BRITISH GEOLOGICAL SURVEY

Hartland Observatory

Monthly Magnetic Bulletin

August 2011

HARTLAND POINTS

I I/08/HA









HARTLAND OBSERVATORY MAGNETIC DATA

1. Introduction

Hartland o bservatory is one of three geomagnetic observatories in the UK operated and maintained by the British Geological Survey (BGS).

This bulletin is published to provide rapid access to the pr ovisional g eomagnetic o bservatory results. The i nformation is f reely available for personal, academic, ed ucational and non-commercial research or use. Magnetic o bservatory data are presented as a series of plots of one-minute, hourly and daily values, followed by tabulations of monthly values, reports of rapid variations and geomagnetic activity indices. The operation of the observatory and presentation of data are described in the rest of this section.

Enquiries about the data should be addressed to:

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British Geological Survey
Murchison House, West Mains Road
Edinburgh EH9 3LA
Scotland, UK

Tel: +44 (0) 131 667 1000 Fax: +44 (0) 131 650 0265 E-mail: enquiries@bgs.ac.uk Internet: www.geomag.bgs.ac.uk

2. Position

The observatory is situated on the NW boundary of the v illage of H artland i n N orth D evon. The observatory co-ordinates are:

Geographic: 50.995°N 355.516°E Geomagnetic: 53.737°N 80.253°E Height above mean sea level: 95 m

The geomagnetic co-ordinates are approximations, calculated using the 11th generation International Geomagnetic Reference Field (IGRF) at ep och 2011.5. On-line access to models (including IGRF), charts and n avigational d ata ar e av ailable at http://www.geomag.bgs.ac.uk/data_service/models-compass/home

3. The Observatory Operation

3.1 GDAS

The observatory operates under the control of the Geomagnetic Data Acquisition System (GDAS), which was developed by BGS staff, i nstalled in

2002, a nd be came f ully ope rational in January 2003. The data a cquisition s oftware, r unning on QNX operated computers, controls the data logging and the communications.

There are t wo set s of sen sors u sed for making magnetic m easurements. A tri-axial l inear-core fluxgate magnetometer, m anufactured by t he Danish Meteorological Institute, is used to measure the variations in the horizontal (H) and vertical (Z) components of the field. The thirds ensor is oriented perpendicular to these, and measures variations, which are proportional to the changes in declination (D). Measurements are made at a rate of 1 Hz

In addition to the fluxgate sensors there is a proton precession m agnetometer (PPM) making measurements of the ab solute total field intensity (*F*) at a rate of 0.1Hz.

The raw unfiltered data are retrieved automatically via I nternet c onnections t o t he B GS of fice i n Edinburgh in near real-time. The fluxgate data are filtered to produce one-minute values using a 61-point cosine filter a nd the to tal field in tensity samples are filtered using a 7-point cosine filter. The one-minute values provide input for various data products, available on-line at

www.geomag.bgs.ac.uk/data_service/home

3.2 Back-up Systems

There are two other fully independent identical systems, GDAS 2 and GDAS 3, ope rating at the observatory. The data from these are also processed in near real-time and used for quality control purposes. They are also used to fill any gaps or replace any corrupt values in the primary system, GDAS 1.

3.3 Absolute Observations

The GDAS fluxgate magnetometers accu rately measure v ariations i in the components of the geomagnetic field, but not the absolute magnitudes. Two sets of absolute measurements of the field are made manually once per week. A fluxgate sensor mounted on a theodolite is used to determine D and inclination (I); the G DAS PPM measurements, with a site difference correction applied, are used for F. The absolute observations are used in conjunction with the G DAS v ariometer measurements to produce a continuous record of the absolute v alues of the geomagnetic field

elements as i ft hey h ad been m easured at t he observatory reference pillar.

4. Observatory Results

The data presented in the bulletin are in the form of plots a nd t abulations described in the following sections.

4.1 Absolute Observations

The absolute observation measurements made during the month are tabulated. Also included are the corresponding baseline values, which are the differences between the absolute measurements and the variometer measurements of D, H and Z (in the sense absolute—variometer). These are also plotted (markers) along with the derived preliminary daily baseline values (line) throughout the year. Daily mean differences between the measured absolute F and the F computed from the baseline corrected H and H values are plotted in the fourth panel (in the sense measured—derived). The bottom panel shows the daily mean temperature in the fluxgate chamber.

4.2 Summary magnetograms

Small-scale magnetograms are plotted which allow the month's data to be viewed at a glance. They are plotted 16 days to a page and show the one-minute variations in D, H and Z. The scales are shown on the right-hand side of the page. On disturbed days the scal es ar e m ultiplied b y a factor, which is indicated above t he pa nel f or t hat da y. T he variations are centred on t he monthly mean value, shown on the left side of the page.

4.3 Magnetograms

The daily magnetograms are plotted using oneminute values of D, H and Z from the fluxgate sensors, with any gaps filled using back-up data. The magnetograms are plotted to a variable scale; scale bars are shown to the right of each plot. The absolute level (the monthly mean value) is indicated on the left side of the plots.

4.4 Hourly Mean Value Plots

Hourly mean values of *D*, *H* and *Z* for the past 12 months a re pl otted i n 27 -day seg ments corresponding to the Bartels solar rotation number. Magnetic d isturbances asso ciated w ith act ive regions and/or coronal holes on the Sun may recur after 27 days: the same is true for geomagnetically quiet intervals. Plotting the d ata in this way highlights t his r ecurrence. D iurnal v ariations are also clear in these plots and the amplitude changes

throughout the year highlight the seasonal changes. Longer term secular variation is also illustrated.

Full lists of t he U K obs ervatory hour ly m ean values from 1983 to the present day are available at www.geomag.bgs.ac.uk/data_service/data/obs_data/hourly_means

4.5 Daily and Monthly Mean Values

Daily mean values of D, H, Z and F are plotted throughout the year. In addition, a table of monthly mean values of all the geomagnetic elements is provided. These values depend on a ccurate specification of the fluxgate sensor baselines. It is anticipated that these provisional values will not be altered by more than a few nT or tenths of arcminutes before being made definitive at the end of the year.

4.6 Rapid Variations

Charged particles stream from the Sun in the solar wind. T he s olar w ind interacts with the geomagnetic field to create a cav ity, t he magnetosphere, in w hich t he f ield i s c onfined. When a region of enhanced velocity and/or density in the solar wind arrives at the dayside boundary of the magnetosphere (at a bout 10 e arth radii) the boundary is pushed towards the Earth. Currents set up on the boundary of the magnetosphere can cause an abrupt c hange i n t he g eomagnetic f ield measured on t he ground and this is recorded on observatory m agnetograms as a su dden i mpulse (si). If, following an si, there is a change in the rhythm of activity, the *si* is termed a storm sudden commencement (ssc). A classical magnetic storm exhibiting initial, main and recovery phases (shown by, for instance, the *Dst* ring current index) c an often occur after a ssc, in which case the start of the storm is taken as the time of the ssc.

Solar f lares, seen at o ptical w avelengths as a sudden brightening of a small region of the Sun's surface, ar e al so r esponsible f or i ncreased X -ray emissions. These X-rays cause increased ionisation in the ionosphere, w hich I eads t o a bsorption of short-wave radio signals. A solar flare effect (*sfe*), or "crochet", may be observed on a magnetogram during g eomagnetically q uiet tim es. I t is a relatively sh ort-term change (tens o f m inutes) t o the normal diurnal variation and can vary in size (tens of n T) depending on I ocal t ime (LT), geomagnetic latitude and solar zenith angle.

4.7 Local geomagnetic activity indices

The Observatory K index. T his su mmarises geomagnetic activity at an observatory by assigning a code, an integer in the range 0 t o 9, t o each 3-hour Universal Time (UT) interval. The index for each 3-hour UT interval is de termined from the maximum range in H or D (scaled in nT), with allowance m ade f or the regular (undisturbed) diurnal variation. The conversion from range to an index value is made u sing a quasi-logarithmic scale, with the scale values dependent on the geomagnetic latitude of the observatory. The lower bounds (in nT) for the classification of each period at Hartland are:

0	1	2	3	4	5	6	7	8	9
0	5	10	20	40	70	120	200	330	500

The *K* index r etains t he L T and seasonal dependence of activity associated with the position of the observatory. The 3-hourly *K* indices for the month are t abulated and al so p lotted as a histogram. All UK observatory *K* indices are available at

www.geomag.bgs.ac.uk/data_service/data/magnetic indices/k indices

4.8 Global geomagnetic activity indices

The aa index. A num ber of 3 -hour g eomagnetic indices are computed by combining *K* indices from networks of observatories to characterise global activity I evels and to el iminate L T and seaso nal effects. T he simplest of these is the aa index, computed using the K indices from two approximately antipodal observatories: Hartland in the UK and Canberra in Australia. The aa index is calculated from linearisations of the Hartland and Canberra K indices, and has units of nT. The 3hourly aa indices are tabulated along with the daily mean value of aa (denoted Aa), the mean values of aa for the intervals 00-12UT (Aa_{am}) and 12-24UT (Aa_{pm}) and the monthly mean value. The 3-hourly aa indices for the month are also plotted as a histogram.

Although the *aa* index is based on da ta from only two observatories, provided averages over 12 hours or longer are used, the index is strongly correlated with the *ap* and *am* indices, which are derived using d ata f rom m ore ex tensive observatory networks.

The *aa* indices listed in this bulletin are available at www.geomag.bgs.ac.uk/data_service/data/magneticindices/aaindex as well as the full data set from 1868.

Definitive *aa* are publ ished by the International Service for Geomagnetic I ndices, LATMOS, 4 Avenue de Neptune, F-94107 S aint M aur C edex, France.

5. Conditions of Use

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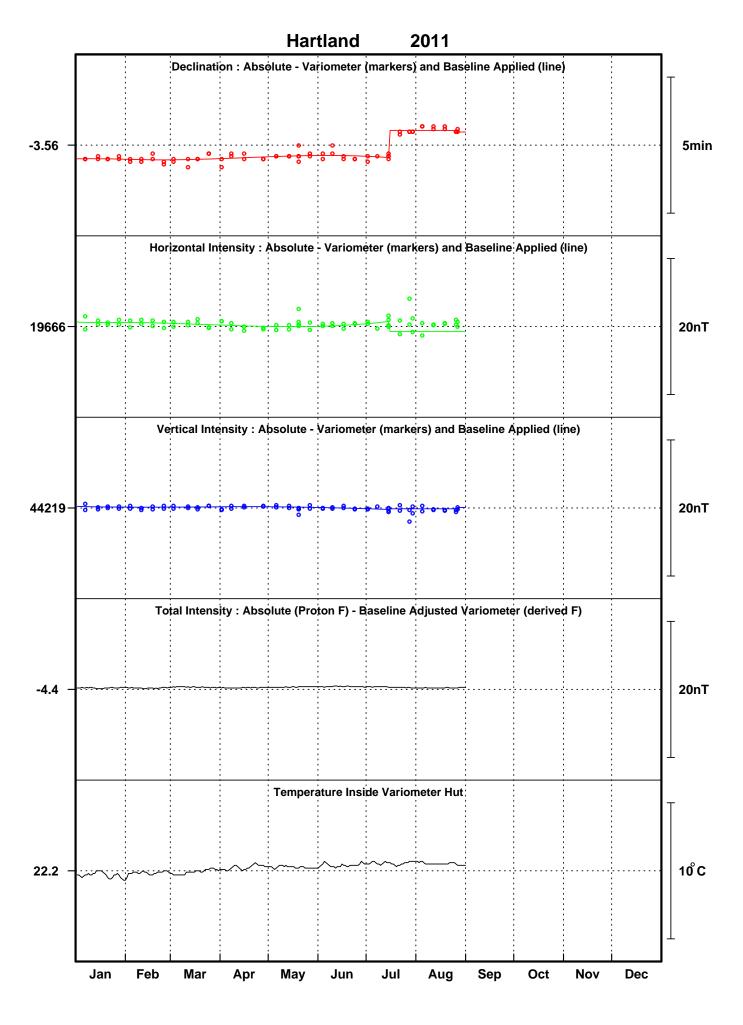
Commercial u sers can contact the geomagnetism team for information on the range of applications and ser vices of fered. Full contact details are available at www.geomag.bgs.ac.uk/contactus/staff

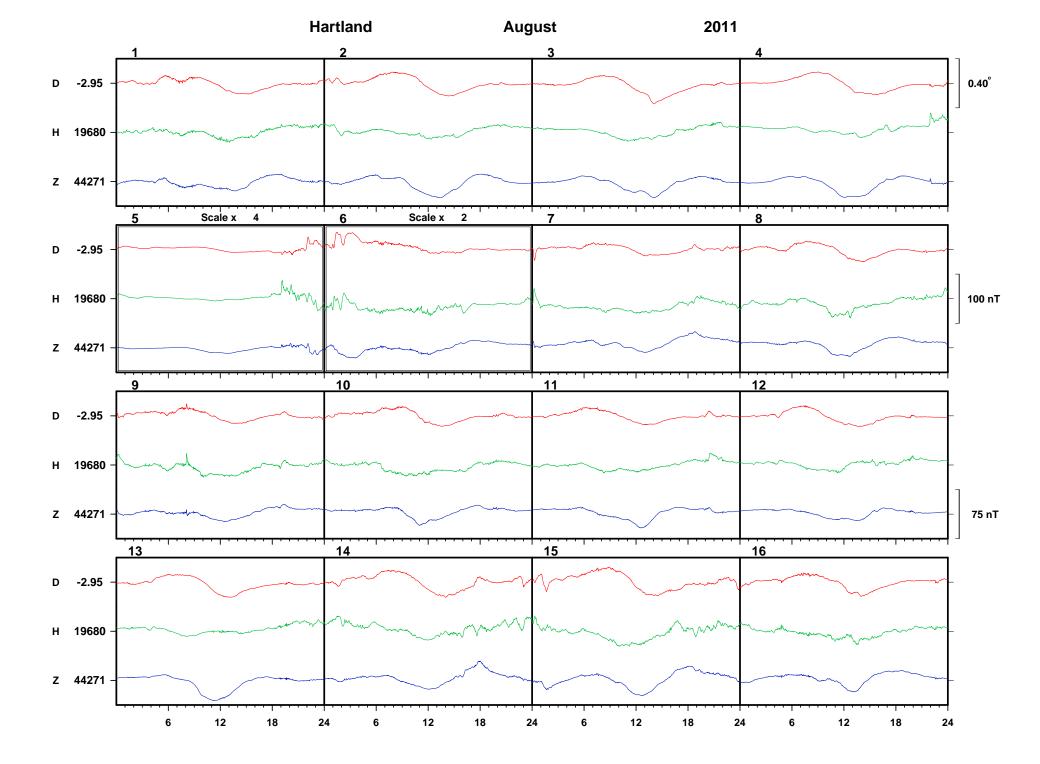
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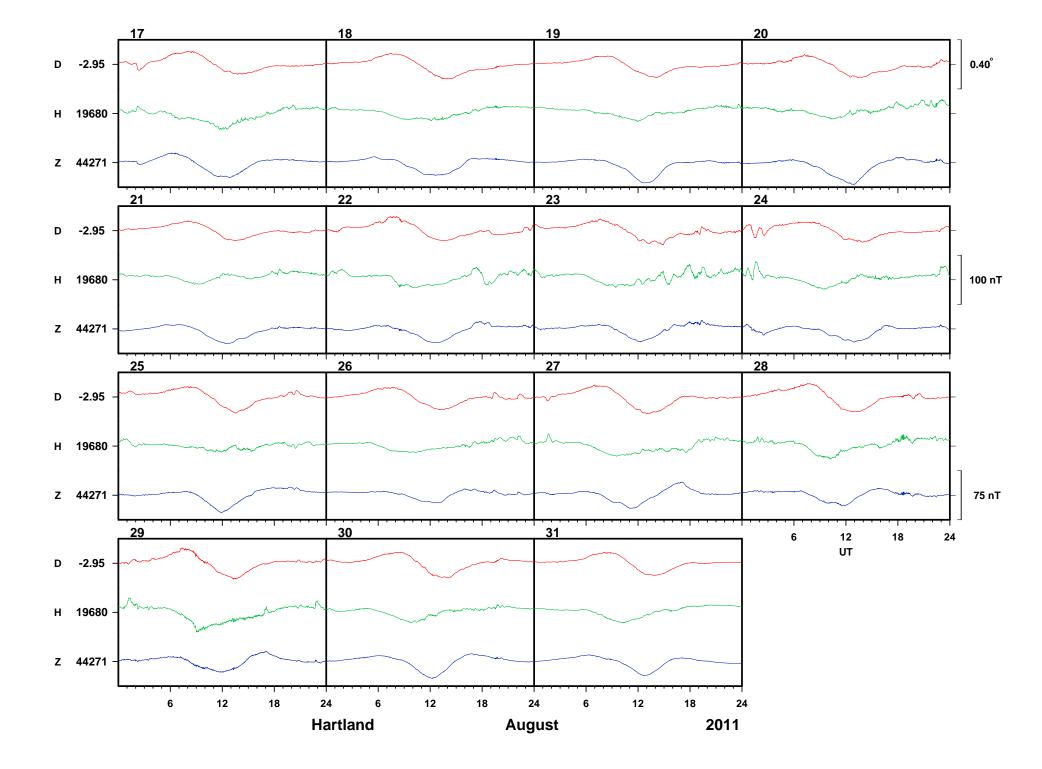
HARTLAND OBSERVATORY

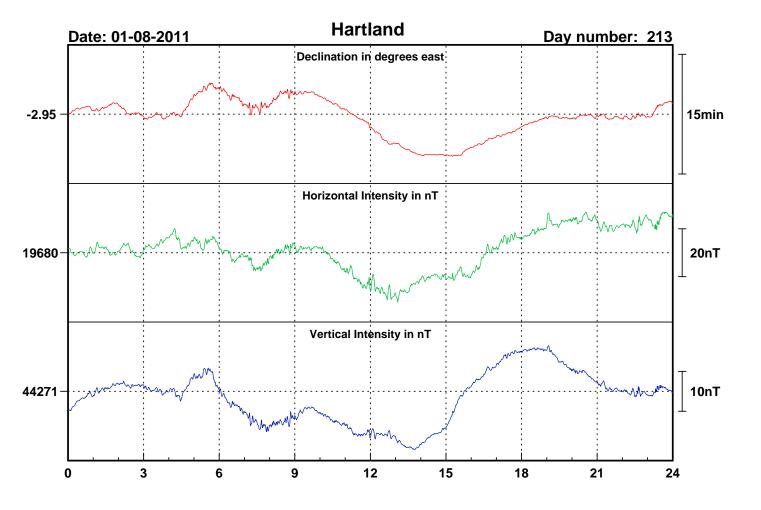
ABSOLUTE OBSERVATIONS

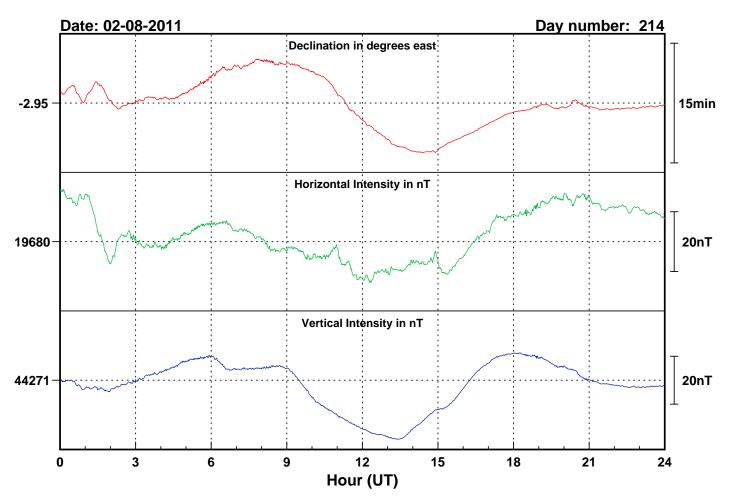
		Declination		Inclination		Total Field		Horizontal Intensity		Vertical Intensity			
Date	Day Number	Time (UT)	Absolute (°)	Baseline (°)	Time (UT)	Absolute (°)	Site difference (nT)	Absolute corrected (nT)	Absolute (nT)	Baseline (nT)	Absolute (nT)	Baseline (nT)	Observer
04-Aug-11	216	07:43	-2.8746	-2.9233	07:50	66.0332	4.4	48455.6	19683.0	19664.9	44277.8	44219.1	ST
04-Aug-11	216	07:56	-2.8708	-2.9233	08:03	66.0318	4.4	48454.1	19683.5	19666.7	44276.0	44218.3	ST
11-Aug-11	223	09:11	-2.9264	-2.9250	09:19	66.0434	4.4	48440.1	19668.8	19666.4	44267.1	44218.6	ST
11-Aug-11	223	09:25	-2.9321	-2.9233	09:34	66.0424	4.4	48440.6	19669.8	19666.5	44267.2	44218.5	ST
18-Aug-11	230	07:38	-2.8684	-2.9250	07:45	66.0402	4.4	48450.7	19675.6	19666.7	44275.7	44218.4	ST
18-Aug-11	230	07:51	-2.8662	-2.9233	07:59	66.0413	4.4	48449.6	19674.3	19666.6	44275.2	44218.5	ST
25-Aug-11	237	06:52	-2.8853	-2.9267	07:15	66.0345	4.4	48452.6	19680.8	19667.2	44275.5	44218.2	ST
25-Aug-11	237	07:06	-2.8827	-2.9267	07:13	66.0359	4.4	48452.5	19679.7	19666.5	44275.9	44218.5	ST
26-Aug-11	238	08:09	-2.8794	-2.9267	08:24	66.0452	4.4	48445.6	19669.7	19666.9	44272.8	44218.6	CWT
26-Aug-11	238	08:37	-2.8920	-2.9250	08:50	66.0459	4.4	48445.8	19669.2	19666.2	44273.2	44218.9	CWT

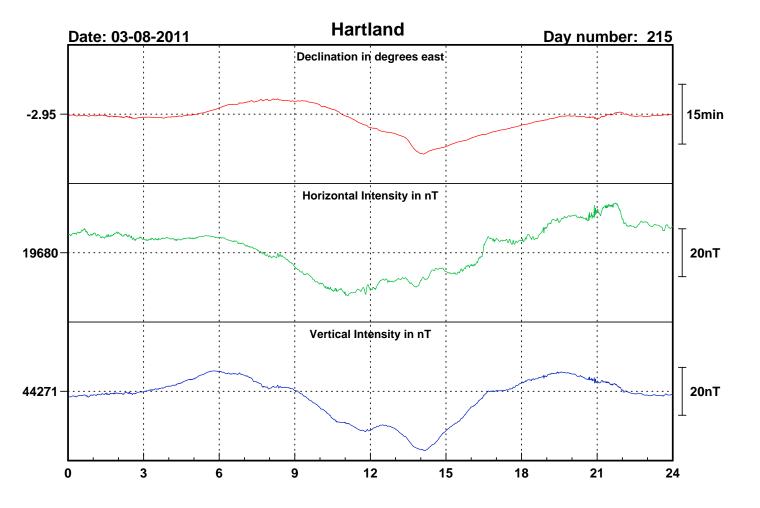


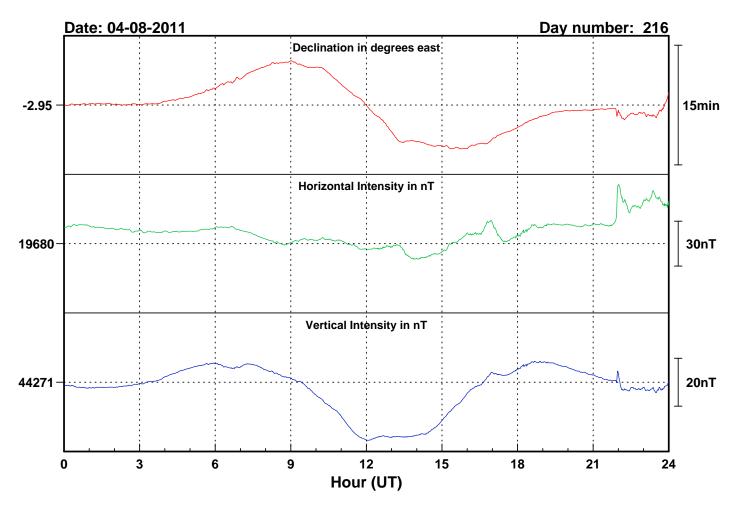


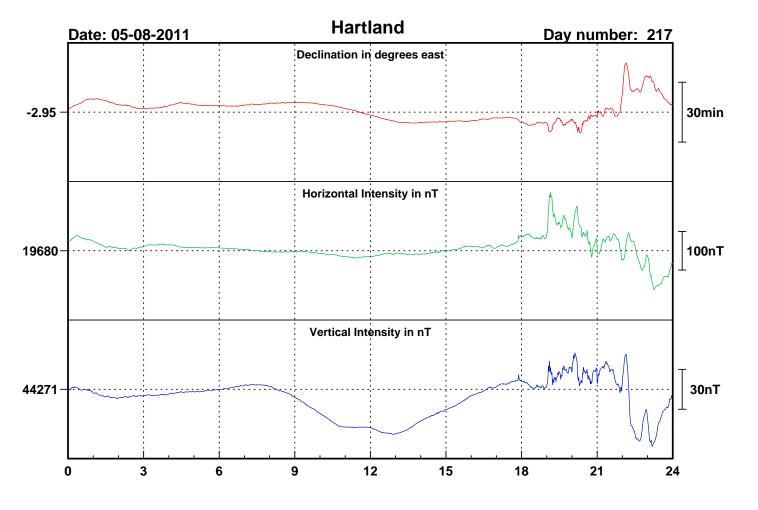


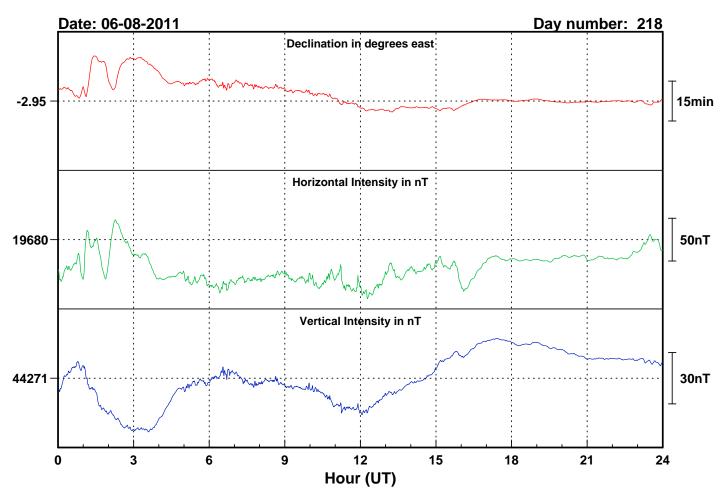


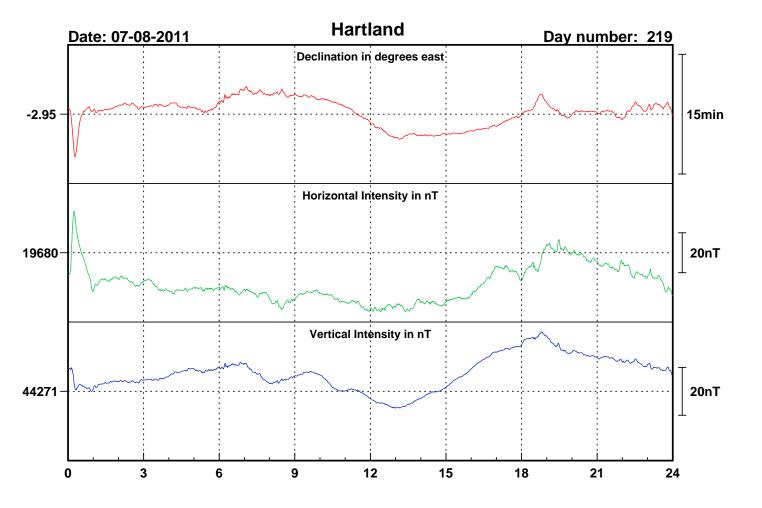


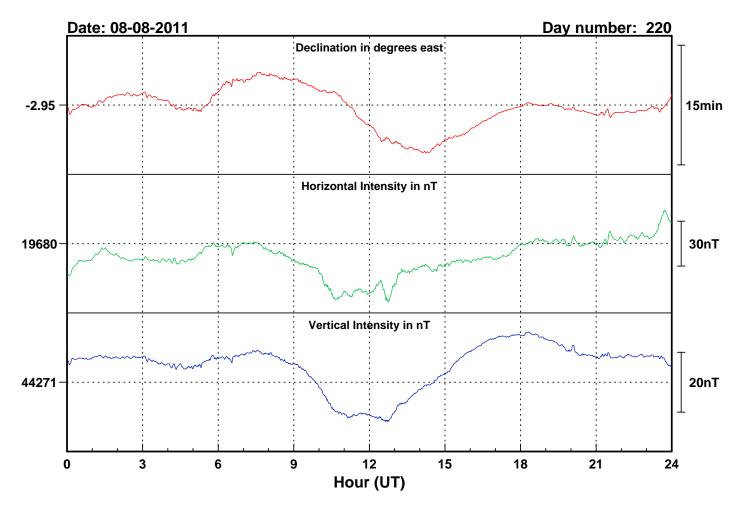


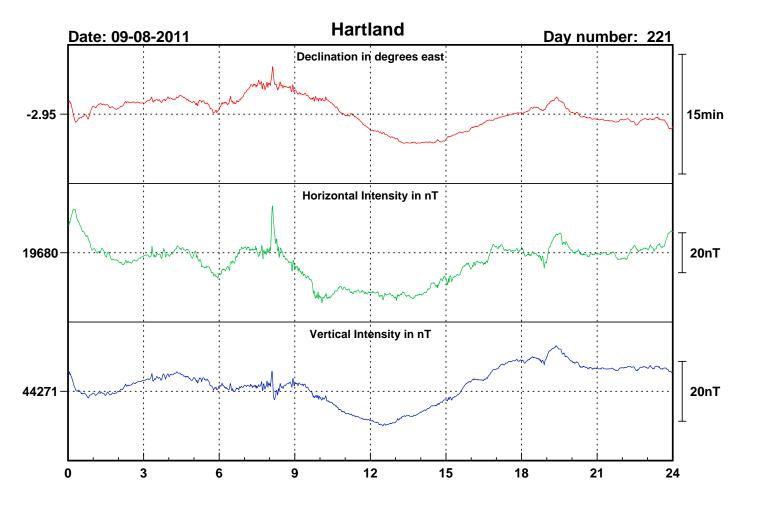


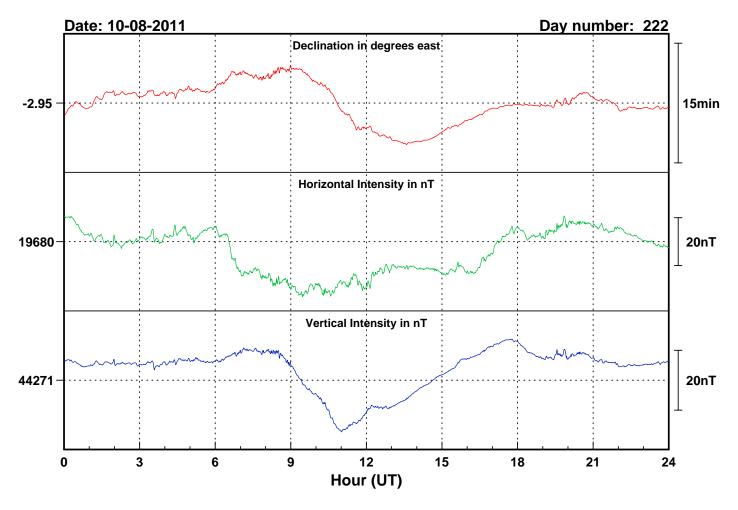


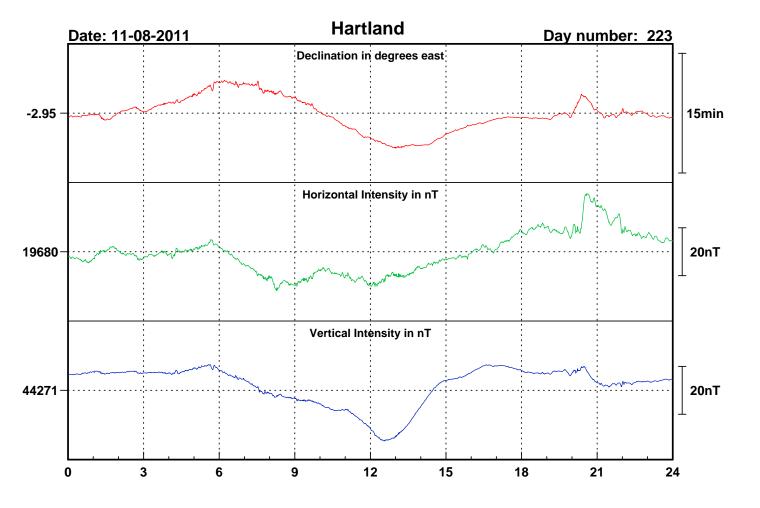


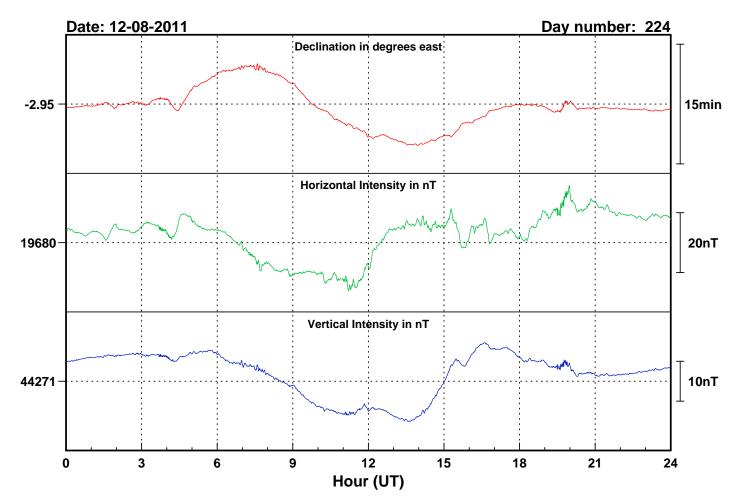


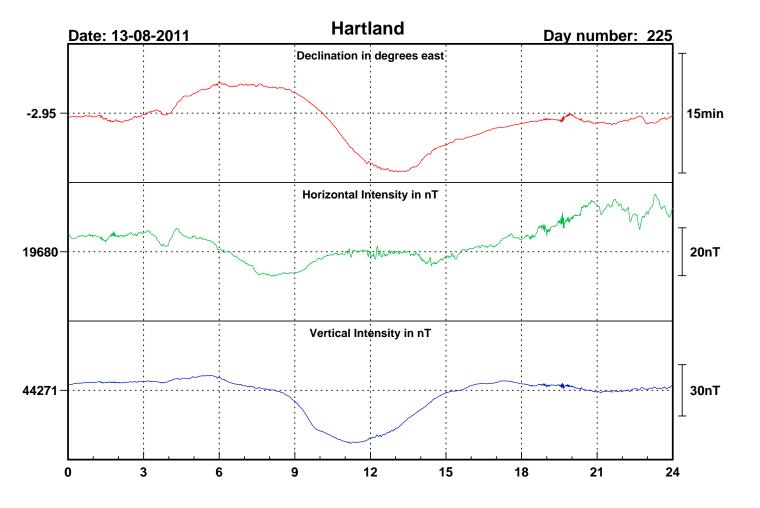


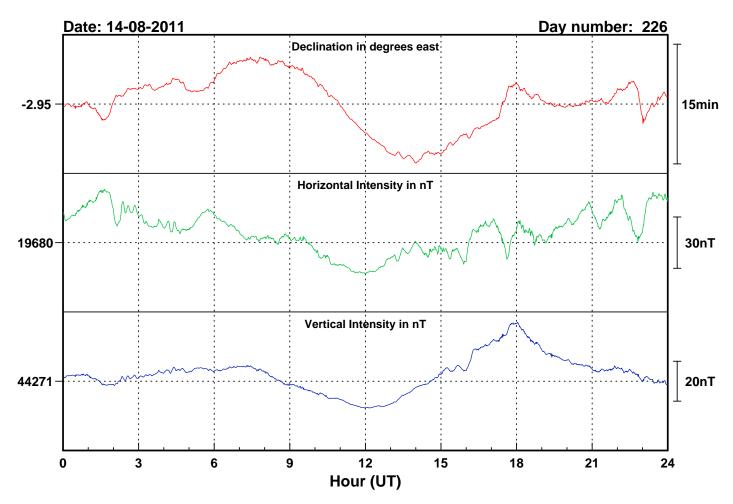


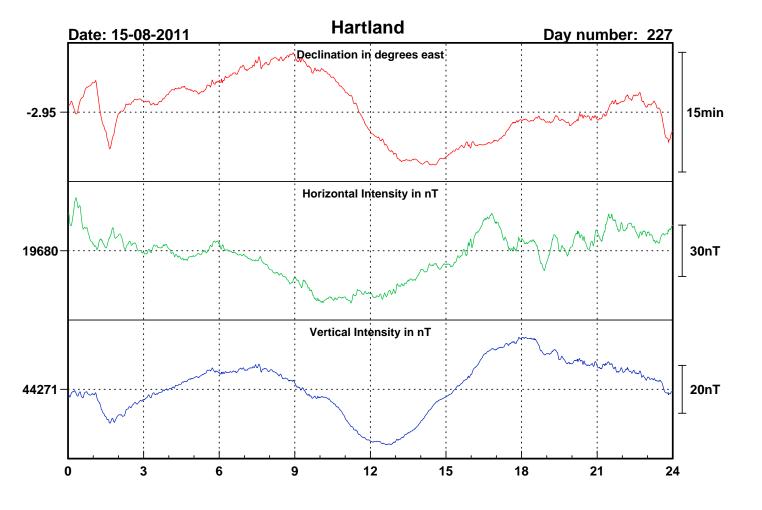


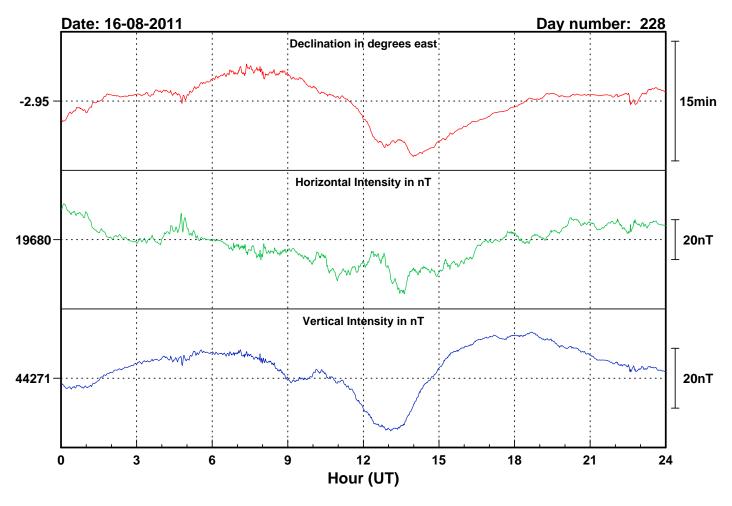


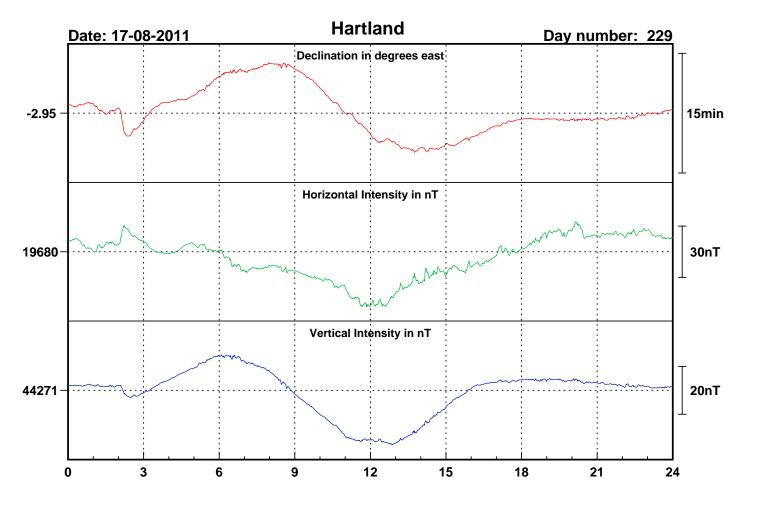


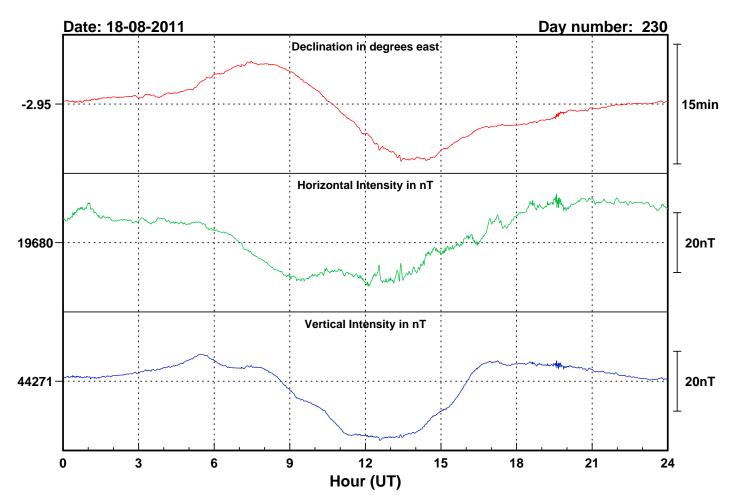


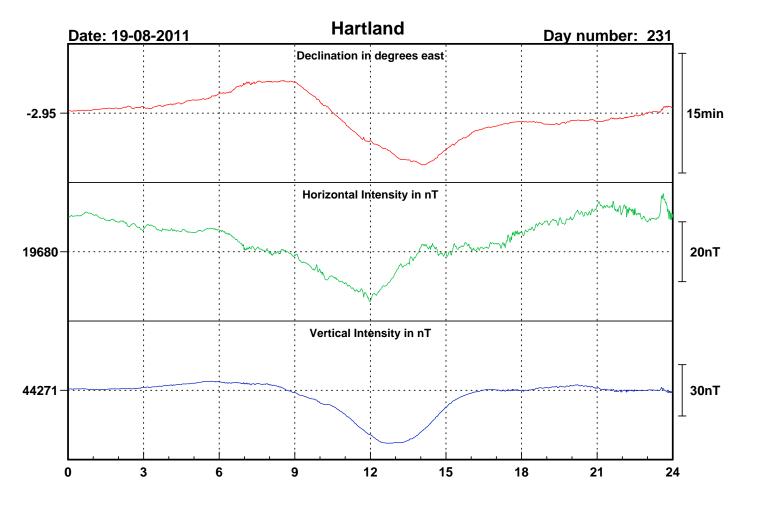


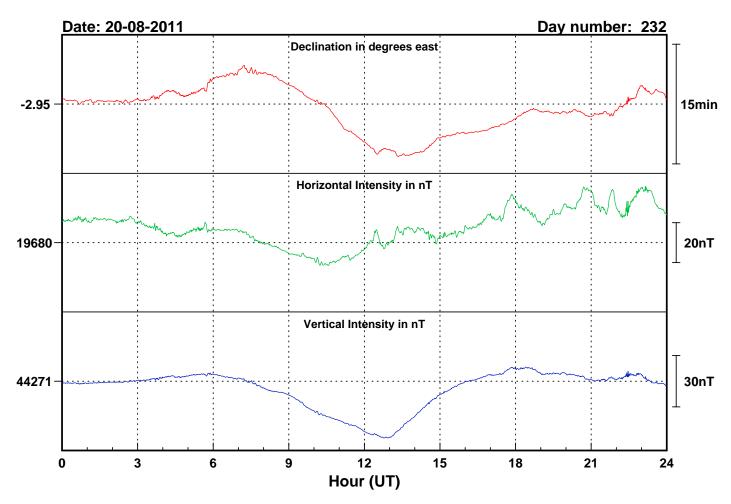


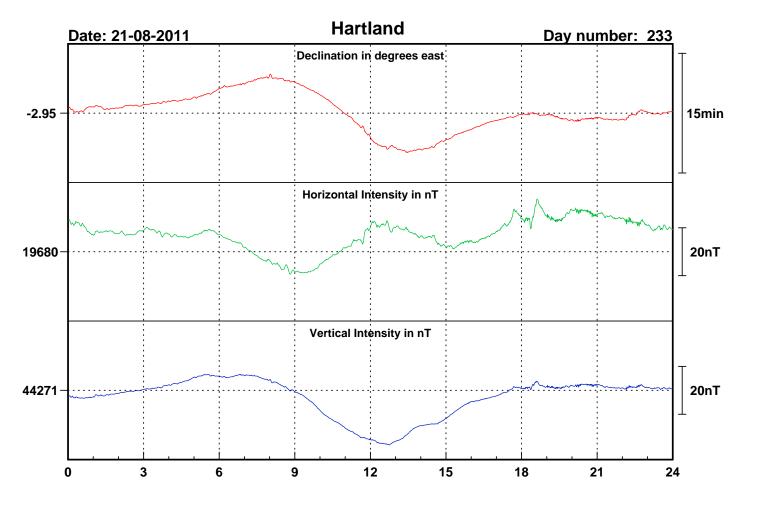


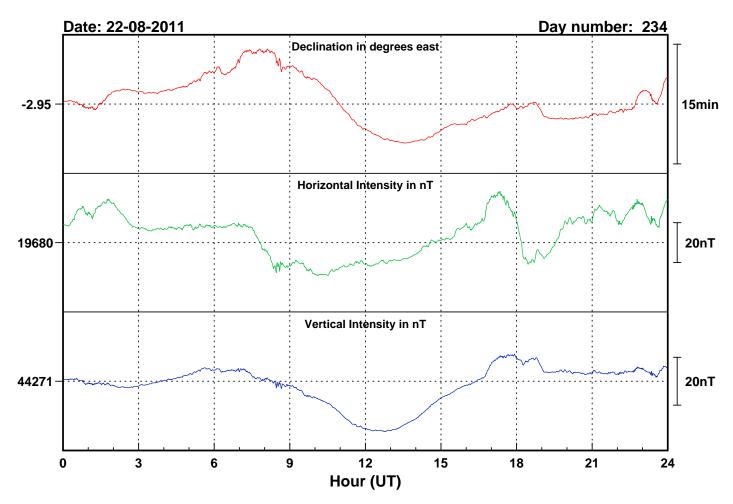


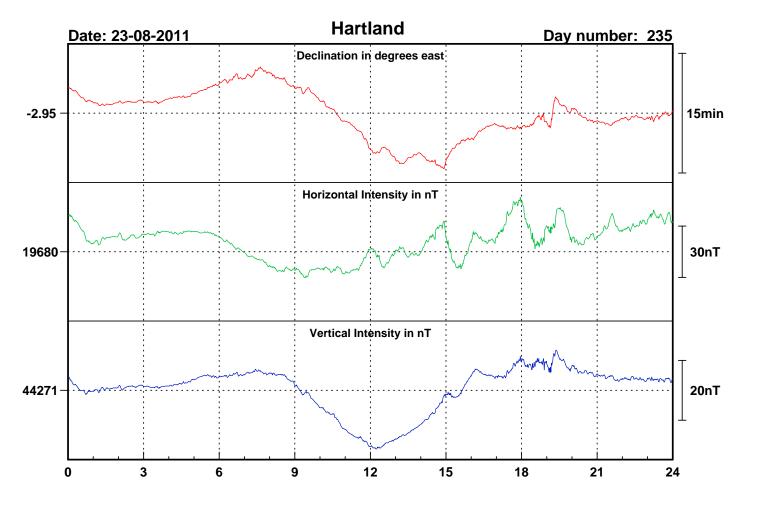


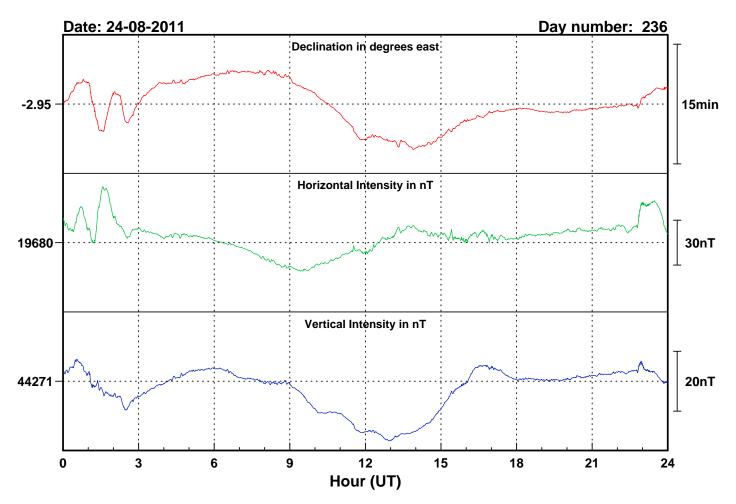


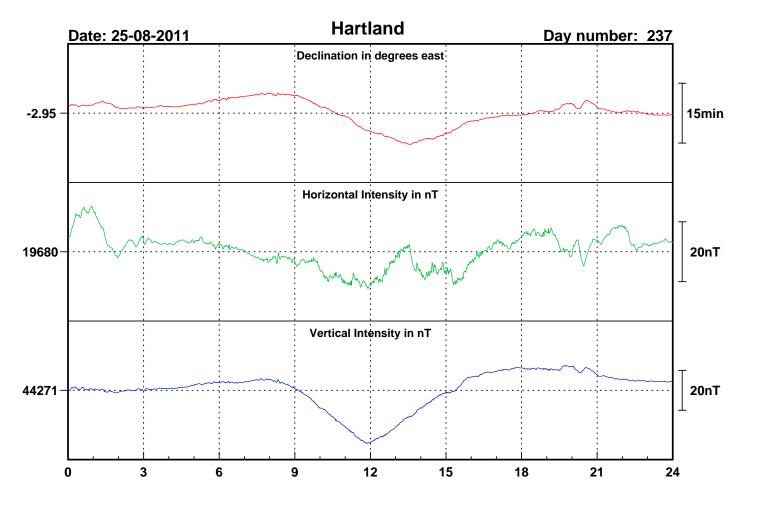


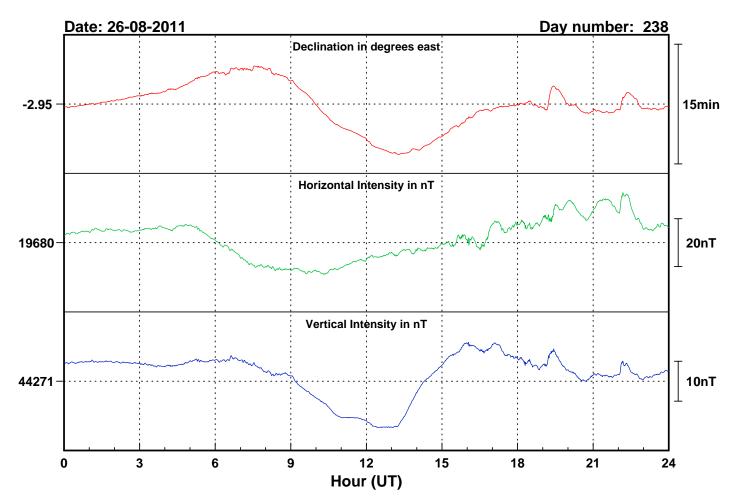


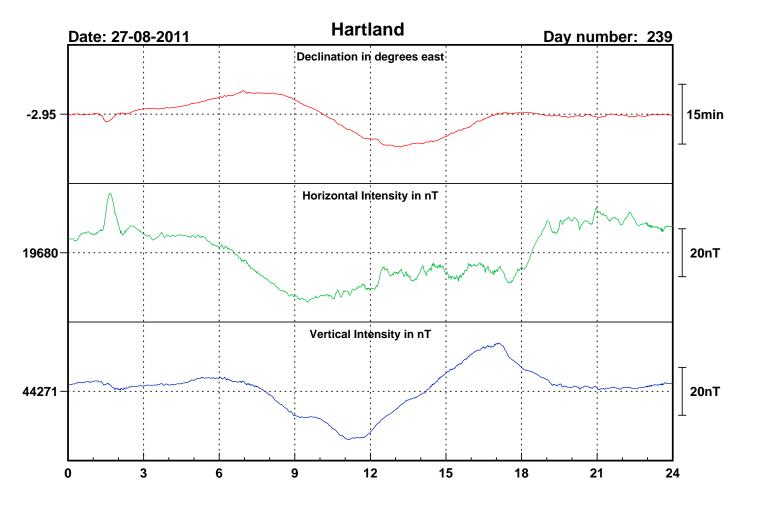


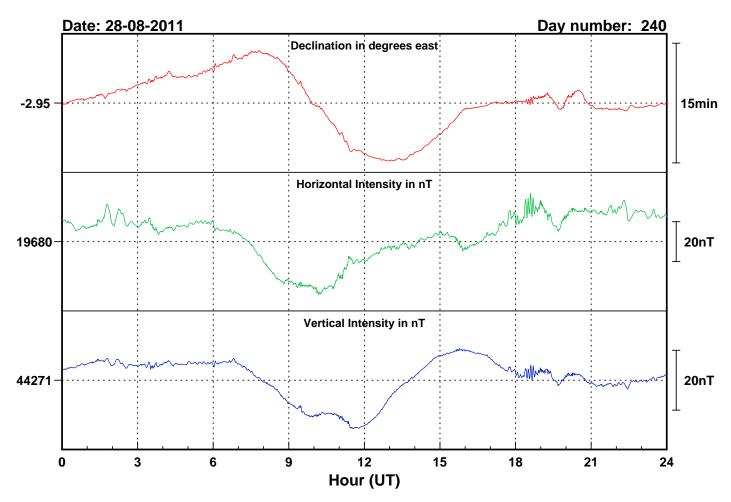


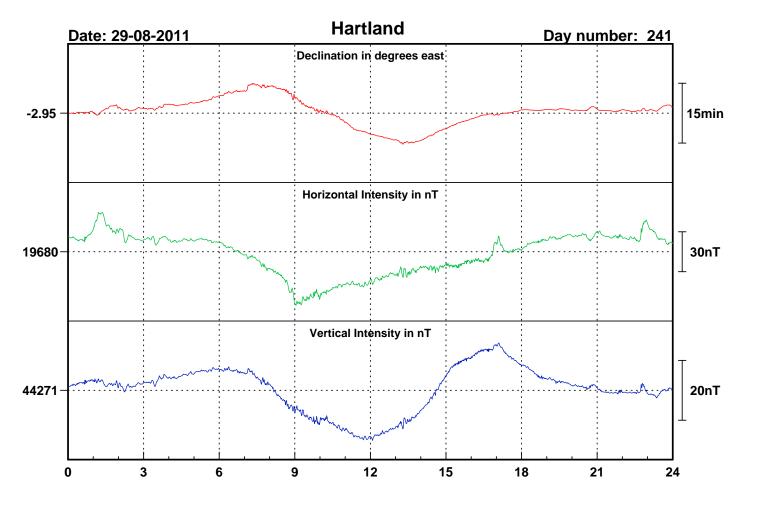


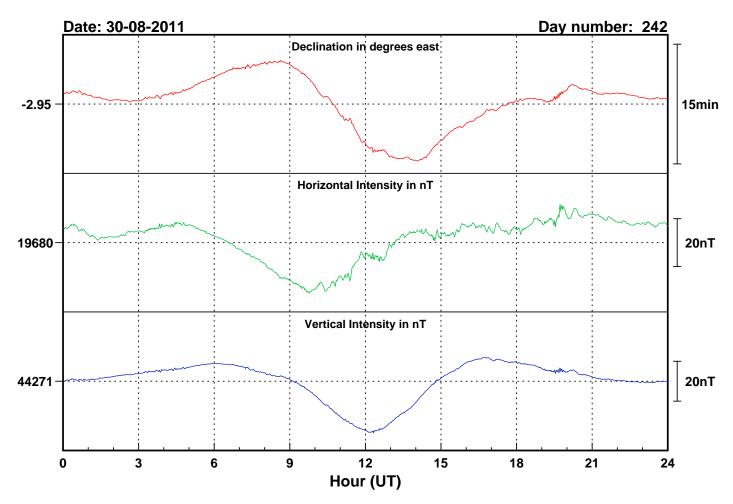


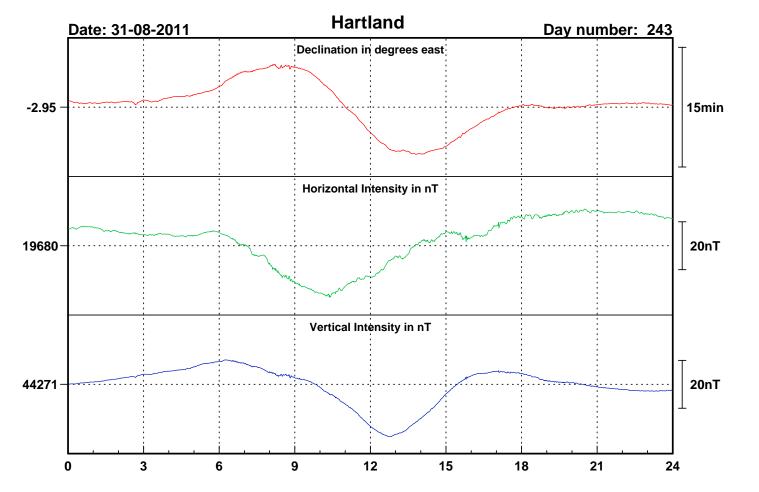




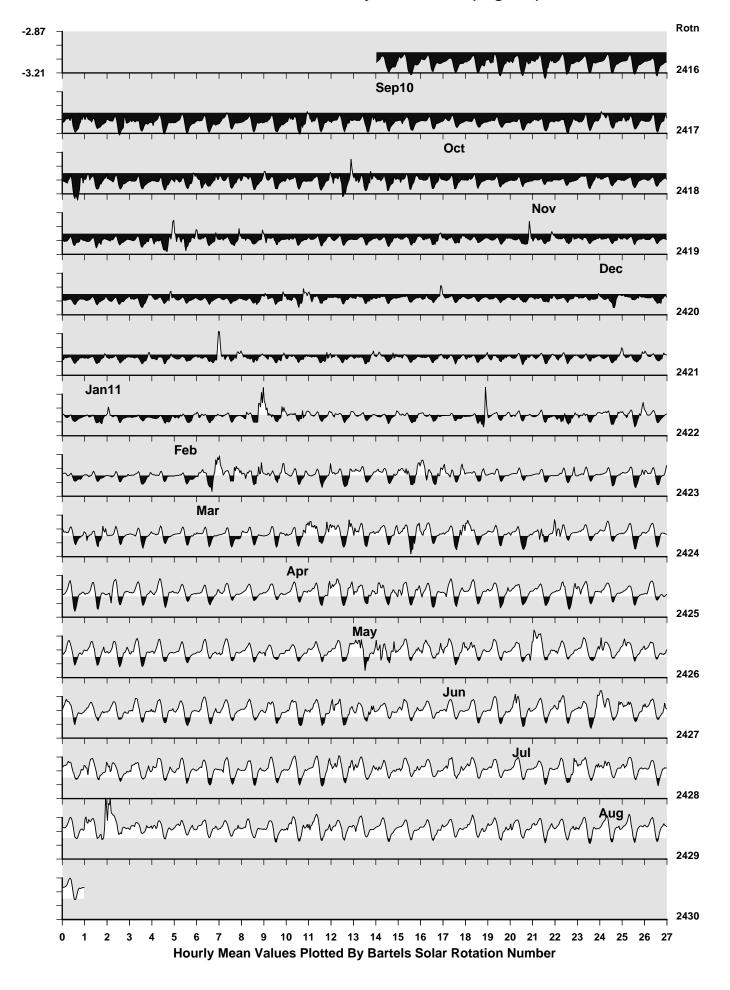




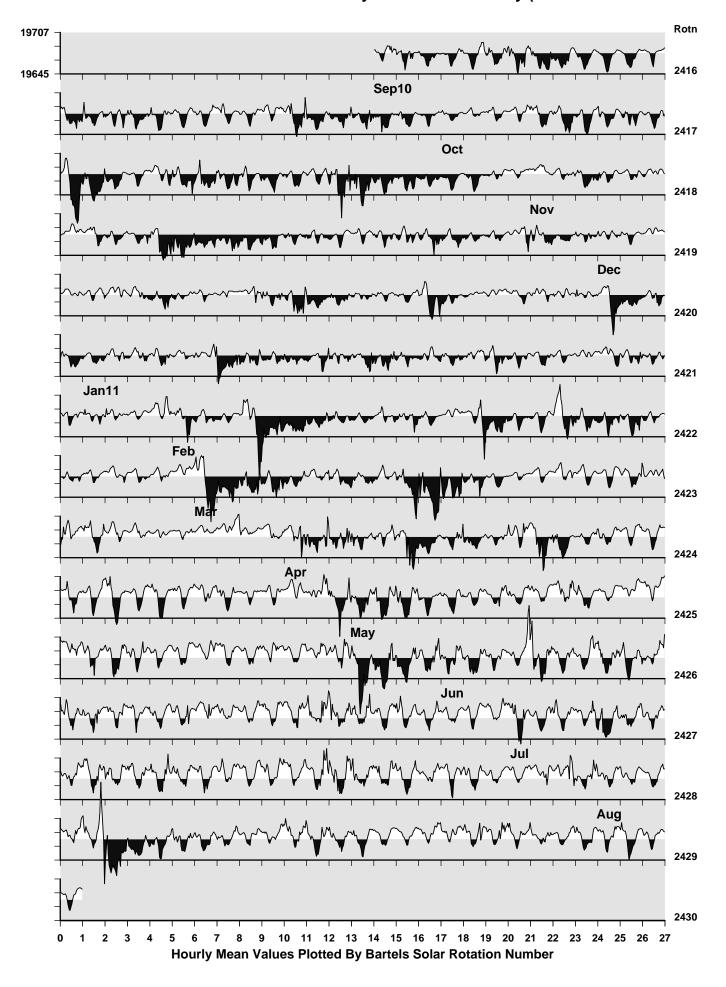




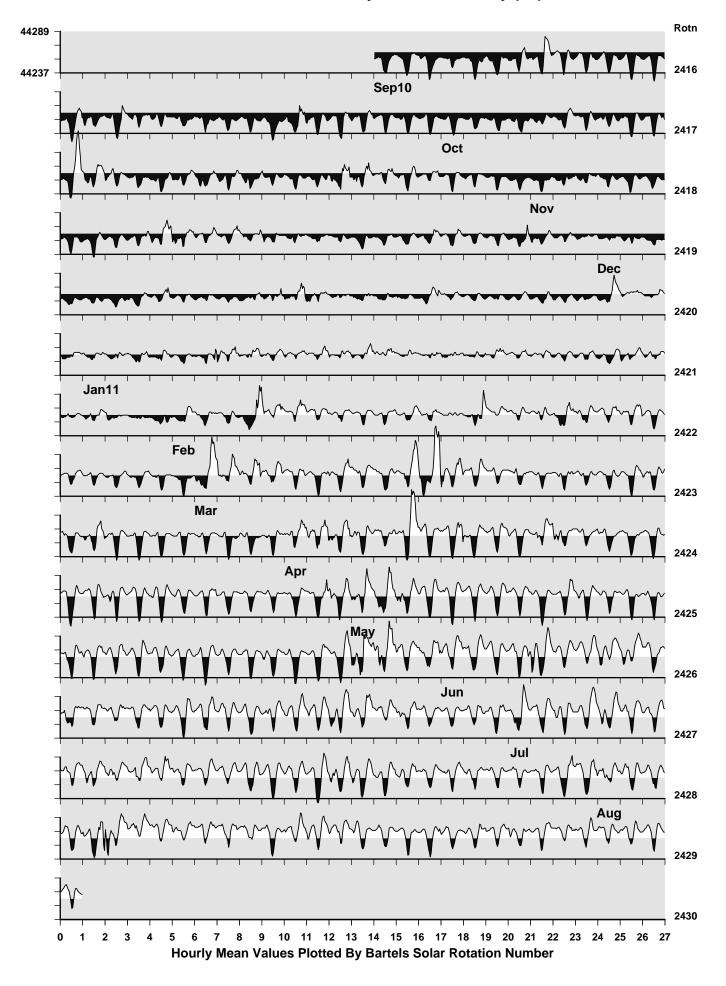
Hartland Observatory: Declination (degrees)

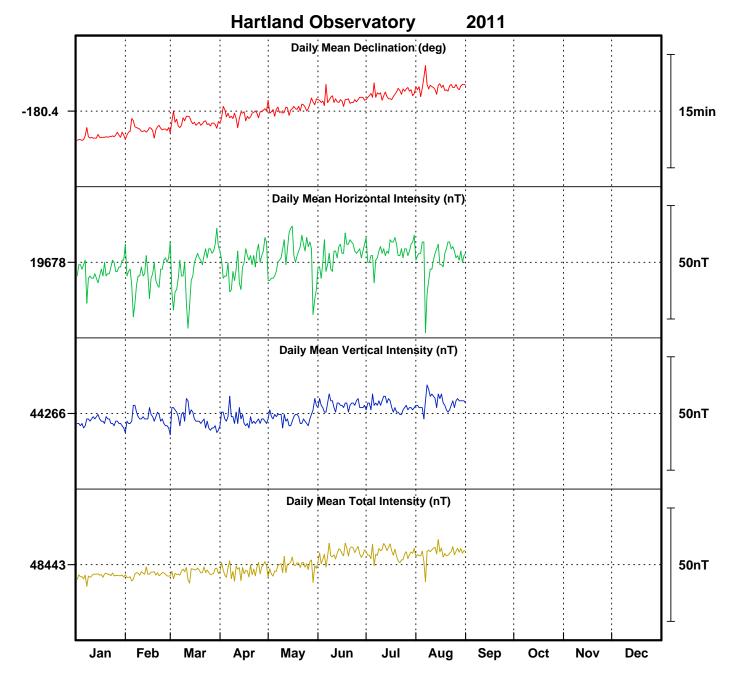


Hartland Observatory: Horizontal Intensity (nT



Hartland Observatory: Vertical Intensity (nT)





Monthly Mean Values for Hartland Observatory 2011

Month	D	H	I	X	Y	Z	F
January	-3° 3.8′	19674 nT	66° 2.1′	19646 nT	-1052 nT	44263 nT	48438 nT
February	-3° 2.8′	19673 nT	66° 2.3′	19645 nT	-1046 nT	44264 nT	48439 nT
March	-3° 1.9′	19675 nT	66° 2.1′	19648 nT	-1041 nT	44264 nT	48440 nT
April	-3° 0.8′	19677 nT	66° 2.0′	19650 nT	-1035 nT	44264 nT	48441 nT
May	-3° 0.0′	19680 nT	66° 1.8′	19653 nT	-1030 nT	44265 nT	48443 nT
June	-2° 59.0′	19682 nT	66° 1.8′	19656 nT	-1024 nT	44270 nT	48448 nT
July	-2° 58.1′	19683 nT	66° 1.8′	19656 nT	-1019 nT	44270 nT	48448 nT
August	-2° 57.2′	19680 nT	66° 2.0′	19653 nT	-1014 nT	44271 nT	48448 nT

Note

i. The values shown here are provisional.

HARTLAND RAPID VARIATIONS

SIs and SSCs

Date	Time (UT)	Туре	Quality	H (nT)	D (min)	Z (nT)
04-08-11	21 53	SSC*	A	23.1	-0.90	4.8
05-08-11	17 50	SSC	В	15.7	-1.60	4.1/-4.1
05-08-11	19 00	SSC	A	95.4	-3.72	13.3

Notes:

An asterisk (*) indicates that the principal impulse was preceded by a smaller reversed impulse. The quality of the event is classified as follows:

A = very distinct

B = fair, ordinary, but unmistakable

C = doubtful

The amplitudes given are for the first chief movement of the event.

SFEs

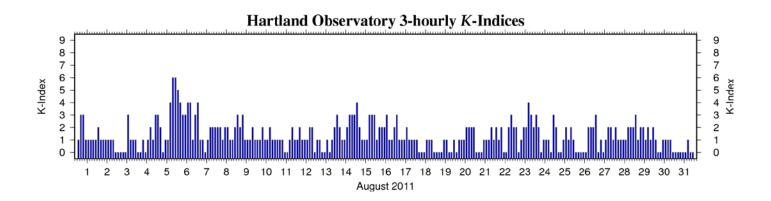
Date		Universal Time		H (nT)	D (min)	Z (nT)
	Start	Maximum	End			
09-08-11	08 01	08 05	08 15	22.4	2.35	5.1

Note:

The amplitudes given are for the first chief movement of the event.

INDICES OF GEOMAGNETIC ACTIVITY

	K - INDICES FOR THREE-HOUR INTERVAL										
Day	00-03	03-06	06-09	09-12	12-15	15-18	18-21	21-24			
1	1	3	3	1	1	1	1	1			
2	2	1	1	1	1	1	1	0			
3	0	0	0	0	3	1	1	1			
4	0	0	1	0	1	2	1	3			
5	3	2	0	1	1	4	6	6			
6	5	4	3	3	4	4	1	3			
7	4	1	1	0	1	2	2	2			
8	2	2	1	2	2	1	1	2			
9	3	2	3	1	1	1	2	1			
10	1	1	2	1	1	2	1	1			
11	1	1	1	0	0	1	2	1			
12	1	2	1	1	1	2	2	0			
13	1	1	0	0	1	0	1	2			
14	3	2	1	1	2	3	3	3			
15	4	2	1	1	1	3	3	3			
16	1	2	2	2	3	1	1	2			
17	3	1	1	1	2	1	1	1			
18	1	0	0	0	1	1	1	0			
19	0	0	0	1	1	0	0	1			
20	0	1	1	1	2	2	2	2			
21	0	0	0	1	1	1	2	1			
22	2	1	2	0	0	2	3	2			
23	2	0	1	2	2	4	3	2			
24	3	2	0	1	1	1	0	3			
25	2	0	0	1	2	1	2	1			
26	0	0	0	0	0	2	2	2			
27	3	0	1	0	1	2	2	1			
28	2	1	1	1	1	2	2	2			
29	3	1	2	2	1	2	1	2			
30	1	0	0	1	1	1	1	0			
31	0	0	0	0	0	1	0	0			



The aa Index

Date	Day	3-hourly aa-indices								Aa _{am}	Aa_{pm}	Aa
01-08-11	213	12	24	24	8	8	5	5	8	17.0	6.5	11.7
02-08-11	214	12	8	8	8	5	5	5	2	8.8	4.5	6.6
03-08-11	215	2	5	2	2	17	5	5	5	3.1	8.1	5.6
04-08-11	216	2	9	5	2	8	9	5	24	4.8	11.5	8.2
05-08-11	217	24	12	2	5	5	37	115	171	10.9	82.1	46.5
06-08-11	218	102	59	32	46	59	45	8	24	59.6	34.0	46.8
07-08-11	219	37	8	8	2	8	12	16	9	13.8	11.2	12.5
08-08-11	220	9	16	8	24	38	8	8	12	14.3	16.3	15.3
09-08-11	221	24	24	32	12	8	5	12	8	23.1	8.2	15.6
10-08-11	222	8	8	24	12	12	12	8	5	13.0	9.2	11.1
11-08-11	223	5	8	8	5	5	5	9	8	6.5	6.8	6.7
12-08-11	224	5	12	8	12	5	12	16	2	9.2	8.9	9.0
13-08-11	225	5	8	2	2	8	2	5	12	4.5	6.8	5.6
14-08-11	226	24	12	8	12	12	32	24	24	13.9	23.0	18.5
15-08-11	227	45	16	8	12	12	32	32	24	20.3	25.1	22.7
16-08-11	228	8	24	16	24	32	8	8	12	18.1	14.9	16.5
17-08-11	229	24	8	12	12	24	5	5	5	14.0	9.9	11.9
18-08-11	230	5	2	2	2	8	12	5	2	3.1	6.8	5.0
19-08-11	231	2	2	2	5	5	2	2	8	3.1	4.5	3.8
20-08-11	232	5	12	5	5	12	9	12	16	6.8	12.3	9.6
21-08-11	233	2	2	2	5	8	5	9	5	3.1	6.8	5.0
22-08-11	234	16	8	24	9	2	12	32	16	14.4	15.6	15.0
23-08-11	235	12	5	12	16	12	45	32	16	11.3	26.4	18.8
24-08-11	236	24	12	5	8	12	12	2	24	12.2	12.6	12.4
25-08-11	237	12	5	5	8	16	20	16	5	7.5	14.3	10.9
26-08-11	238	5	5	2	2	2	12	12	16	3.8	10.6	7.2
27-08-11	239	24	5	5	2	12	16	12	8	9.2	11.9	10.6
28-08-11	240	12	8	8	12	8	12	16	12	9.9	11.9	10.9
29-08-11	241	20	8	24	16	8	12	5	12	17.0	9.2	13.1
30-08-11	242	5	5	2	12	8	5	5	2	6.2	5.1	5.7
31-08-11	243	5	2	5	2	2	5	2	2	3.8	3.1	3.5
	Mean Monthly Value										13.0	

Notes

The units of the aa index are nT. i.

The 3-hour aa values are rounded to the nearest integer. Where aa = *.5, aa is rounded down. Daily values $(Aa_{am}, Aa_{pm} \text{ and } Aa)$ are computed from aa values of original resolution. The monthly mean value is computed from the daily mean values, Aa. ii.

iii.

iv.

Definitive aa indices are derived and published by the International Service for Geomagnetic Indices. v.

