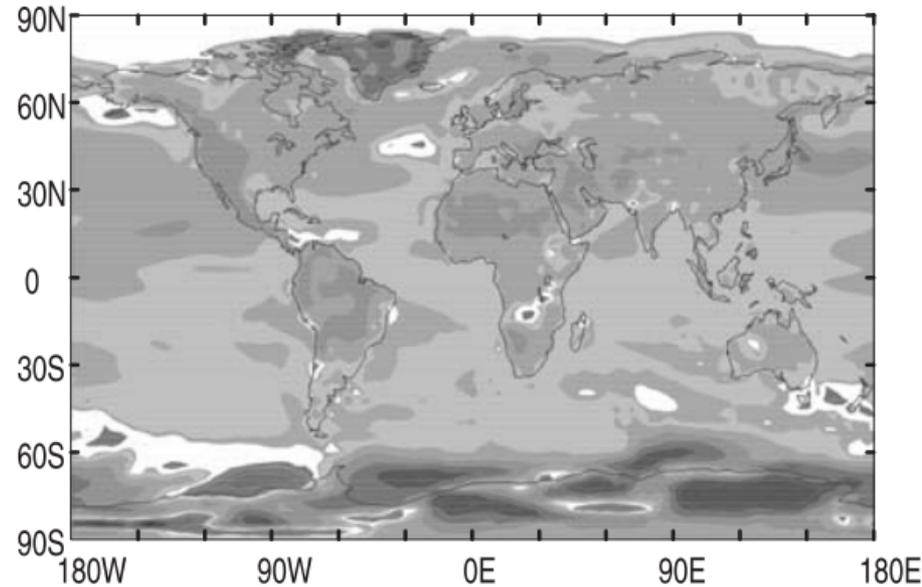
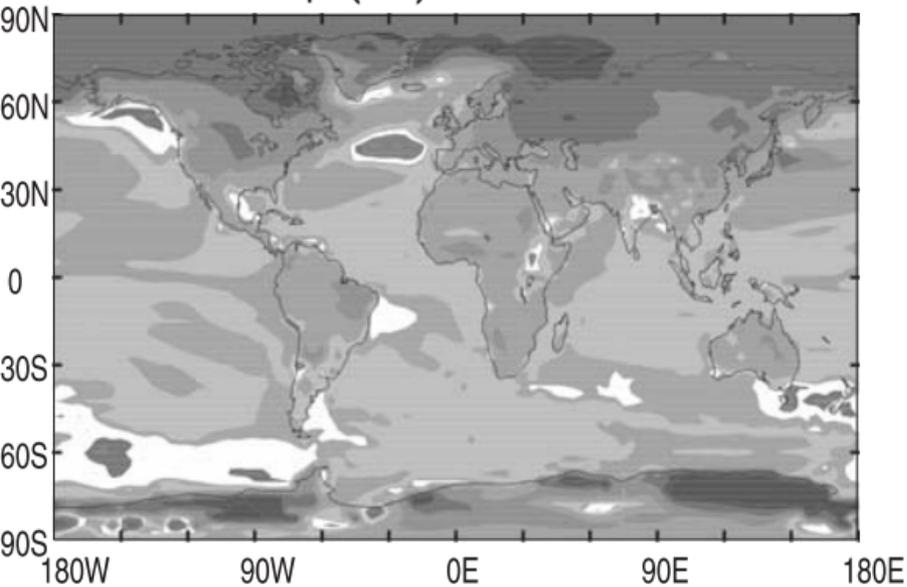
	Cd/P ratio in phytoplankton particulate matter	Pleistocene–Recent	Cullen & Sherrell (2005)
	$\delta^{15}\text{N}$	Early Jurassic–Recent	Robinson <i>et al.</i> (2005)
	$\delta^{30}\text{Si}$ of diatom silica	Pleistocene–Recent	Reynolds <i>et al.</i> (2006)
	Transfer functions based on lake diatom assemblages	Early Holocene–Recent, applicable only to lakes	Hausmann & Kienast (2006)
<i>Carbonate alkalinity/weathering fluxes</i>	Foraminiferal Ba/Ca ratios	Pleistocene–Recent	Lea (1993)
	$^{87}\text{Sr}/^{86}\text{Sr}$ in carbonate	Cambrian–Recent	Dessert <i>et al.</i> (2001)
	$^{187}\text{Os}/^{186}\text{Os}$ in mudrocks	Jurassic–Recent	Cohen <i>et al.</i> 2004
	$^{87}\text{Sr}/^{86}\text{Sr}$ curve	Cambrian–Recent	McArthur <i>et al.</i> (2001)
	Ge/Si ratios in diatom silica	Pleistocene–Recent	Jones <i>et al.</i> (2002)
	Hf and Nd isotope ratio time trajectories in manganese crusts	?Miocene–Recent	van de Flierdt <i>et al.</i> (2002)
	Clay mineralogy in ocean sediments	Triassic–Recent	Thiry (2000)
	Mass of individual foraminifera of a particular size	Late Pleistocene–Recent	Lohmann (1995)
<i>pH</i>	$\delta^{11}\text{B}$ of foraminiferal and other carbonate	Neoproterozoic–Recent, but in practice only reliable from Palaeocene–Recent	Pearson & Palmer (2000)
	calcium-ion concentration of seawater and modelled atmospheric CO_2 concentrations	Cambrian–Recent	Royer <i>et al.</i> (2004)
<i>Atmospheric CO_2</i>	$\delta^{13}\text{C}$ of alkenones	Eocene–Recent	Pagani <i>et al.</i> (2005b)
	$\delta^{13}\text{C}$ of other organic materials	Jurassic–Recent	Hesselbo <i>et al.</i> (2000)
	$\delta^{13}\text{C}$ of marine, freshwater or pedogenic carbonate	Devonian–Recent	Buggisch (1991), Yemane & Kelts (1996), Royer <i>et al.</i> (2001)
	$\delta^{11}\text{B}$ and $\delta^{44}\text{Ca}$ of carbonate calibrated with $\delta^{18}\text{O}$ and $\delta^{13}\text{C}$	Neoproterozoic–Recent	Demicco <i>et al.</i> (2003)
	Density of stomata in leaf cuticle	Carboniferous–Recent	Thorn & DeConto (2006)
<i>Precipitation/evaporation</i>	Evaporite facies analysis	Archaean–Recent	Ziegler (2003)
	Coal, lignite and peat facies analysis	Devonian–Recent	Sellwood <i>et al.</i> (1993)
	Fusain (fossil charcoal) in clastic sediments	Devonian–Recent	Scott (2000)
	Palaeosols	Cambrian–Recent	Retallack (2001)
	Transfer function from depth to nodular, pedogenic carbonate horizon	Palaeocene–Recent	Retallack (2005)
	Clay mineralogy in ocean sediments	Triassic–Recent	Ahlberg <i>et al.</i> (2003)
	Loess deposits	Palaeogene–Recent	Sun & An (2005)
	Pollen analysis using mutual climatic range and climatic amplitude techniques	Eocene–Recent	Jimenez-Moreno <i>et al.</i> (2005), Klotz <i>et al.</i> (2006)
	Ecological diversity spectra of mammals	Pleistocene–Recent, but potentially applicable back to the Miocene	Fernandez & Pelaez-Campomanes (2005)
<i>Atmospheric circulation</i>	Mammal cenograms	Eocene–Recent	Legendre <i>et al.</i> (2005)
	Tempestite facies analysis	Neoproterozoic–Recent	Agustsdottir <i>et al.</i> (1999)
	Windblown dust in marine sediments, loess deposits	Palaeogene–Recent	Kohfeld & Harrison (2001)
	Dunes, windblown trees	Palaeozoic–Recent	Segalen <i>et al.</i> (2004)
<i>Glaciation</i>	Aeolianite facies analysis	Devonian–Recent	Le Guern & Davaud (2005)
	Glacial sediments	Archaean–Recent	Sellwood <i>et al.</i> (1993)
	Correlation of $\delta^{18}\text{O}$ and $\delta^{13}\text{C}$ isotope excursions and sea-level falls	Neoproterozoic–Recent	Miller <i>et al.</i> (2005)
	Glendonite carbonate nodules	Neoproterozoic–Recent	Swainson & Hammond (2001)
	Sub-glacial volcanic deposits	Miocene–Recent	Smellie <i>et al.</i> (2006)
	Gibbsite concentration in soils	Pleistocene–Recent	Ballantyne <i>et al.</i> (2006)
<i>Sea-ice cover</i>	C-25 highly branched isoprenoid alkenes IP25	Holocene	Belt <i>et al.</i> in press
	Transfer functions based on dinoflagellate cyst assemblages	Late Pleistocene–Recent	Peyron & De Vernal (2001)
	Diatom faunal assemblages	Late Pleistocene–Recent	Gersonde & Zielinski (2000)
	Similarity maximum modern analog techniques on foraminiferal assemblages	Late Pleistocene–Recent	Sarnthein <i>et al.</i> (2003)
	Modern analogue techniques based on diatom assemblages	Late Pleistocene–Recent	Crosta <i>et al.</i> (1998)

See text for explanation



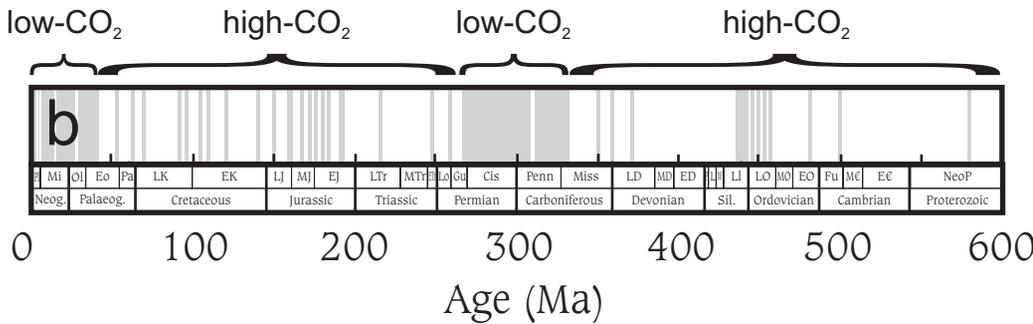
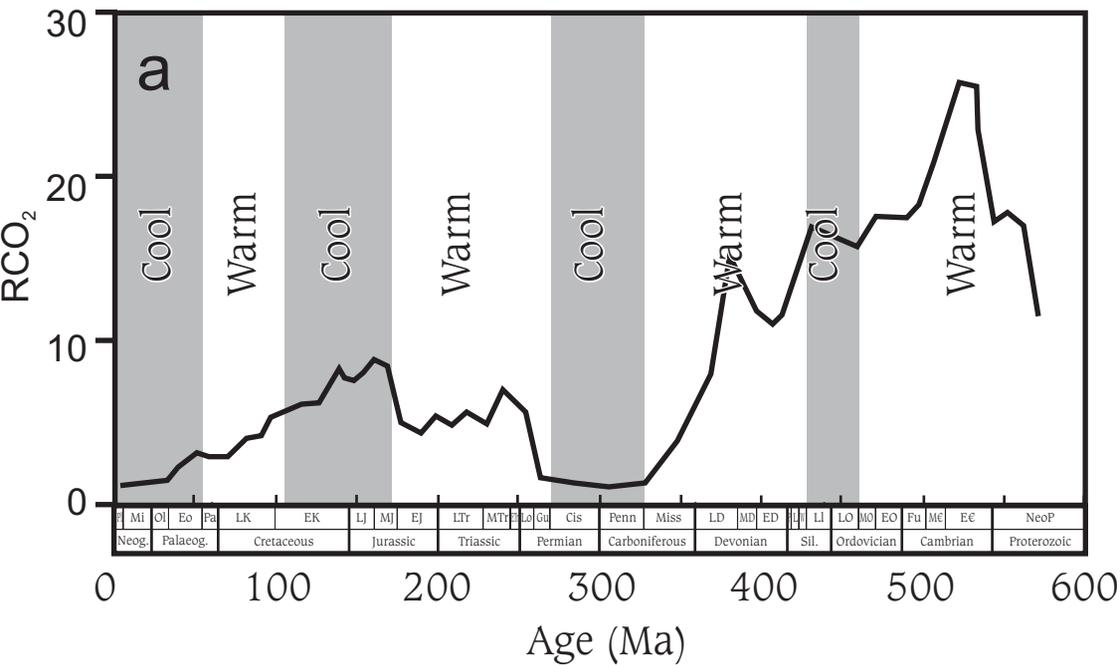
Vaughan Figure 2

Surface Temp ($^{\circ}\text{C}$) Plio^{Control} minus Pre-Ind DJF & JJA

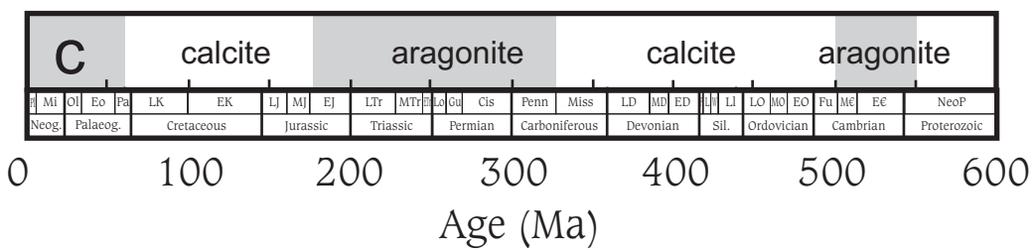
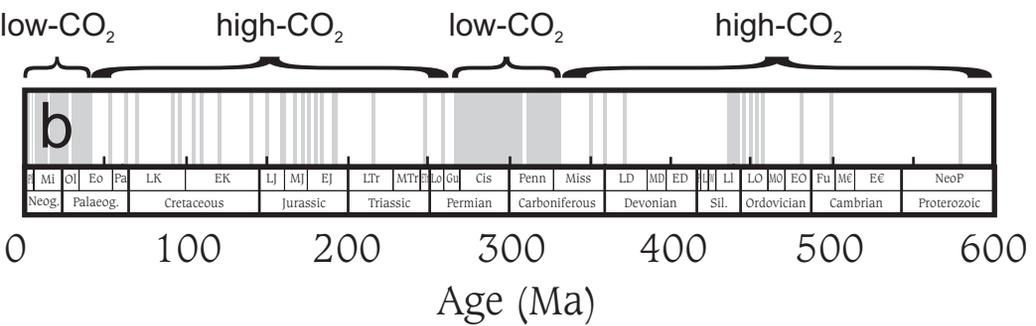
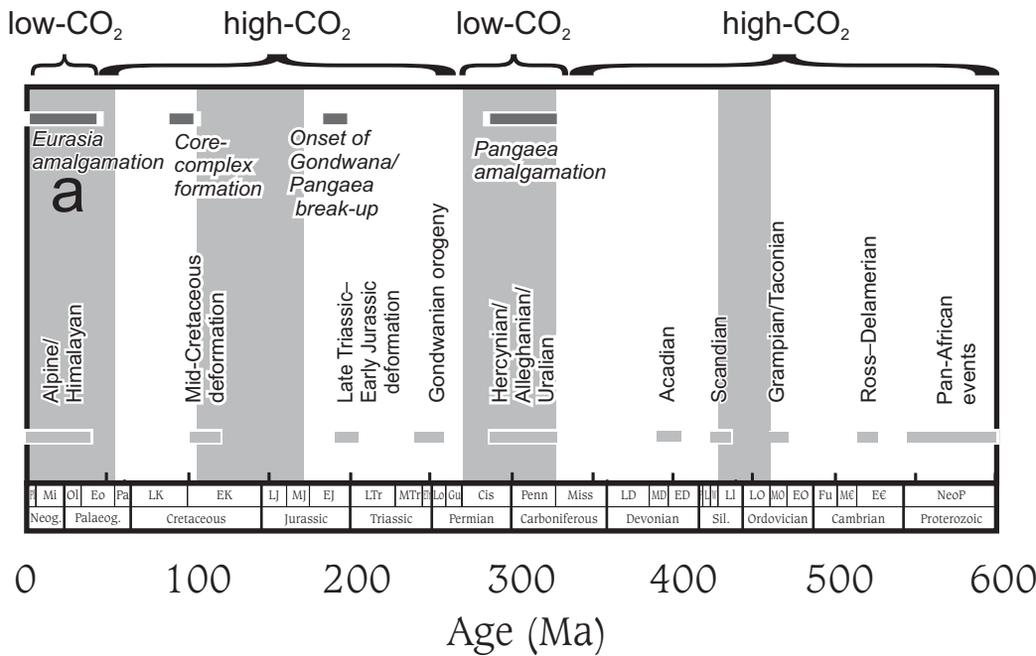


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Vaughan Figure 3

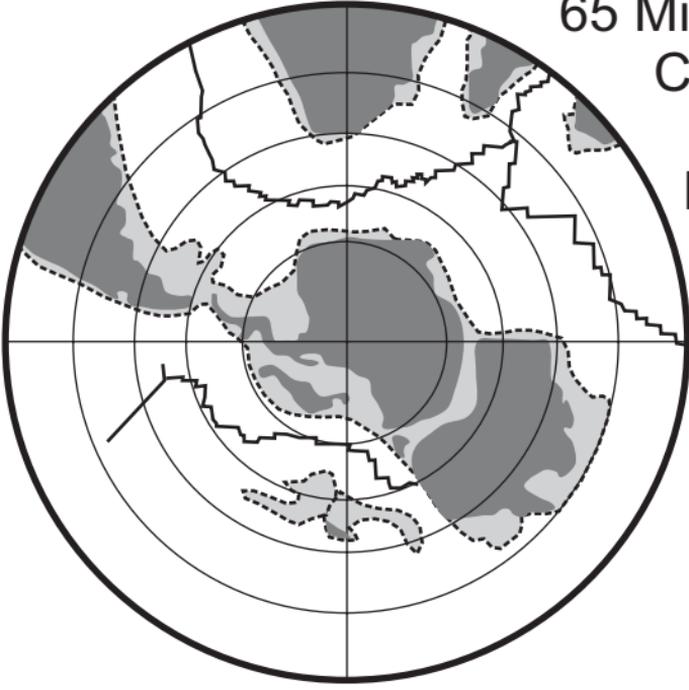


Vaughan Figure 4

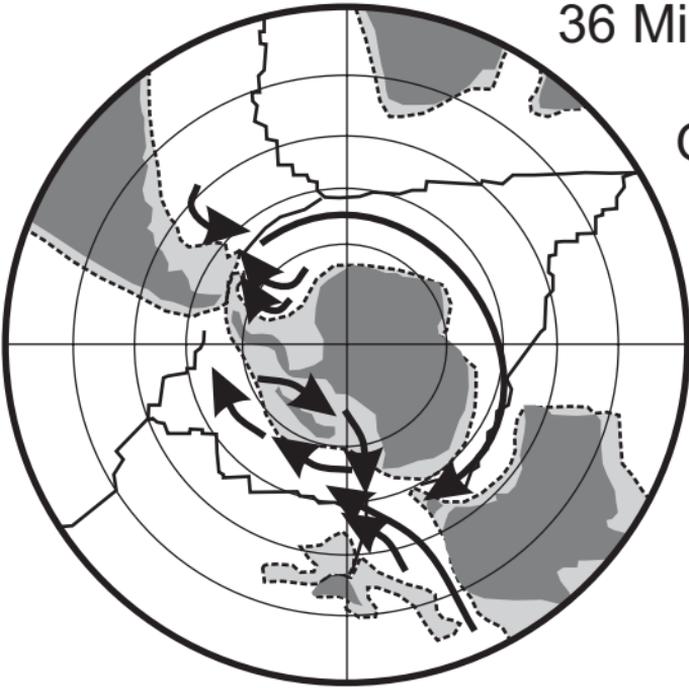


Vaughan Figure 5

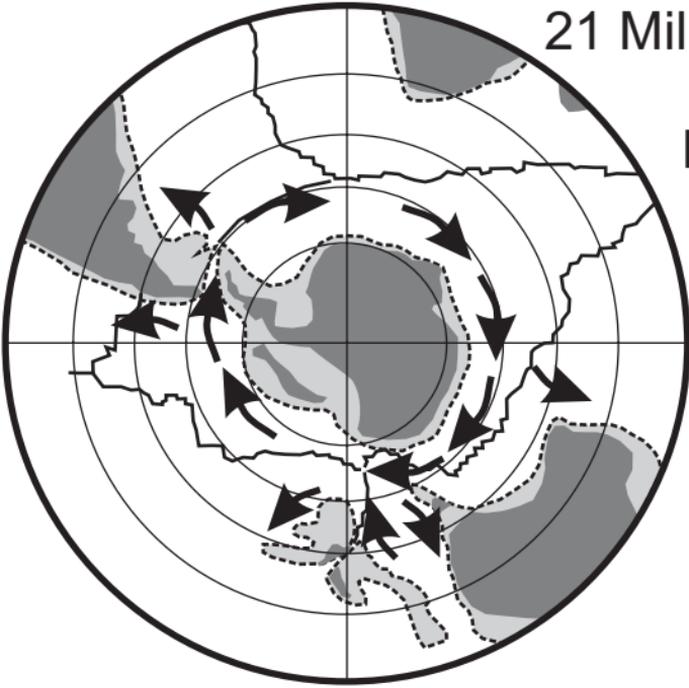
65 Million years ago
Cretaceous/
Tertiary
Boundary



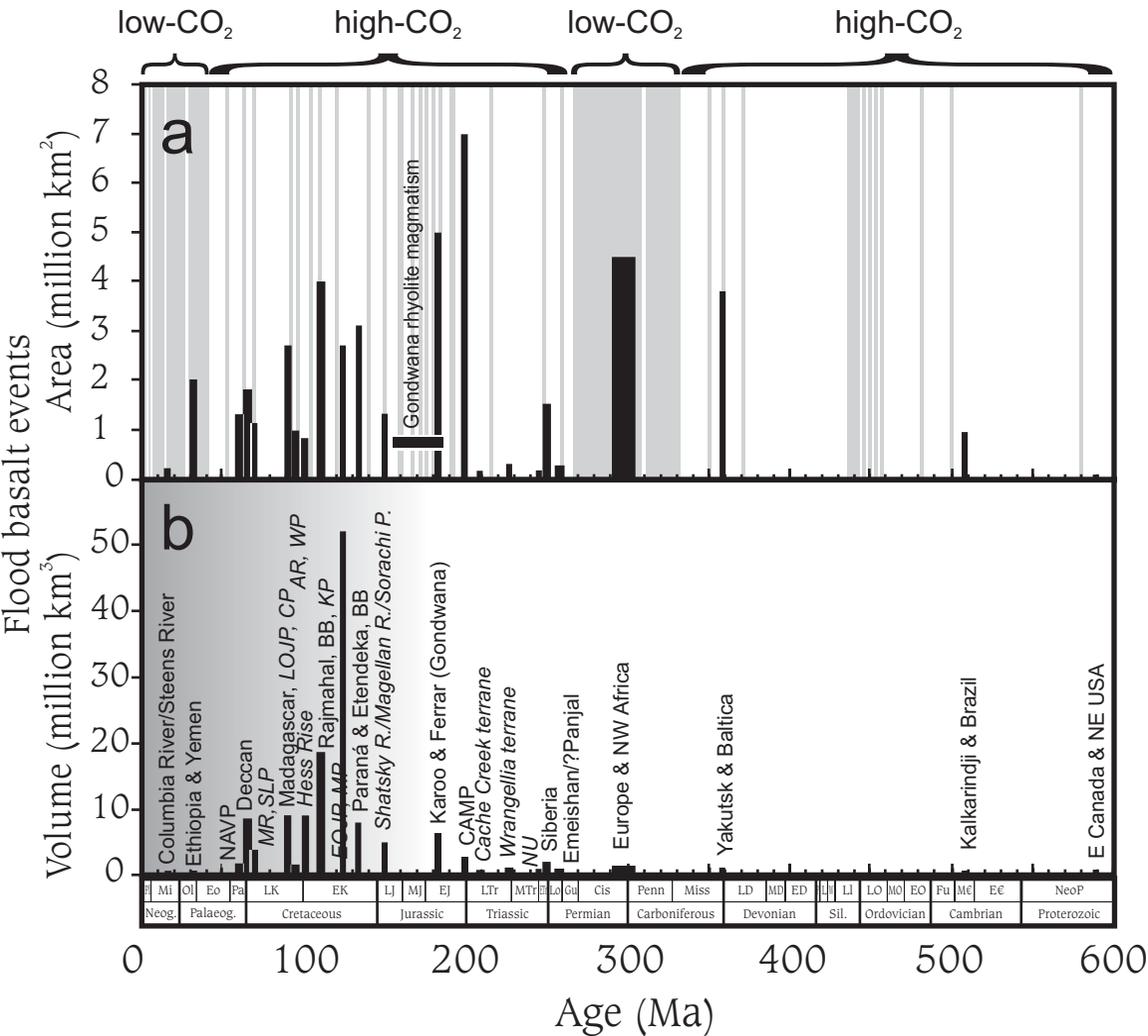
36 Million years ago
early
Oligocene



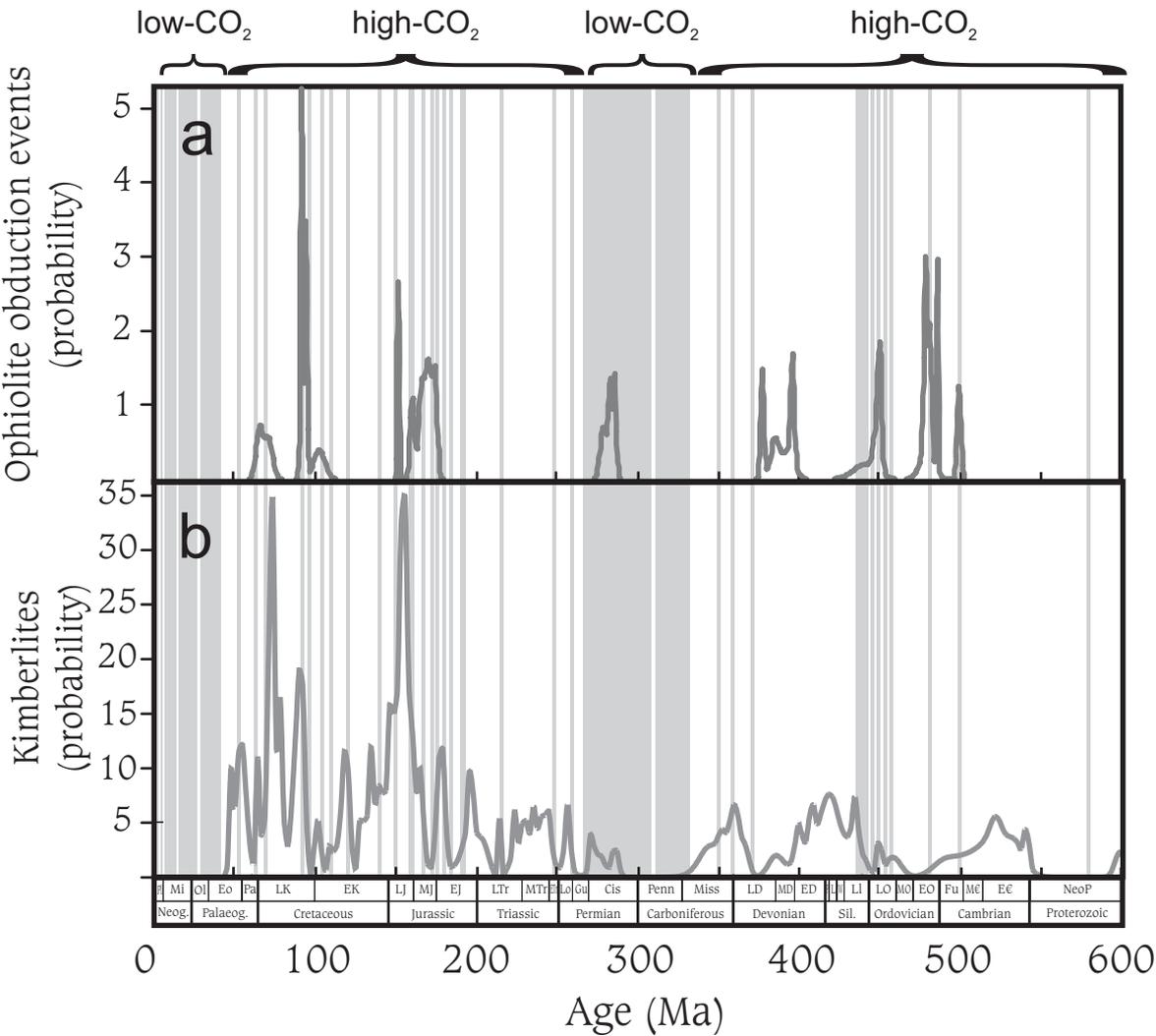
21 Million years ago
earliest
Miocene



Vaughan Figure 6



Vaughan Figure 7



Vaughan Figure 8

